

Strongly coupled dynamics with a phase transition – some surprises

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RJ, M. Järvinen, H. Soltanpanahi, J. Sonnenschein, PRL '22 [2205.06274]
RJ, M. Järvinen, J. Sonnenschein, to appear...

Outline

Long term motivation

Domain wall velocities

Conventional picture and a key question

Results of holographic simulations

A simple formula for domain wall velocity

Boost-invariant expansion and hot remnants

Boost-invariant expansion and cooling

Reheating of plasma remnants

Entropy considerations

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Question: How to understand phase transitions in real time within holography?

- ▶ Phase transitions in equilibrium in holography correspond to switching between **two spacetime geometries**
- ▶ It is nontrivial what happens in real time...
- ▶ In some cases there is a classical gravity description, in others probably not...

Theoretically interesting even for its own sake!

But also interesting real-world applications:

- ▶ What does holography tell us about hadronization??
caution for $N_c = 3$: crossover!
- ▶ 1st order phase transition reappears at nonzero density...
- ▶ Some physics in early universe??

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1. Bubble wall velocities

Introduction

- ▶ At a 1st order phase transition $T = T_c$, we can have domains of coexisting phases separated by domain walls
- ▶ The pressures on both sides are balanced and the domain wall can be static...

Question: What happens when we move away from $T = T_c$?

- ▶ This can occur for nucleated bubbles of a stable phase within an supercooled medium
- ▶ At an interface between phases away from $T = T_c$
- ▶ Or an interface between phases at different temperatures

Domain walls start to move...

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Goal:

Understand bubble wall velocities at strong coupling...

- ▶ The pressures on both sides are not balanced so one would expect accelerated motion...
- ▶ ...but this does not happen — the domain wall ultimately moves with a constant velocity...
- ▶ **Common lore:** friction in the second phase balances the net force — challenging to calculate...

Our claim: At **strong coupling** (+ some entropy ratio assumptions), the domain wall velocity can be understood in a much simpler way using essentially only the equation of state...

- ▶ The pressures on both sides are not balanced so one would expect accelerated motion...



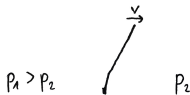
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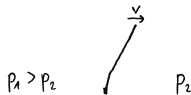
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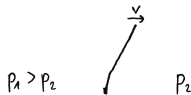


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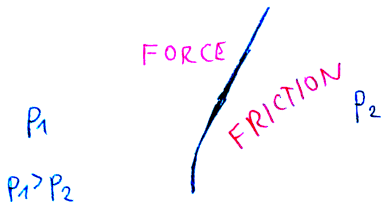


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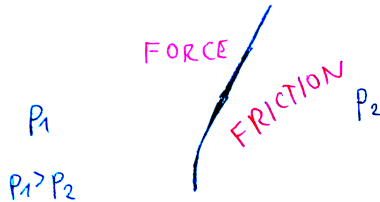
Recall the standard picture:



- ▶ The net force across the domain wall implies that the pressure difference is localized close to the domain wall...
- ▶ It is not obvious *a-priori* if this is always the case...

Perform **holographic** simulation and read off the pressure profile...

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1. Witten model in 3D – confinement/deconfinement transition
use **simplified hydrodynamics+scalar field fitted to holography**

2. Holographic gravity+scalar model with a transition between two deconfined phases

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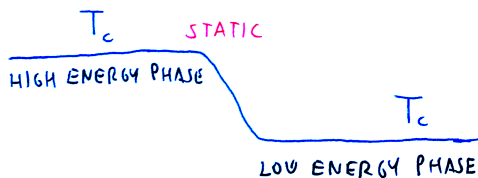
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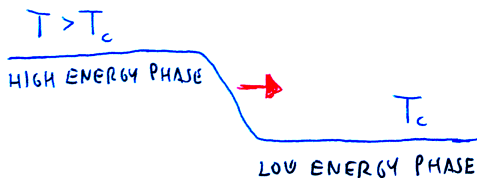
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Simplest scenario



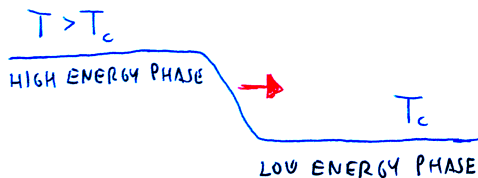
- ▶ We increase the temperature of the high energy phase...
- ▶ We perform time evolution from the above static initial conditions...
- ▶ The pressure in each phase may be read off from the T^{yy} component of the energy-momentum tensor

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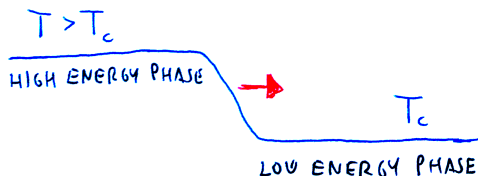
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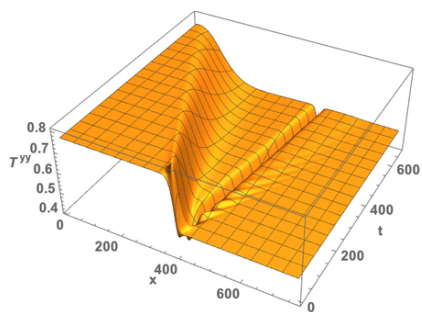
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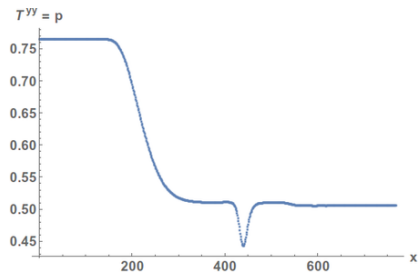
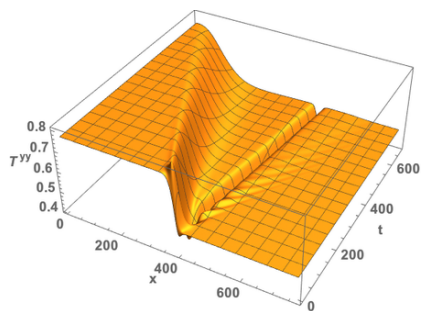
Spacetime pressure profiles



Key features:

- ▶ The large pressure difference appears **away** from the domain wall
- ▶ The change in pressures occurs in the high energy density phase
→ **hydrodynamic description**
- ▶ The pressure is essentially constant across the domain wall, and very close to $\rho(T_c)$...

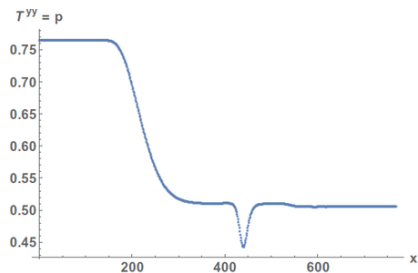
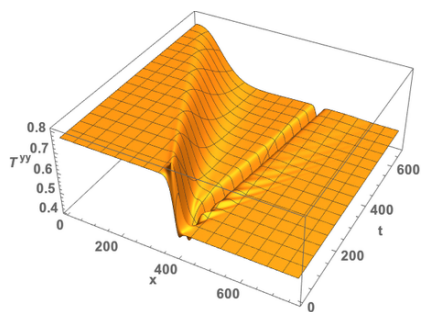
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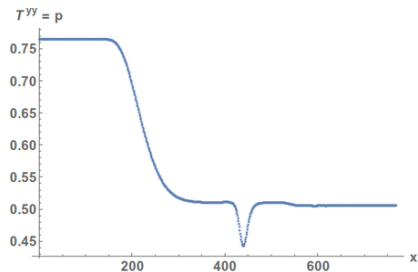
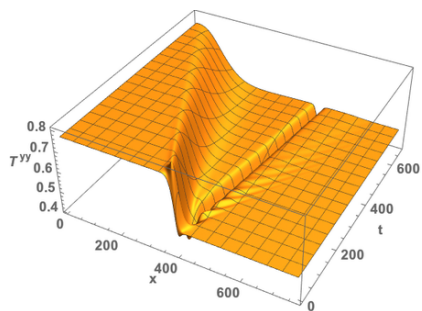
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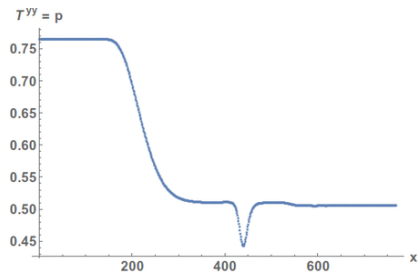
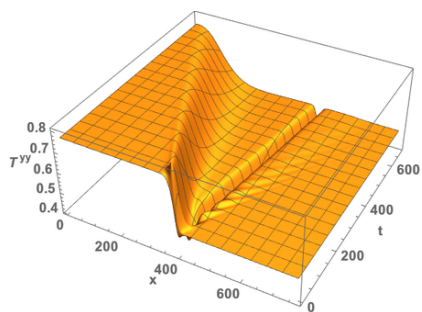
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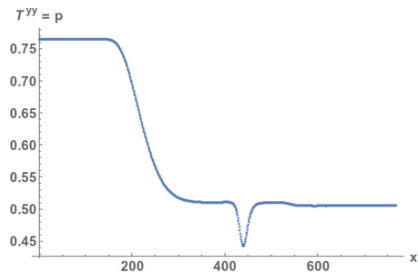
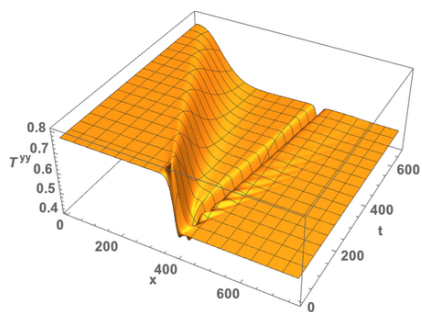
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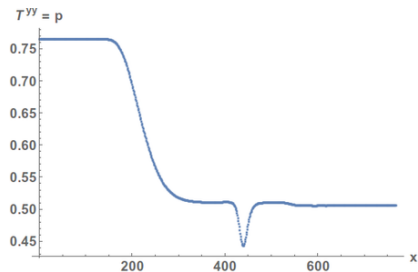
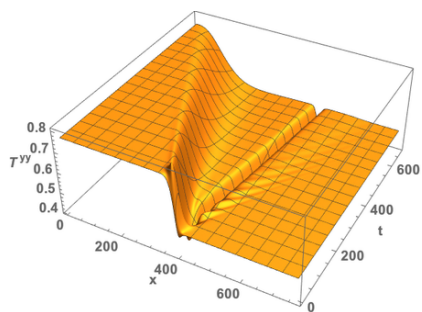
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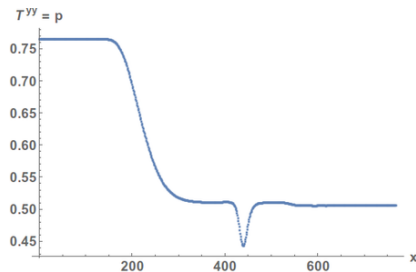
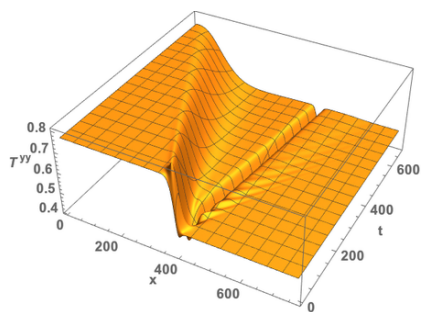
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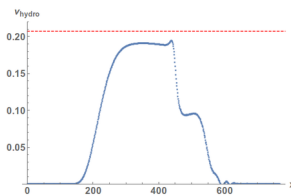
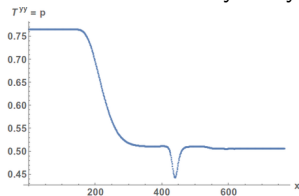


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What happens in the high energy phase?

We can extract the hydrodynamic velocity from T^{xt}



- ▶ The hydrodynamic velocity is quite close to the domain wall velocity v_{dw}
- ▶ This gets better with increasing ratio of entropies in the two phases...
- ▶ We can formulate finding v_{dw} as a hydrodynamic problem:

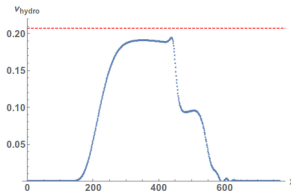
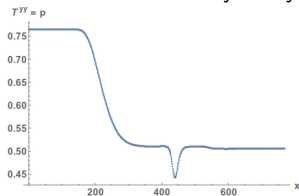
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$$v_{dw} = \tanh \int_{p_c}^{p_A} \frac{1}{(\varepsilon + p)c_s} dp$$

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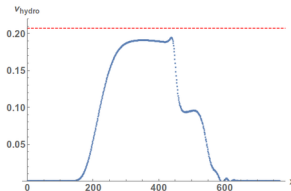
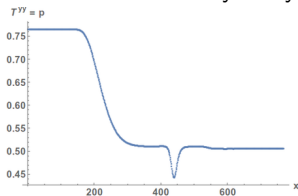
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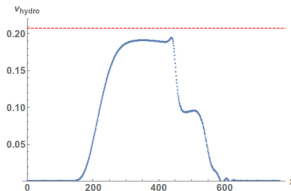
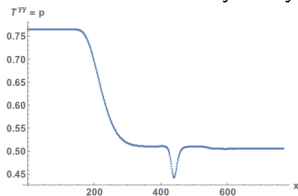
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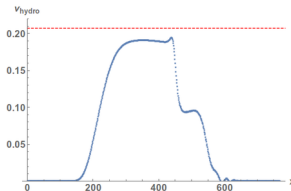
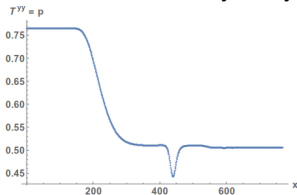
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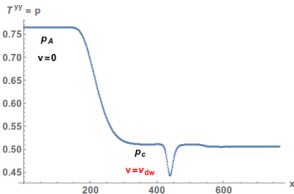
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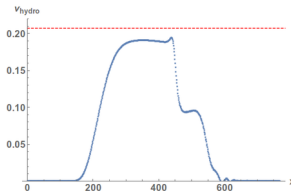
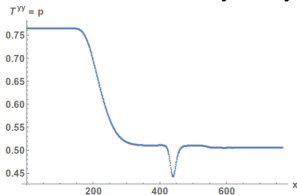
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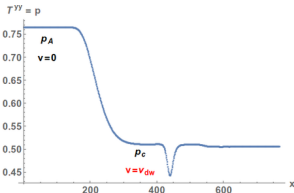
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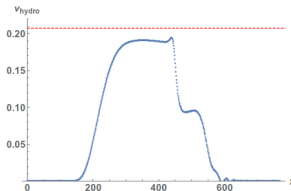
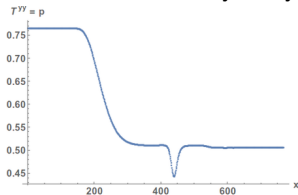
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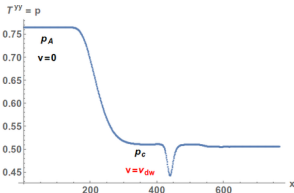
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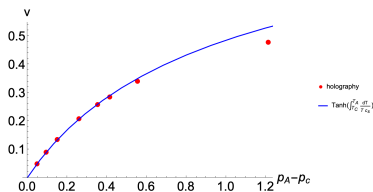
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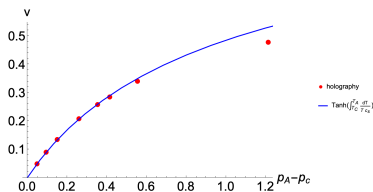


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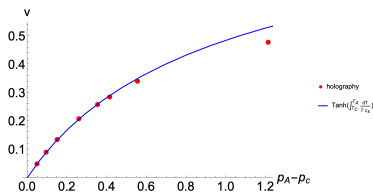


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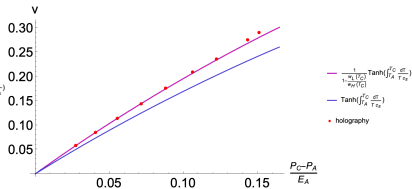
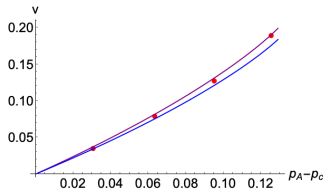
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Key takeaways:

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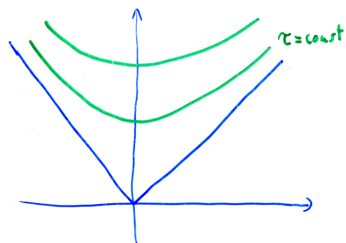
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2. Boost-invariant expansion and hot remnants

Boost-invariance and 1st order phase transition



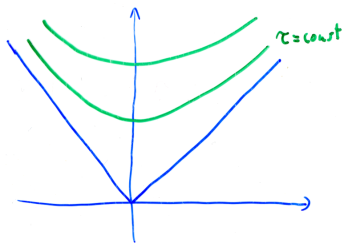
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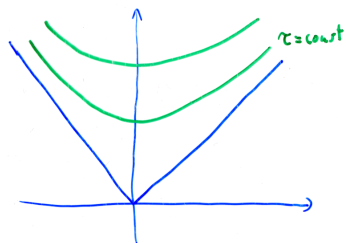
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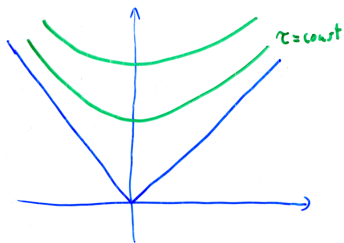
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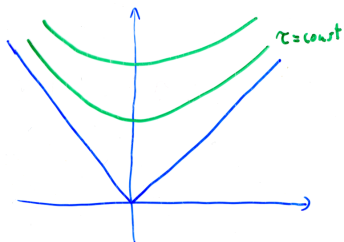
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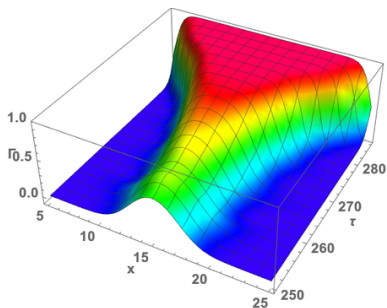
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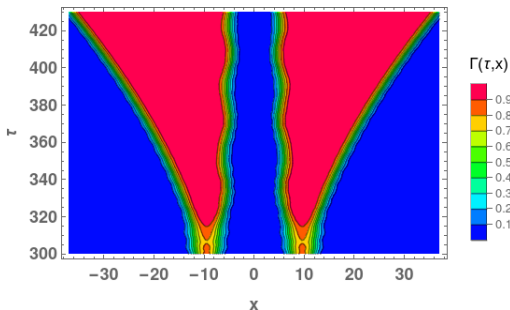
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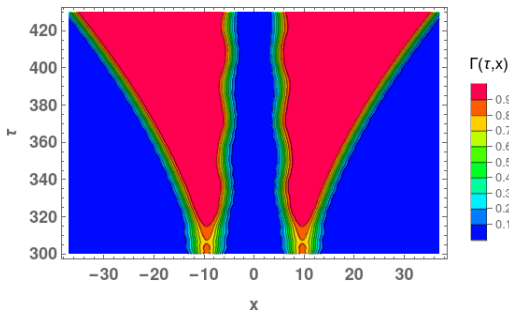
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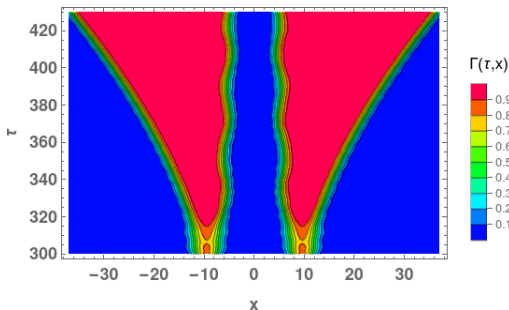
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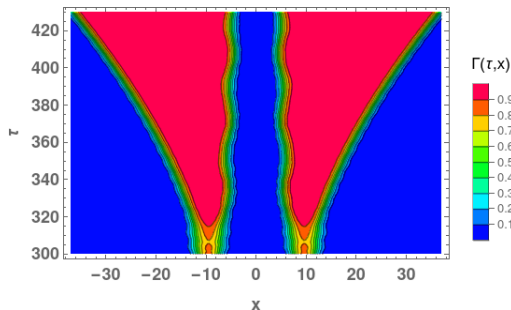
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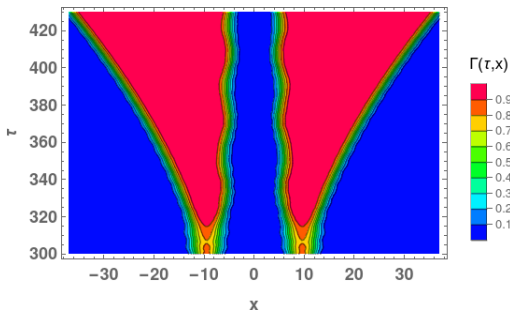
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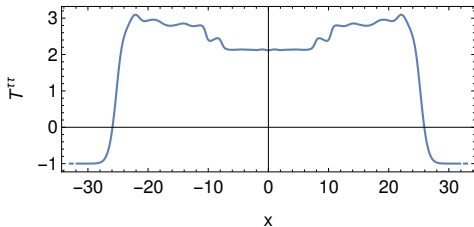


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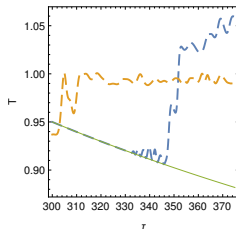
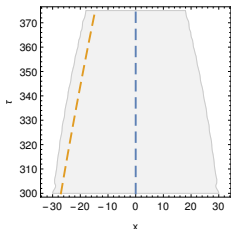
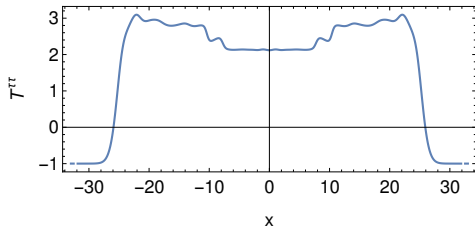
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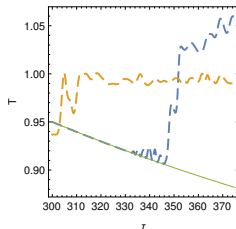
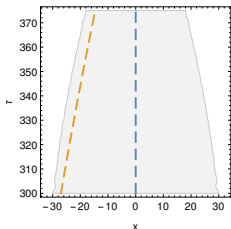
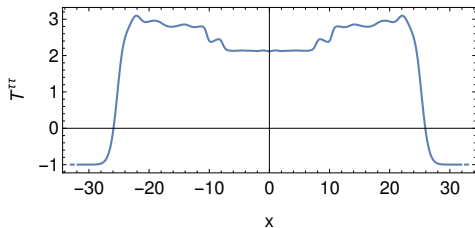
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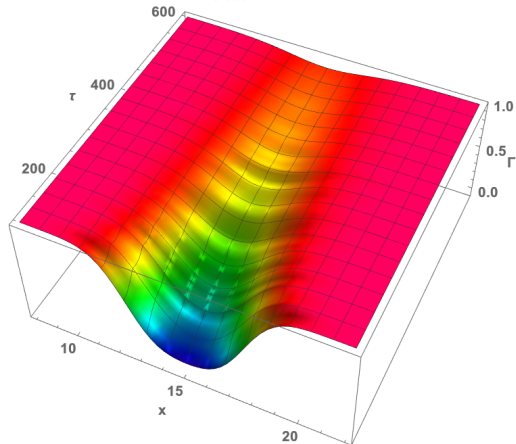
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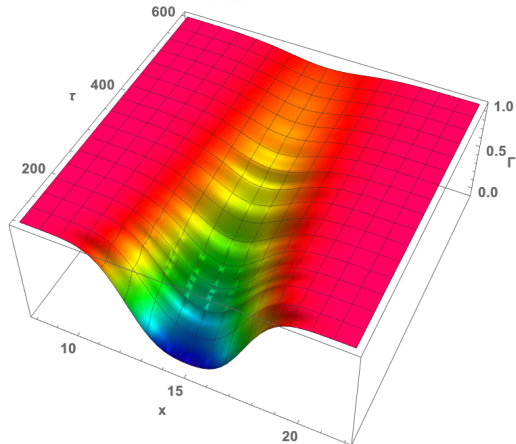
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