# Model-independent analysis of nearly Lévy source



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### OUTLINE

### Model-independent shape analysis:

- General introduction
- Edgeworth, Laguerre
- Lévy expansions
- Applications

### Summary

## **MODEL - INDEPENDENT SHAPE ANALYIS I.**

Model-independent method, proposed to analyze Bose-Einstein correlations IF experimental data satisfy

- The measured data *tend to a constant* for large values of the observable Q.
- There is a *non-trivial structure* at some definite value of Q, shift it to Q = 0.

Model-independent, but experimentally testable:

- t = Q R
- dimensionless scaling variable
- approximate form of the correlations *w(t)*
- Identify w(t) with a measure in an abstract Hilbert-space

$$\int dt w(t) h_n(t) h_m(t) = \delta_{n,m},$$
  
$$f(t) = \sum_{n=0}^{\infty} f_n h_n(t),$$
  
$$f_n = \int dt w(t) f(t) h_n(t).$$

e.g. 
$$t = Q_I R_I$$

T. Csörgő and S: Hegyi, hep-ph/9912220, T. Csörgő, hep-ph/001233

### **MODEL - INDEPENDENT SHAPE ANALYIS II.**

$$C_2(\mathbf{k}_1, \mathbf{k}_2) = \frac{N_2(\mathbf{k}_1, \mathbf{k}_2)}{N_1(\mathbf{k}_1) N_1(\mathbf{k}_2)},$$

$$R_2(\mathbf{k}_1, \mathbf{k}_2) = C_2(\mathbf{k}_1, \mathbf{k}_2) - 1.$$

Let us assume, that the function  $g(t) = R_2(t)/w(t)$  is also an element of the Hilbert space H. This is possible, if

$$\int dt \, w(t)g^2(t) = \int dt \, \left[ R_2^2(t)/w(t) \right] < \infty,\tag{6}$$

Then the function *g* can be expanded as

From the completeness of the Hilbert space, if g(t) is also in the Hilbert space:

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$$g(t) = \sum_{n=0}^{\infty} g_n h_n(t),$$
$$g_n = \int dt R_2(t) h_n(t).$$

$$R_2(t) = w(t) \sum_{n=0}^{\infty} g_n h_n(t).$$

## **MODEL - INDEPENDENT SHAPE ANALYIS III.**

$$C_2(\mathbf{k}_1, \mathbf{k}_2) = \frac{N_2(\mathbf{k}_1, \mathbf{k}_2)}{N_1(\mathbf{k}_1) N_1(\mathbf{k}_2)},$$

$$C_2(t) = \mathcal{N}\left\{1 + \lambda_w w(t) \sum_{n=0}^{\infty} g_n h_n(t)\right\}$$

### Model-independent AND experimentally testable:

- method for any approximate shape w(t)
- the core-halo intercept parameter of the CF is

$$\lambda_* = \lambda_w \sum_{n=0}^{\infty} g_n h_n(0)$$

- coefficients by numerical integration (fits to data)  $g_n = \int dt R_2(t) h_n(t)$
- condition for applicability: experimentally testable

$$\int dt \, \left[ R_2^2(t)/w(t) \right] < \infty$$

## **GAUSSIAN w(t): EDGEWORTH EXPANSION**

$$t = \sqrt{2}QR_E,$$
  

$$w(t) = \exp(-t^2/2),$$
  

$$\int_{-\infty}^{\infty} dt \, \exp(-t^2/2) H_n(t) H_m(t) \propto \delta_{n,m},$$
  

$$H_1(t) = t,$$
  

$$H_2(t) = t^2 - 1,$$
  

$$H_3(t) = t^3 - 3t,$$
  

$$H_4(t) = t^4 - 6t^2 + 3, \dots$$
  

$$C_2(Q) = \mathcal{N} \left\{ 1 + \lambda_E \exp(-Q^2 R_E^2) \times \left[ 1 + \frac{\kappa_3}{3!} H_3(\sqrt{2}QR_E) + \frac{\kappa_4}{4!} H_4(\sqrt{2}QR_E) + \dots \right] \right\}.$$

### 3d generalization straightforward

• Applied by NA22, L3, STAR, PHENIX, ALICE, CMS (LHCb)

## **EXPONENTIAL w(t): LAGUERRE EXPANSIONS**

 $\infty$ 

### Model-independent but experimentally tested:

- w(t) exponential
- t. dimensionless
- Laguerre polynomials

$$t = QR_L,$$
  

$$w(t) = \exp(-t)$$
  

$$dt \ \exp(-t)L_n(t)L_m(t) \propto \delta_{n,m},$$

$$L_n(t) = \exp(t) \frac{d^n}{dt^n} (-t)^n \exp(-t). \quad \begin{array}{l} L_0(t) = 1, \\ L_1(t) = t - 1, \end{array}$$

$$C_2(Q) = \mathcal{N}\left\{1 + \lambda_L \exp(-QR_L)\left[1 + c_1L_1(QR_L) + \frac{c_2}{2!}L_2(QR_L) + \dots\right]\right\}$$

### **First successful tests**

- NA22, UA1 data
- convergence criteria satisfied

$$\int_{0}^{\infty} dt \, R_2^2(t) \exp(+t) < \infty,$$

### **EXAMPLE, LAGUERRE EXPANSIONS**

Laguerre expansion fit



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## **STRETCHED w(t): LÉVY EXPANSIONS**

$$w(t|\alpha) = \exp(-t^{\alpha}) = \exp(-Q^{\alpha}R^{\alpha})$$

### **Model-independent but:**

- Lévy: stretched exponential
- generalizes exponentials and Gaussians
- ubiquitous in nature
- How far from a Lévy?
- Need new set of polynomials orthonormal to a Lévy weight

$$L_0(t \mid \alpha) = \mu_{0,\alpha},$$
  

$$L_1(t \mid \alpha) = \det \begin{pmatrix} \mu_{0,\alpha} & \mu_{1,\alpha} \\ 1 & t \end{pmatrix},$$
  

$$L_2(t \mid \alpha) = \det \begin{pmatrix} \mu_{0,\alpha} & \mu_{1,\alpha} & \mu_{2,\alpha} \\ \mu_{1,\alpha} & \mu_{2,\alpha} & \mu_{3,\alpha} \\ 1 & t & t^2 \end{pmatrix},$$

$$L_{3}(t \mid \alpha) = \det \begin{pmatrix} \mu_{0,\alpha} & \mu_{1,\alpha} & \mu_{2,\alpha} & \mu_{3,\alpha} \\ \mu_{1,\alpha} & \mu_{2,\alpha} & \mu_{3,\alpha} & \mu_{4,\alpha} \\ \mu_{2,\alpha} & \mu_{3,\alpha} & \mu_{4,\alpha} & \mu_{5,\alpha} \\ 1 & t & t^{2} & t^{3} \end{pmatrix}$$
$$\mu_{r,\alpha} = \int_{0}^{\infty} dx \ x^{r} f(x \mid \alpha) = \frac{1}{\alpha} \Gamma(\frac{r+1}{\alpha})$$

STRETCHED w(t) =  $exp(-t^{\alpha})$ : LÉVY EXPANSIONS

In case of  $\alpha = 1$ , in 1 dimension Laguerre expansion is recovered

$$L_0(t \mid \alpha = 1) = 1,$$
  

$$L_1(t \mid \alpha = 1) = t - 1,$$
  

$$L_2(t \mid \alpha = 1) = t^2 - 4t + 2,$$
  

$$L_3(t \mid \alpha = 1) = 4t^3 - 36t^2 + 72t - 24...$$

These reduce to the Laguerre expansions and Laguerre polynomials. **STRETCHED** w(t)=exp( $-t^{\alpha}$ ): LEVY EXPANSIONS

## In case of $\alpha = 2$ , a new formulae for one-sided Gaussians:

$$L_{0}(t \mid \alpha = 2) = \frac{\sqrt{\pi}}{2},$$

$$L_{1}(t \mid \alpha = 2) = t\frac{\sqrt{\pi}}{2} - \frac{1}{2},$$

$$L_{2}(t \mid \alpha = 2) = t^{2} \left[\frac{\pi}{8} - \frac{1}{4}\right] - t\frac{\sqrt{\pi}}{8} - \frac{\pi}{16} + \frac{1}{4},$$

$$L_{3}(t \mid \alpha = 2) = t^{3} \left[\frac{\pi^{3/2}}{32} - \frac{3\sqrt{\pi}}{32}\right] + t^{2} \left[\frac{1}{8} - \frac{3\pi}{64}\right] + t \left[\frac{5\sqrt{\pi}}{32} - \frac{3\pi^{3/2}}{64}\right] + \frac{5\pi}{128} - \frac{1}{8}$$

Provides a new expansion around a Gaussian shape that is defined for the non-negative values of *t* only. Edgeworth expansion different, its around two-sided Gaussian, includes non-negative values of *t* also.

arXiv:1604.05513 [physics.data-an]

## **EXAMPLE, LÉVY EXPANSIONS**

### Model-independent but:

- Levy generalizes exponentials and Gaussians
- ubiquitous in nature
- How far from a Lévy?
- Not necessarily positive definit !

$$\begin{split} L_0(t \mid \alpha) &= \frac{1}{\alpha} \Gamma\left(\frac{1}{\alpha}\right), \\ L_1(t \mid \alpha) &= \frac{1}{\alpha} \left\{ \Gamma\left(\frac{1}{\alpha}\right) t - \Gamma\left(\frac{2}{\alpha}\right) \right\}, \\ L_2(t \mid \alpha) &= \frac{1}{\alpha^2} \left\{ \left[ \Gamma\left(\frac{1}{\alpha}\right) \Gamma\left(\frac{3}{\alpha}\right) - \Gamma^2\left(\frac{2}{\alpha}\right) \right] t^2 - \right. \\ &\left. - \left[ \Gamma\left(\frac{1}{\alpha}\right) \Gamma\left(\frac{4}{\alpha}\right) - \Gamma\left(\frac{3}{\alpha}\right) \Gamma\left(\frac{2}{\alpha}\right) \right] t + \right. \\ &\left. + \left[ \Gamma\left(\frac{2}{\alpha}\right) \Gamma\left(\frac{4}{\alpha}\right) - \Gamma^2\left(\frac{3}{\alpha}\right) \right] \right\}. \end{split}$$



Lévy polynomials of first and third order times the weight function  $e^{-x^{\alpha}}$  for  $\alpha = 0.8, 1.0, 1.2, 1.4$ .

1st-order Lévy polynomial  $\gamma \left[ 1 + \lambda e^{-R^{\alpha}Q^{\alpha}} [1 + c_1 L_1(Q|\alpha, R)] \right]$ 3rd-order Lévy polynomial  $\gamma \left[ 1 + \lambda e^{-R^{\alpha}Q^{\alpha}} [1 + c_1 L_1(Q|\alpha, R) + c_3 L_3(Q|\alpha, R)] \right]$ M. de Kock, H. C. Eggers, T. Cs: arXiv:1206.1680v1 [nucl-th]

#### **BE correlation function** Example

FIT7  $\chi^2$ /DoF=91.3/93 CL=52% .0<Q<4.0, .10.10< $m_t$ <4.02 4.02 4.02,  $\Delta m_t$ <9.0 this plot: 14.14< $m_t$ <4.02 4.02  $\Delta m_t$ <9.00  $\chi^2$ =93.6, 100 pts

FIT7  $\chi^2$ /DoF=170.4/94 CL= 2.5E-06 .0<Q<4.0, .10.10< $m_i$ <4.024.02,  $\Delta m_i$ <9.0 this plot: .14.14< $m_i$ <4.024.02  $\Delta m_i$ <9.00  $\chi^2$ =172.4, 100 pts

 $\alpha = 2.000 \pm .000$ 

 $R = .740 \pm .024$ 

 $\lambda = .507 \pm .019$ 

 $\gamma = .954 \pm .002$ 

 $\delta = .015 \pm .001$ 

з

c1=.073±.078

c=5.862±.686



## HB et al. observed oscillations

Mon. Not. R. astr. Soc. (1974) 167, 121-136.

#### THE ANGULAR DIAMETERS OF 32 STARS

#### R. Hanbury Brown, J. Davis and L. R. Allen

The normalized correlation also depends upon whether a star is single or multiple. It was shown in Paper II that, if a star is binary and the angular separation of the two components is completely resolved by the interferometer at the shortest baseline, then the normalized zero-baseline correlation  $\overline{c_N(0)}'$  averaged over a range of position angles is reduced relative to a single star  $\overline{c_N(0)}$  by the factor,

$$\overline{c_{\rm N}({\rm o})'}/\overline{c_{\rm N}({\rm o})} = (I_1^2 + I_2^2)/(I_1 + I_2)^2 \tag{9}$$

where  $I_1$ ,  $I_2$  are the brightness of the two components. It is simple to extend this analysis to a multiple star with n components and to show that, if the angular separation between all the components is resolved, the zero-baseline correlation is reduced relative to a single star by the factor,

$$\overline{c_{\mathrm{N}}(\mathrm{o})'/c_{\mathrm{N}}(\mathrm{o})} = \sum_{n} I^{2} / \left(\sum_{n} I\right)^{2}.$$
 (10)

It follows that if a star yields a correlation which is significantly less than that expected from a single star, then it must be multiple.

## LÉVY EXPANSIONS for POSITIVE DEFINIT FORMS

experimental conditions:

(i) The correlation function tends to a constant for large values of the relative momentum Q.

(ii) The correlation function deviates from its asymptotic, large Q value in a certain domain of its argument.

(iii) The two-particle correlation function is related to a Fourier transformed space-time distribution of the source.

### **Model-independent but:**

- Assumes that Coulomb can be corrected
- No assumptions about analyticity yet
- For simplicity, consider 1d case first
- For simplicity, consider factorizable x k
- Normalizations :
  - density
  - multiplicity
  - single-particle spectra

T. Cs, S. Hegyi, W.A. Zajc, EPJ C36, 67 (2004)

$$C_2(\mathbf{k}_1, \mathbf{k}_2) = \frac{N_2(\mathbf{k}_1, \mathbf{k}_2)}{N_1(\mathbf{k}_1) N_1(\mathbf{k}_2)}$$

$$S(x,k) = f(x) g(k)$$

$$\int \mathrm{d}x \, f(x) \,=\, 1, \qquad \qquad \int \mathrm{d}k \, g(k) = \langle n \rangle,$$

$$N_1(k) = \int \mathrm{d}x \, S(x,k) = g(k).$$

## **MINIMAL MODEL ASSUMPTION: LÉVY**

### **Model-independent but:**

- not assumes analyticity
- C<sub>2</sub> measures a modulus squared
   Fourier-transform vs relative momentum
- Correlations non-Gaussian
- Radius not a variance
- $0 < \alpha \leq 2$

$$C_2(k_1, k_2) = 1 + |\tilde{f}(q_{12})|^2,$$

$$\tilde{f}(q_{12}) = \int \mathrm{d}x \, \exp(\mathrm{i}q_{12}x) \, f(x),$$

$$C(q;\alpha) = 1 + \lambda \exp\left(-|qR|^{\alpha}\right).$$

T. Cs, S. Hegyi, W.A. Zajc, EPJ C36, 67 (2004)

## UNIVARIATE LEVY EXAMPLES

Include some well known cases:

•  $\alpha = 2$ 

•  $\alpha = 1$ 

• Gaussian source, Gaussian C<sub>2</sub>

• Lorentzian source, exponential C<sub>2</sub>

$$f(x) = \frac{1}{(2\pi R^2)^{1/2}} \exp\left[-\frac{(x-x_0)^2}{2R^2}\right]$$
$$C(q) = 1 + \exp\left(-q^2 R^2\right)$$

$$f(x) = \frac{1}{\pi} \frac{R}{R^2 + (x - x_0)^2},$$
  
$$C(q) = 1 + \exp(-|qR|).$$

- asymmetric Lévy:
  - asymmetric support
  - Streched exponential

$$f(x) = \sqrt{\frac{R}{8\pi}} \frac{1}{(x - x_0)^{3/2}} \exp\left(-\frac{R}{8(x - x_0)}\right)$$
$$x_0 < x < \infty,$$
$$C(q) = 1 + \exp\left(-\sqrt{|qR|}\right).$$

T. Cs, hep-ph/0001233, T. Cs, S. Hegyi, W.A. Zajc, EPJ C36, 67 (2004)

# Multivariate, nearly Lévy correlations

## If the BE correlation function is $C_2(k_1,k_2) = 1 + |\tilde{f}(q_{12})|^2$ , then

$$t = \left(\sum_{i,j=\text{side,out,long}} R_{i,j}^2 q_i q_j\right)^{1/2},$$
  
$$C_2(t) = N \left\{ 1 + \lambda \exp(-t^{\alpha}) \left| 1 + \sum_{n=1}^{\infty} (a_n + ib_n) L_n(t|\alpha) \right|^2 \right\}$$

where  $\{c_n = a_n + ib_n\}_{n=1}^{\infty}$  are now complex valued expansion coefficients,

# Lévy expansion fit to pp elastic scattering at sqrt(s) = 7 TeV



# **Towards imaging**

$$C(q) = 1 + \exp(-(qR)^{\alpha}) = 1 + |\tilde{f}(q)|$$

$$\tilde{f}(q) = \exp\left(-\frac{1}{2}(qR)^{\alpha}\right) \to \tilde{f}(q) = \exp\left(-\frac{1}{2}(qR)^{\alpha}\right)\left(\sum_{j} c_{j}L_{j}(q|\alpha)\right)$$

$$c_j = a_j + ib_j$$

$$f(r) = \frac{1}{2\pi} \int \mathrm{d}q e^{-iqr} \tilde{f}(q)$$

$$f(r) = \frac{1}{2\pi} \int dq e^{-iqr} \exp\left(-\frac{1}{2}(qR)^{\alpha}\right) \left(\sum_{j} d_{j} \frac{q^{j}}{j!}\right)$$
$$f(r) = \frac{1}{2\pi} \sum_{j} \frac{d_{j}}{j!} \int dq e^{-iqr} q^{j} \exp\left(-\frac{1}{2}(qR)^{\alpha}\right)$$
$$f(r) = \frac{1}{2\pi} \sum_{j} \frac{d_{j}}{j!} (-i)^{j} \left(\frac{d}{dr}\right)^{j} \operatorname{LS}\left(\frac{r}{R},\alpha\right)$$

$$\sum_{j} c_{j} L_{j}(q|\alpha) = \sum_{j} d_{j} \frac{q^{j}}{j!}$$
$$d_{j} = \left(\frac{\mathrm{d}}{\mathrm{d}q}\right)^{j} \left(\sum_{j} c_{j} L_{j}(q|\alpha)\right) |_{q=0}$$

## **SUMMARY AND CONCLUSIONS**

### **Several model-independent methods:**

- Based on matching an abstract measure in *H* to the approximate shape of data
- Gaussian: Edgeworth expansions
- Exponential: Laguerre expansions
- Lévy (0 <  $\alpha \leq 2$ ): Lévy expansions
- Lévy expansions for positive definit functions

## Thank you for your attention!