

# Non-identical particle femtoscopy at STAR

Paweł Szymański (for the STAR collaboration)

Warsaw University  
of Technology



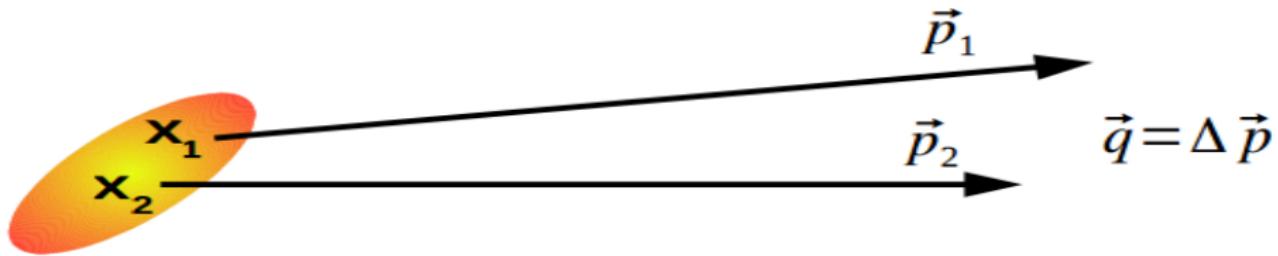
XIII Workshop on Particle Correlation and Femtoscopy  
Cracow, 22.05.18

## HBT interferometry (Hanbury-Brown, Twiss)

Intensity interferometry → allows to study size of the emitting source by measuring a momentum distribution of emitted particles.  
Originally this effect was used to study properties of stars

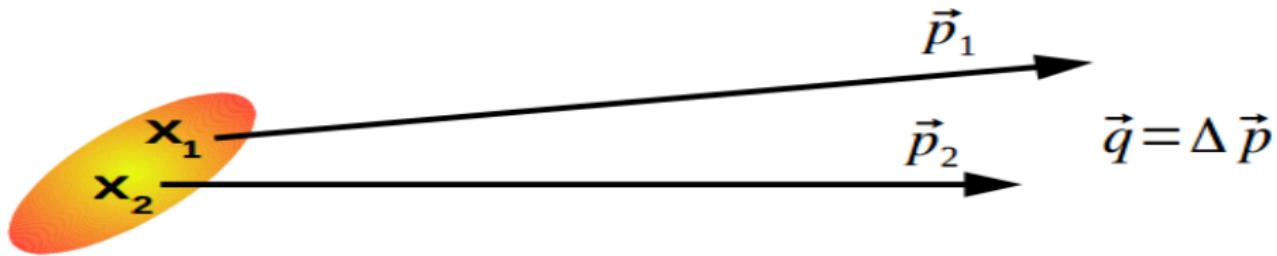
We can use a two-particle correlation to measure one of the smallest sizes in the nature ( $\sim 10^{-15} m$ )

# Correlation Function



- analyze many pairs of particles  $(\vec{p}_1, \vec{x}_1)$  and  $(\vec{p}_2, \vec{x}_2)$  with relative momentum  $\vec{q} = \vec{p}_1 - \vec{p}_2$

# Correlation Function



- analyze many pairs of particles  $(\vec{p}_1, \vec{x}_1)$  and  $(\vec{p}_2, \vec{x}_2)$  with relative momentum  $\vec{q} = \vec{p}_1 - \vec{p}_2$
- calculate correlation function (CF) of pairs:

$$CF(\vec{p}_1, \vec{p}_2) = \frac{P_2(\vec{p}_1, \vec{p}_2)}{P_1(\vec{p}_1)P'_1(\vec{p}_2)}$$

$P_2(\vec{p}_1, \vec{p}_2)$  — probability of observing two particles with momentum  $\vec{p}_1$  and  $\vec{p}_2$  at the same time and the same place

$P_1(\vec{p}_1), P'_1(\vec{p}_2)$  — probability of observing two particles with momentum  $\vec{p}_1$  and  $\vec{p}_2$  separately

# Correlation Function



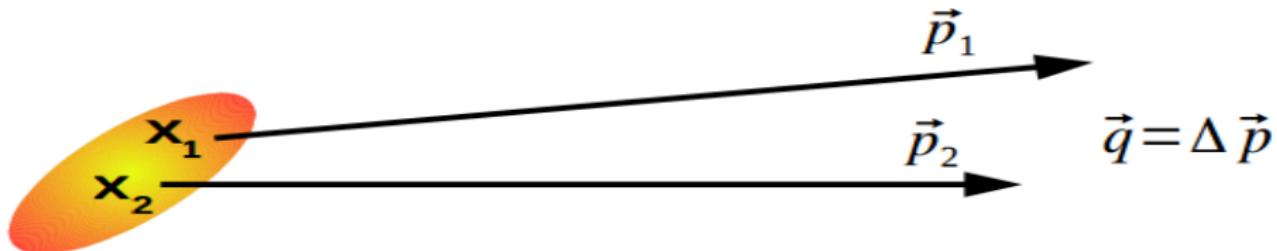
- analyze many pairs of particles  $(\vec{p}_1, \vec{x}_1)$  and  $(\vec{p}_2, \vec{x}_2)$  with relative momentum  $\vec{q} = \vec{p}_1 - \vec{p}_2$
- calculate correlation function (CF) of pairs:

$$CF(\vec{p}_1, \vec{p}_2) = \frac{P_2(\vec{p}_1, \vec{p}_2)}{P_1(\vec{p}_1)P'_1(\vec{p}_2)}$$

experimental  
correlation function:

$$CF(\vec{q}) = \frac{A(\vec{q})}{B(\vec{q})}$$

# Correlation Function



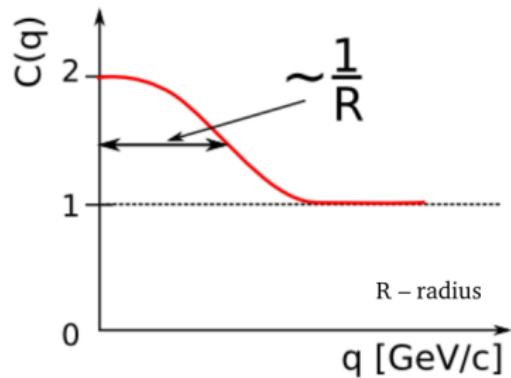
- analyze many pairs of particles  $(\vec{p}_1, \vec{x}_1)$  and  $(\vec{p}_2, \vec{x}_2)$  with relative momentum  $\vec{q} = \vec{p}_1 - \vec{p}_2$
- calculate correlation function (CF) of pairs:

$$CF(\vec{p}_1, \vec{p}_2) = \frac{P_2(\vec{p}_1, \vec{p}_2)}{P_1(\vec{p}_1)P'_1(\vec{p}_2)}$$

- calculate size of the source

$P_2(\vec{p}_1, \vec{p}_2)$  — probability of observing two particles with momentum  $\vec{p}_1$  and  $\vec{p}_2$  at the same time and the same place

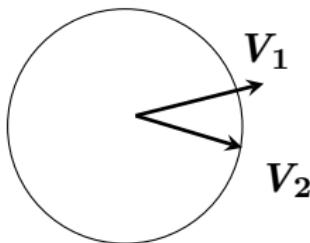
$P_1(\vec{p}_1), P'_1(\vec{p}_2)$  — probability of observing two particles with momentum  $\vec{p}_1$  and  $\vec{p}_2$  separately



# Non-identical particle combinations

Time asymmetry

$$t_1 \neq t_2 \\ \Delta r = 0$$



$t_1 > t_2$  - Catching up  
 $t_1 < t_2$  - Run away

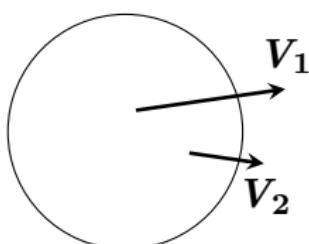
$t$  — emission time

$r$  — emission point distance from the center

R. Lednický, et al.,  
Phys. Lett. B373,  
30-34 (1996)

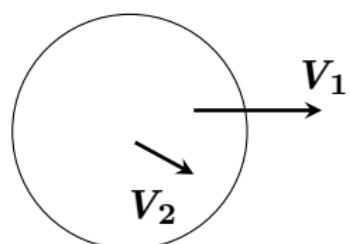
Space asymmetry

$$t_1 = t_2 \\ \Delta r \neq 0$$



Catching up

$$t_1 = t_2 \\ \Delta r \neq 0$$



Run away

**Catching up**  
longer interaction,  
strong correlation

**Running away**  
shorter interaction,  
weaker correlation

# Spherical Harmonics

$$C(\mathbf{q}) = \sum_{l,m} C_l^m(q) Y_l^m(\theta, \phi)$$

$$C_l^m(q) = \int_{\Omega} C(q, \theta, \phi) Y_l^m(\theta, \phi) d\Omega$$

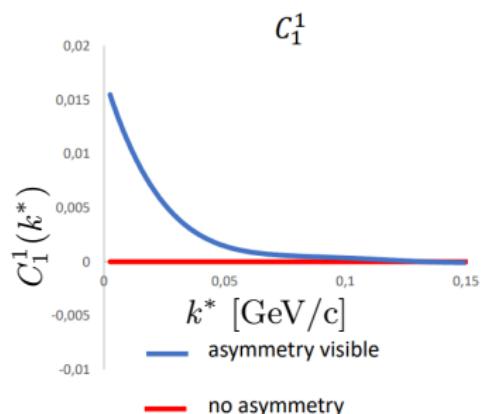
$\Omega$  - full solid angle

$Y_l^m(\theta, \phi)$  - spherical harmonic function

$q = |\mathbf{q}|$ ,  $\theta$  and  $\phi$  - spherical coordinates

$C_0^0 \rightarrow$  sensitive to the size of the emitting source  
(shapes same as correlation function)

$C_1^1 \rightarrow$  sensitive to the spacetime emission asymmetry



P. Danielewicz and S. Pratt.  
Phys. Lett B618: 60 2005

A. Kisiel and D.A. Brown  
Phys. Rev. C80:064911 2009

P. Danielewicz and S. Pratt.  
Phys. Rev. C75:034907 2007

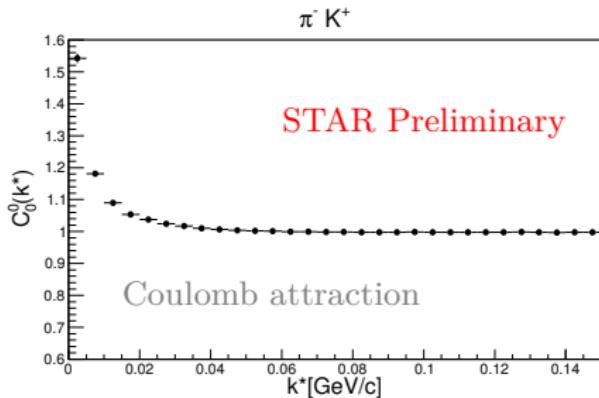
A. Kisiel  
Phys. Rev. C81:064906 2010

# Final State Interactions (FSI)

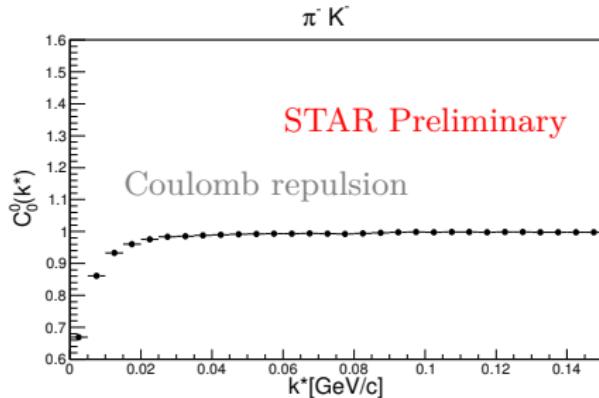
The shape of non-identical particle CF depends on FSI:

- Strong Interaction
- *Coulomb force*
- ~~Quantum Statistics effect~~

Correlation between unlike-sign pairs

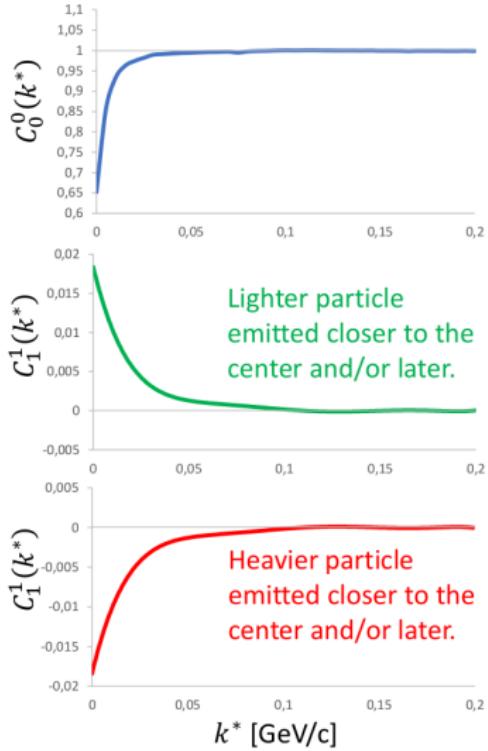


Correlation between like-sign pairs

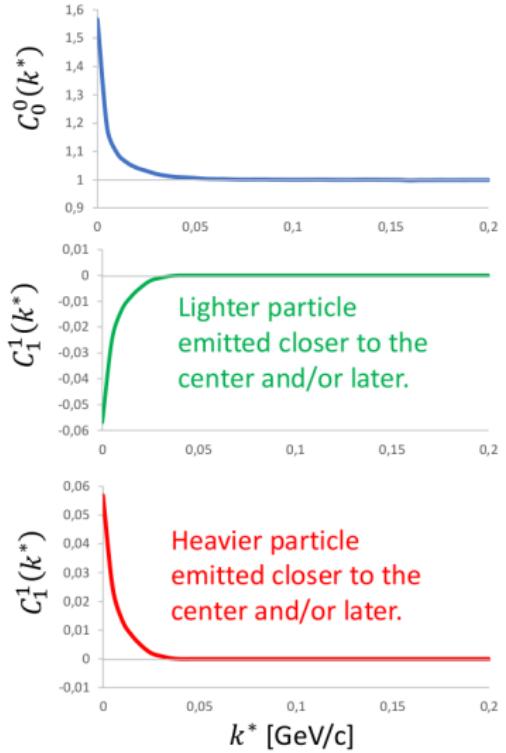


# Which particle...?

## Like-sign particle combinations



## Unlike-sign particle combinations

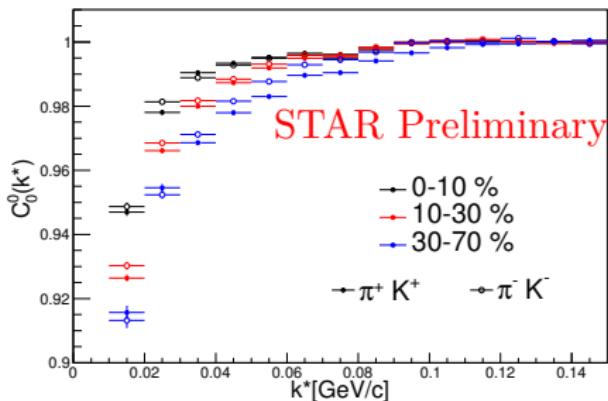


# Results

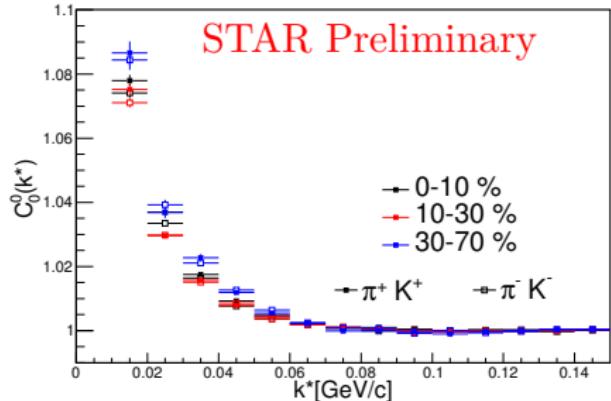
# Centrality dependence

$C_0^0 \rightarrow$  sensitive to the size of the emitting source

like-sign  $\pi - K$  @ Au+Au 39 GeV



unlike-sign  $\pi - K$  @ Au+Au 39 GeV



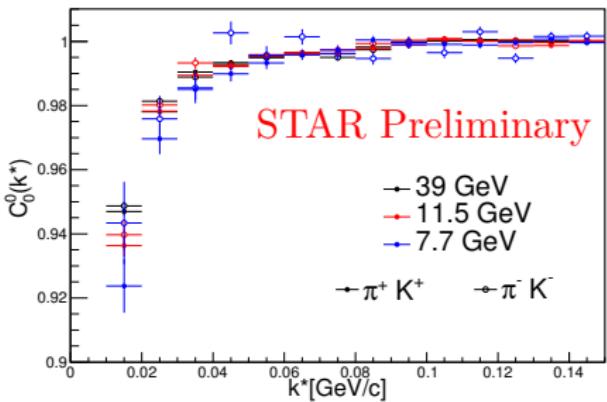
Clear centrality dependence

$R(0-10\%) > R(10-30\%) > R(30-70\%)$

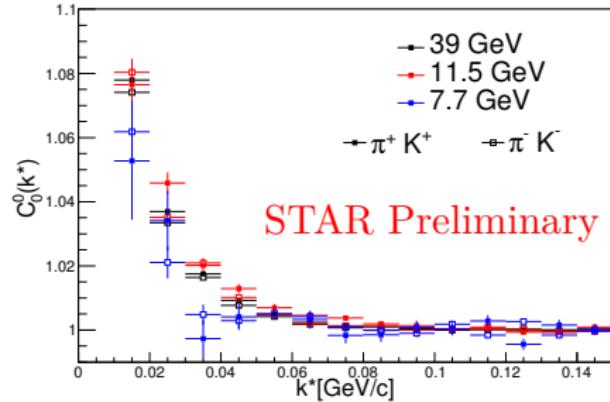
# Energy dependence

$C_0^0 \rightarrow$  sensitive to the size of the emitting source

like-sign  $\pi - K$ : 0-10%



unlike-sign  $\pi - K$ : 0-10%



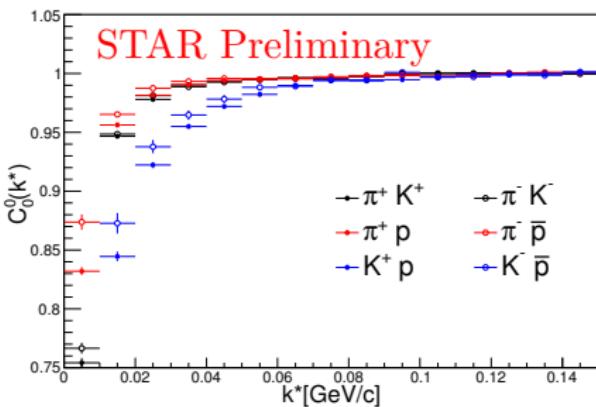
No significant dependence on energy.

Required source parameters to compare.

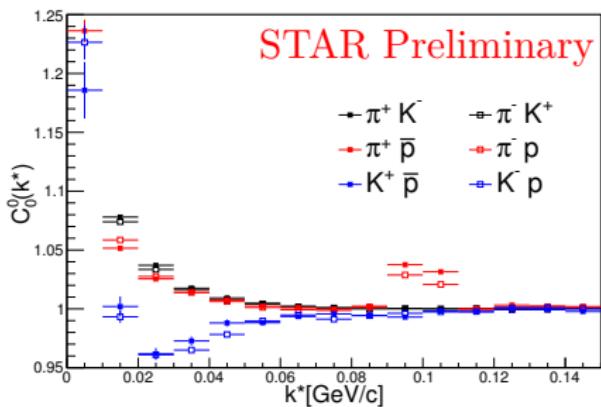
# System dependence

## Clear system dependence

like-sign @ Au+Au 39 GeV 0-10%



unlike-sign @ Au+Au 39 GeV 0-10%



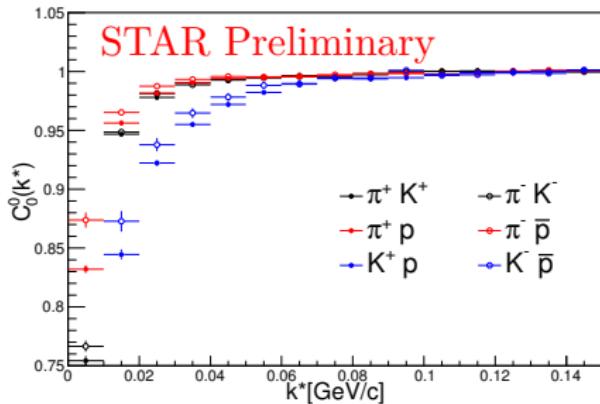
Like sign: correlations are dominated by Coulomb interaction

Coulomb strength depends on Bohr radius of the pair  
K-p — lowest Bohr radius → strongest correlation

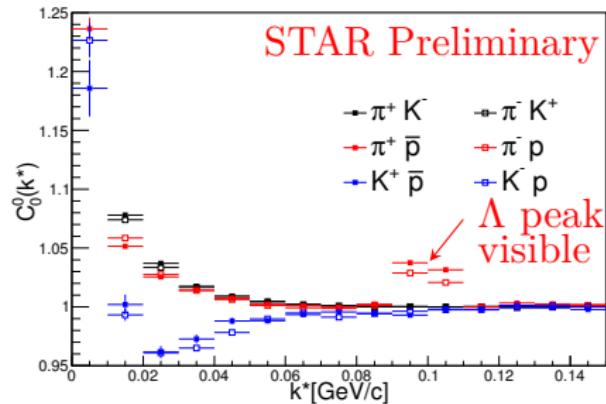
# System dependence

## Clear system dependence

like-sign @ Au+Au 39 GeV 0-10%



unlike-sign @ Au+Au 39 GeV 0-10%

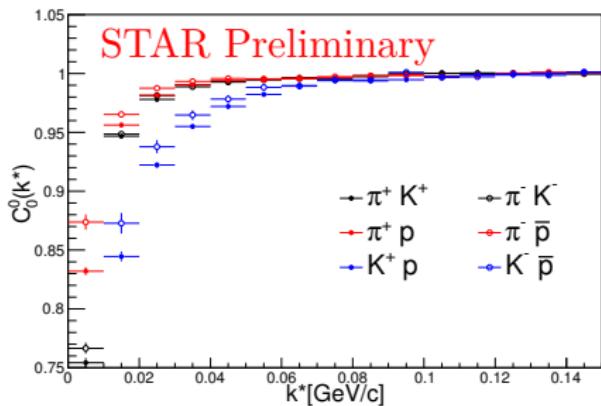


Unlike sign: interaction more complicated  
Lambda peak visible in pion-proton

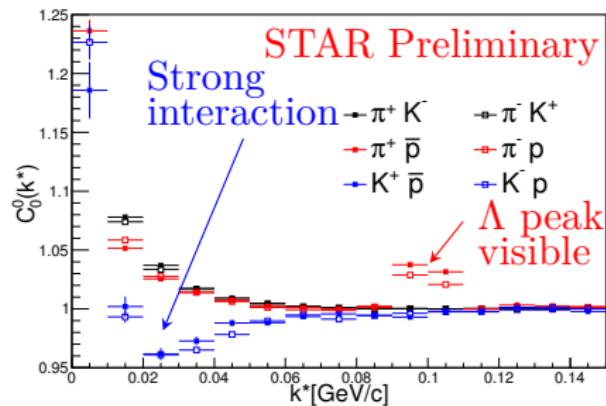
# System dependence

## Clear system dependence

like-sign @ Au+Au 39 GeV 0-10%



unlike-sign @ Au+Au 39 GeV 0-10%

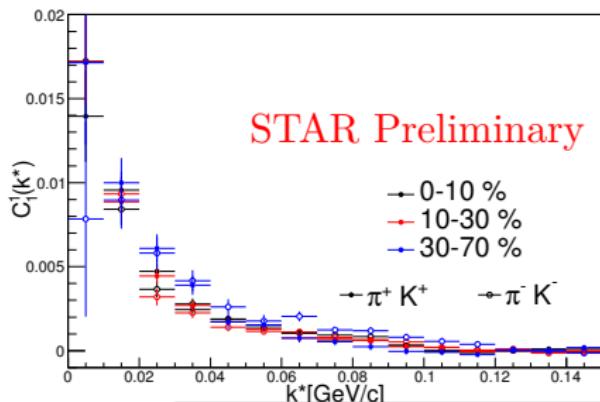


Unlike sign: interaction more complicated  
Lambda peak visible in pion-proton  
strong interaction not negligible in kaon-proton

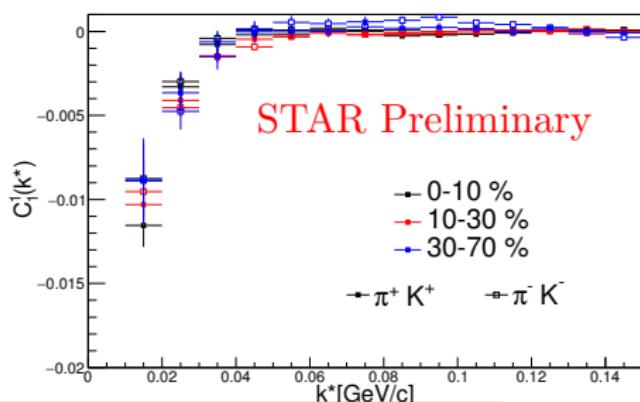
# Emission asymmetry — centrality dependence

$C_1^1 \rightarrow$  sensitive to the space-time emission asymmetry

like-sign  $\pi - K$  @ Au+Au 39 GeV



unlike-sign  $\pi - K$  @ Au+Au 39 GeV



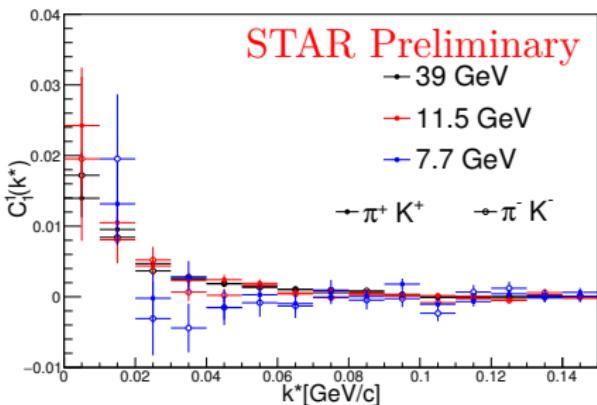
Clear signal of emission asymmetry

pions are emitted closer to center and/or later than kaons

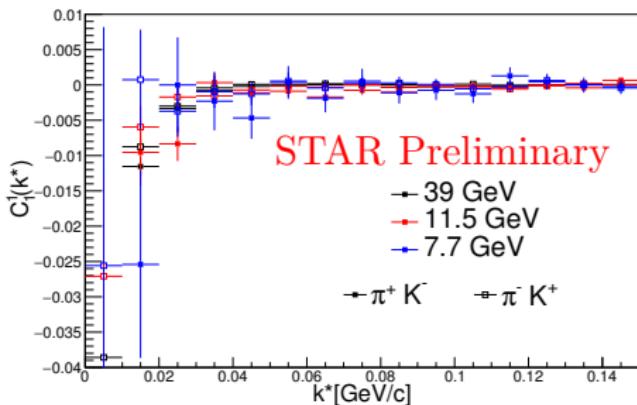
# Emission asymmetry — energy dependence

$C_1^1 \rightarrow$  sensitive to the space-time emission asymmetry

like-sign  $\pi - K$ : 0-10%



unlike-sign  $\pi - K$ : 0-10%



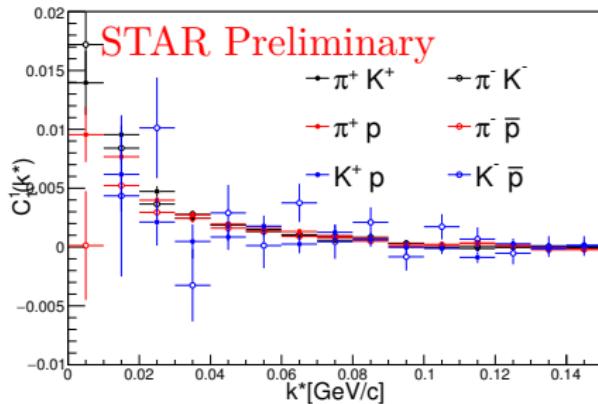
Clear signal of emission asymmetry

asymmetry does not disappear for low energies

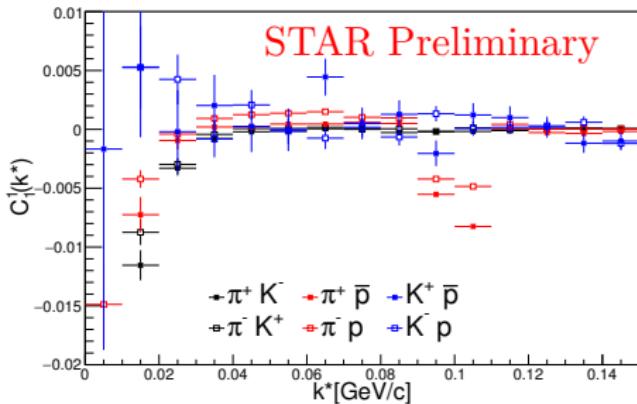
# Emission asymmetry — system dependence

$C_1^1 \rightarrow$  sensitive to the space-time emission asymmetry

like-sign @ Au+Au 39 GeV 0-10%



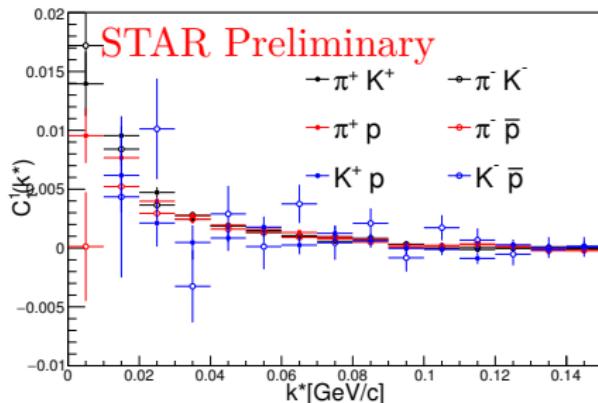
unlike-sign @ Au+Au 39 GeV 0-10%



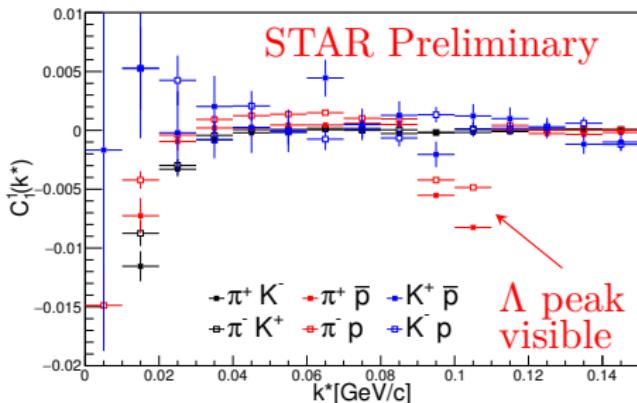
# Emission asymmetry — system dependence

$C_1^1 \rightarrow$  sensitive to the space-time emission asymmetry

like-sign @ Au+Au 39 GeV 0-10%



unlike-sign @ Au+Au 39 GeV 0-10%



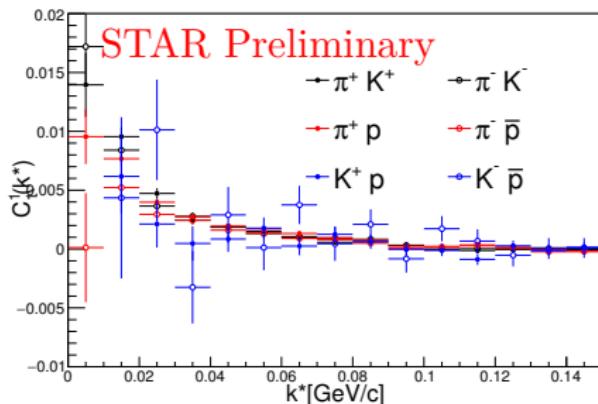
Expected ordering of particles:  
Lighter particle is emitted closer to the center  
and/or later.

A. Kisiel Phys. Rev. C81:064906 2010

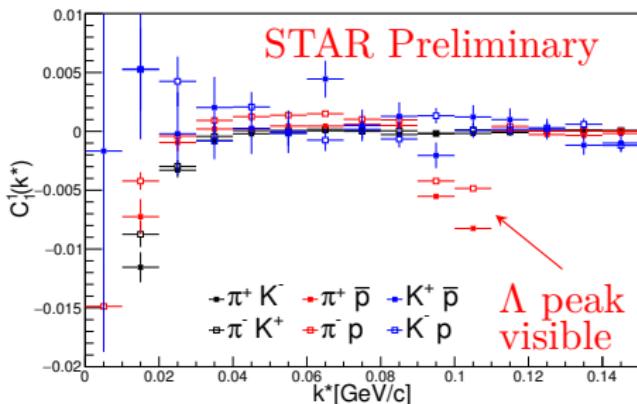
# Emission asymmetry — system dependence

$C_1^1 \rightarrow$  sensitive to the space-time emission asymmetry

like-sign @ Au+Au 39 GeV 0-10%



unlike-sign @ Au+Au 39 GeV 0-10%



Heavier particles pushed by flow towards the edge  
of the source more strongly than lighter particles.  
Heavier particles freeze-out earlier.

A. Kisiel Phys. Rev. C81:064906 2010

# Summary

## Geometry:

- Significant centrality and system dependence of the source size at BES energies
- **Strong interaction is not negligible in kaon-proton**

## Dynamics:

- Clear signal of emission asymmetry for particles with different masses at BES energies
- Asymmetry does not disappear for low energies
- **Lighter particles are emitted closer to the center of the source and/or later than heavier particles — flow pushes heavier particles harder to the edge**

# Summary

## Geometry:

- Significant centrality and system dependence of the source size at BES energies
- **Strong interaction is not negligible in kaon-proton**

## Dynamics:

- Clear signal of emission asymmetry for particles with different masses at BES energies
- Asymmetry does not disappear for low energies
- **Lighter particles are emitted closer to the center of the source and/or later than heavier particles — flow pushes heavier particles harder to the edge**

**Thank you for your attention!**