Oscillons in gapless scalar field theories

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Plan

Plan:

Oscillons: overview and the gap

Oscillons are spatially-localised, long-lived, oscillatory solutions of the field equation(s) of classical field theories

- nonlinearity of the field equation is essential
- no topological charge
- it is surprising that oscillons do not couple more strongly to radiation
- the fundamental reason for their existence remains somewhat mysterious
- sometimes the oscillon can be thought of as a decaying **sphaleron** of the field theory ¹.

¹N. Manton and TR, The Simplest Oscillon and its Sphaleron, Phys. Rev. D 107, 085012 (2023)

- They are found in many models and in various dimensions,
 - phase transitions in condensed matter
 - astrophysics (dark matter candidates)
 - cosmology (inflation)^{2 3}
 - fundamental theory of electroweak interactions
 - graphene ribbons
 - can be created directly from radiation, as an intermediate stage for kink-antykink creation $^{\rm 4}$

²J. Sakstein and M. Trodden, Oscillons in higher-derivative effective field theories, Phys. Rev. D 98, 123512 (2018)

³Kaloian D. Lozanov and Mustafa A. Amin, End of inflation, oscillons, and matter-antimatter asymmetry, Phys. Rev. D 90, 083528 (2014)

⁴TR and Ya. Shnir, Phys. Rev. Lett. 105, 081601, (2010)

- A basic oscillon is periodic, with a fundamental frequency ω < m, where m is the threshold for radiation modes.
- A model has to have a frequency gap (counterexamples in the future).
- The oscillon has an arbitrary amplitude lying in some finite range.
- As the amplitude increases, the frequency ω decreases away from the threshold *m*.
- In 1+1d $\omega \rightarrow m$ but in higher dimensions $\omega \rightarrow \omega_{\max} < m$.
- The main channel of losing energy to radiation is usually through 2ω (but there are exceptions):
 - large oscillons with $\omega < m/2$
 - models without quadratic terms in the e.o.m.
 - breathers in the sine-Gordon model.

Oscillons: overview and the gap: Example: ϕ^4

Let us consider the standard ϕ^4 model:

$$L = \int \frac{1}{2}\phi_t^2 - \frac{1}{2}\phi_x - \frac{1}{2}(\phi^2 - 1)^2 dx$$

Starting from gaussian initial conditions

$$\phi(x,0) = 1 - Ae^{-cx^2}, \ \phi_t(x,0) = 0$$

for A = 0.5 we see different behaviour for different values of *c*.



Oscillon profile and small radiation tails for A = 0.5, c = 0.14 and $t \approx 2000$



Similar behaviour is observed for arbitrary small value of A, with a certain scaling c(A).

Small oscillations around vaccua obey (?) Klein-Gordon equation

$$\psi_{tt} - \psi_{xx} - 4\psi + \mathcal{O}(\psi^2) = 0$$

for $\phi = \pm 1 + \psi$

· For large times solutions of KG equations behave as

$$\psi(0,t) \sim \frac{\cos(2t+\delta)}{\sqrt{t}}$$

- This was true for *c* = 1 and *c* = 0.5, but for *c* ≈ 0.14 oscillations did not change visibly.
- · Sometimes almost quasi-periodic oscillations were visible.
- Scalability of oscillons means that linearization can be incorrect even as $A \rightarrow 0$.

Oscillons: overview and the gap: Example: ϕ^4

Naive explanation: Assuming that the field profile does not change with time we can plug:

$$\phi(x,t) = 1 - A(t)e^{-cx^2}$$

into the field theory lagrangian and integrate over *x*. This gives us mechanical lagrangian for anharmonic oscillator:

$$L = \sqrt{\frac{\pi}{8c}} \left[\dot{A}^2 - A^2(c+4) + 4\sqrt{\frac{2}{3}}A^3 - \frac{A^4}{\sqrt{2}} \right]$$

Small oscillations oscillate with frequency $\omega_0 = \sqrt{c+4} > 2$. In field theory such oscillations frequencies are above the mass threshold and disperse.

However, nonlinear corrections lower the frequency below the mass threshold

$$\omega = \omega_0 - \frac{80 + 3\sqrt{2}\omega_0^2}{8\omega_0^3}A_0^2 + \mathcal{O}(A_0^4)$$

Critical amplitude, for $\omega < m$:

$$A_{crit} = \sqrt{\frac{\omega_0 - 2}{\omega}}$$

Oscillons: overview and the gap: Example: ϕ^4

Power spectrum of the field at centre for A = 0.5, c = 0.1



- Basic frequency is below the mass threshold $m_{\phi^4} = 2$
- Higher harmonics are in the continuous spectrum (radiation).
- Smaller peaks (for example at the threshold) are responsible for the modulations.

In the integrable sine-Gordon model

$$L = \int \frac{1}{2}\phi_t^2 - \frac{1}{2}\phi_x^2 - V_{sG}(\phi) \, dx \,, \qquad V_{sG}(\phi) = 1 - \cos(\phi)$$

there exists a periodic finite-energy breather solution

$$\phi_B(x,t) = 4 \operatorname{atan}\left(\frac{\sqrt{1-\omega^2}}{\omega} \frac{\cos(\omega t)}{\cosh\left(\sqrt{1-\omega^2} x\right)}\right)$$

with $\omega = \cos(A/4) \in (0, 1)$ with mass threshold $m_{sG} = 1$.

Potentials in many field theory models around vacua (after shift and rescaling) can be expressed as

$$V(\phi) = 1 - \cos(\phi) + \mathcal{O}(\phi^n)$$
 where $n > 2$

therefore, in some sense, small amplitude oscillons owe their existence to sine-Gordon integrability.

Massless model

Let us consider a real scalar field theory in (d + 1) dimensions

$$L = \int \left(\frac{1}{2}\phi_t^2 - \frac{1}{2}(\nabla\phi)^2 - U(\phi)\right) d^d x.$$

Here we will assume d = 1 or 3. Let $U(\phi)$ have a form very similar to the simplest, most prototypical, double vacuum potential, i.e., ϕ^4 but *flattened* near the vacua, i.e. so that the potential is *quartic* near the vacua. This can be achieved with a rational modification

$$U(\phi) = \frac{W^2}{W + \epsilon}, \qquad W = \frac{1}{2}(1 - \phi^2)^2,$$

where ϵ is a small positive parameter, which controls the size of the deformed region.

Massless model: Construction



where the shift $U_0(\epsilon) = -1/(4\epsilon + 2) + 1/2$ is just for better visualization. Near vacua:

$$U \approx \frac{1}{4\epsilon} (1-\phi)^4 (1+\phi)^4$$

so the mass of (infinitesimally) small perturbations vanishes,

$$m^2 = \left. \frac{d^2 U}{d\phi^2} \right|_{\phi=\pm 1} = 0.$$

This model has no mass gap, so oscillons should not exist.

Massless model: Numerical results

However, large amplitude oscillons "see" the potential as if it was the ϕ^4 model, and only the low amplitude tails "feel" that the model is massless.



Massless model: Numerical results

Oscillons in the modified model definitely exist, although they live shorten than usual oscillons.



Massless model: Numerical results

- similar spectra, but
- ϕ^4 : radiation through 2nd harmonic
- φ⁴: slow threshold radiation (residue from bad IC)



- mod: radiation through 1st harmonic (speed of light)
- mod: no threshold radiation



- In 3+1D oscillons in ϕ^4 are less stable
- No integrable sine-Gordon model (no 3d breathers)
- Additional upper frequency $0 < \omega_{cr,1} < \omega < \omega_{cr,2} < m$
- Modulations grow with time
- Lifetimes form resonant patters (with modulation⁵)
- After reaching ω_{cr,2} fast decay is observed
- In the deformed model modulations vanish

 $^{{}^{5}}$ E. P. Honda, M. W. Choptuik, Fine structure of oscillons in the spherically symmetric ϕ^{4} Klein-Gordon model, Phys. Rev. D 65 084037 (2002).

Massless model: Oscillons in 3+1d

Lifetimes of oscillons from gaussian initial conditions:

$$\phi(x,0) = 1 - 2e^{-r^2/r_0^2}$$



- Sometimes oscillons in the modified model live longer than in the non-modified model.
- More radiation in the deformed model
- But no instability due to the modulations.

Massless model: Oscillons in 3+1d



For fixed r_0 but changing ϵ we can see another resonant (fractal-like?) structures

Summary

- · Oscillons are interesting and important objects
- Nonlinearities are essential for their existence
- In standard theories their frequency is below the mass threshold
- But they also can exits in gap-less models
 - only finite amplitude oscillons exist (similar to usual 3+1d)
 - can have more applications then previously expected

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