

# Dynamics and Symmetry Energy at Intermediate Energies

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Particle Correlations and Femtoscopy

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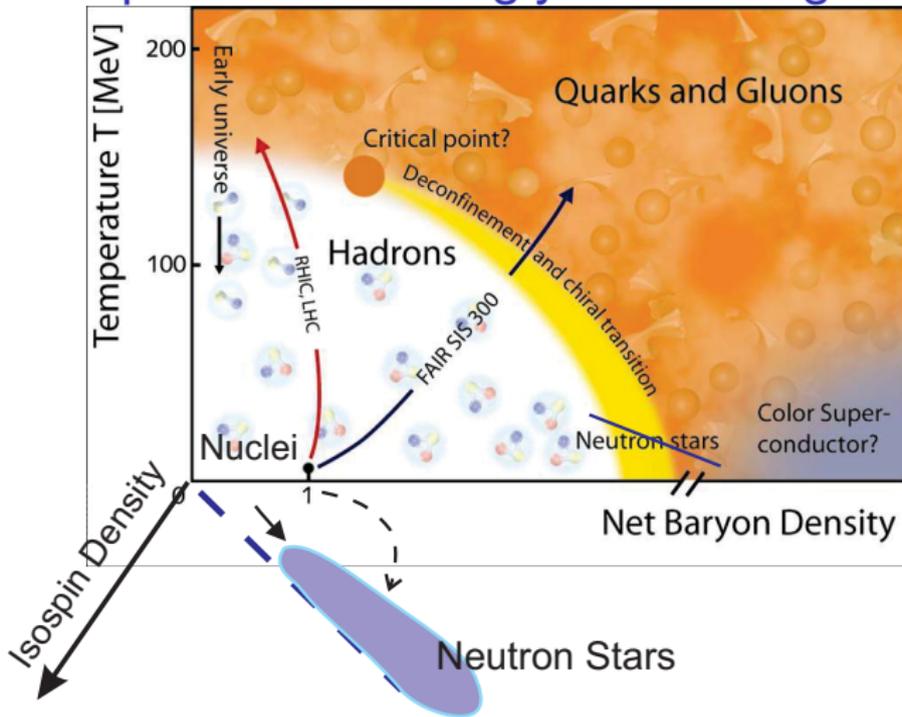


# Outline

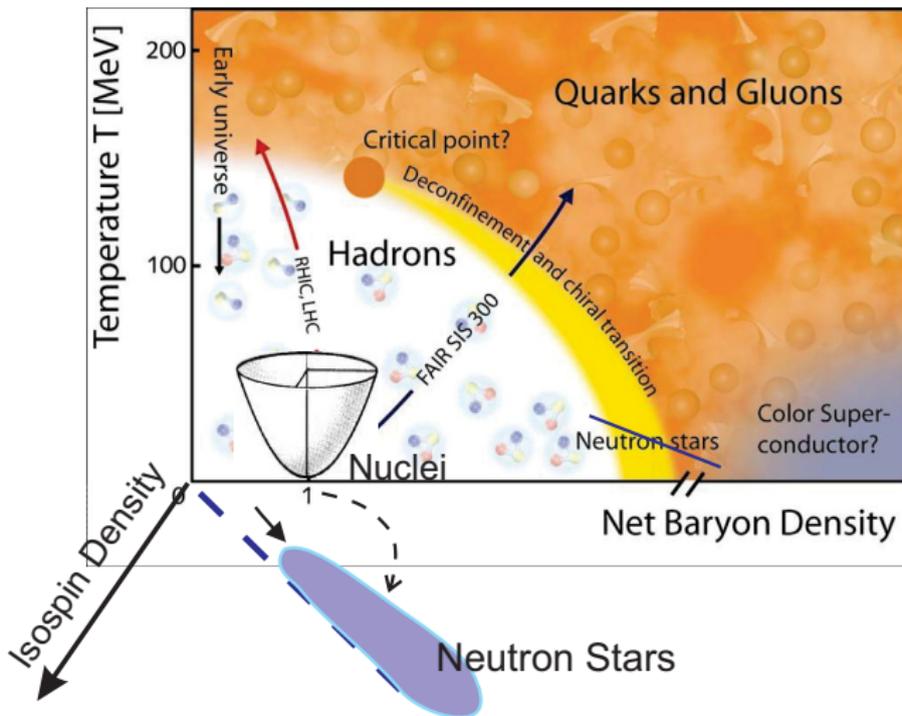
- 1 Introduction
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- 3 n-Star Merger
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- 5 Comparison Project
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# Bulk Properties of Strongly-Interacting Matter



# Equation of State



Reactions - coarse, but all  $\rho$ .

Structure -  $\rho < \rho_0$  only



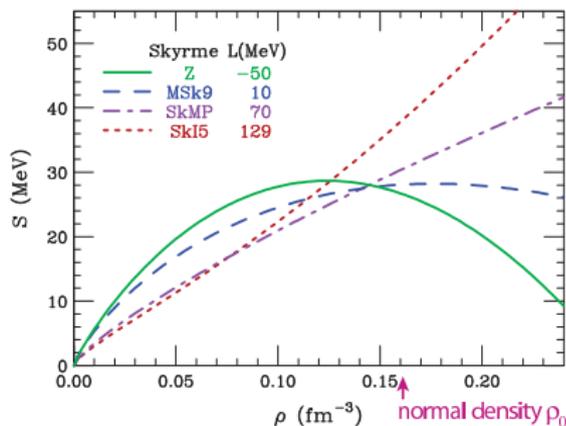
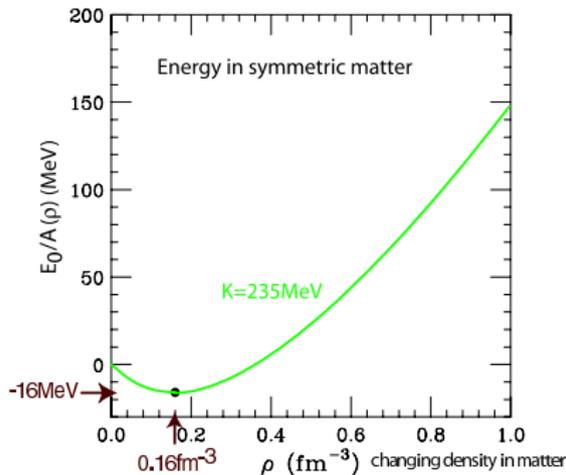
# Energy in Uniform Matter

$$\frac{E}{A}(\rho_n, \rho_p) = \frac{E_0}{A}(\rho) + S(\rho) \left( \frac{\rho_n - \rho_p}{\rho} \right)^2 + \mathcal{O}(\dots^4)$$

symmetric matter

(a)symmetry energy

$$\rho = \rho_n + \rho_p$$



$$\frac{E_0}{A}(\rho) = -a_v + \frac{K}{18} \left( \frac{\rho - \rho_0}{\rho_0} \right)^2 + \dots$$

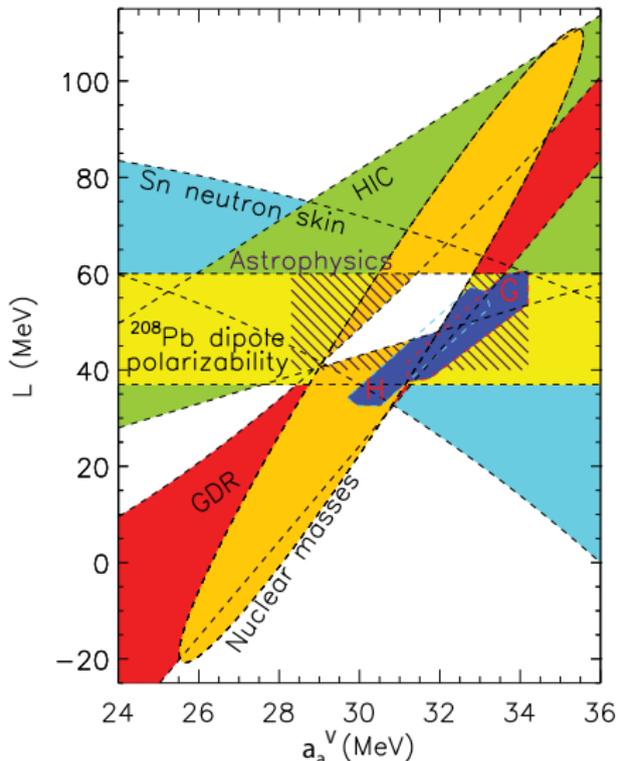
Known:  $a_a \approx 16$  MeV  $K \sim 235$  MeV

$$S(\rho) = a_a^V + \frac{L}{3} \frac{\rho - \rho_0}{\rho_0} + \dots$$

Unknown:  $a_a^V?$   $L?$



# Importance of Slope



Lattimer&Lim *ApJ*771(2013)51

$$\frac{E}{A} = \frac{E_0}{A}(\rho) + S(\rho) \left( \frac{\rho_n - \rho_p}{\rho} \right)^2$$

$$S \simeq a_a^V + \frac{L}{3} \frac{\rho - \rho_0}{\rho_0}$$

In neutron matter:

$$\rho_p \approx 0 \text{ \& \ } \rho_n \approx \rho.$$

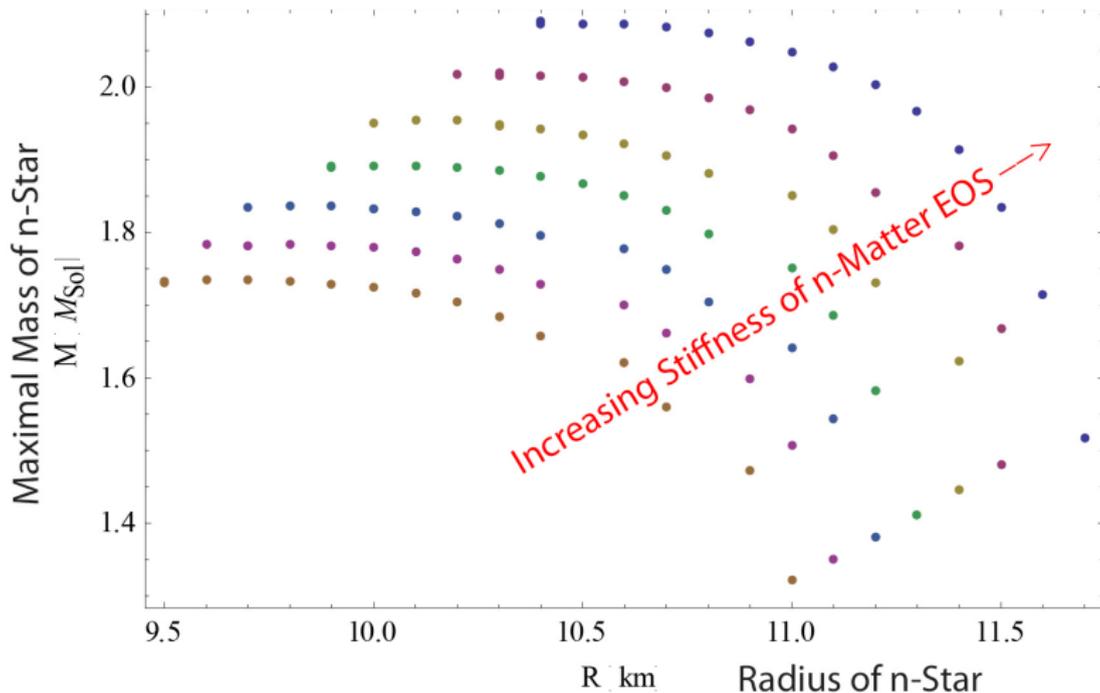
$$\text{Then, } \frac{E}{A}(\rho) \approx \frac{E_0}{A}(\rho) + S(\rho)$$

Pressure:

$$P = \rho^2 \frac{d}{d\rho} \frac{E}{A} \simeq \rho^2 \frac{dS}{d\rho} \simeq \frac{L}{3\rho_0} \rho^2$$



# Both Radius & Max Mass Increase w/Stiffness



Schematic Calculation by Stephen Portillo (Harvard U)



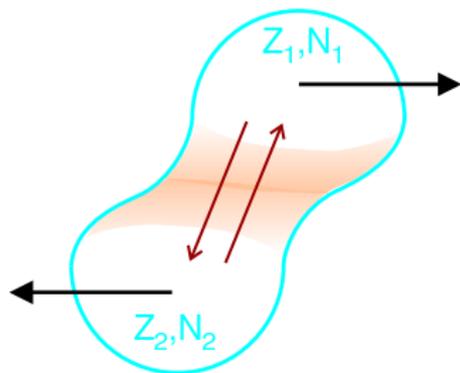
## n-p Asymmetry Diffusion

Different projectile-target combinations out of  $^{112}_{50}\text{Sn}$  and  $^{124}_{50}\text{Sn}$  at 50 MeV/nucl

Tsang *et al*

$$\delta_1 = \frac{N_1 - Z}{A_1} = 0.11$$

$$\delta_2 = \frac{N_2 - Z}{A_2} = 0.19$$



Irreversible flux of n-p asymmetry, according to Fick's law:

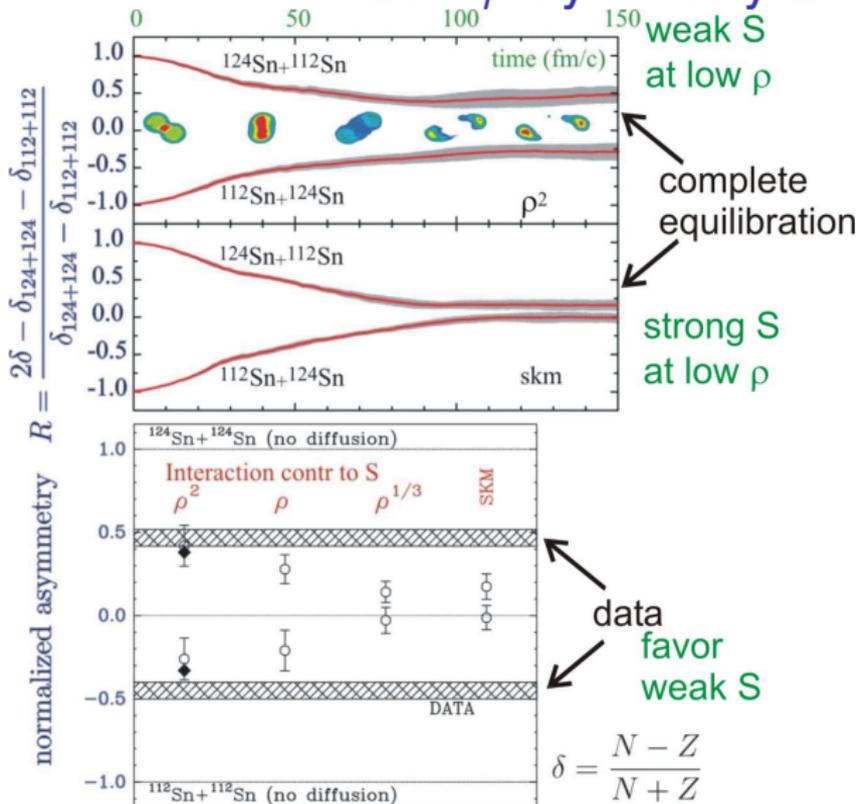
$$\vec{j}_{np} = \vec{j}_n - \vec{j}_p = -\nu \vec{\nabla} [\mu_n - \mu_p] \simeq -4\nu \vec{\nabla} \left[ S \frac{\rho_n - \rho_p}{\rho} \right]$$

where  $\nu > 0$ , independent of the symmetry energy  $S$ .  
Analog of the electric conductivity eq:  $\vec{j} = \sigma \vec{E} = -\sigma \vec{\nabla} \Phi$ .

Shi&PD



# Low- $\rho$ Symmetry Energy



Linear theory:  
 $\vec{J}_{np} \simeq -4\nu \vec{\nabla} [S \delta]$

Transport simulations on n-p asymmetry equilibration at velocities close to a Sn projectile, by Shi (*Tsang et al*)

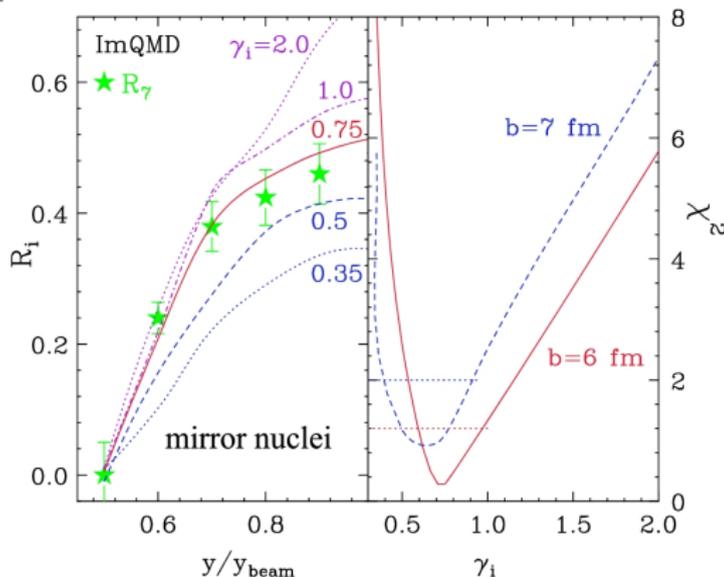
$\Rightarrow$  slow change of S with  $\rho$  eliminated



# Impact-Parameter Filtered Isospin-Equilibration

Tsang, Zhang *et al*

Equilibration ratio  $R$  from measured yields of  $A = 7$  mirror nuclei, compared to ImQMD



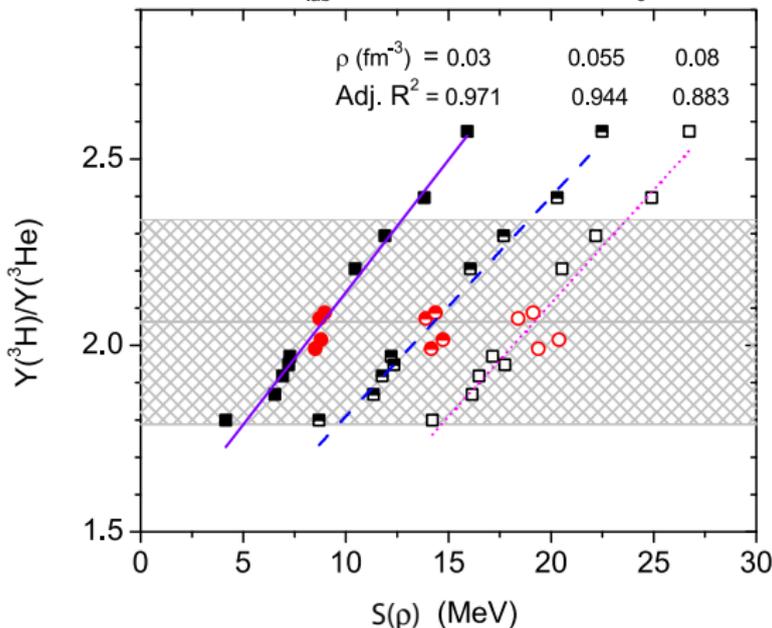
$$S(\rho) = 12.3 \text{ MeV} (\rho/\rho_0)^{2/3} + 17.6 \text{ MeV} (\rho/\rho_0)^{\gamma_i}$$



## ${}^3\text{H}/{}^3\text{He}$ Yield Ratio

Clusters form at moderate  $\rho$ . Stiffness of  $S(\rho)$  decides how  $n-p$  imbalance gets partitioned across densities & what yield ratios of clusters with different isospin become

Au+Au  $E_{\text{lab}}=400$  MeV/nucleon  $b_0 < 0.15$

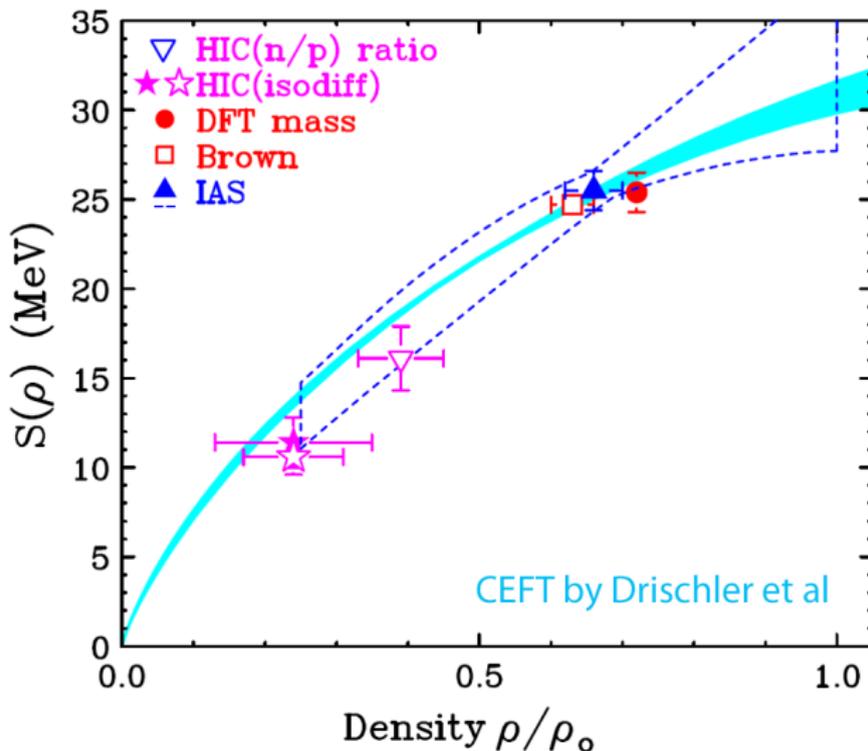


UrQMD +  
phase-space  
coalescence results by  
Yongjia Wang,  
Qingfeng Li *et al* for  
different Skyrme ints,  
plotted vs  
symmetry-energy  
values at different  $\rho$ ,  
vs FOPI results from  
Au+Au collisions

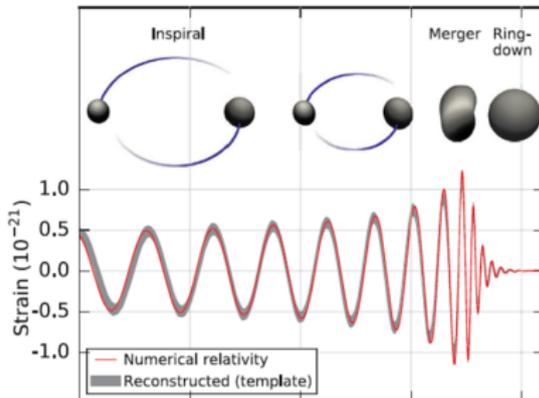
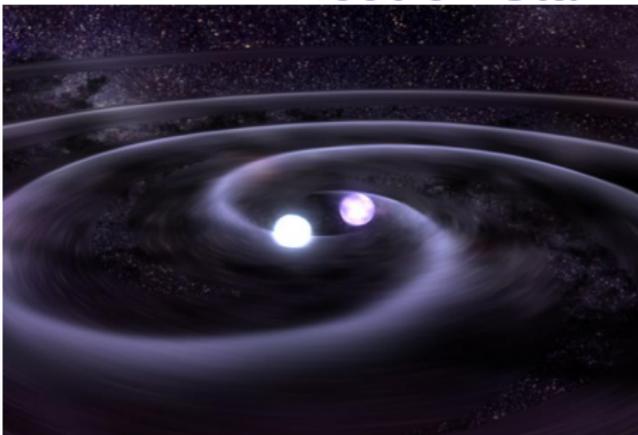


# Low- $\rho$ Summary of Symmetry Energy

from work of Tsang & Lynch



# Neutron Star Merger Event

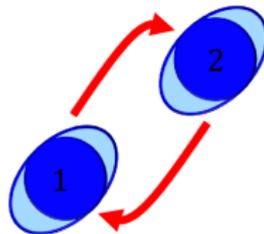


LIGO signal amenable to perturbative analysis -  
*t*-resolution insufficient for late stage.

First-order point masses, radiated power  
 $P \propto Q^2$ , where  $Q$  - quadrupole moment.

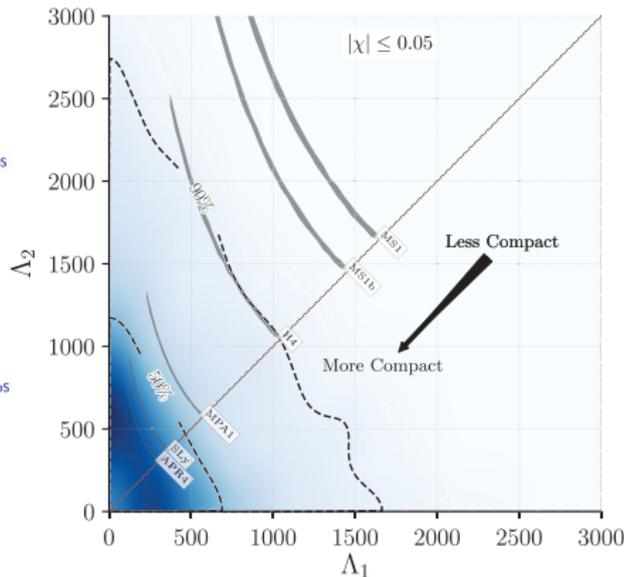
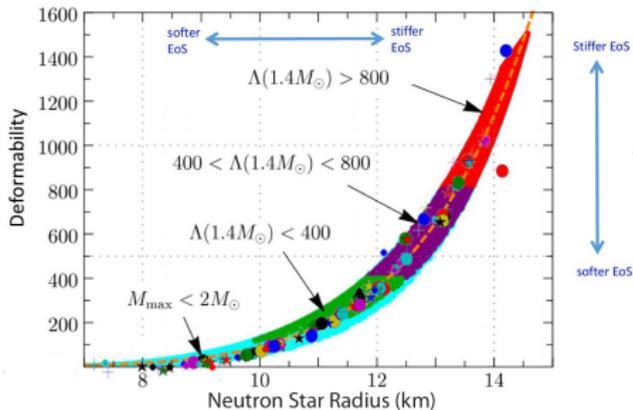
Next order distortion of extended masses  
 matters,  $Q_{ij} = \Lambda \partial^2 V / \partial r_i \partial r_j$ , where  $\Lambda$  - tidal  
 deformability

⇒ No significant additional contribution seen!



# Implications for Stars & EOS

Large  $\Lambda \Leftrightarrow$  large star

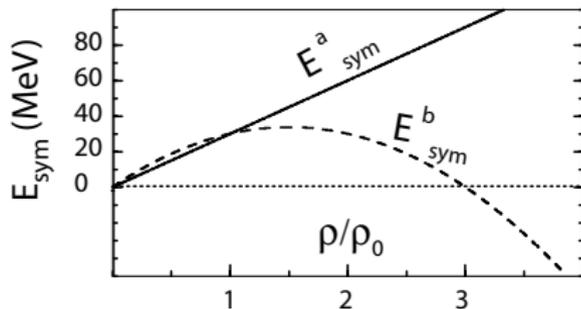


$\Rightarrow$  Largest deformabilities/largest stars/stiffest EOS excluded by merger event!

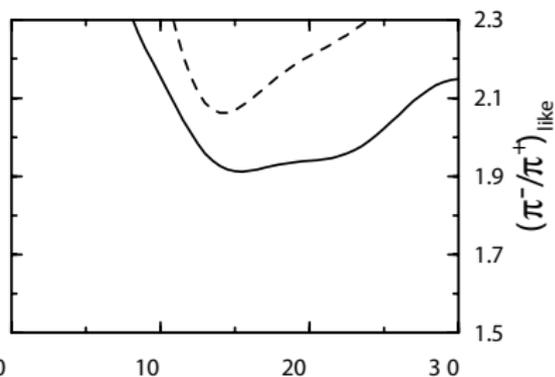
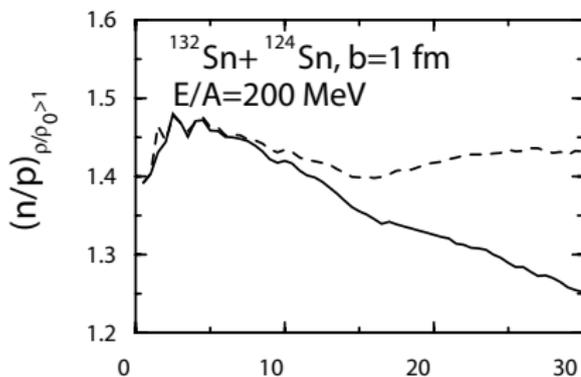


# Pions as Probe of High- $\rho$ Symmetry Energy

B-A Li PRL88(02)192701:  $S(\rho > \rho_0) \Rightarrow n/p_{\rho > \rho_0} \Rightarrow \pi^-/\pi^+$



Pions originate from high  $\rho$



$t$  (fm/c)

## Dedicated Experimental Efforts

**SAMURAI-TPC Collaboration** (data taken; 8 countries and 43 researchers): comparisons of near-threshold  $\pi^-$  and  $\pi^+$  and also  $n$ - $p$  spectra and flows at RIKEN, Japan.

NSCL/MSU, Texas A&M U

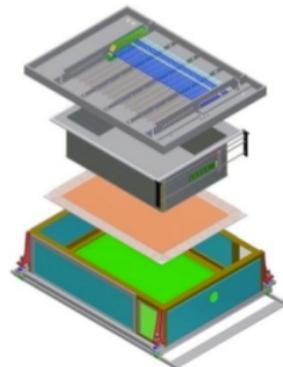
Western Michigan U, U of Notre Dame

GSI, Daresbury Lab, INFN/LNS

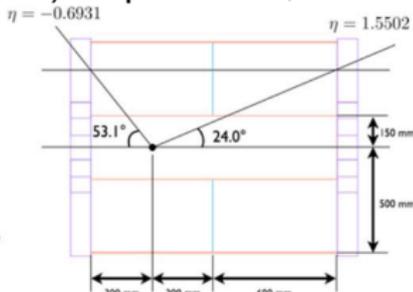
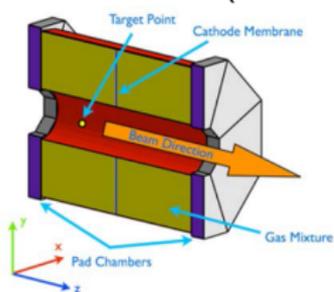
U of Budapest, SUBATECH, GANIL

China IAE, Brazil, RIKEN, Rikkyo U

Tohoku U, Kyoto U

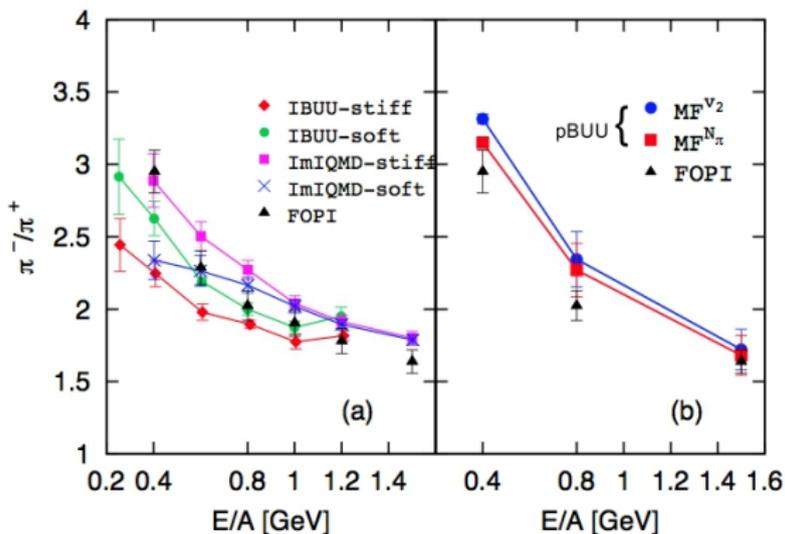


**LAMPS TPC at RAON** (S Korea): triple GEM,  $3\pi$  sr



# FOPI Au+Au $\pi^-/\pi^+$ Data?

Reisdorf *et al.* (FOPI) NPA781(07)459



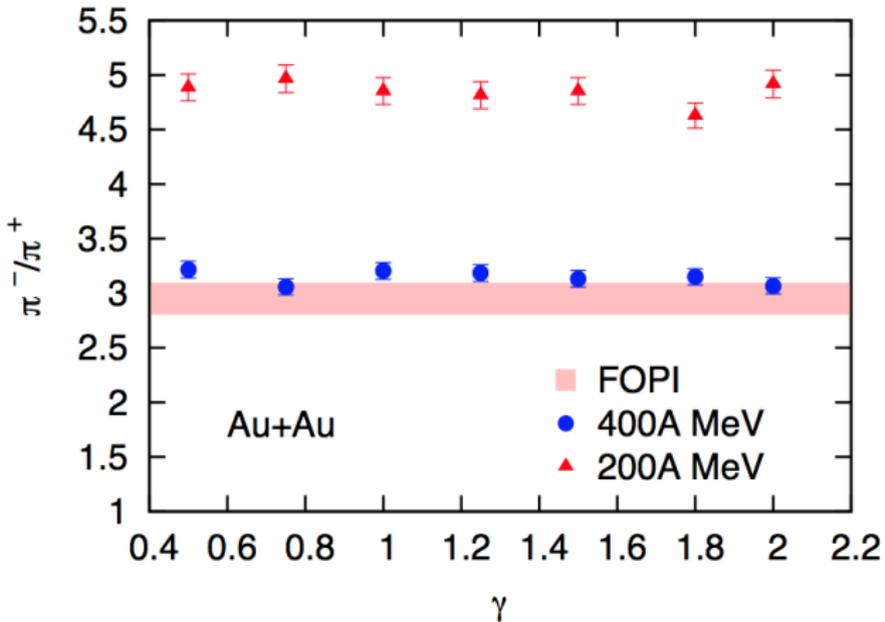
data: black symbols

theory: colored symbols

Opposing claims on  $S(\rho)$  needed to explain data!

# FOPI $\pi^-/\pi^+$ Reproduced by pBUU

... irrespectively of  $S_{int}(\rho) = S_0 (\rho/\rho_0)^\gamma$ :



Jun Hong & PD PRC90(14)024605

... Other probes possible, but general problem of model ambiguity remains!



# Code Comparison Project

BUU type	Code correspondents	Energy range	QMD type	Code correspondents	Energy range
BLOB	P. Napolitani, M. Colonna	0.01 0.5	AMD	A. Ono	0.01 0.3
GIBUU-RMF	J. Weil	0.05 40	IQMD-BNU	J. Su, F. S. Zhang	0.05 2
GIBUU-Skyrme	J. Weil	0.05 40	IQMD	C. Hartnack, J. Aichelin	0.05 2
IBL	W. J. Xie, F. S. Zhang	0.05 2	CoMD	M. Papa	0.01 0.3
IBUU	J. Xu, L. W. Chen, B. A. Li	0.05 2	ImQMD-CIAE	Y. X. Zhang, Z. X. Li	0.02 0.4
pBUU	P. Danielewicz	0.01 12	IQMD-IMP	Z. Q. Feng	0.01 10
RBUU	K. Kim, Y. Kim, T. Gaitanos	0.05 2	IQMD-SINAP	G. Q. Zhang	0.05 2
RVUU	T. Song, G. Q. Li, C. M. Ko	0.05 2	TuQMD	D. Cozma	0.1 2
SMF	M. Colonna, P. Napolitani	0.01 0.5	UrQMD	Y. J. Wang, Q. F. Li	0.05 200

Leaders in the effort:

Jorg Aichelin, Evgeni Kolomeitsev, Betty Tsang + others

Jun Xu *et al.* PRC93(16)044609, Yingxun Zhang *et al.* PRC97(18)034625

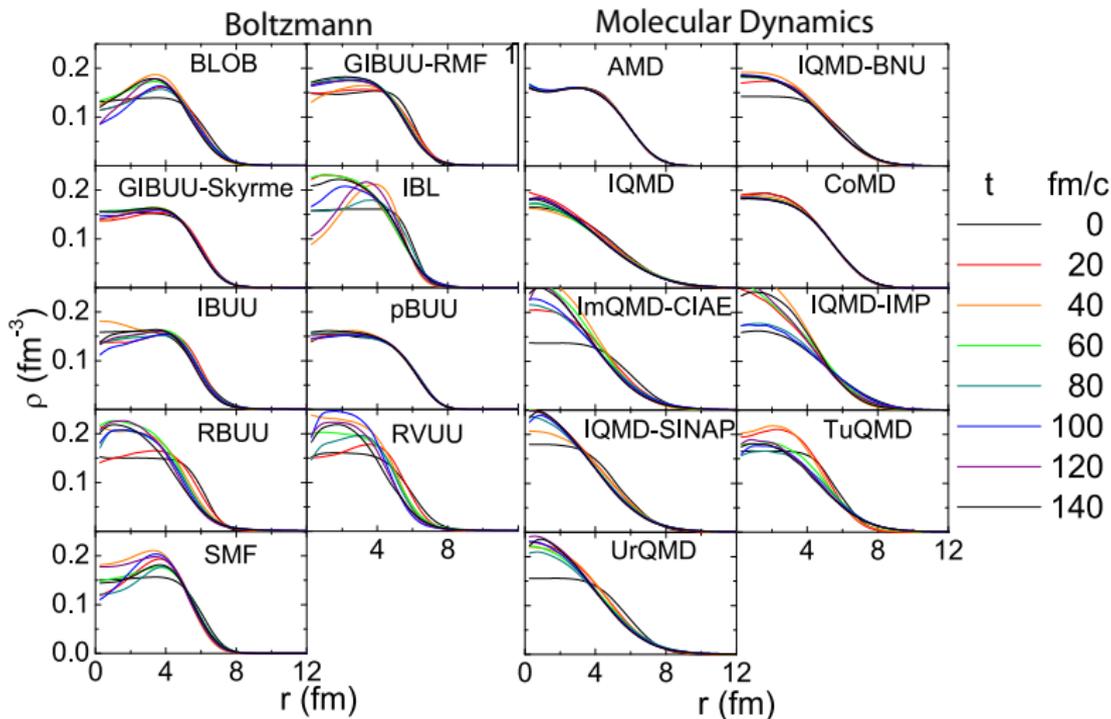


# Premise

- Specify the same physics inputs for different transport codes
- Compare outputs
  - elastic collisions only
  - constant isotropic cross section  $\sigma = 40$  mb
  - soft EOS + momentum-independent mean-field
- Full-run comparisons
- Controlled simplified conditions
  - \* collisions in a box ← approach to equilibrium
  - \* mean field in a box
  - \* Next:  $\Delta + \pi$  production in a box. . .



# Full Runs: Stability of Initial Density



Jun Xu *et al.* PRC93(16)044609

Isolated Au nucleus

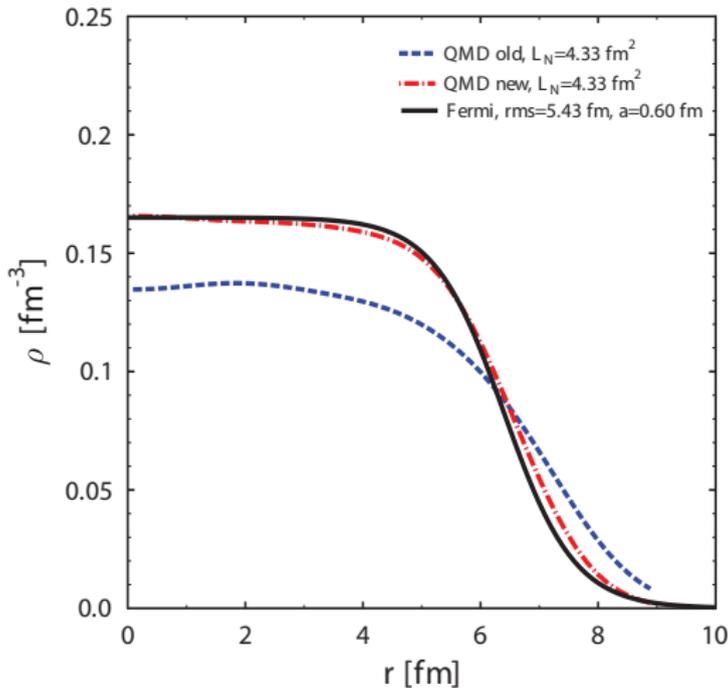
⇒ Initial state must be constructed consistently with dynamics



# Example: Rebuilt TuQMD

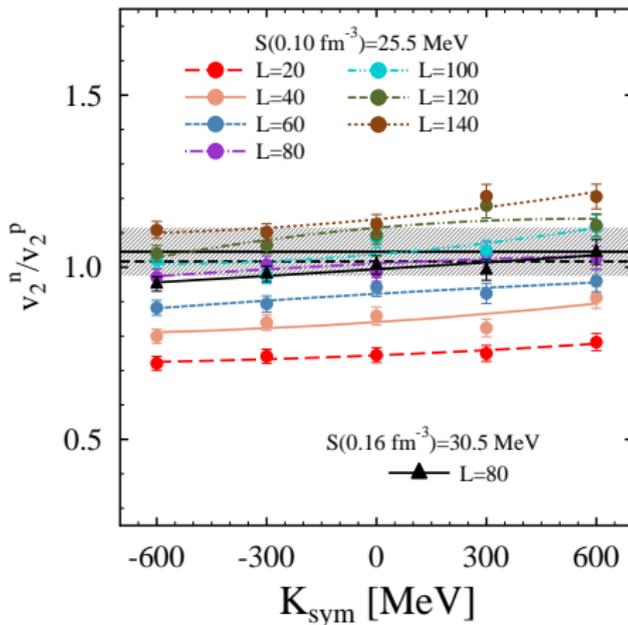
Dan Cozma arXiv: 1706.01300

Rebuilt density initializations and Pauli principle



# FOPI-LAND & ASYEOS Elliptic-Flow Data

Data Cozma PRC88(13)044912

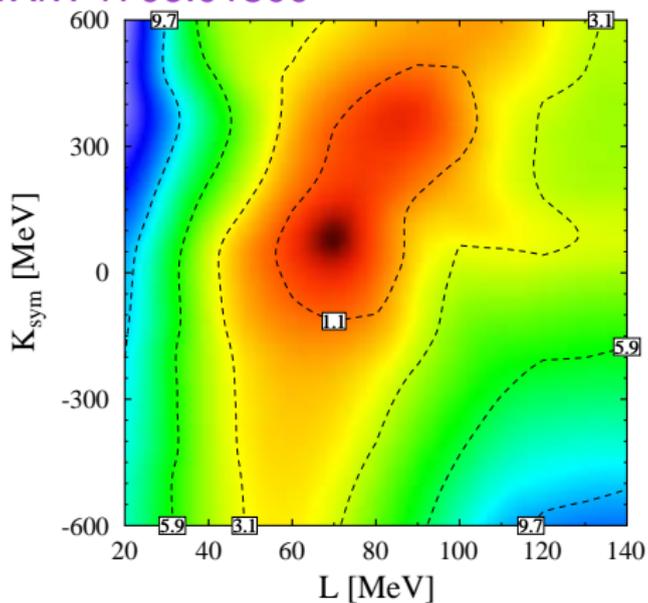


400 MeV/mucl Au + Au data above + other, particularly more differential



# Constraints on Symmetry Energy Parameters

Dan Cozma arXiv: 1706.01300



Linear slope parameter  $L$  & curvature  $K_{\text{sym}}$  vs density



# Conclusions

- Convergence on symmetry-energy conclusions at  $\rho \lesssim \rho_0$ . Slope parameter  $L$  at  $\rho_0$  still elusive.
- Neutron-star merger constrains stiffness of EOS from above.
- Code comparison project aims at improving firmness of conclusions drawn from comparing data to transport. First benefits begin to emerge.

