

Freezing-in Cannibal Dark Sectors

Esau Cervantes

Based on: [arXiv:2407.12104](https://arxiv.org/abs/2407.12104)

Polish Particle and Nuclear Theory

Summit

IFJ-PAN

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NATIONAL
CENTRE
FOR NUCLEAR
RESEARCH
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Our starting point

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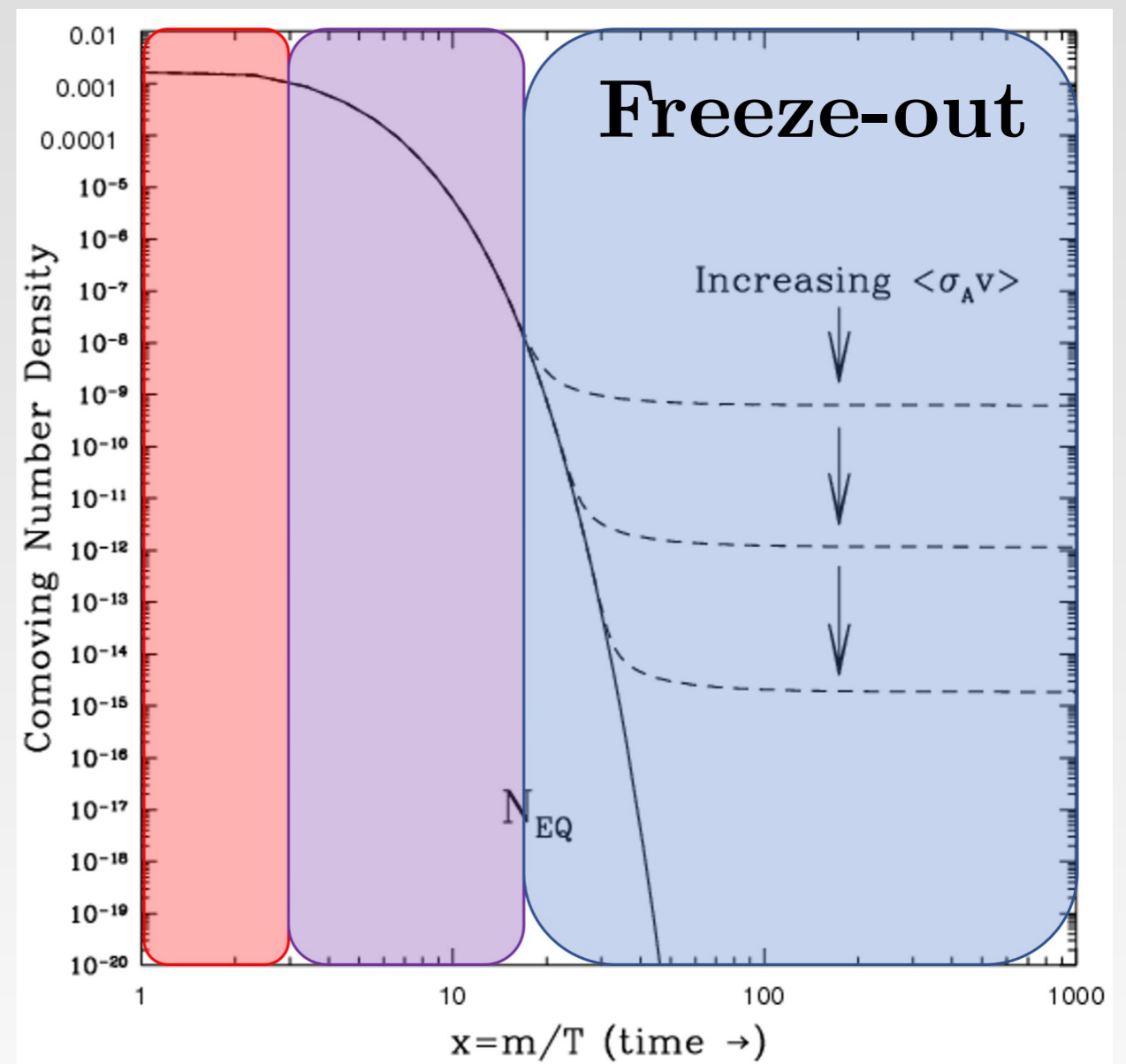
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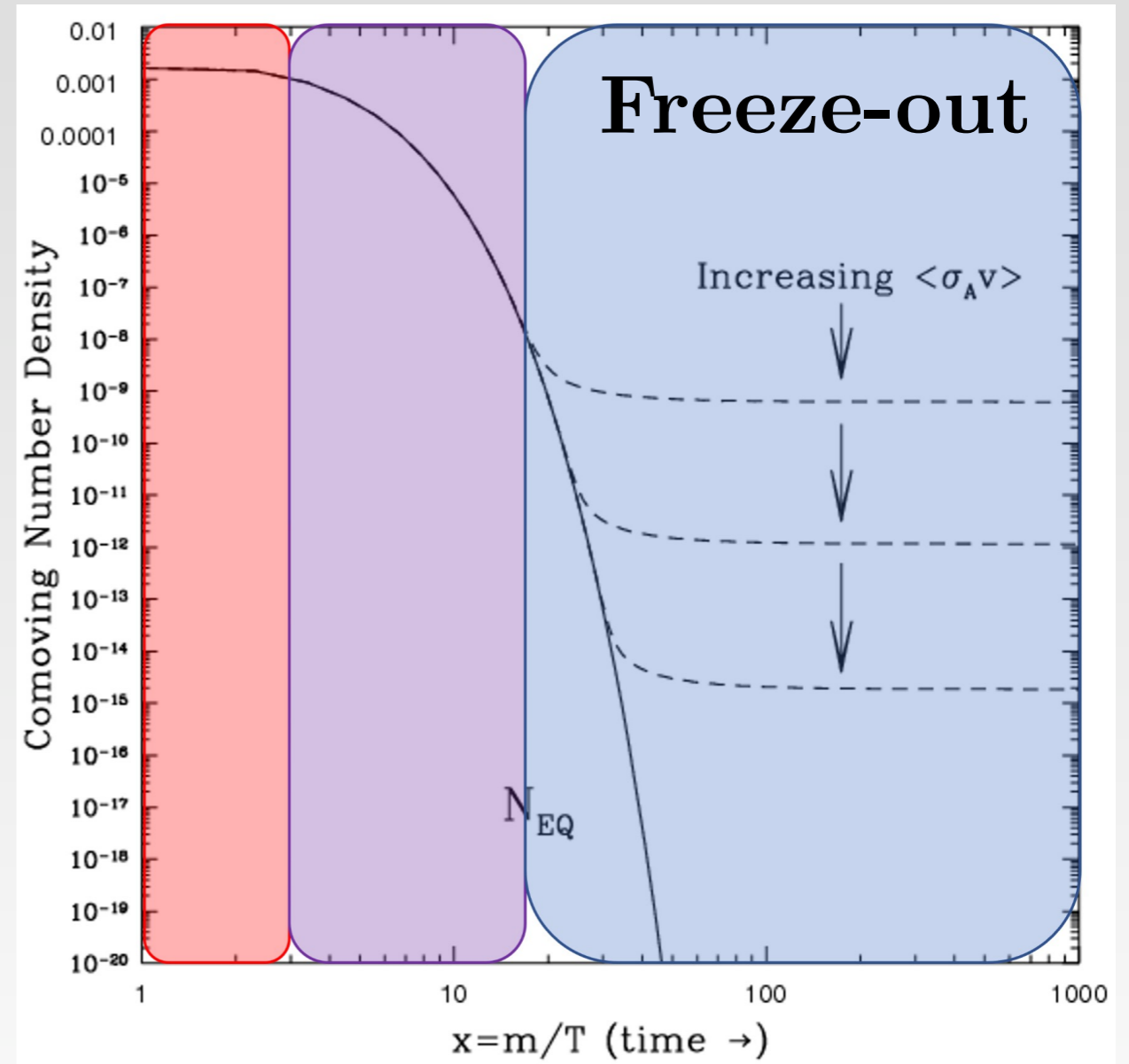
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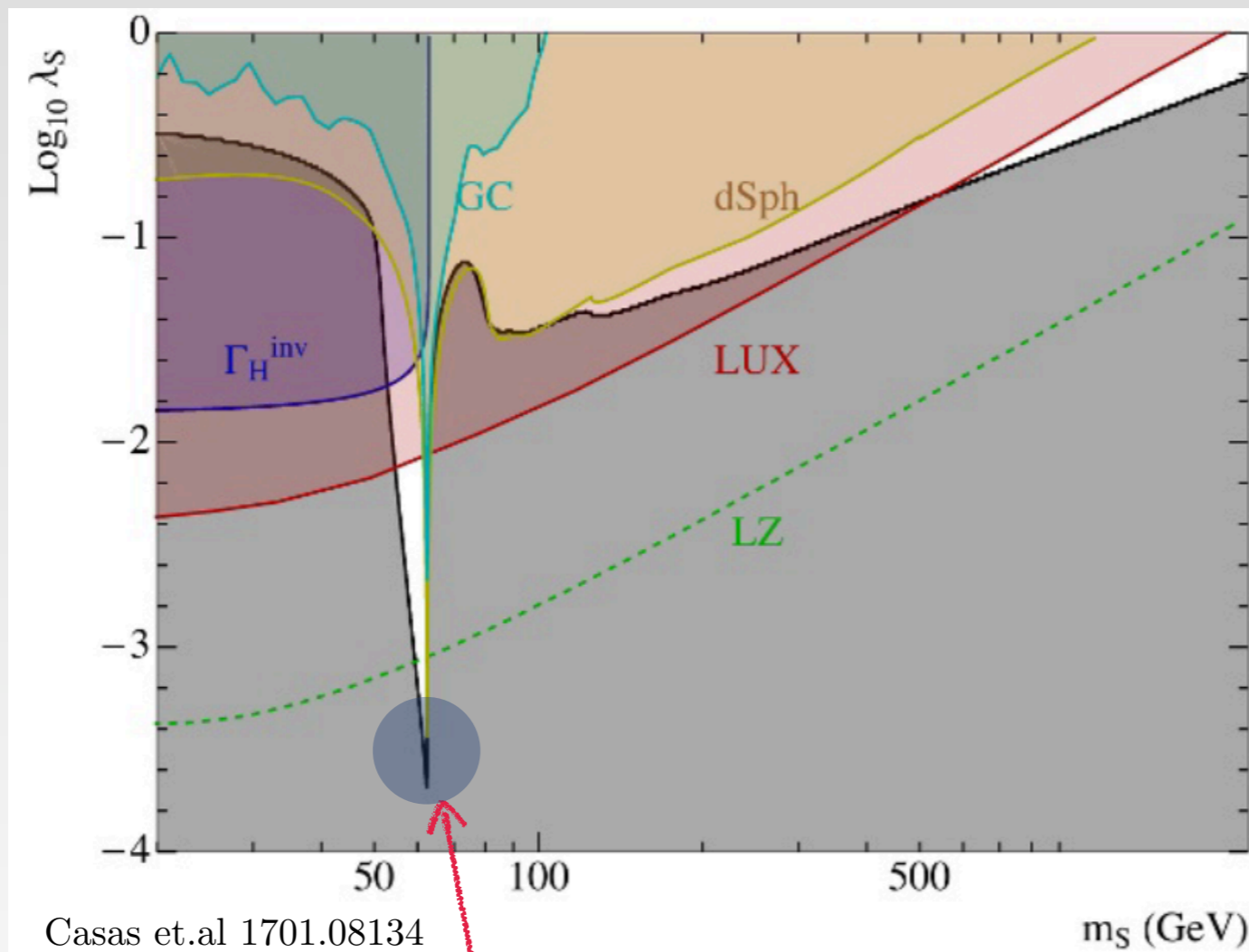


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Implicit assumptions: $f_i = (e^{-(E_i - \mu_i)/T_i} \pm 1)^{-1}$ and $T_i = T_{SM}$ (kinetic eq.)

The WIMP dilution

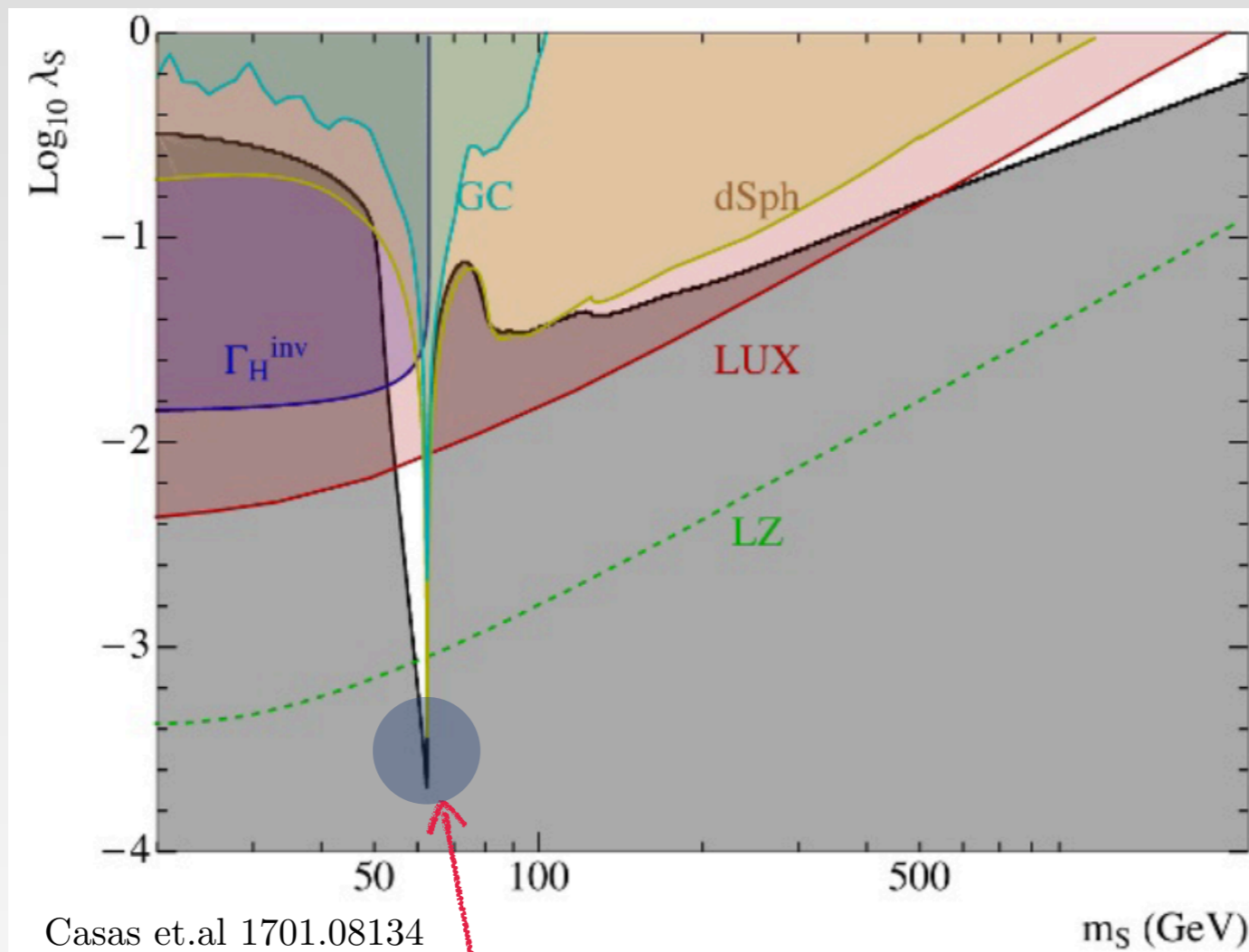
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‘Remaining’ parameter space around
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What **now**?

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Self Interacting Dark Matter

SELF-INTERACTING DARK MATTER

ERIC D. CARLSON

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Received 1992 March 17; accepted 1992 April 20

Explains amount of DM in the universe **solely** through **self** number changing reactions!

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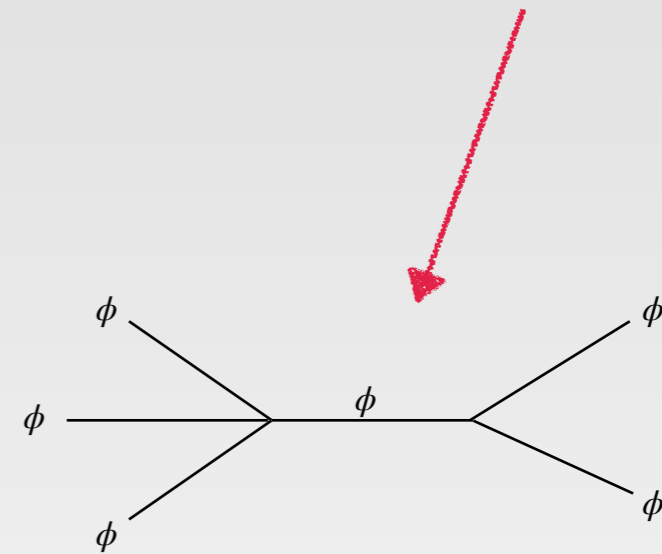
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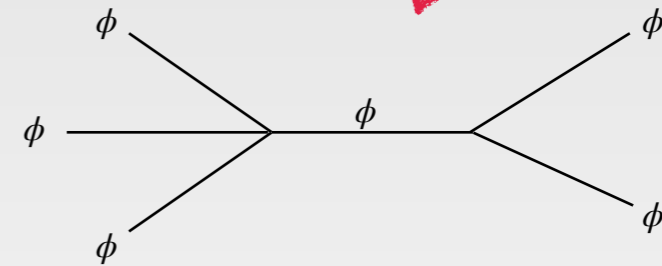
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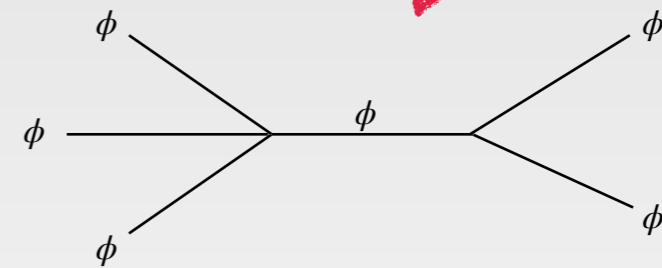
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Conservation of energy leads to *faster moving* (hotter) DM states;

i.e.; the DM fluid **exchanges** number density for kinetic energy!

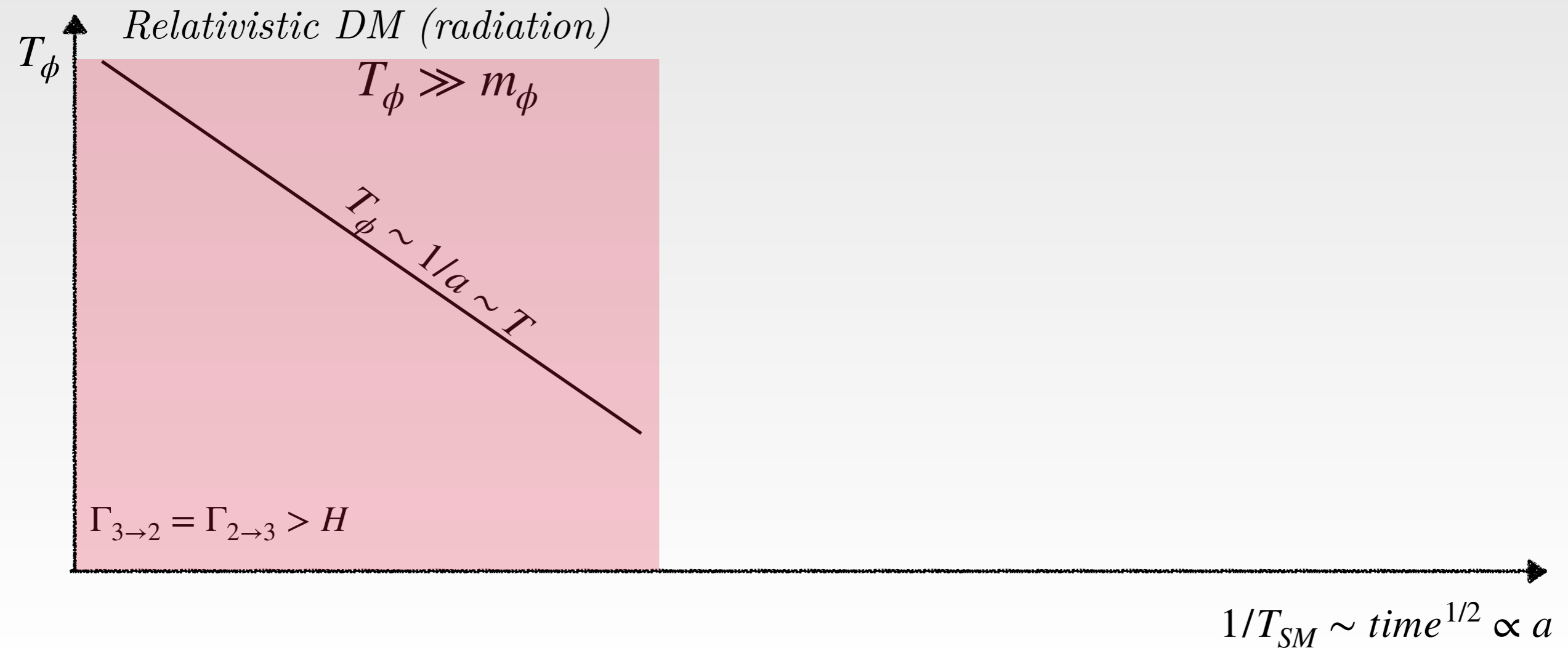
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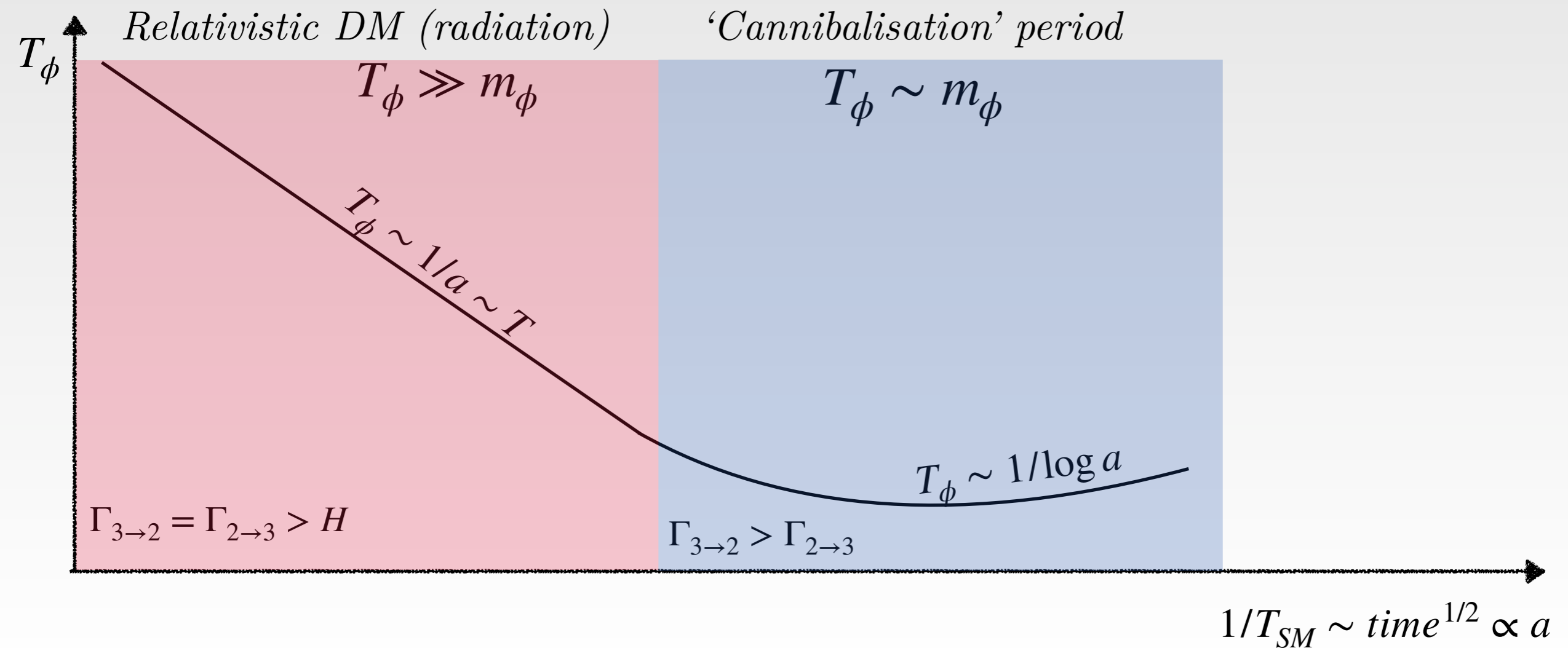
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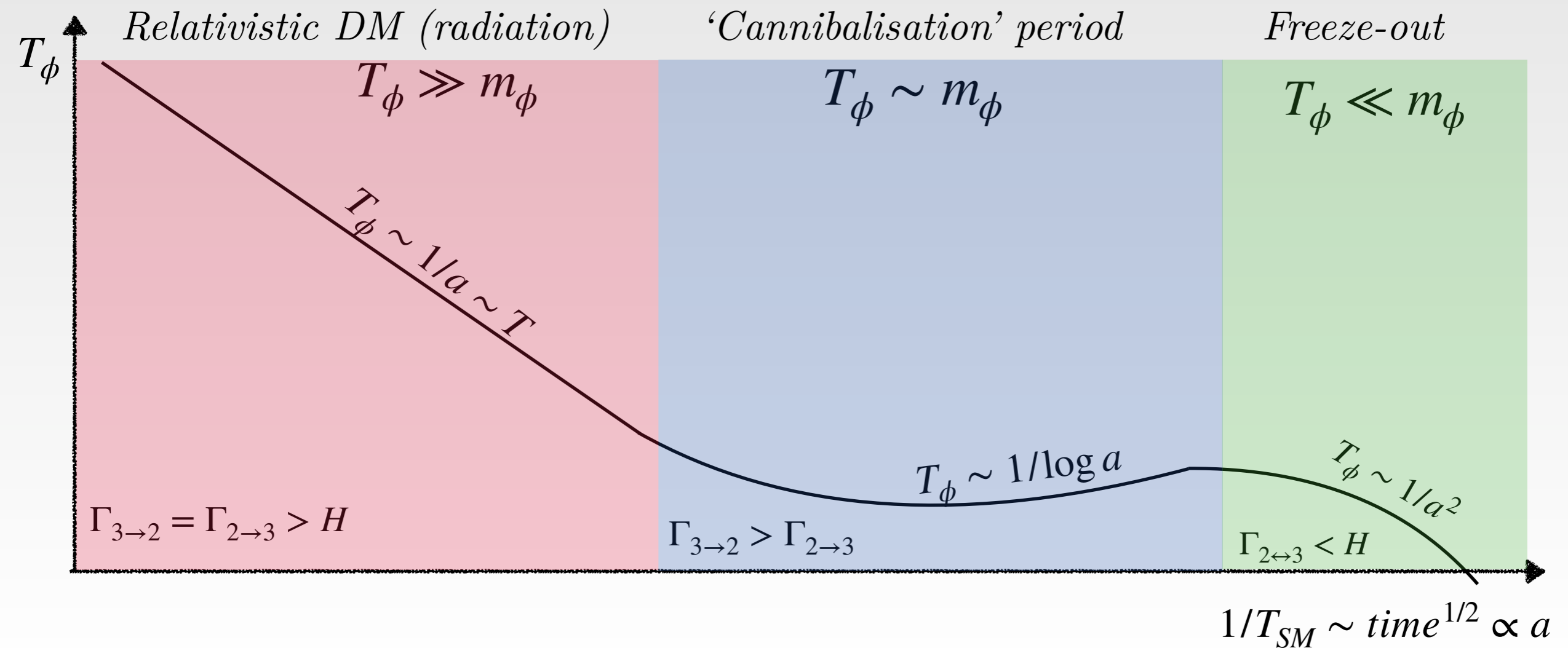
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- Initially DM is *relativistic* ($T_{DM} \gg m_{DM}$);
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- The system decouples and behaves as a non-relativistic gas.



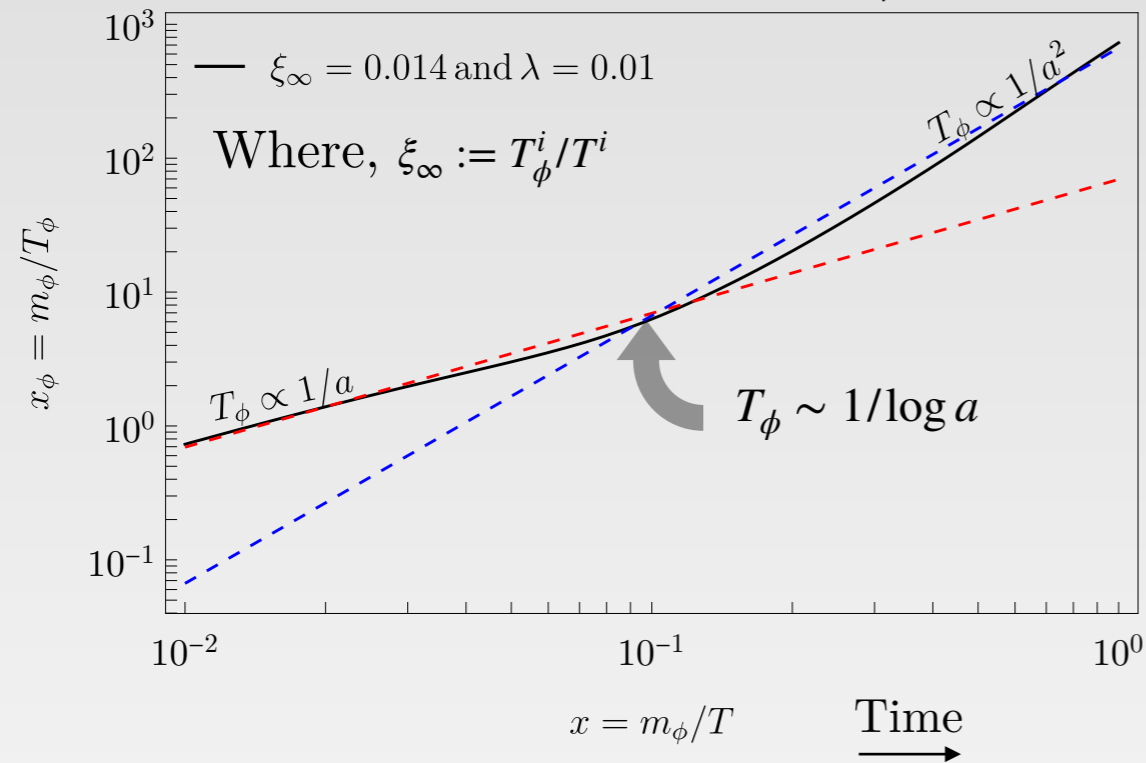
Internal Freeze-out

Secluded singlet scalar DM: $\mathcal{L}_{SM} \supset -\frac{m^2}{2}\varphi^2 - \frac{\lambda w}{3!}\varphi^3 - \frac{\lambda}{4!}\varphi^4 - \lambda_{H\varphi}\varphi^2|H|^2$

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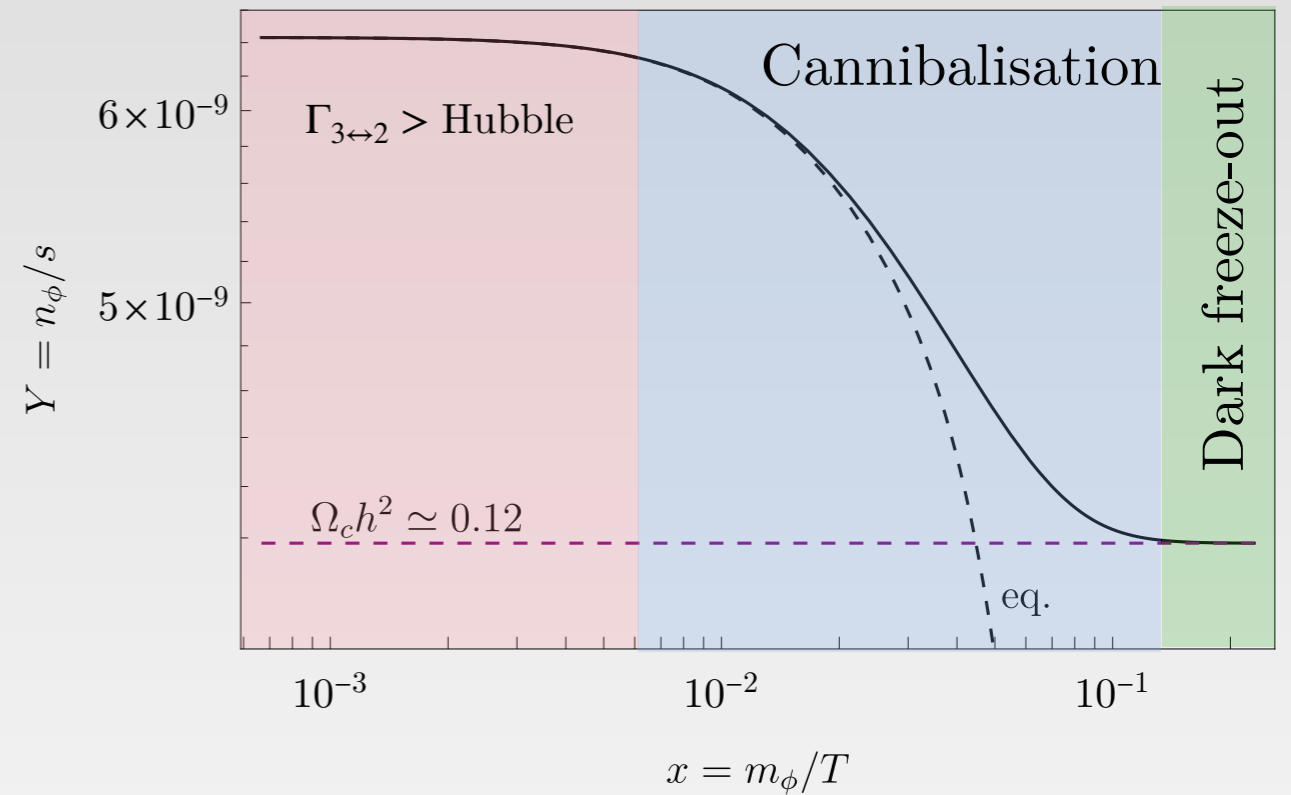
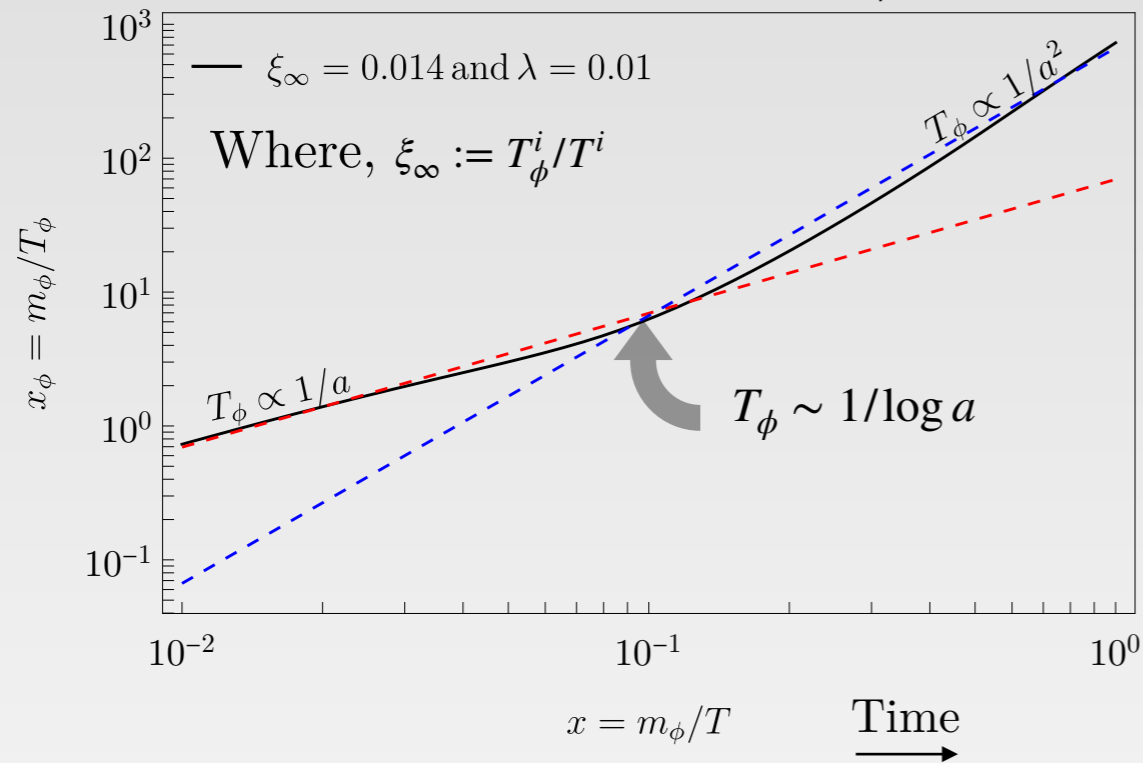
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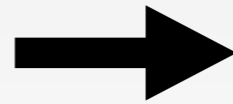
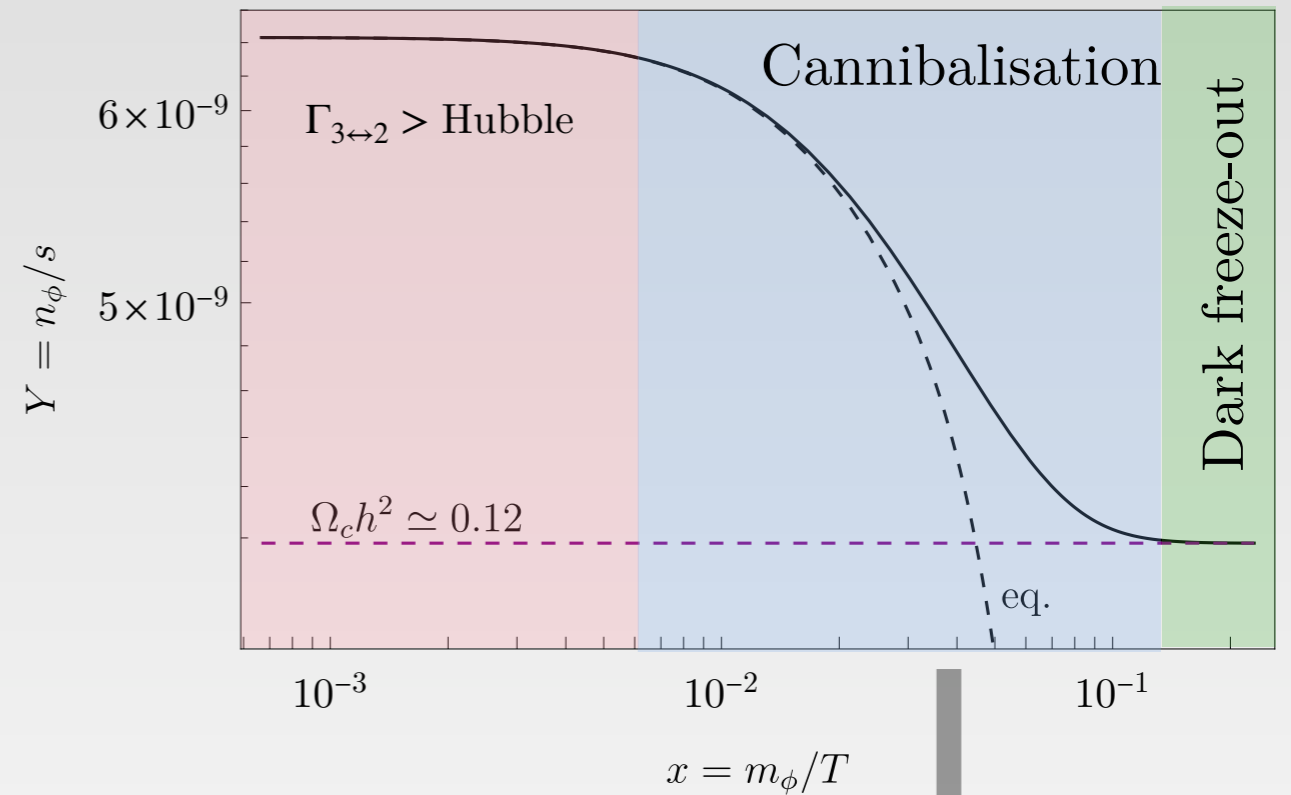
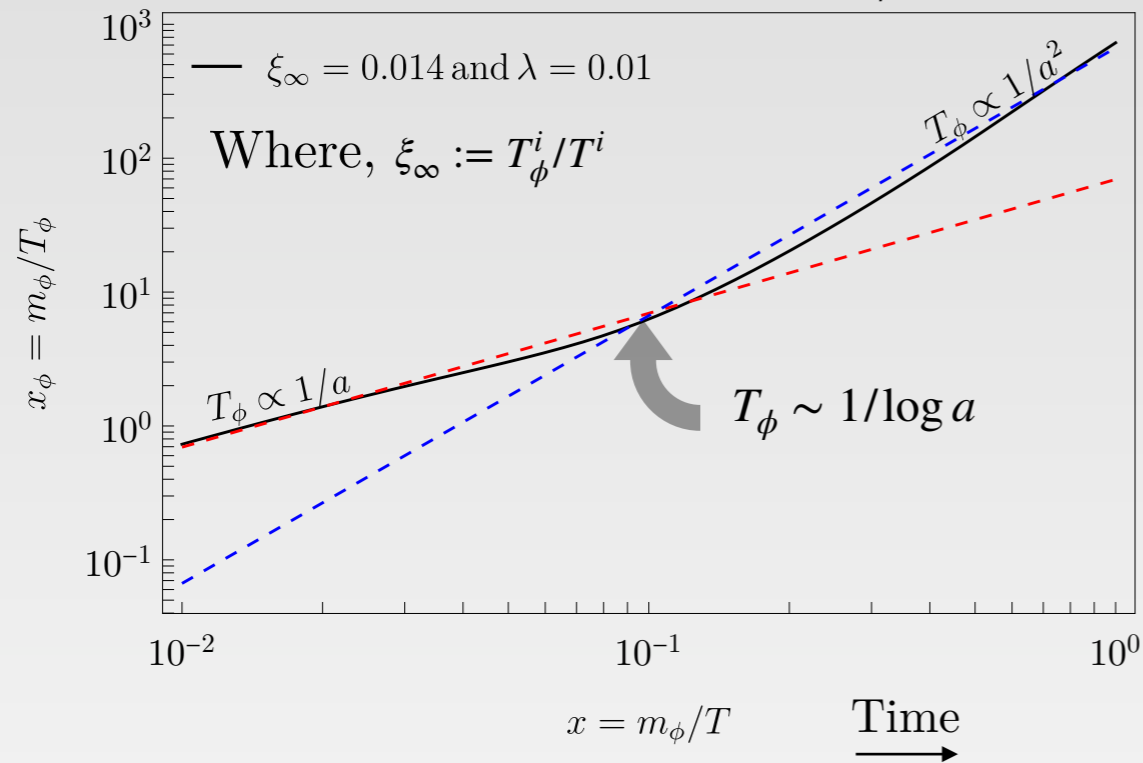
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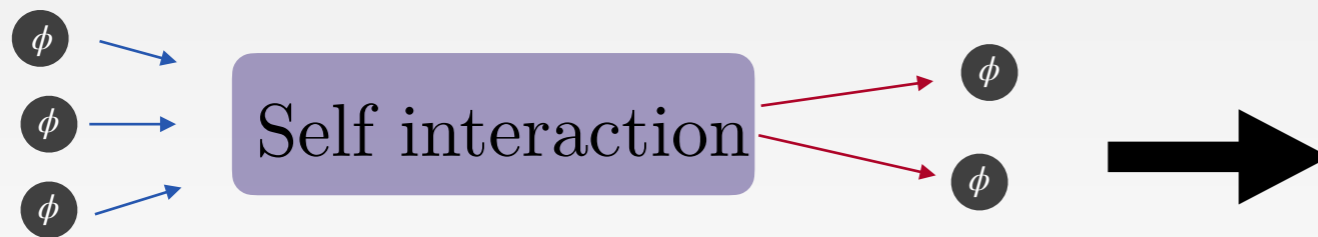
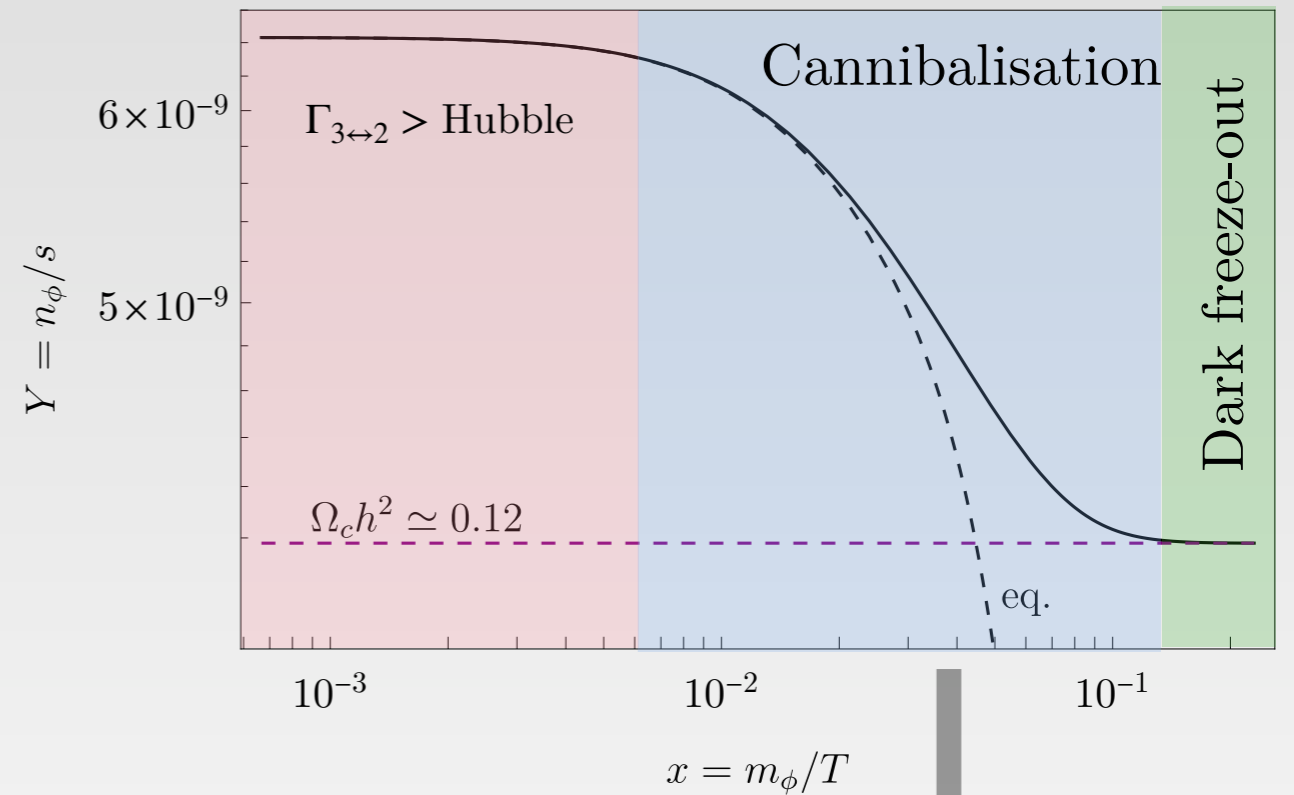
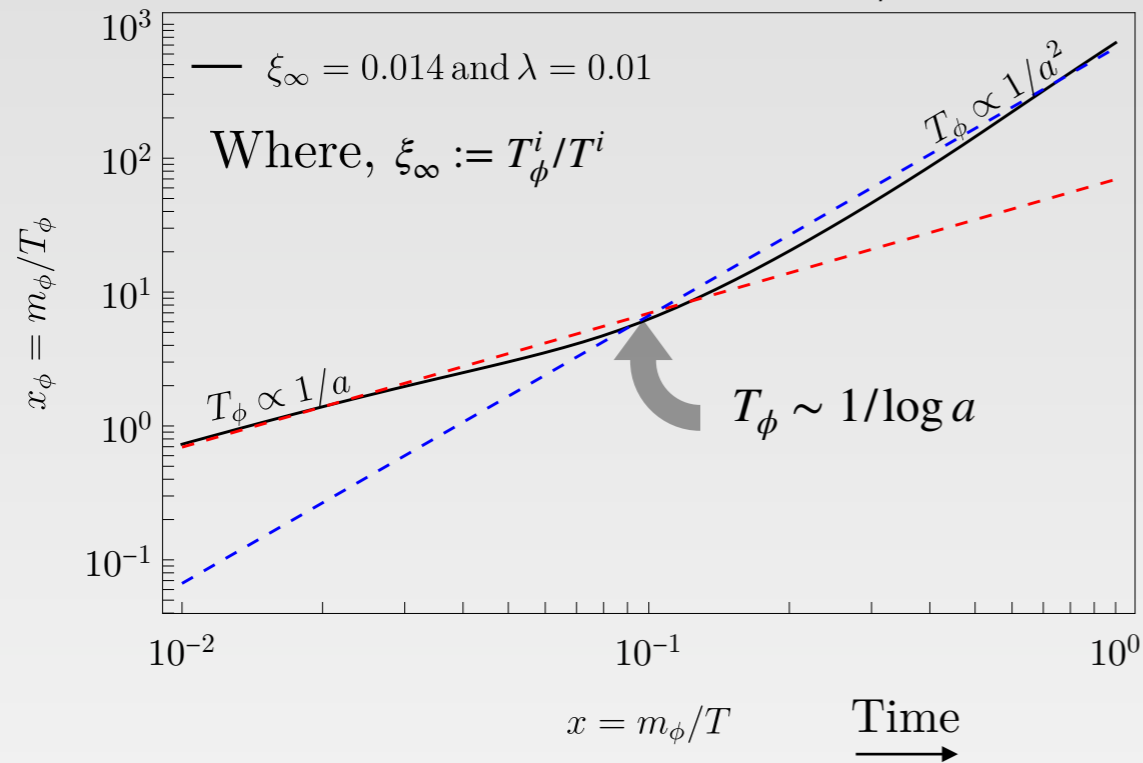


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the DM fluid **exchanges** number density for kinetic energy!

→ Is the **reverse** possible?

Reconciling portals to matter

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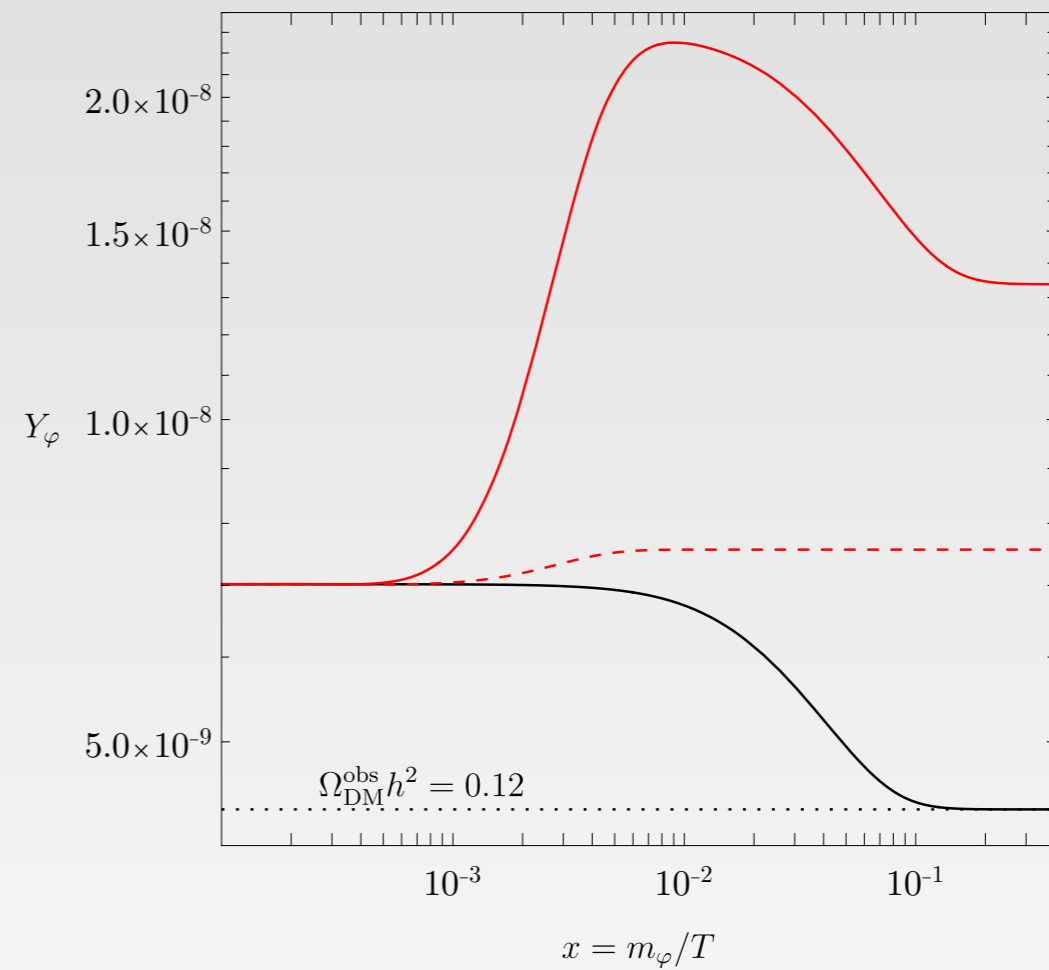
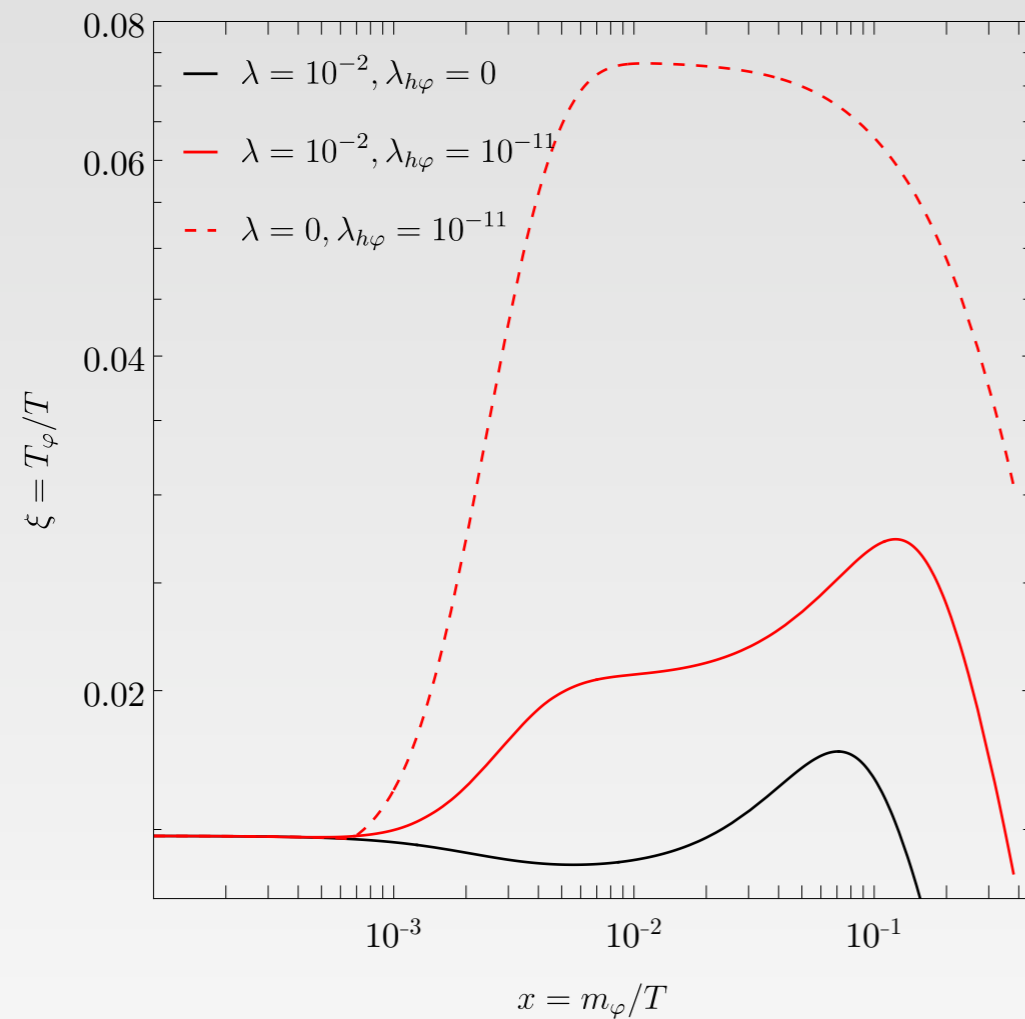
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$$V(\varphi + v_\varphi) \supset \frac{\lambda v_\varphi}{3!} \varphi^3 + \frac{\lambda}{4!} \varphi^4 + \text{Higgs mix.} \longrightarrow \tau_\varphi \text{ strongly constrained by observations.}$$

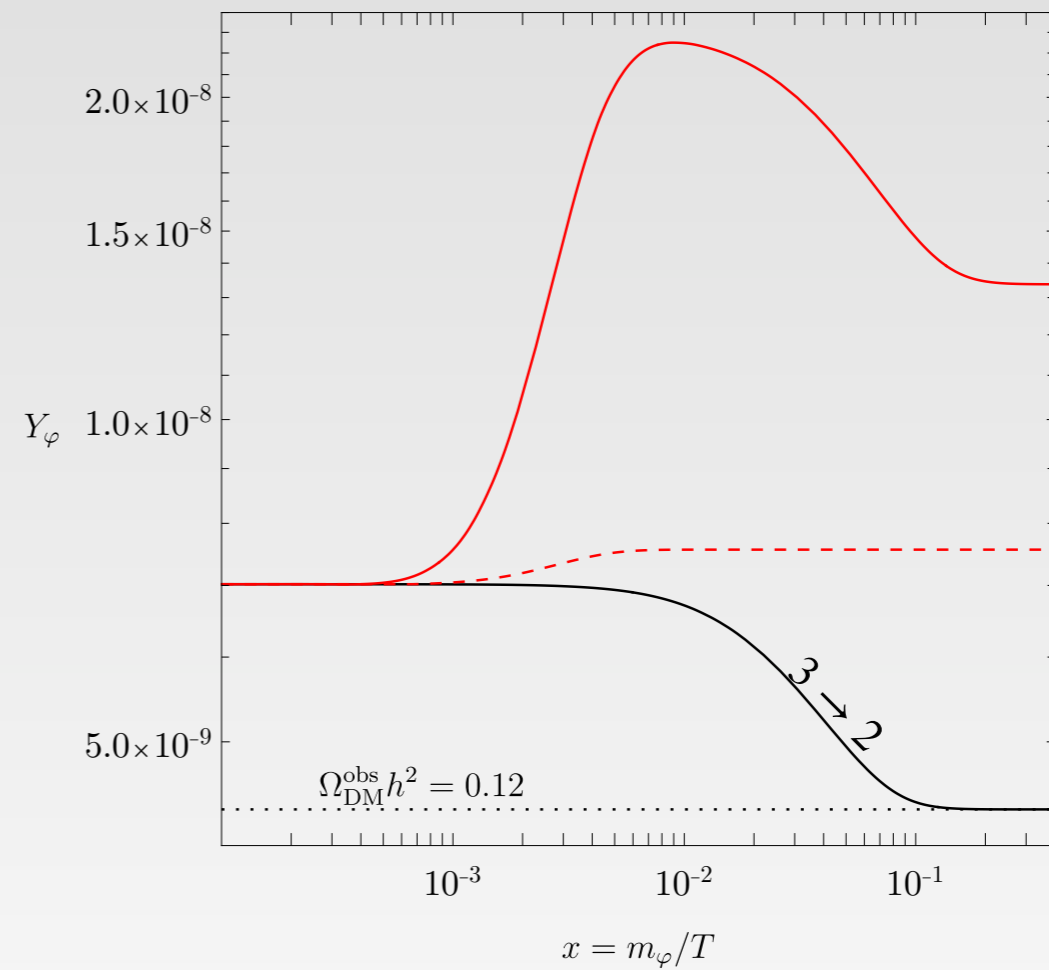
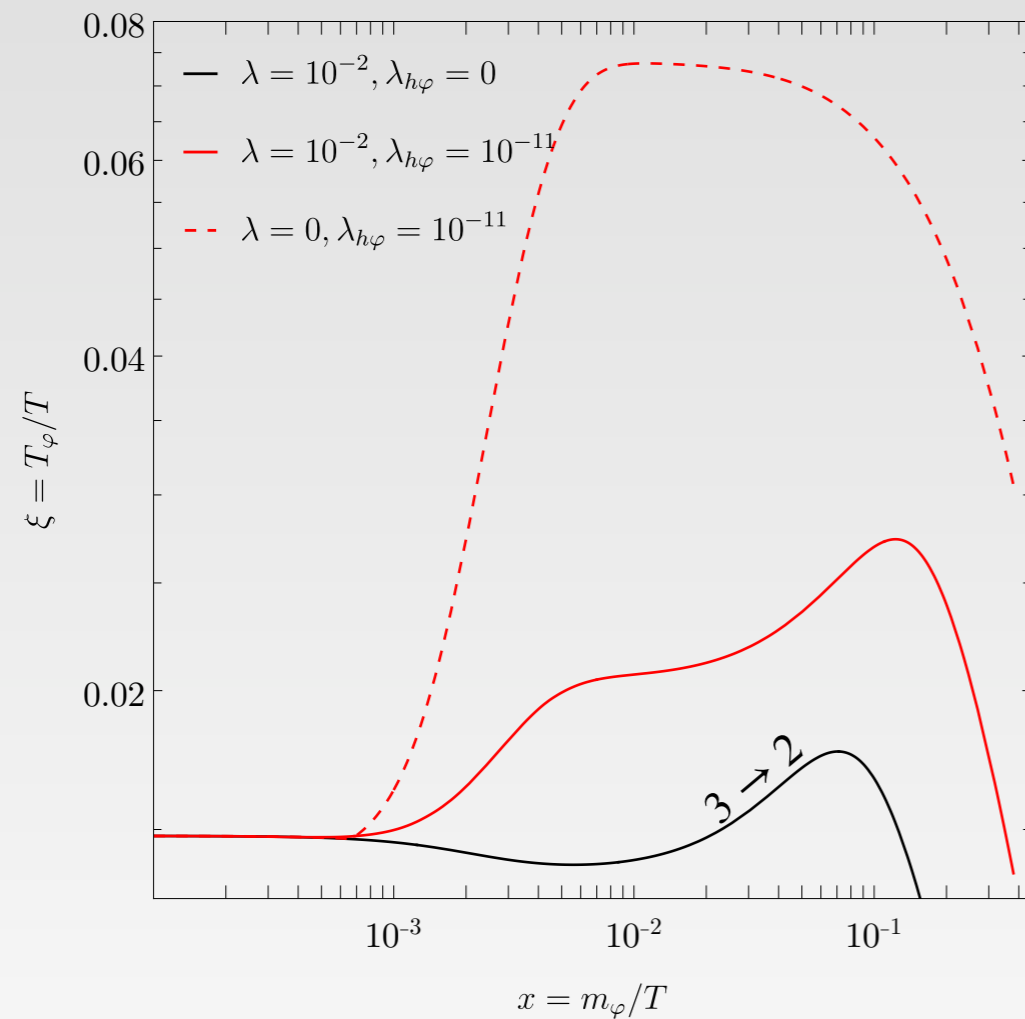
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Consider an initial cold-low populated dark sector; $T_{SM}^i = 100 T_{DM}^i$



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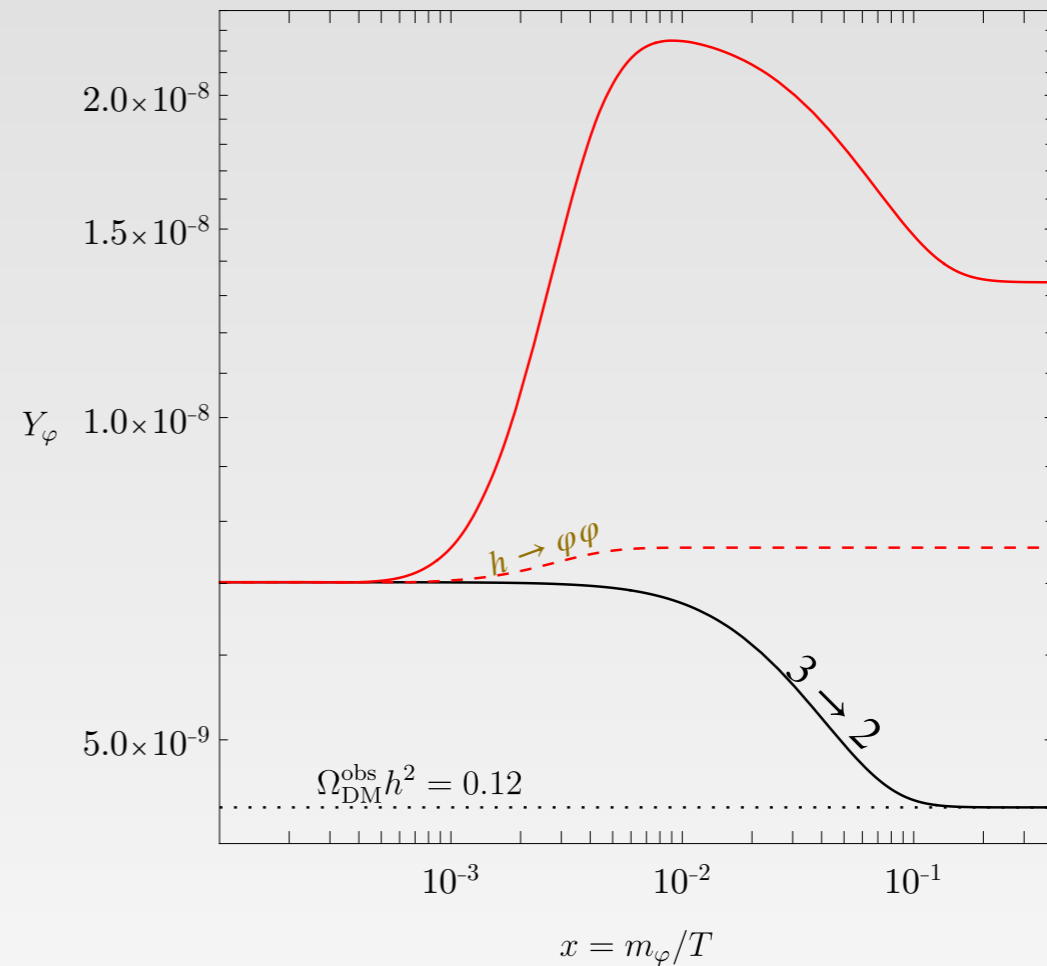
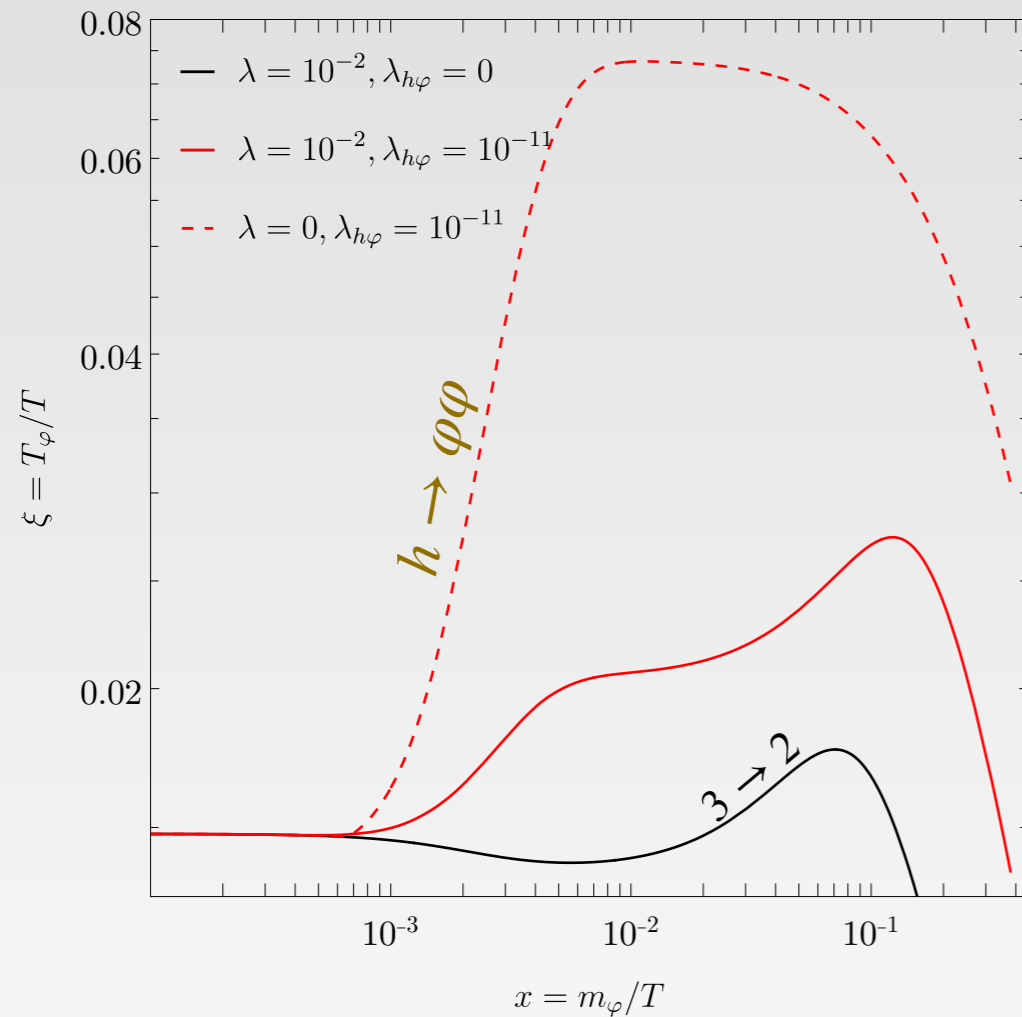
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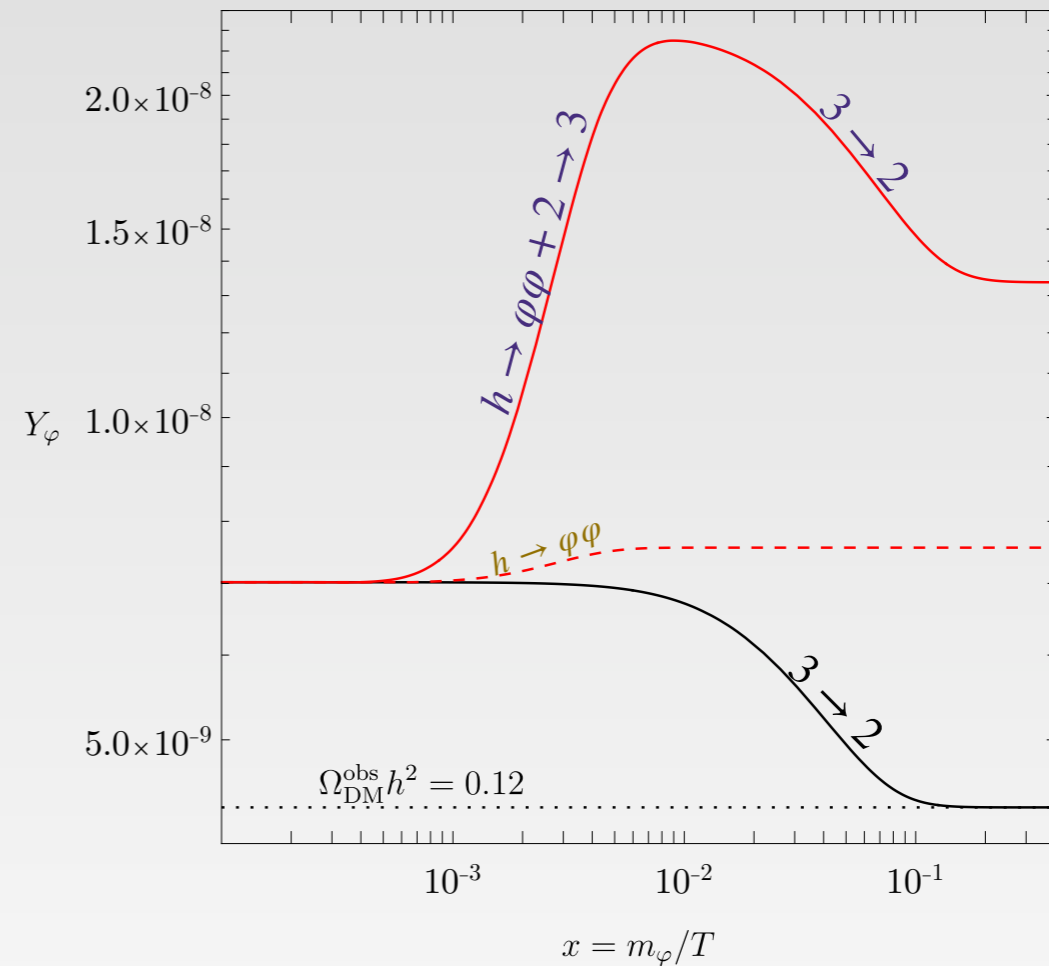
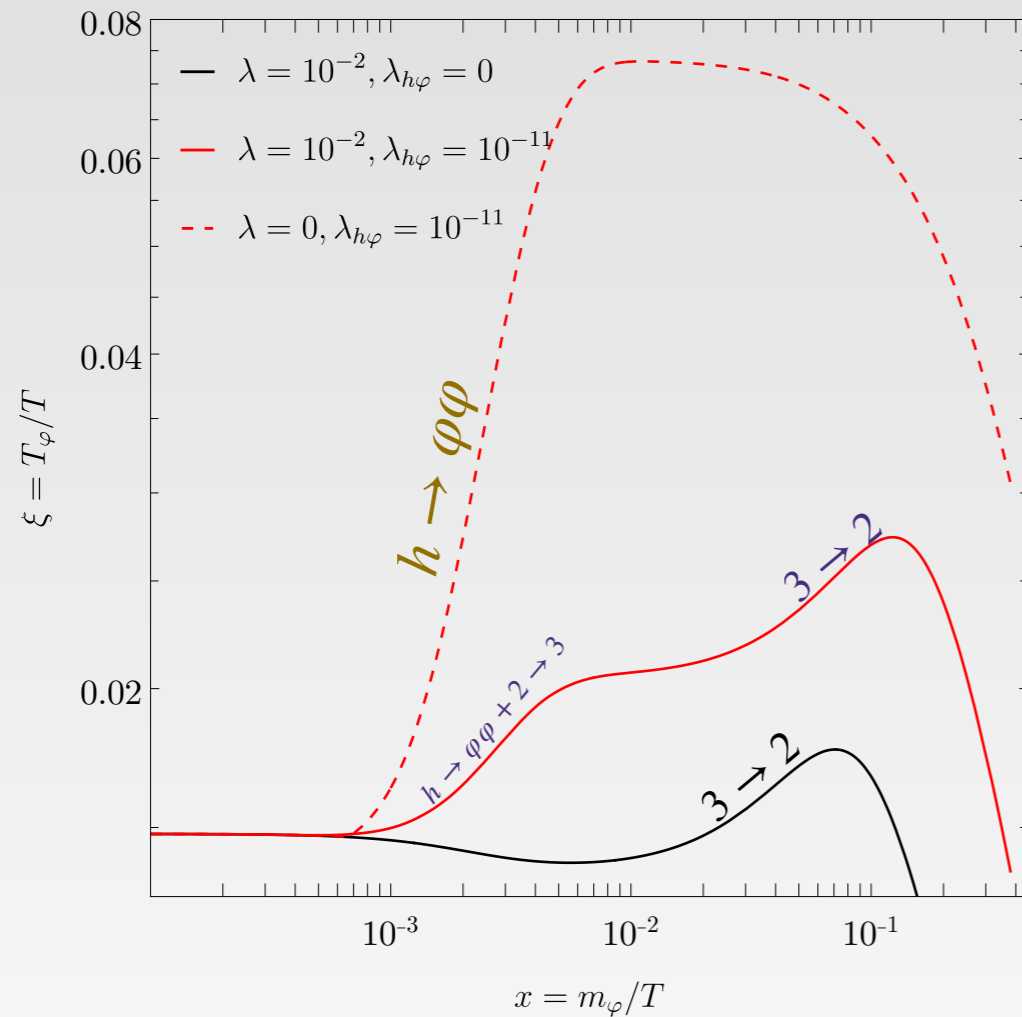


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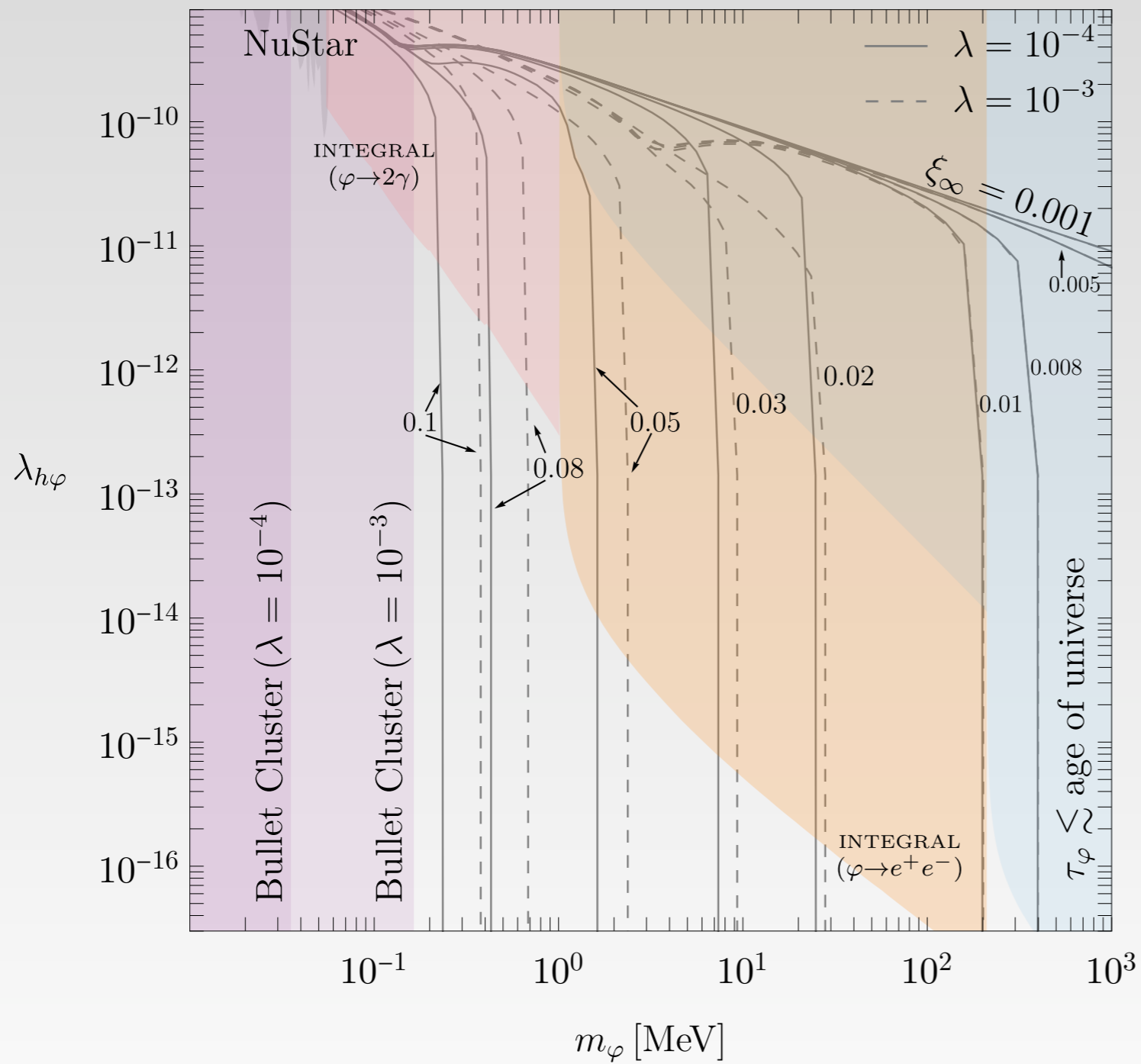


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Both

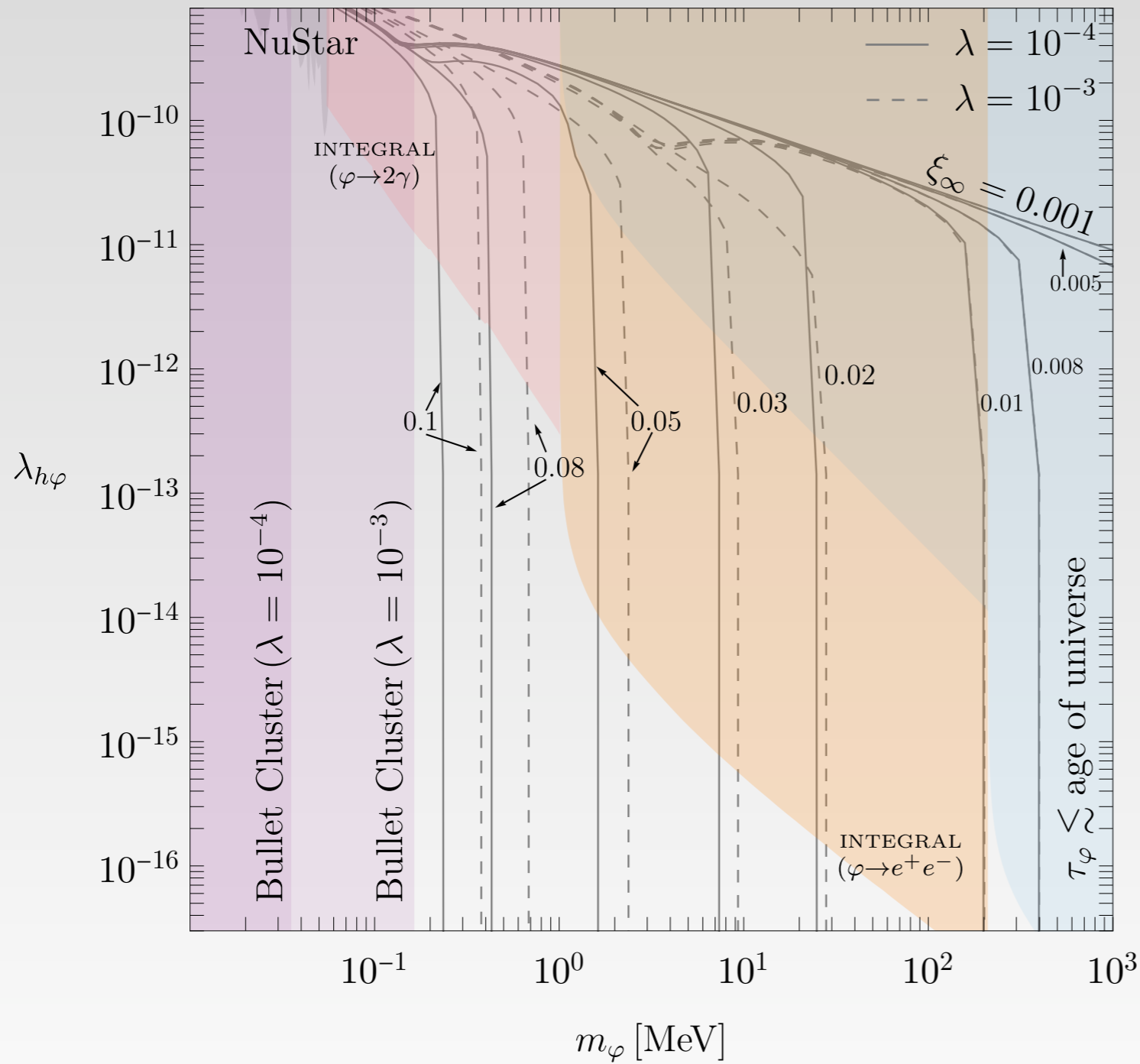
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Gray lines match observations;

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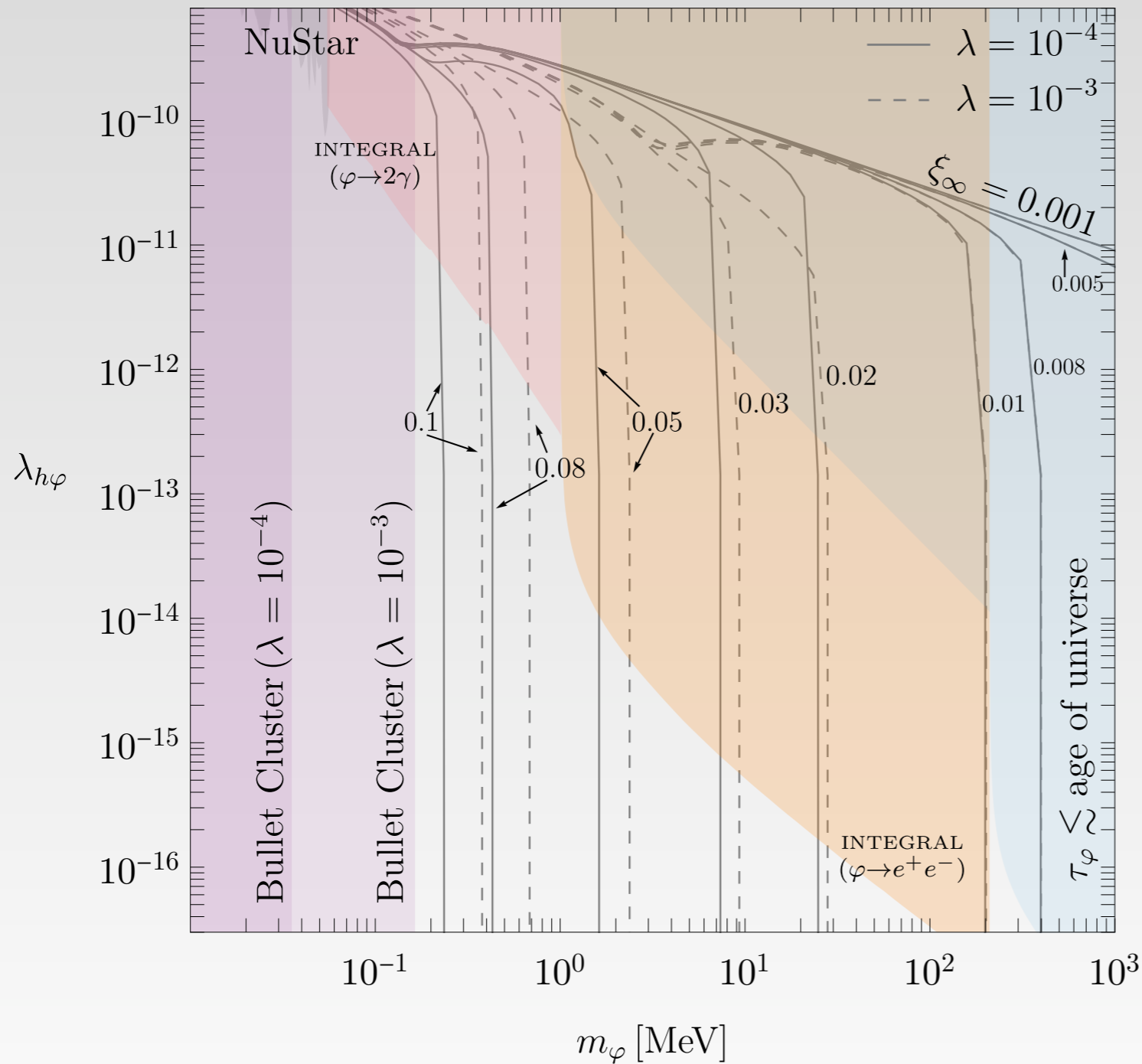


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Expected results... this mark the starting point for a more interesting scenario.

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thermalisation between
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 ϕ couples to matter as a second Higgs

Dynamics

DM dynamical evolution

$$\frac{Y'_S}{Y_S} = \frac{1}{x \tilde{H}} \left(\langle C_{h \rightarrow \phi SS^*} \rangle + \langle C_{h \rightarrow SS^*} \rangle + \langle C_{\phi\phi \leftrightarrow SS^*} \rangle + \langle C_{3 \leftrightarrow 2} \rangle \right), \text{ Amount of DM}$$

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DM temperature

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ϕ temperature

Dynamics

Freeze-in (populating the dark sector)

DM dynamical evolution

$$\frac{Y'_S}{Y_S} = \frac{1}{x \tilde{H}} \left(\langle C_{h \rightarrow \phi SS^*} \rangle + \langle C_{h \rightarrow SS^*} \rangle + \langle C_{\phi \phi \leftrightarrow SS^*} \rangle + \langle C_{3 \leftrightarrow 2} \rangle \right), \text{ Amount of DM}$$

$$-\frac{x'_S}{x_S} = \frac{1}{x \tilde{H}} \left(\langle C_{h \rightarrow \phi SS^*} \rangle_2 + \langle C_{h \rightarrow SS^*} \rangle_2 + \langle C_{\phi S \leftrightarrow \phi S} \rangle_2 + \langle C_{3 \leftrightarrow 2} \rangle_2 \right) - \frac{Y'_S}{Y_S} + \frac{H}{x \tilde{H}} \frac{\langle p^4/E^3 \rangle}{3T_S} + \frac{2s'}{3s}$$

DM temperature

Mediator dynamical evolution

$$\frac{Y'_\phi}{Y_\phi} = \frac{1}{x \tilde{H}} \left(\langle C_{h \rightarrow \phi SS^*} \rangle + \langle C_{sm \ sm \rightarrow sm \ \phi} \rangle + \langle C_{\phi \phi \leftrightarrow SS^*} \rangle \right), \text{ Amount of mediators}$$

$$-\frac{x'_\phi}{x_\phi} = \frac{1}{x \tilde{H}} \left(\langle C_{h \rightarrow \phi SS^*} \rangle_2 + \langle C_{sm \ sm \rightarrow sm \ \phi} \rangle_2 + \langle C_{\phi S \leftrightarrow \phi S} \rangle_2 \right) - \frac{Y'_\phi}{Y_\phi} + \frac{H}{x \tilde{H}} \frac{\langle p^4/E^3 \rangle}{3T_\phi} + \frac{2s'}{3s}.$$

ϕ temperature

Dynamics

DM dynamical evolution

Freeze-in (populating the dark sector)

DM-mediator interactions

$$\frac{Y'_S}{Y_S} = \frac{1}{x \tilde{H}} \left(\langle C_{h \rightarrow \phi SS^*} \rangle + \langle C_{h \rightarrow SS^*} \rangle + \langle C_{\phi \phi \leftrightarrow SS^*} \rangle + \langle C_{3 \leftrightarrow 2} \rangle \right), \text{ Amount of DM}$$

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DM temperature

Mediator dynamical evolution

$$\frac{Y'_\phi}{Y_\phi} = \frac{1}{x \tilde{H}} \left(\langle C_{h \rightarrow \phi SS^*} \rangle + \langle C_{sm \ sm \rightarrow sm \ \phi} \rangle + \langle C_{\phi \phi \leftrightarrow SS^*} \rangle \right), \text{ Amount of mediators}$$

$$-\frac{x'_\phi}{x_\phi} = \frac{1}{x \tilde{H}} \left(\langle C_{h \rightarrow \phi SS^*} \rangle_2 + \langle C_{sm \ sm \rightarrow sm \ \phi} \rangle_2 + \langle C_{\phi S \leftrightarrow \phi S} \rangle_2 \right) - \frac{Y'_\phi}{Y_\phi} + \frac{H}{x \tilde{H}} \frac{\langle p^4/E^3 \rangle}{3T_\phi} + \frac{2s'}{3s}.$$

ϕ temperature

Dynamics

DM dynamical evolution

- Freeze-in (populating the dark sector)
- DM-mediator interactions
- DM self-interactions

$$\frac{Y'_S}{Y_S} = \frac{1}{x \tilde{H}} \left(\langle \mathcal{C}_{h \rightarrow \phi SS^*} \rangle + \langle \mathcal{C}_{h \rightarrow SS^*} \rangle + \langle \mathcal{C}_{\phi \phi \leftrightarrow SS^*} \rangle + \langle \mathcal{C}_{3 \leftrightarrow 2} \rangle \right), \text{ Amount of DM}$$

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DM temperature

Mediator dynamical evolution

$$\frac{Y'_\phi}{Y_\phi} = \frac{1}{x \tilde{H}} \left(\langle \mathcal{C}_{h \rightarrow \phi SS^*} \rangle + \langle \mathcal{C}_{sm \ sm \rightarrow sm \ \phi} \rangle + \langle \mathcal{C}_{\phi \phi \leftrightarrow SS^*} \rangle \right), \text{ Amount of mediators}$$

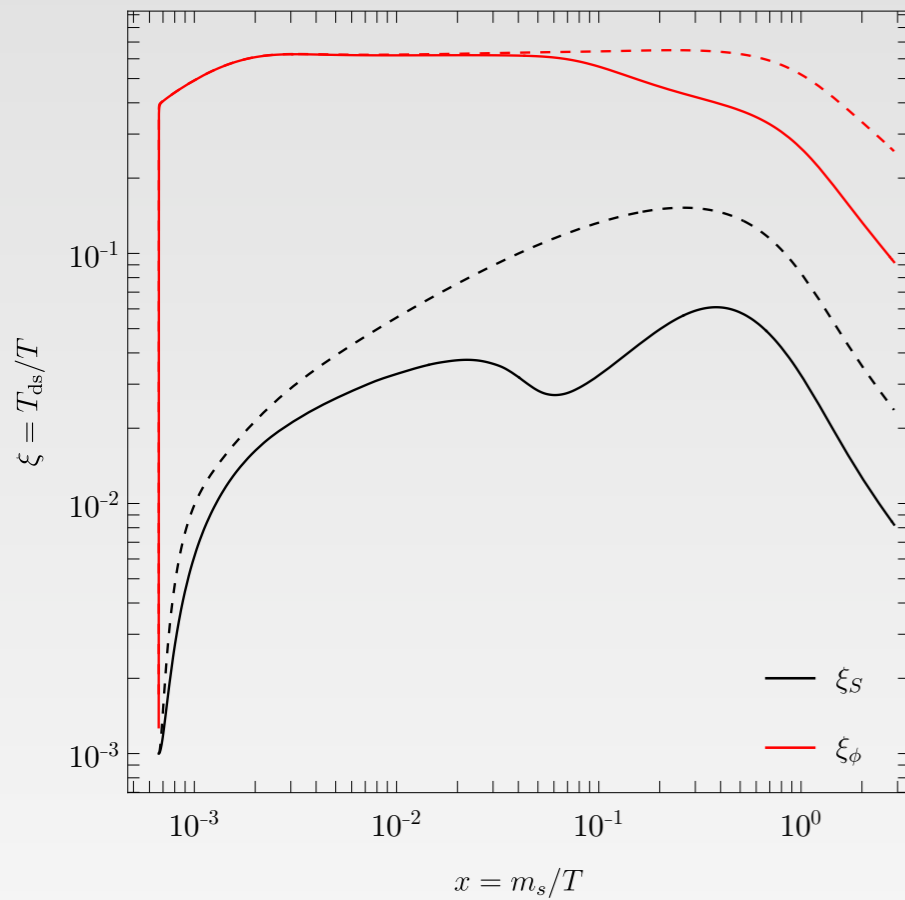
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ϕ temperature

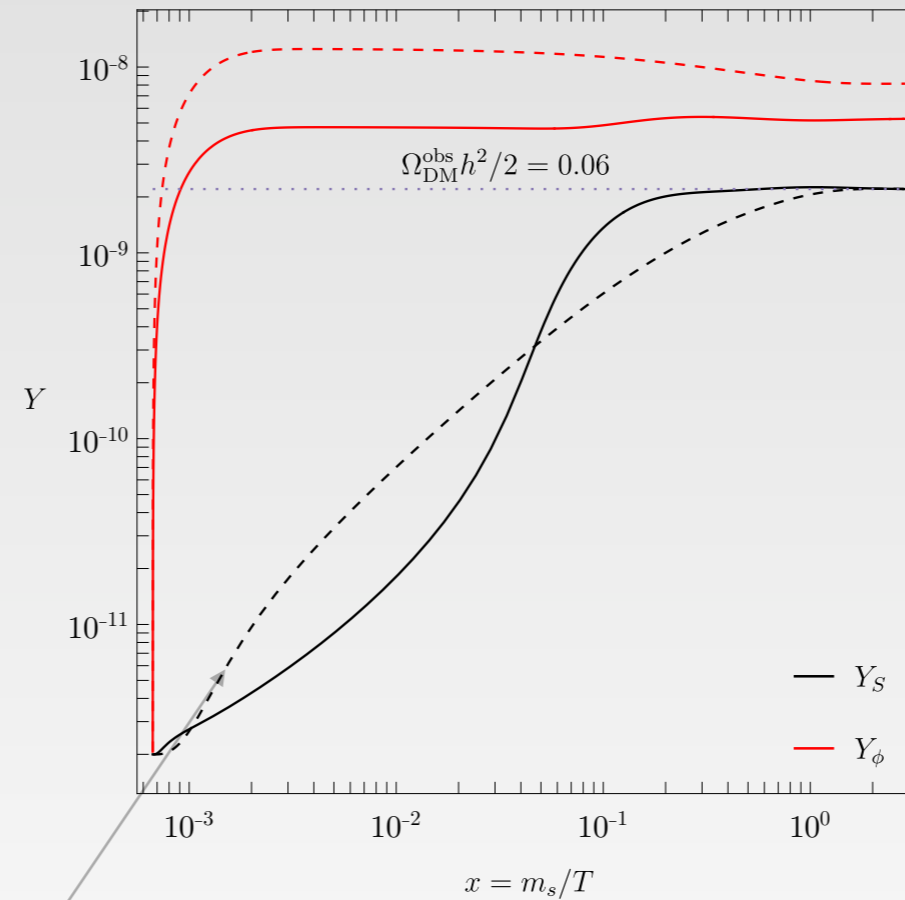
Dynamics

Benchmark point with
 $m_\phi < m_s$ and parameters that
match Ω_{DM}^{obs}

Temperature evolution



Yield evolution

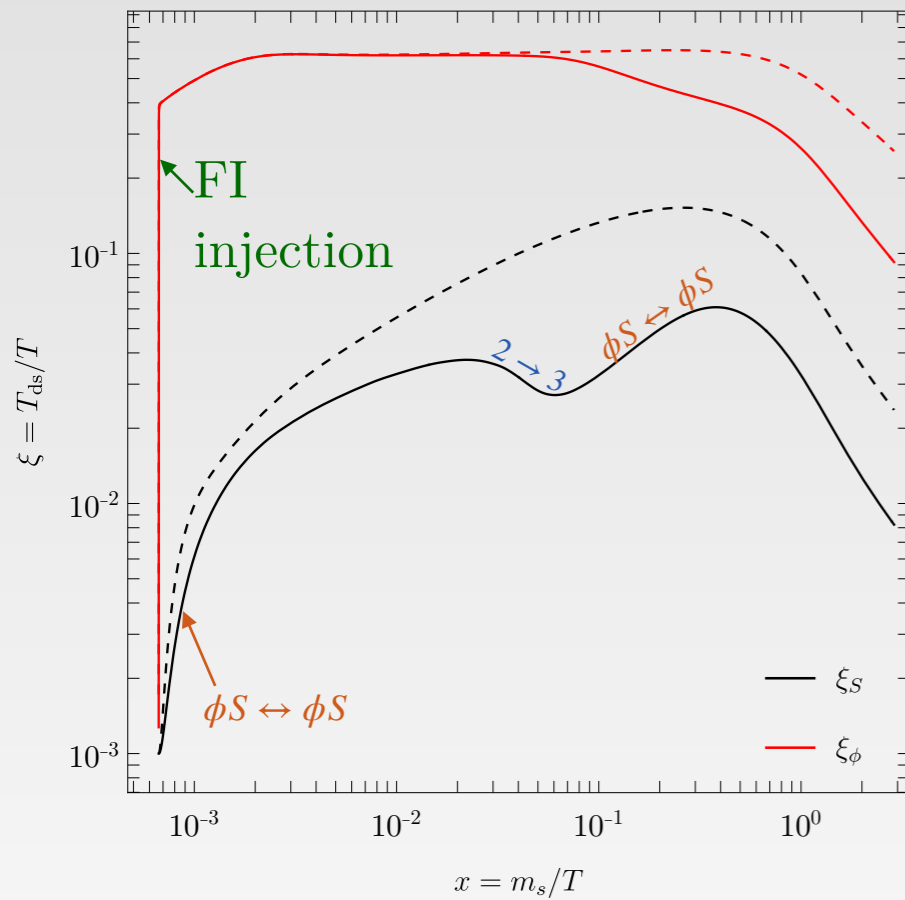


Dashed lines display the evolution **without** self interactions

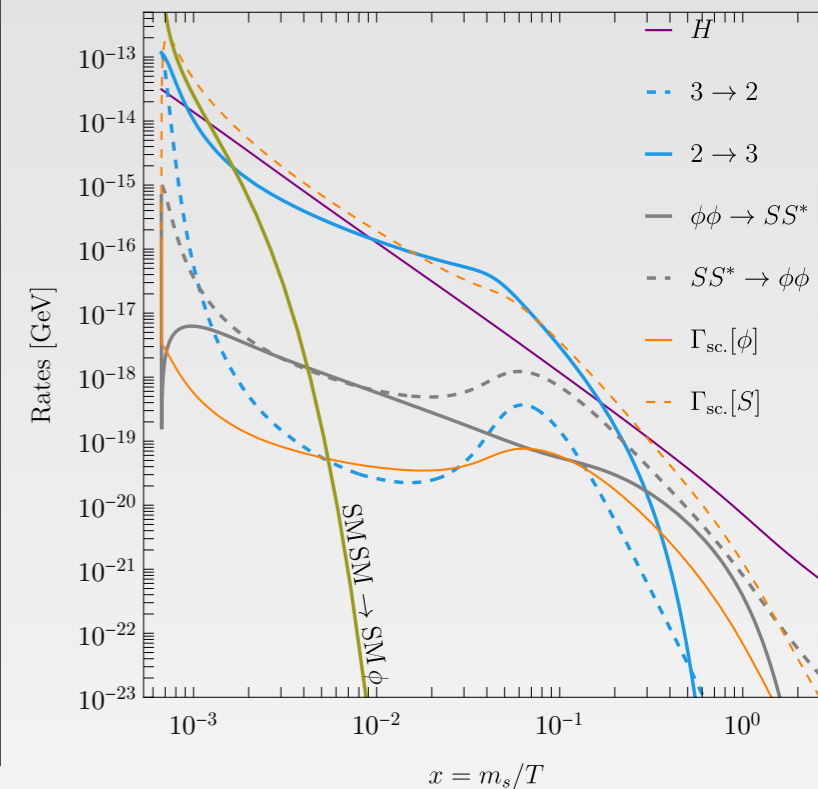
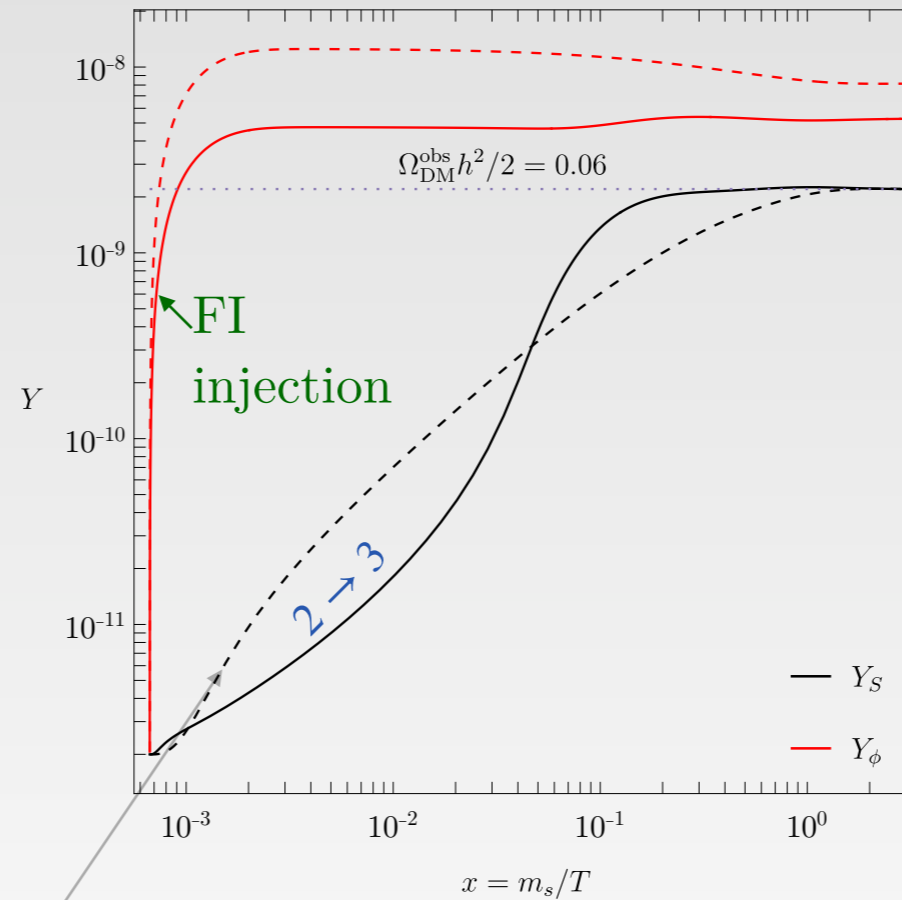
Dynamics

Benchmark point with
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Yield evolution

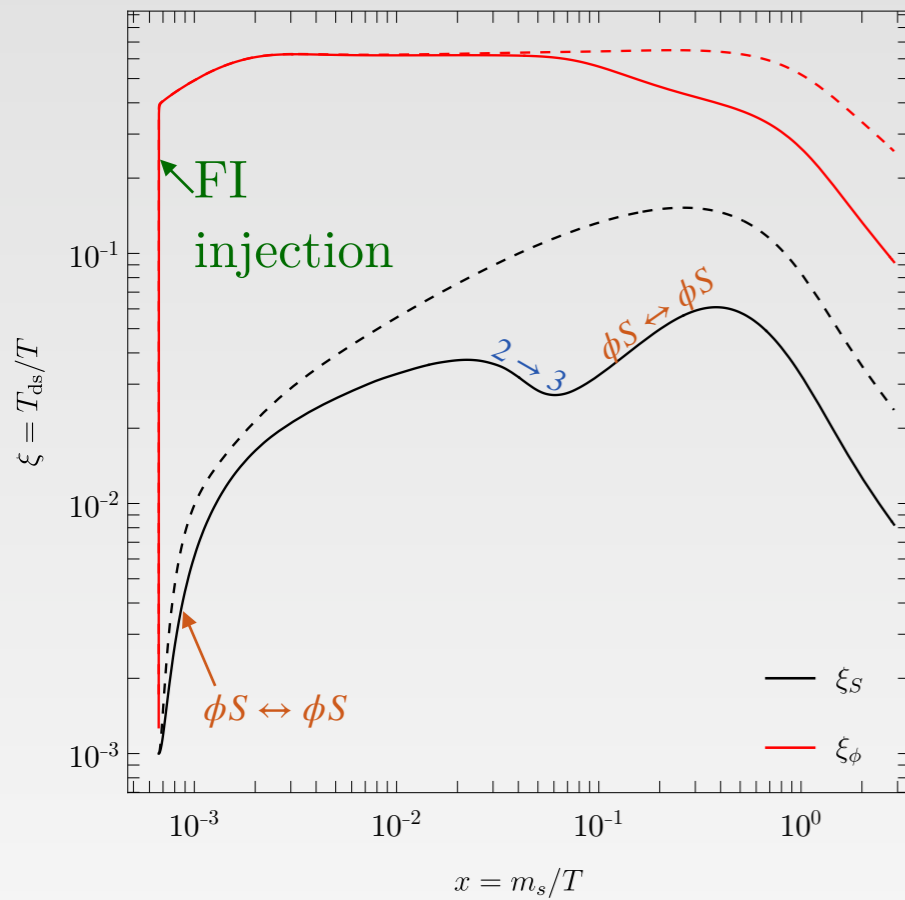


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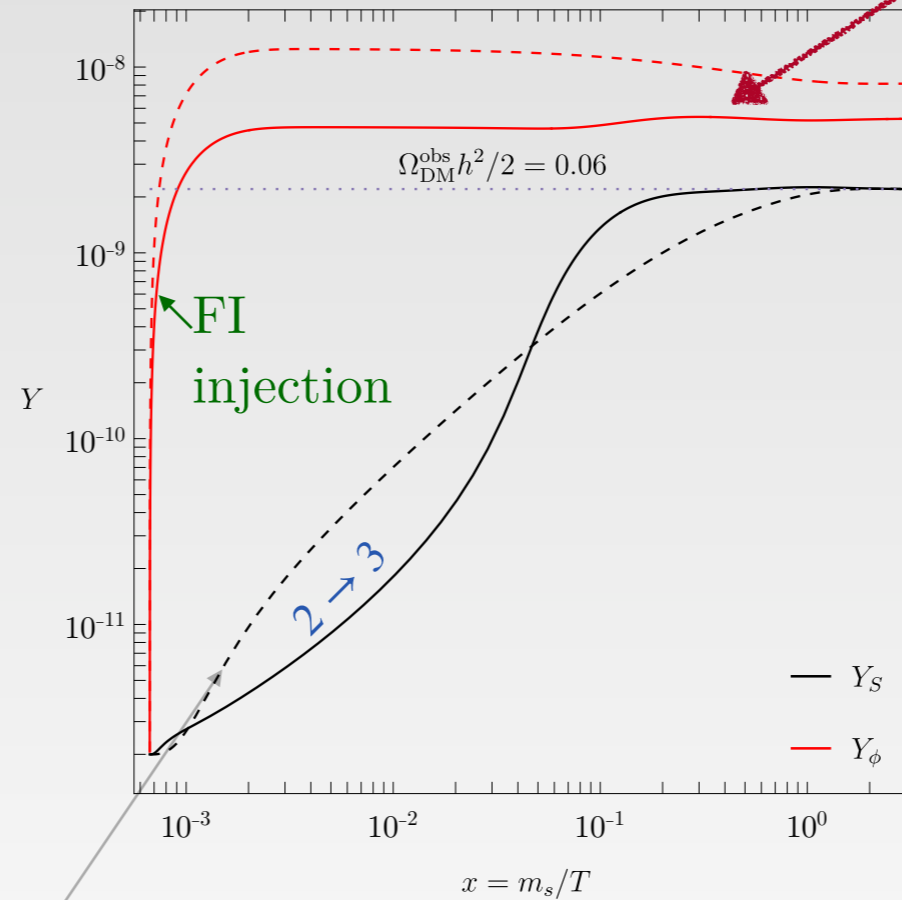
Dynamics

Benchmark point with $m_\phi < m_s$ and parameters that match Ω_{DM}^{obs}

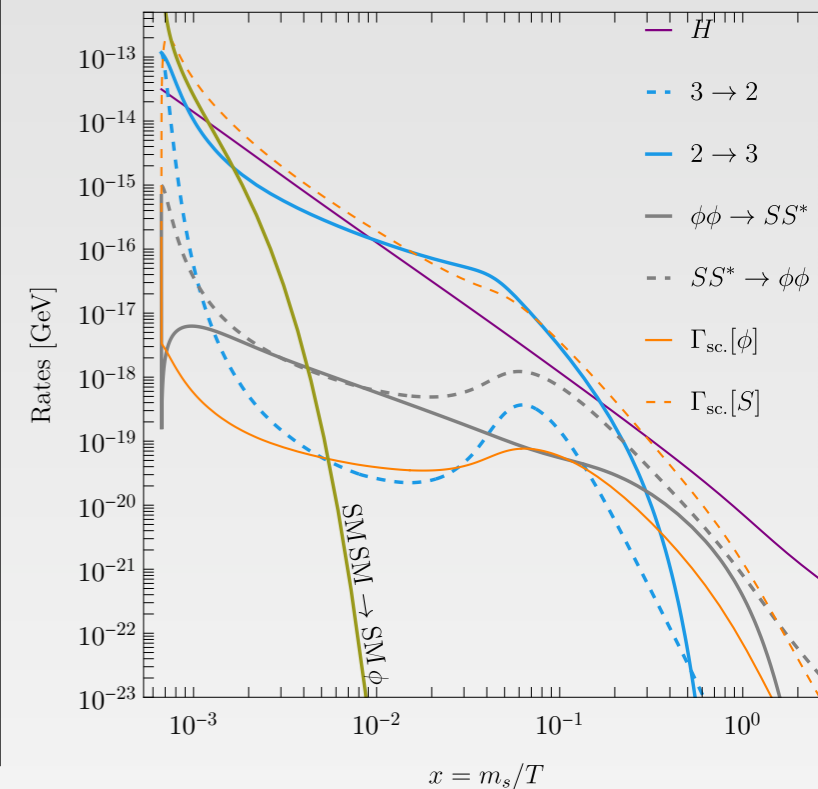
Temperature evolution



Yield evolution



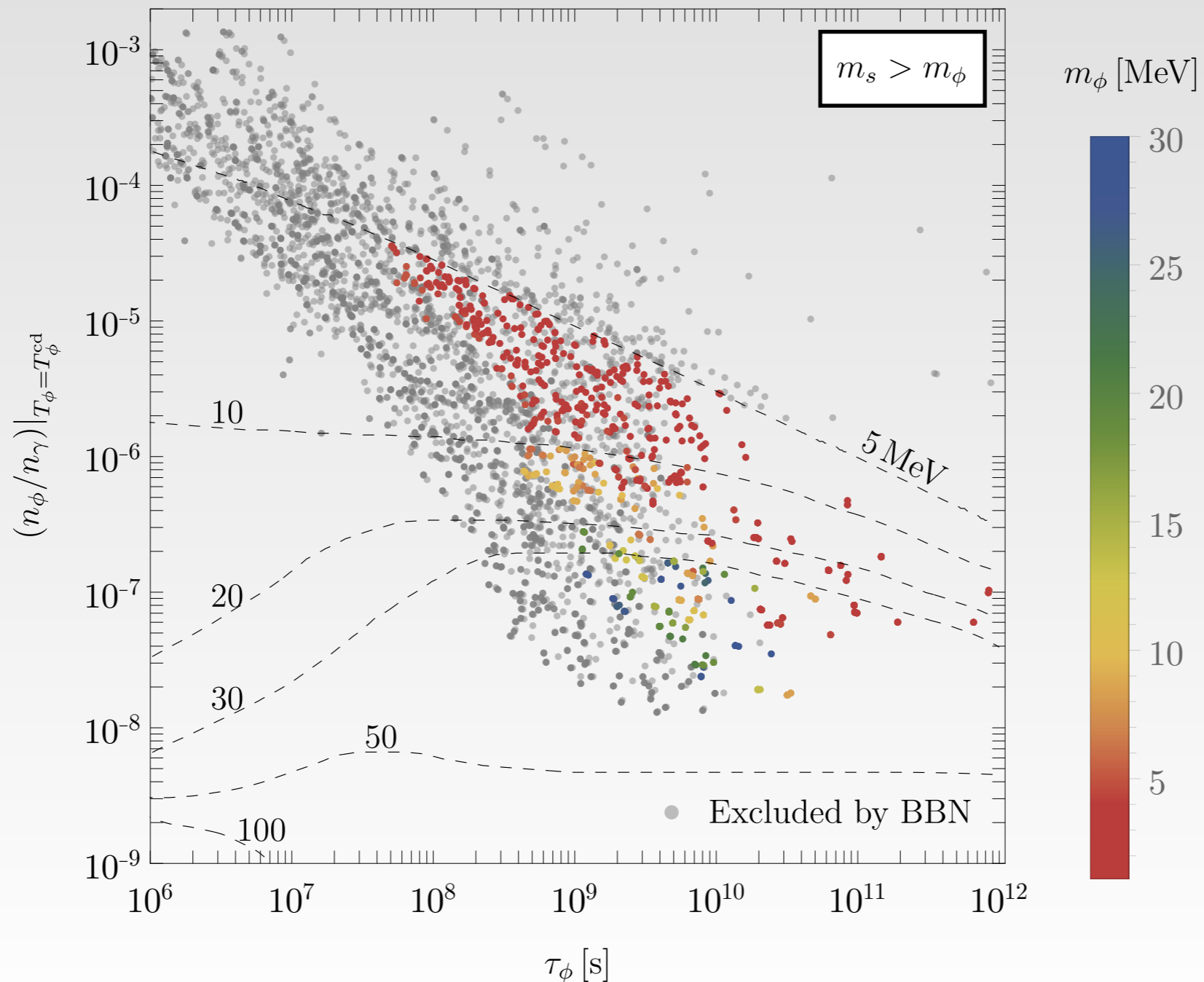
Significant final abundance of ϕ



Dashed lines display the evolution **without** self interactions

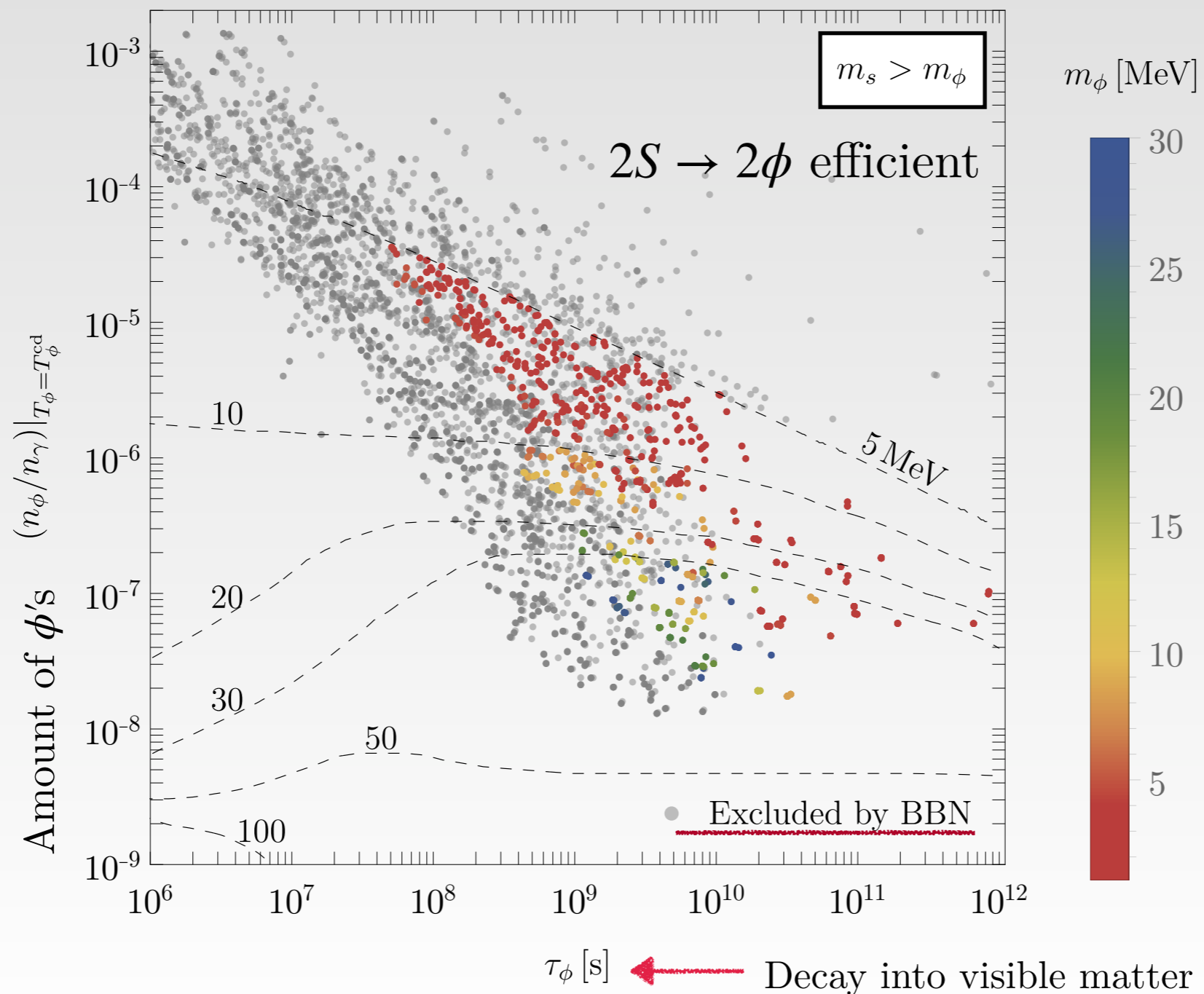
Constraints on mediator

During Big Bang Nucleosynthesis ($T_{SM} \sim 150$ MeV) nucleons combined. Presence of mediators may **expoil** BBN observations:



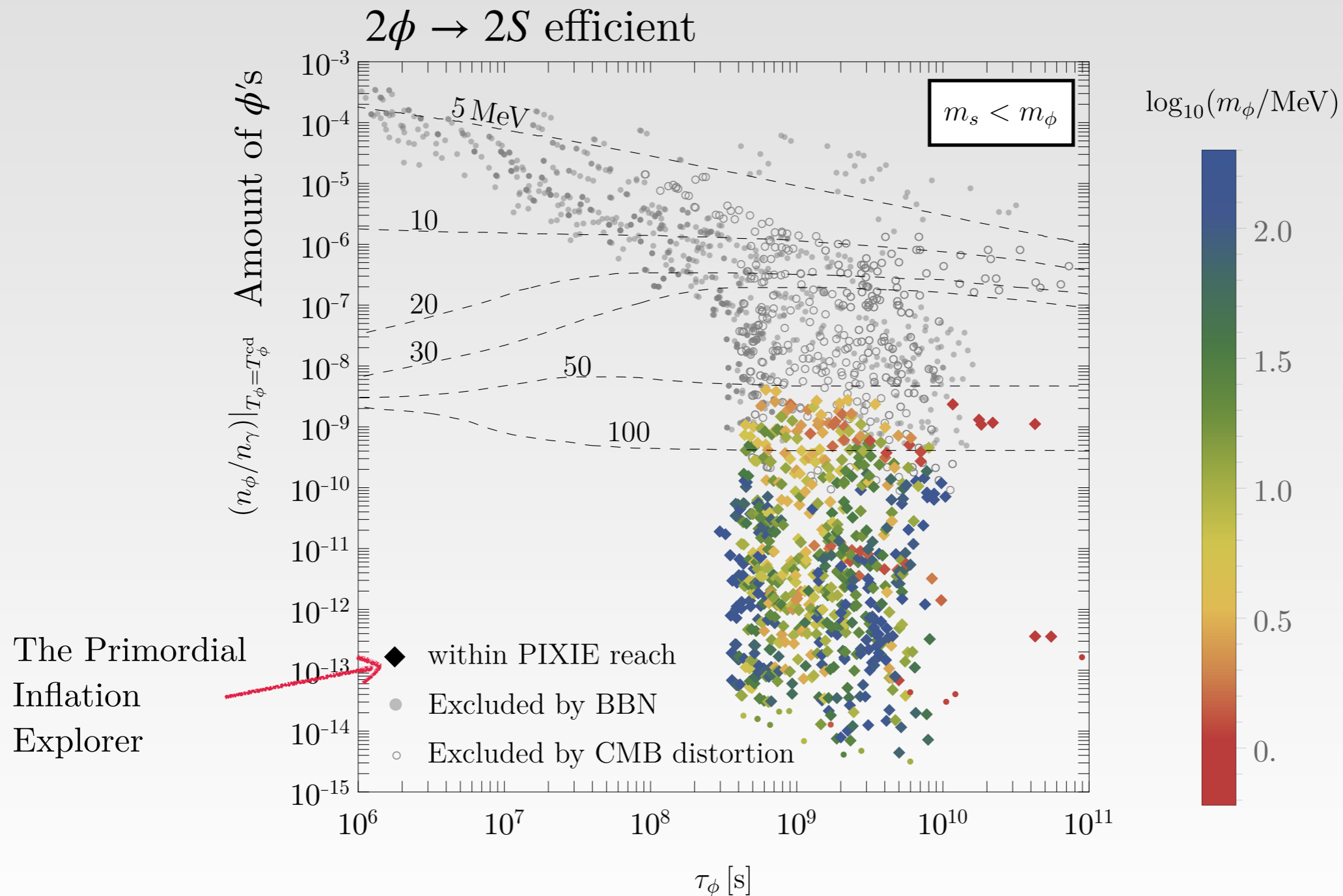
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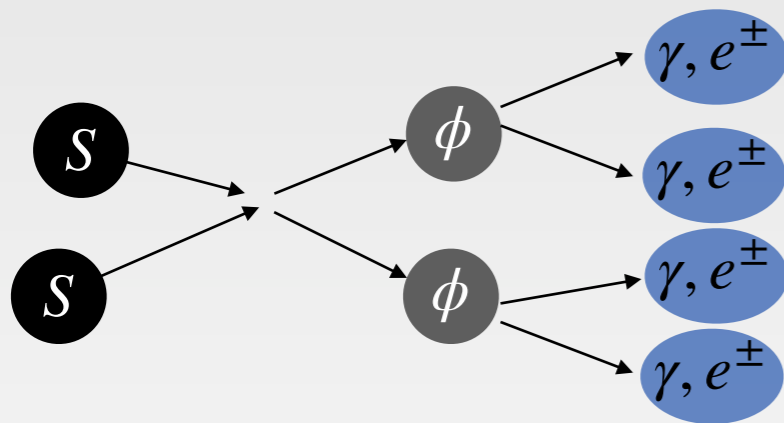
Constraints on mediator

Reversing the mass hierarchy ($m_S < m_\phi$) leads to **less constrained** results



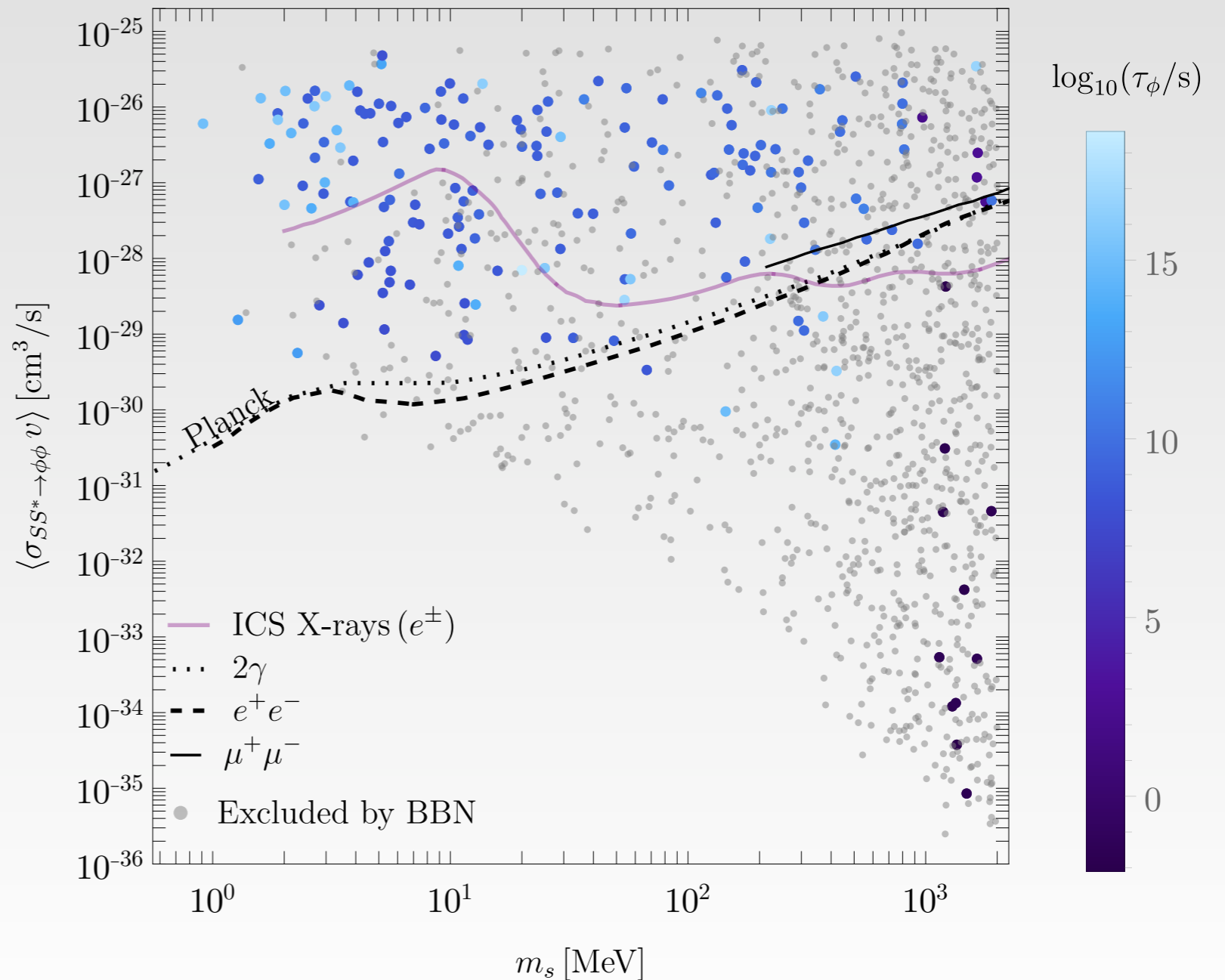
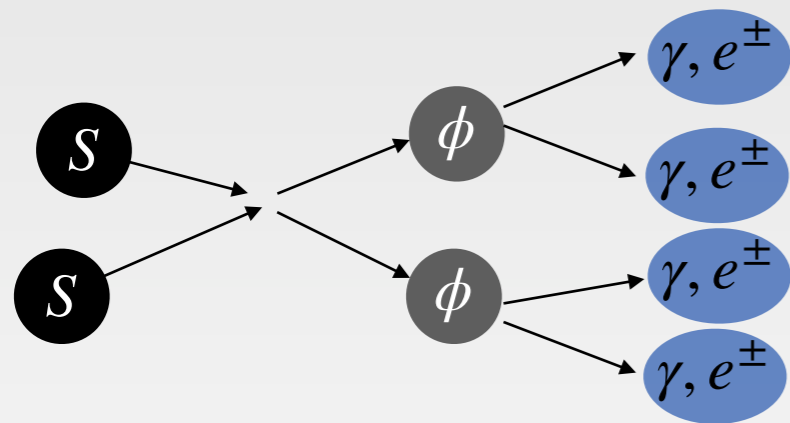
Indirectly observing DM

If DM is heavier, it can efficiently annihilate into mediators leading to detectability via telescopes:



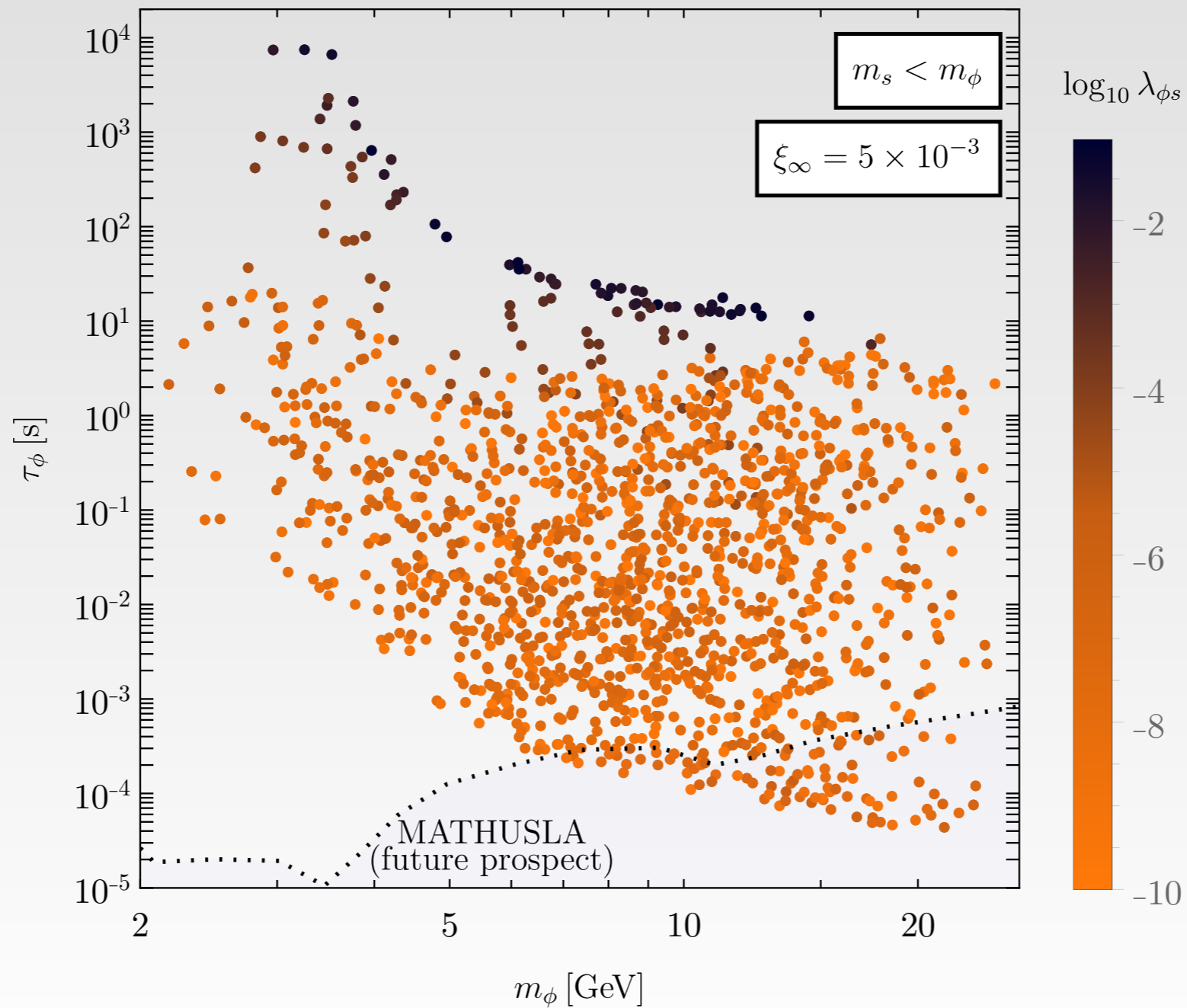
Indirectly observing DM

If DM is heavier, it can efficiently annihilate into mediators leading to detectability via telescopes:



Collider's signals

GeV DM could be detected at the **LHC**:



Summary

- In the Freeze-in mechanism, we have explicitly shown that self interactions lead to a **colder** dark sector, while **boosting** the production of DM;

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Summary

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- $m_s > m_\phi$ leads to signatures in **Planck**, the latter in **PIXIE** and **MATHUSLA**;
- Interactions between the DM and the mediator lead to **different phenomenology**.

Thank you for your attention:)