



CREDO science case

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CREDO » Funding » H2020:
Exploring and supporting citizen science, 10.04.2018

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Horizon 2010: Working Package 1 (Science)

Added by [Dariusz Gora](#) 19 days ago. Updated 5 days ago.

Status: New
Priority: Normal
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 [WP1_science_v0p1.pdf](#) (123 KB)  [Dariusz Gora](#), 03/12/2018 11:50 AM

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Subtasks

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Related issues

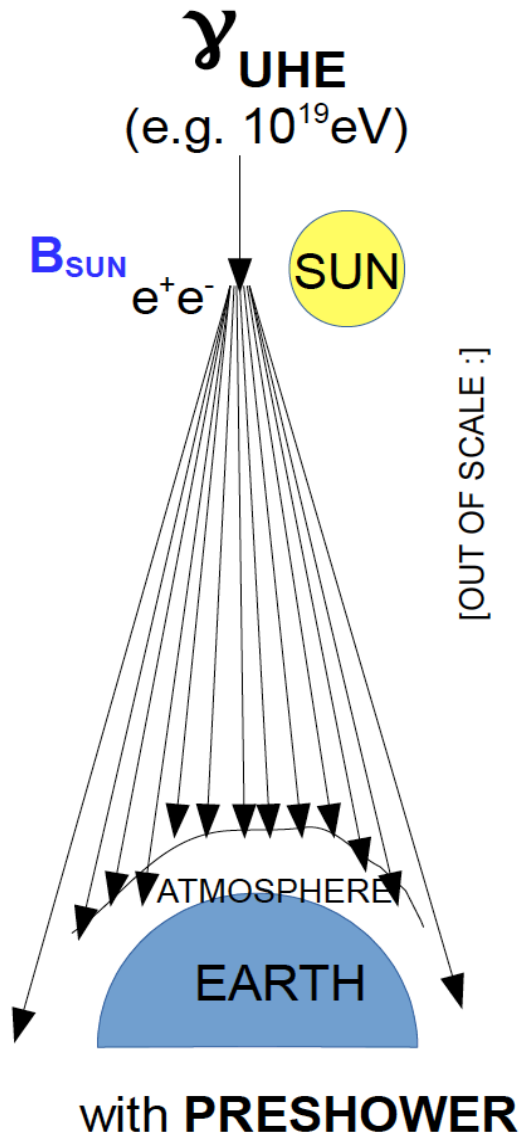
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History

Topics

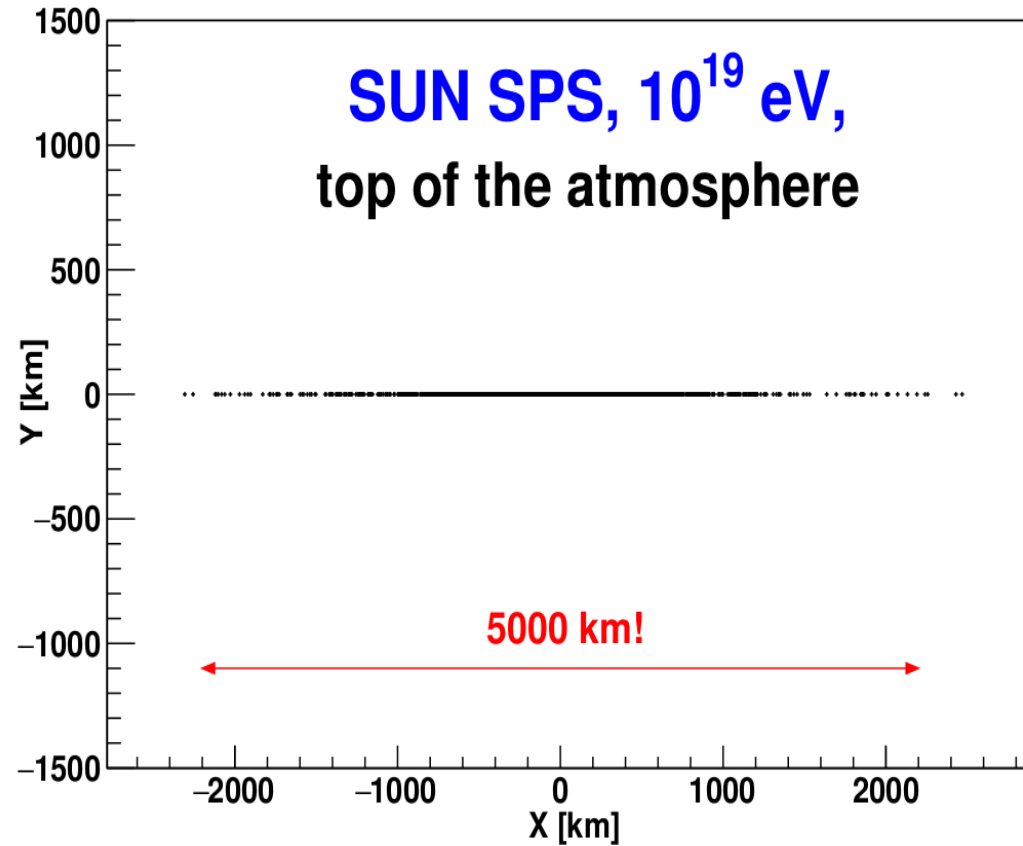
- 1) CRPropa standard / exotic CRE: conditions
- 2) looking for Sun Cosmic Ray Ensembles
- 3) clusters in time: any time scale suspected
- 4) Integration of data from different experiments: Credo monitor
- 5) search for correlations: cosmic / GMF / weather / sun activity / seismic effects

Topic 1: Preshower in vicinity of Sun



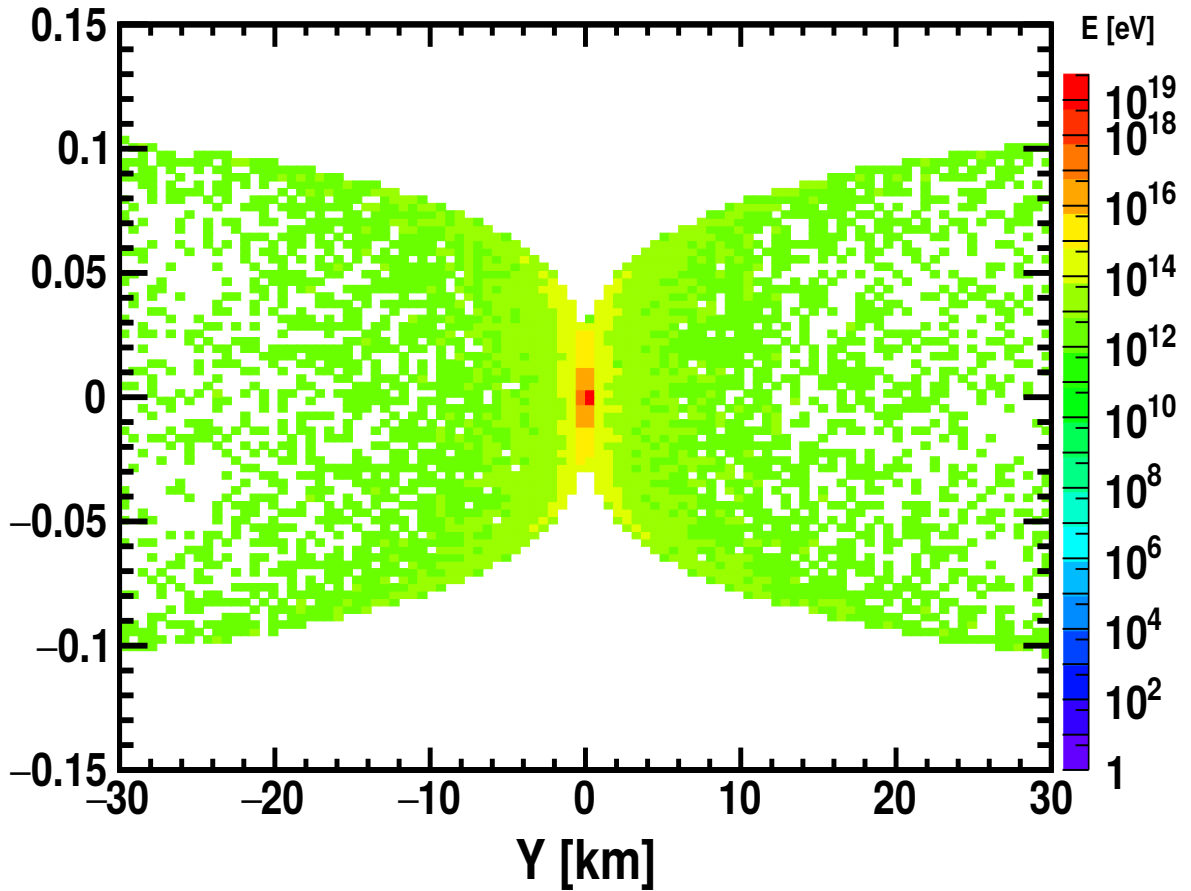
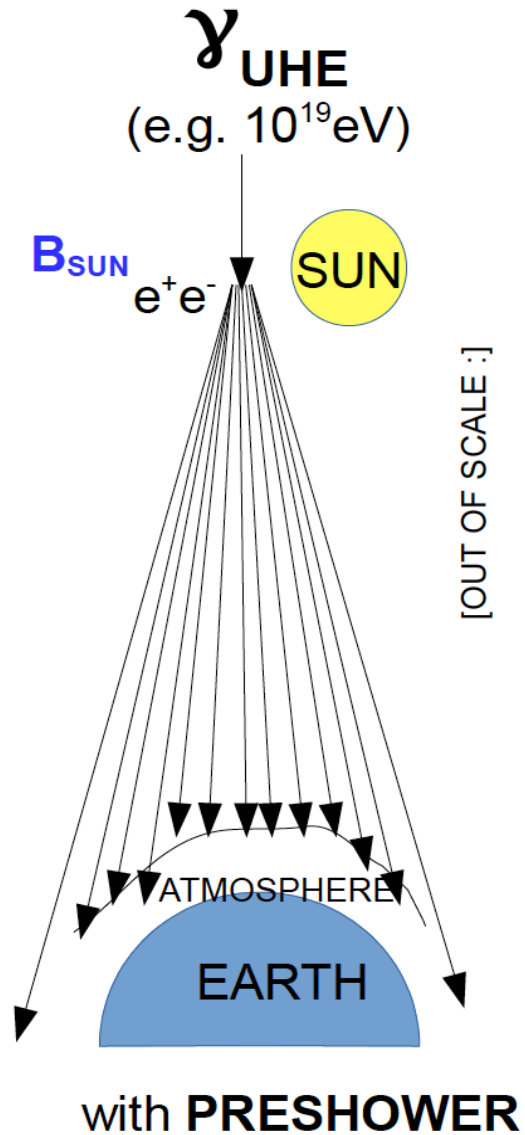
[OUT OF SCALE :]

Distribution of photons at the top of the Atm



Distribution of photons ($E > 10^{13}$ eV) at the top of the atmosphere.
 $E_{\gamma} = 10$ EeV, Impact parameter = $2.5R_{\text{S}}$.

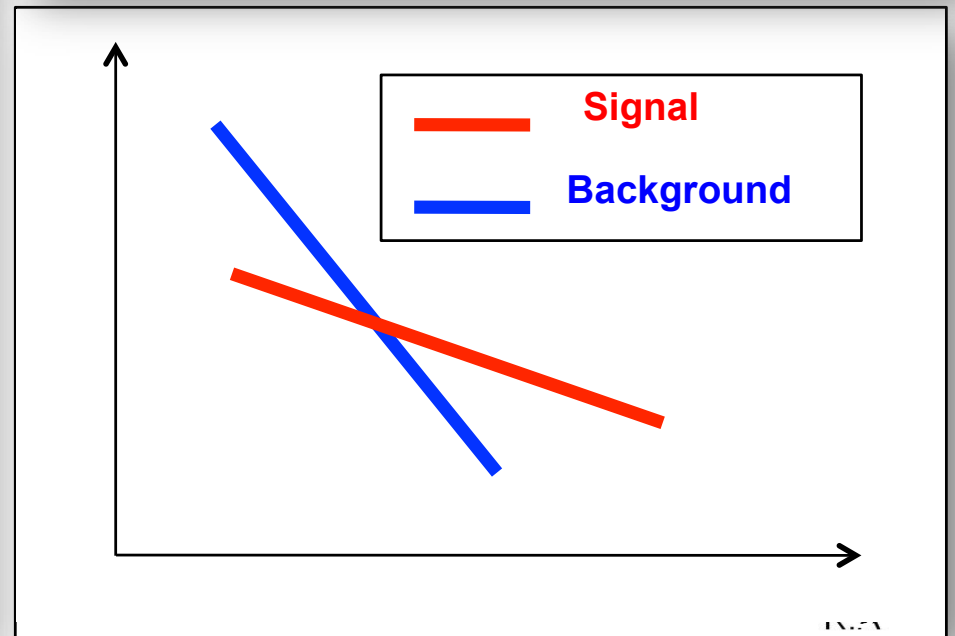
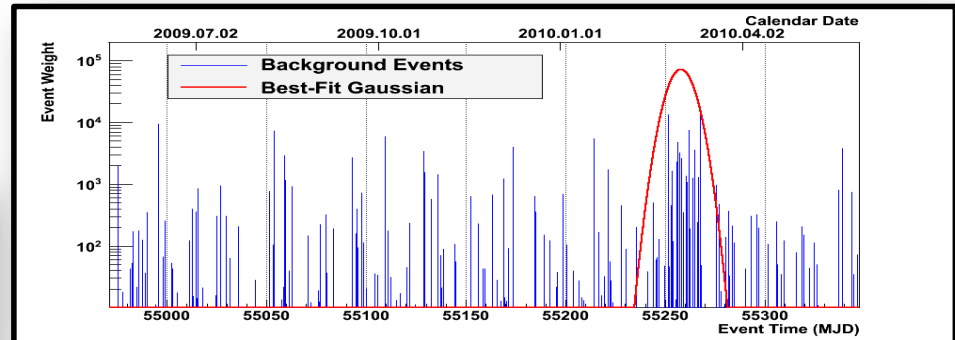
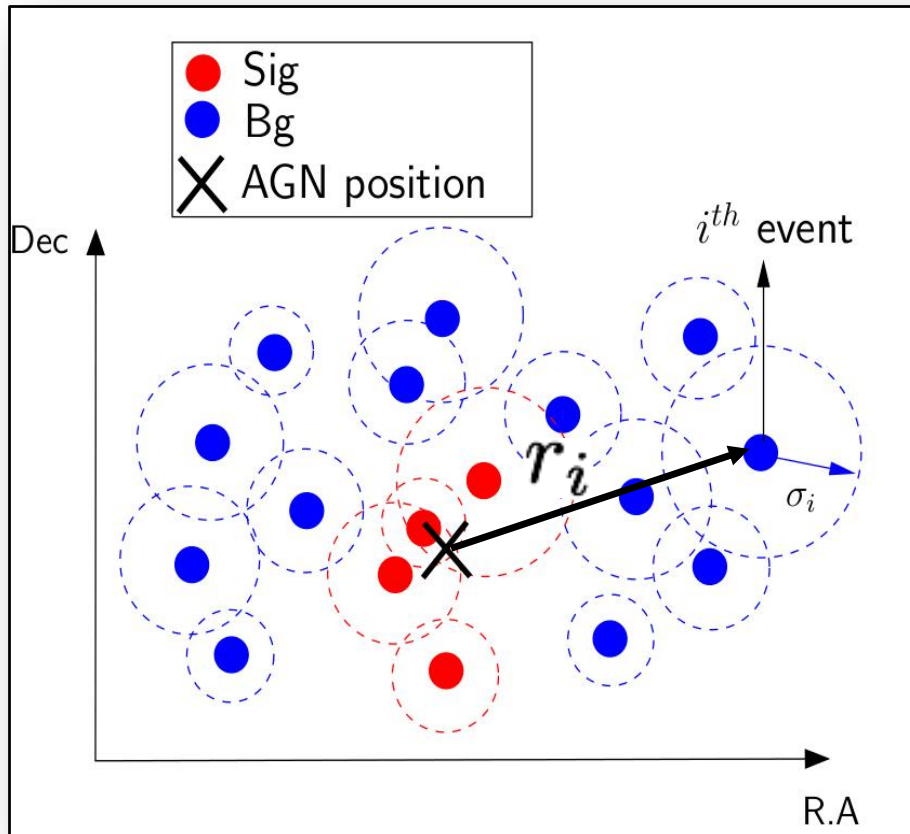
Example of CRE : Preshower in vicinity of Sun



Topic3: Basic concept of point-source search filter

σ_i - the angular uncertainty of event

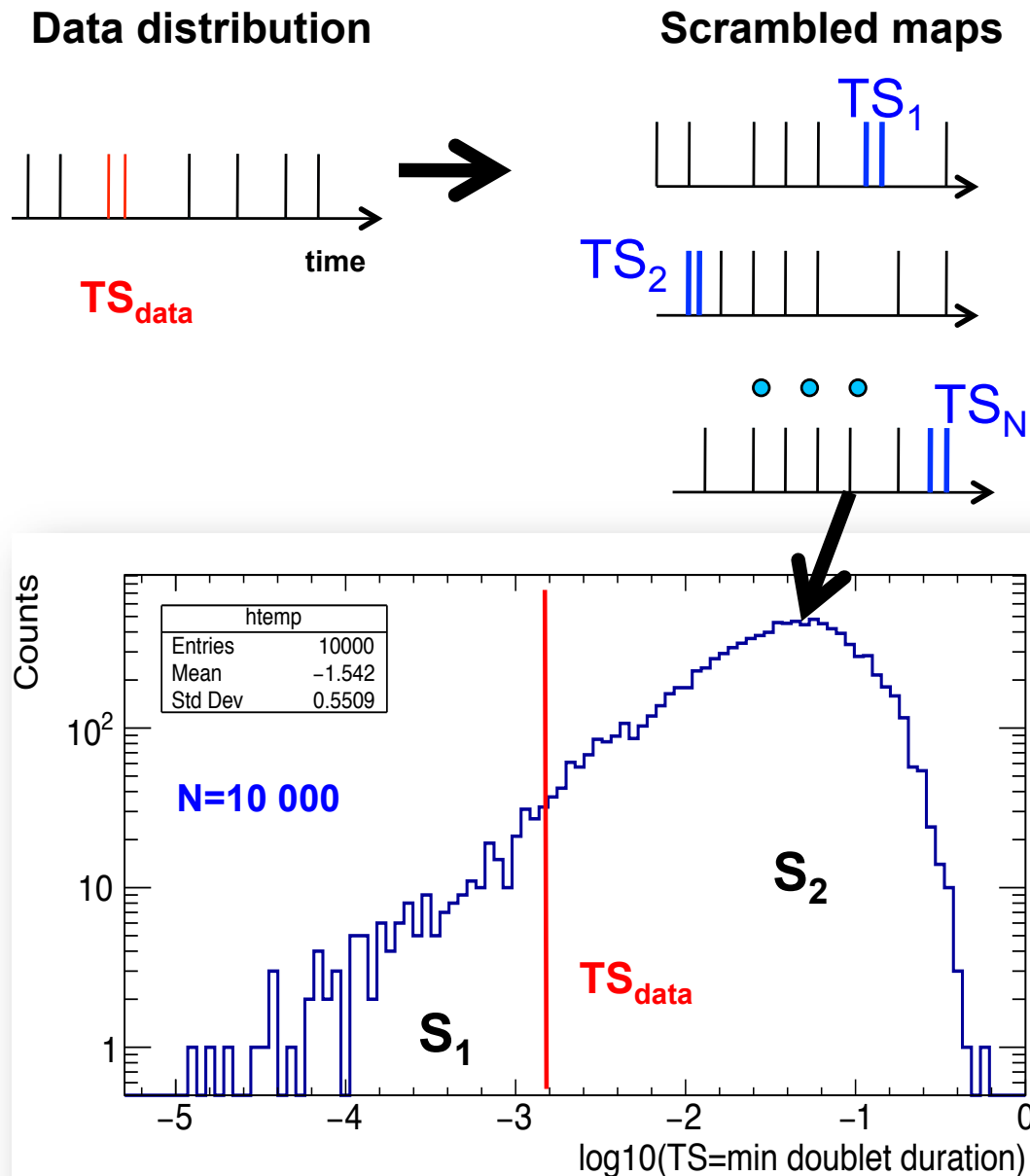
r_i - the angular distance of event from source



Finding point sources in the sky means to locate *an excess of events from a particular direction* over the background.

The signals events may present additional features: *different energy spectrum or time structure*

Simple method: search for time cluster



Test Statistics (TS)

$TS := \text{doublet duration}^*$

*for other multiplets different definitions assumed

RESULT:

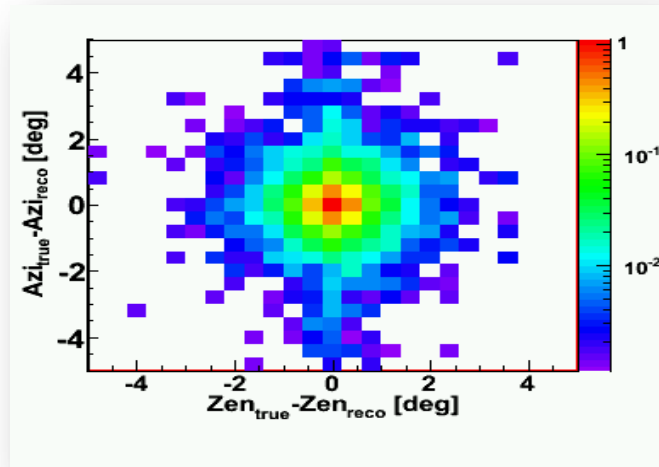
p-value as the fraction of events below TS_{data}

$$\text{P-value} = \frac{S_1}{S_1 + S_2}$$

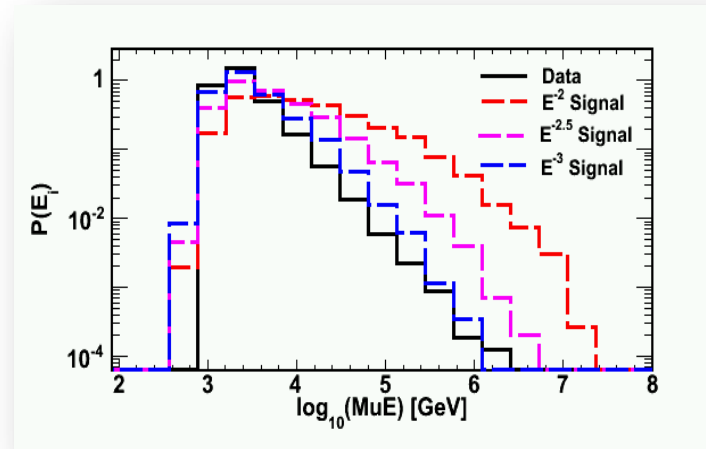
General method: an unbinned maximum likelihood

Method (J. Braun et al., Astropart. Phys.33:175,2010)

Space probability density function



Energy probability density function

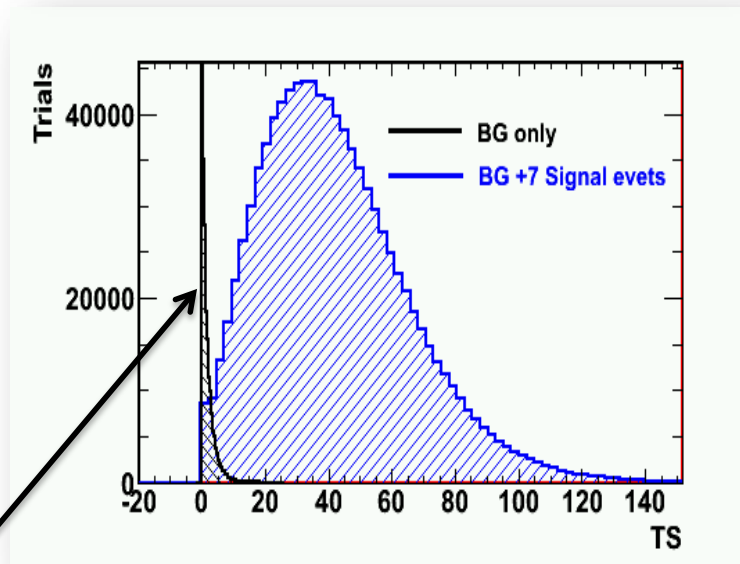


$$\mathcal{L}(\mathbf{n}_s, \gamma) = \prod_{i=1}^N \left[\frac{n_s}{N} \mathcal{S}_i + \left(1 - \frac{n_s}{N}\right) \mathcal{B}_i \right]$$

$$\mathcal{S}_i = P^{\text{space}}(|x_i - x_s|, \sigma_i) \times P^{\text{energy}}(E_i | \gamma) \times P^{\text{time}}(\Delta t_j)$$

$$\mathcal{B}_i = \frac{1}{d\Omega} \times P_i(E_i) \times P^{\text{time}}(\theta, \phi, t_i)$$

$$\text{TS} = -2 \log(\lambda) = -2 \log \frac{\mathcal{L}(\mathbf{n}_s = 0)}{\mathcal{L}(\mathbf{n}_s^f, \gamma^f)}$$



TS distribution for no signal events in data sample

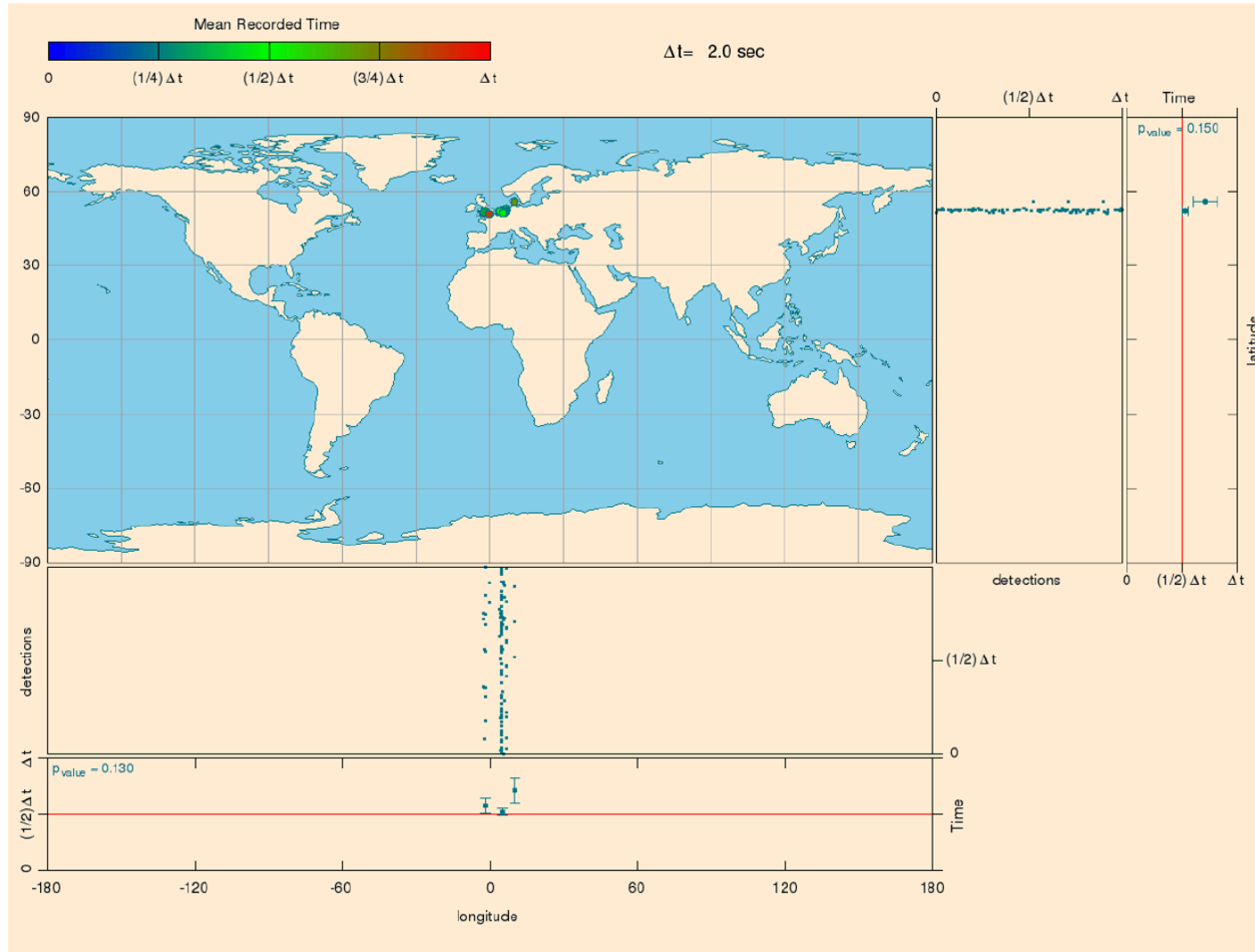
CREDO Monitor

CREDO monitor is a daily automated list of tasks whose ultimate objective is a search of Cosmic Ray Ensembles (CRE).

- **Data Acquisition/Migration:** Acquisition of data and transfer to the central server.
- **Data Conversion:** Conversion of data from different detectors/sources into a common format.
- **Data Preprocessing:** Data sorted in time and merged into the final form (daily sets).
- **Analysis:** A preliminary scanning done to look for possible CRE signatures.
- **Map Production:** Creation of maps for cosmic ray arrival time as a function of geographic locations.
- **Map Sharing:** Sharing the map for analysis/further classification.

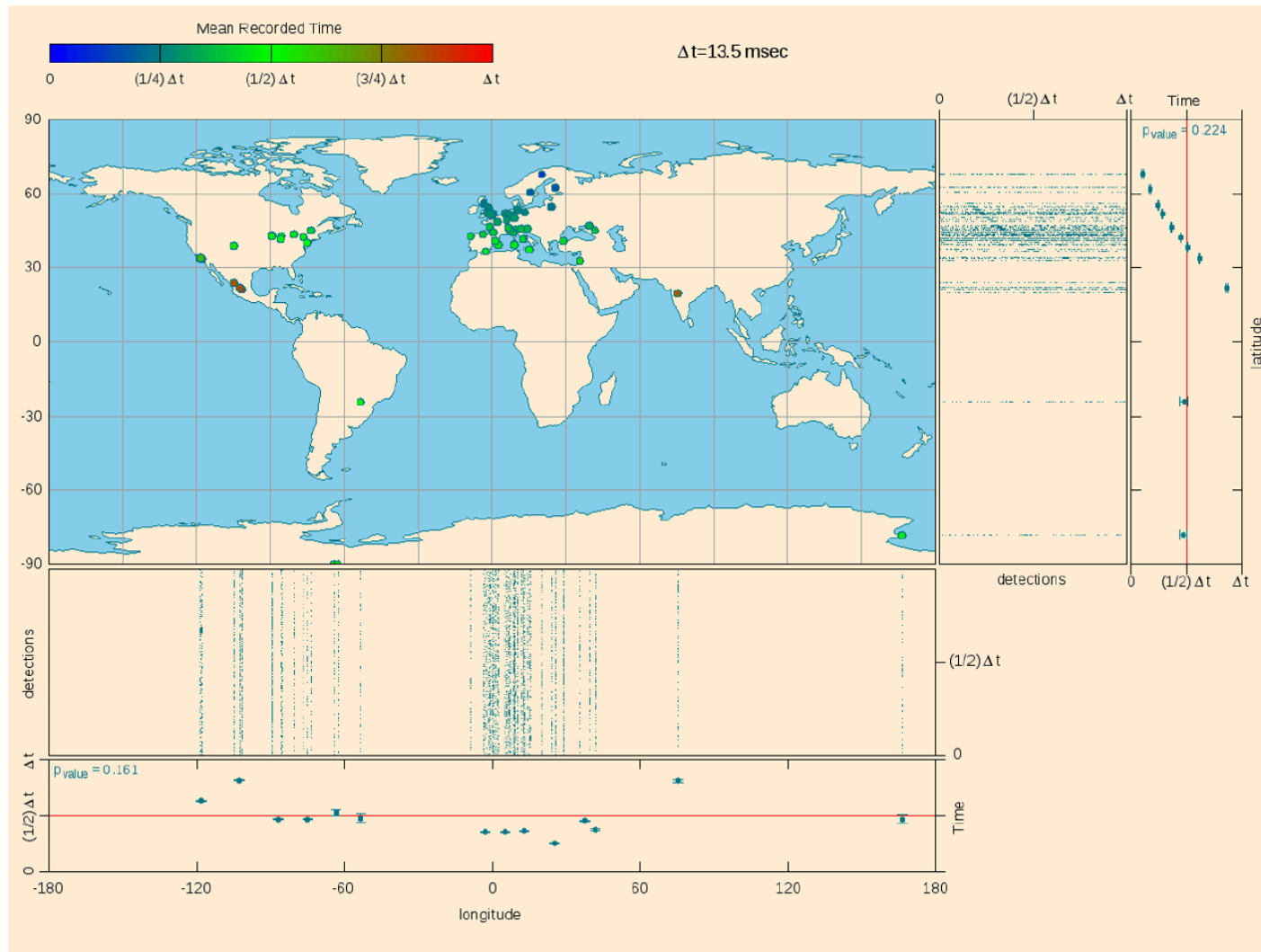
The storage, monitoring and computing tasks of CREDO are hosted by ACC Cyfronet AGH-UST.

CREDO Monitor



Real data
analysis

CREDO Monitor

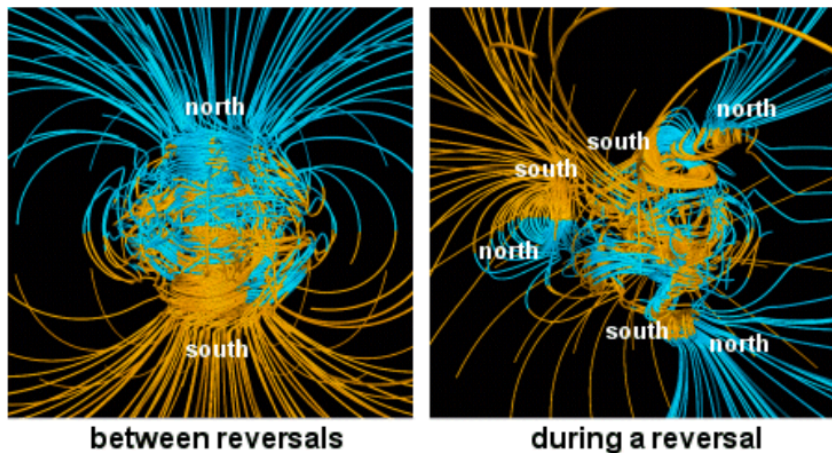


How a CRE signature can look like

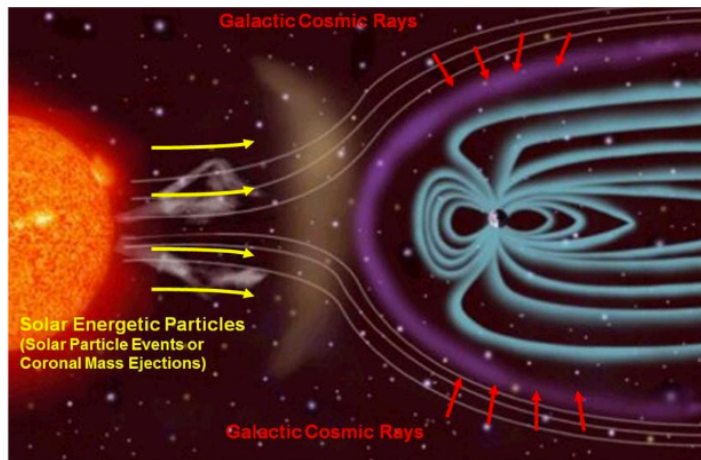
Simulation

Topic 5:

Wikipedia: „Geomagnetic reversal”



Wikipedia: „Health threat from cosmic rays”



Earth outer core: Liquid (molten iron)

→ geomagnetism



Impulse (tidal forces)

→ hydrodynamics: waves



→ Mechanical wave upwards (slow, hours?)

→ Electromagnetic wave („instant”, ms)



Local geomagnetic field vector changes
AND seismic effect might occur!



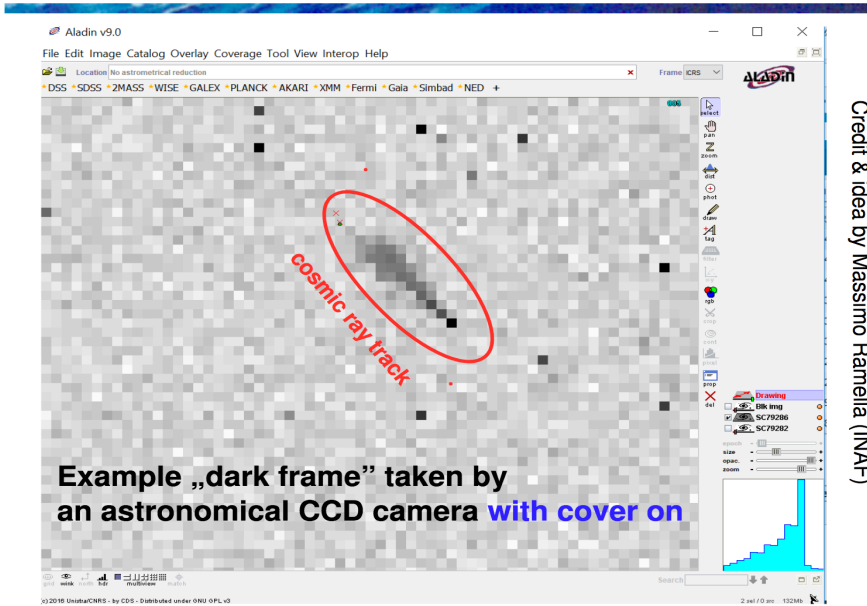
Variation of the CR rate!



Earthquake precursors?

Varia:

CREDO attracts... astronomers!



... and gamma-ray astronomy