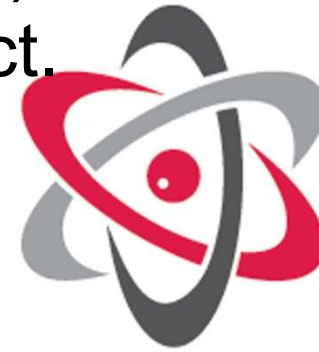
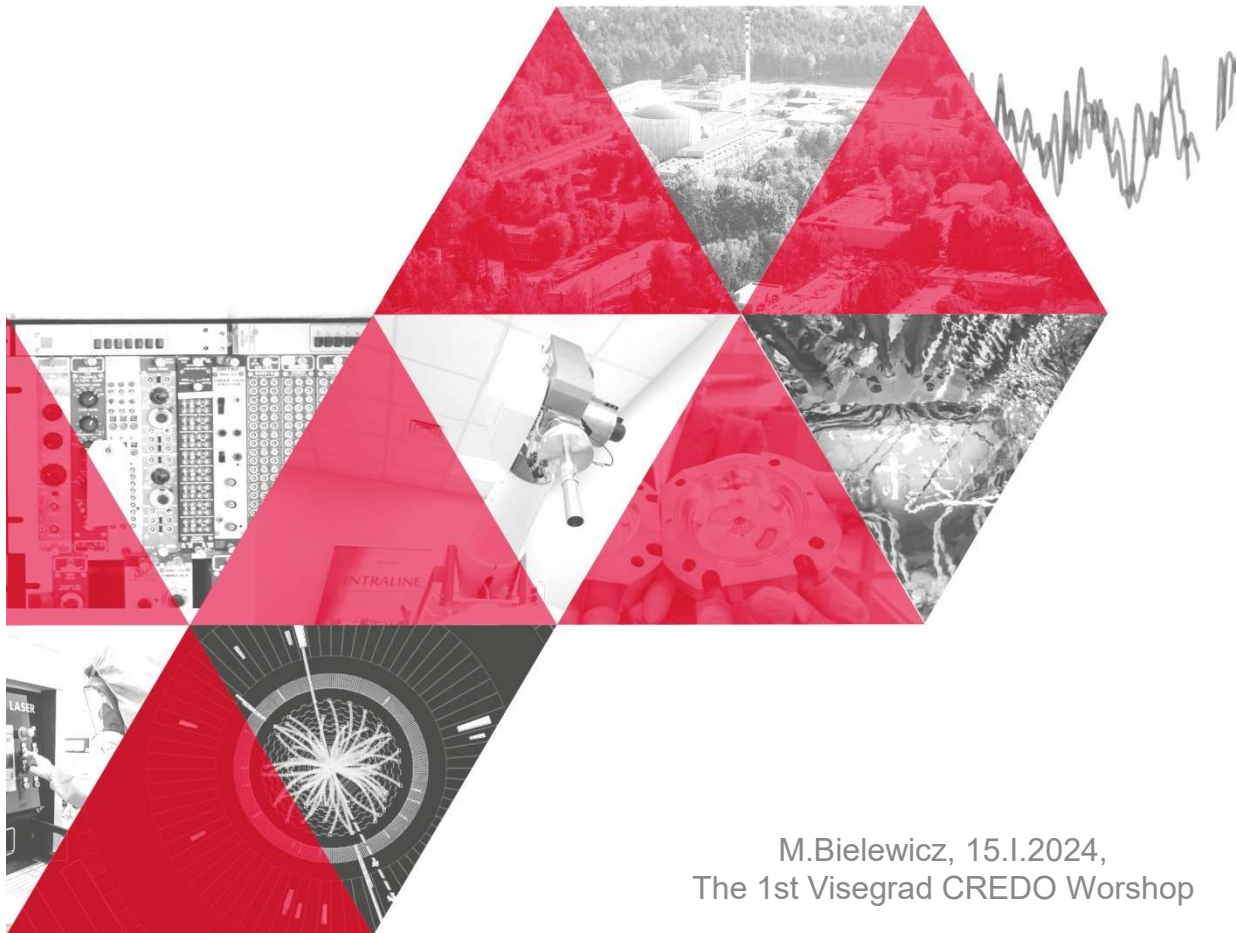


# Modular Cosmic Ray Detector (MCORD) possible use in the cosmo-seismic project.

**Dr Marcin Bielewicz**  
Nacional Centre for Nuclear Research



**NARODOWE  
CENTRUM  
BADAŃ  
JĄDROWYCH  
ŚWIERK**

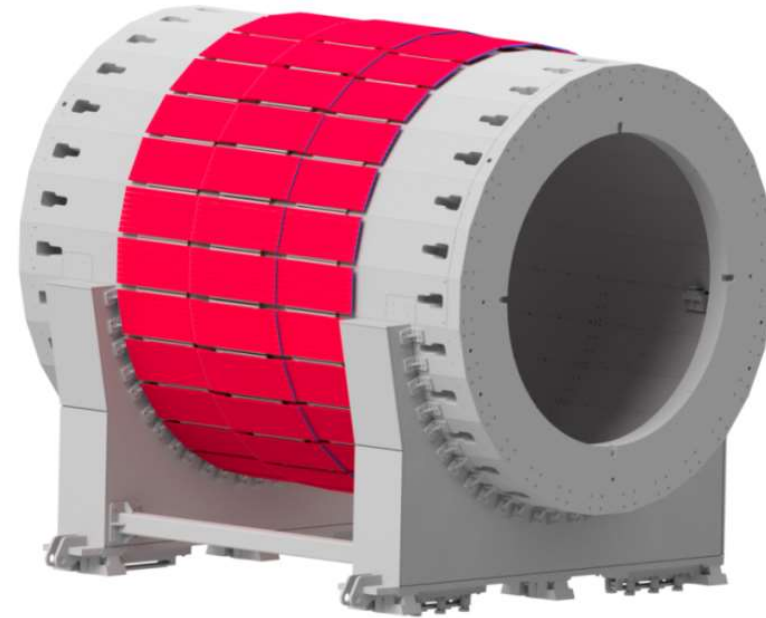


M.Bielewicz, 15.I.2024,  
The 1st Visegrad CREDO Workshop

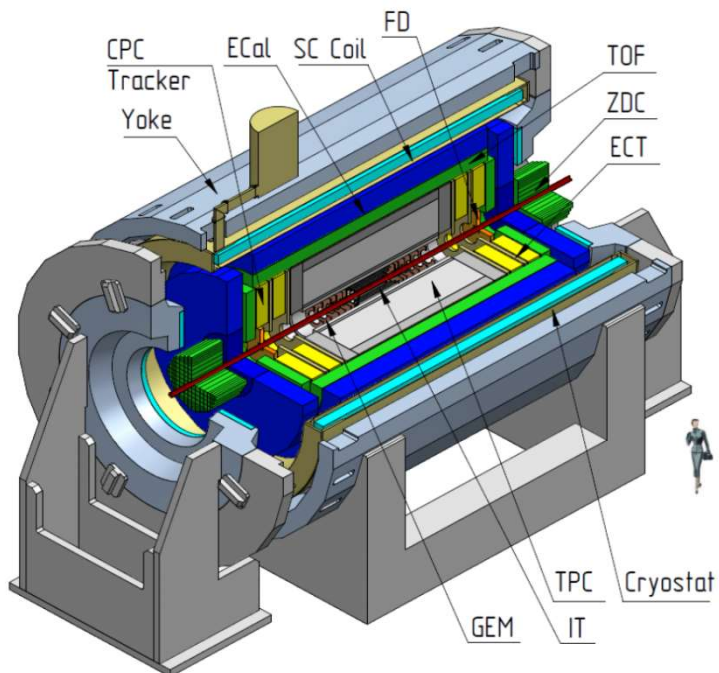
# Outline



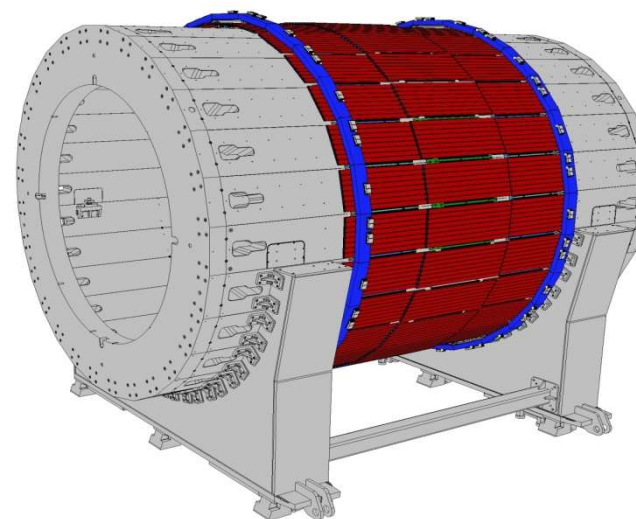
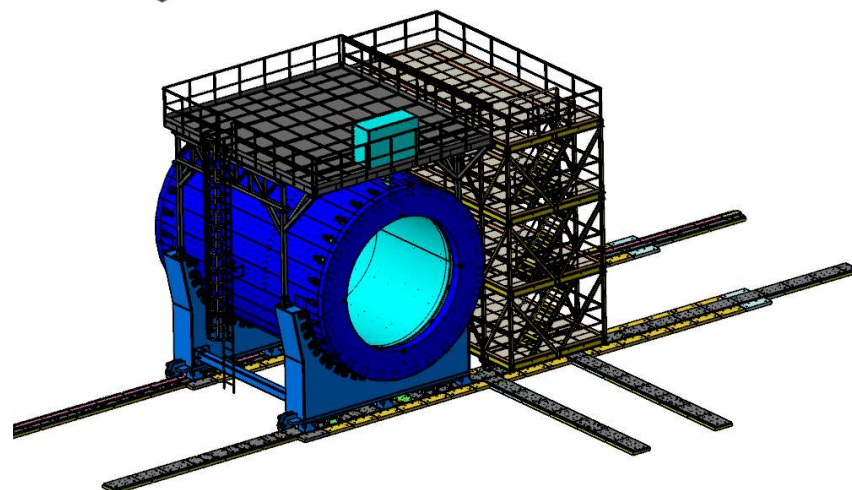
- 1. Introduction**
- 2. Detector and Electronic**
- 3. Laboratory tests**
- 4. CREDO - Mexico**
- 5. Summary**



# 1. Genesis of MCORD



- FD Forward detector
- Superconductor solenoid (SC Coil)
- Inner Tracker (IT)
- Straw-tube Tracker (ECT)
- **Time-projection chamber (TPC)**
- **Time-of-Flight system (TOF)**
- **Electromagnetic calorimeter (EMC - ECal)**
- Zero degree calorimeter (ZDC).
- Cosmic Ray Detector (MCORD)



# 1. Examples of MCORD application



1. Trigger for cosmic muons for:
  - laboratory tests of different subsystems (2 separate MCORD sections)
  - Cosmic calibration in off-beam time
2. Muon identifier (from collision):
  - pions and kaons decays
  - rare mesons decays ( $\eta$ ,  $\rho$ )
3. Astrophysics (muon showers and bundles)
  - Identification of extremely high energy particle sources
  - Sensitivity for horizontal events
  - Earthquake correlation
4. Modular construction – easy upgrade and/or alternative use

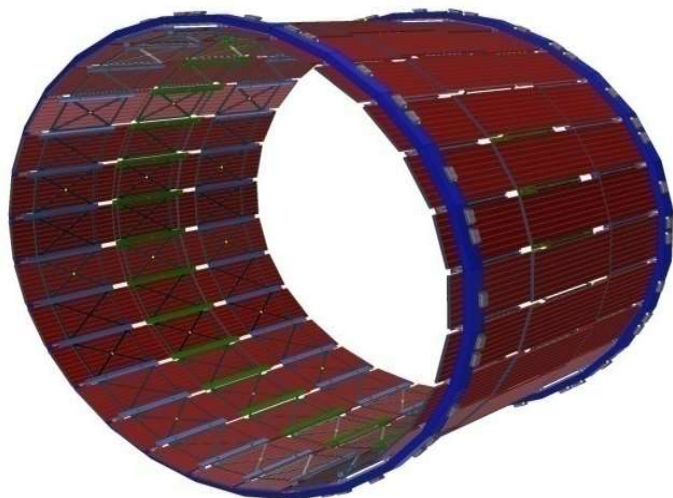


# 1. Modular construction

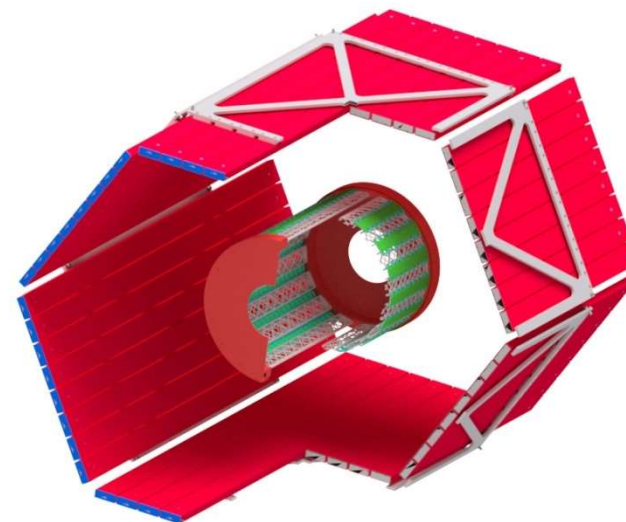


**Big cylindrical detector +  
28 Modules (3 section each)**

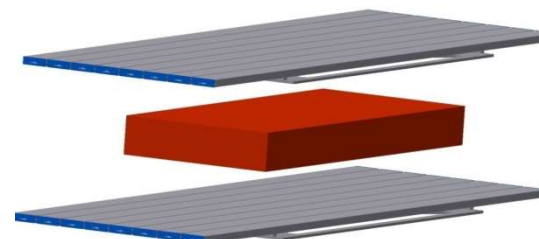
Size: 4784 x 735 x 140 mm



Phase Zero conceptions  
6 MCORD section + miniBeBe



2 MCORD section + other detector



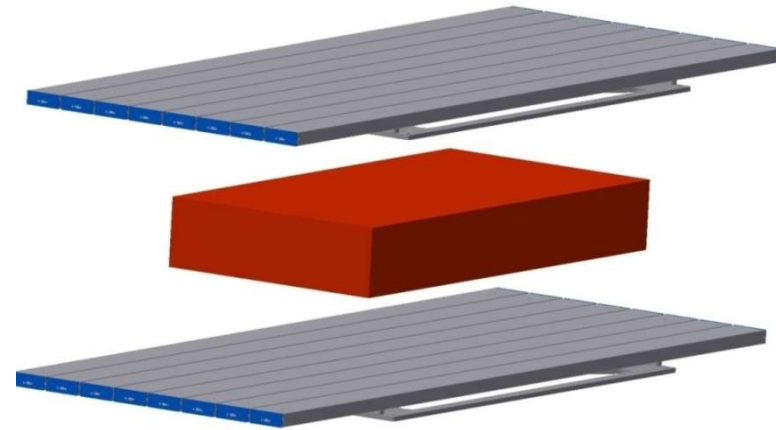
# 1. Introduction



MCORD HUB



Mini MTCA (FPGA)



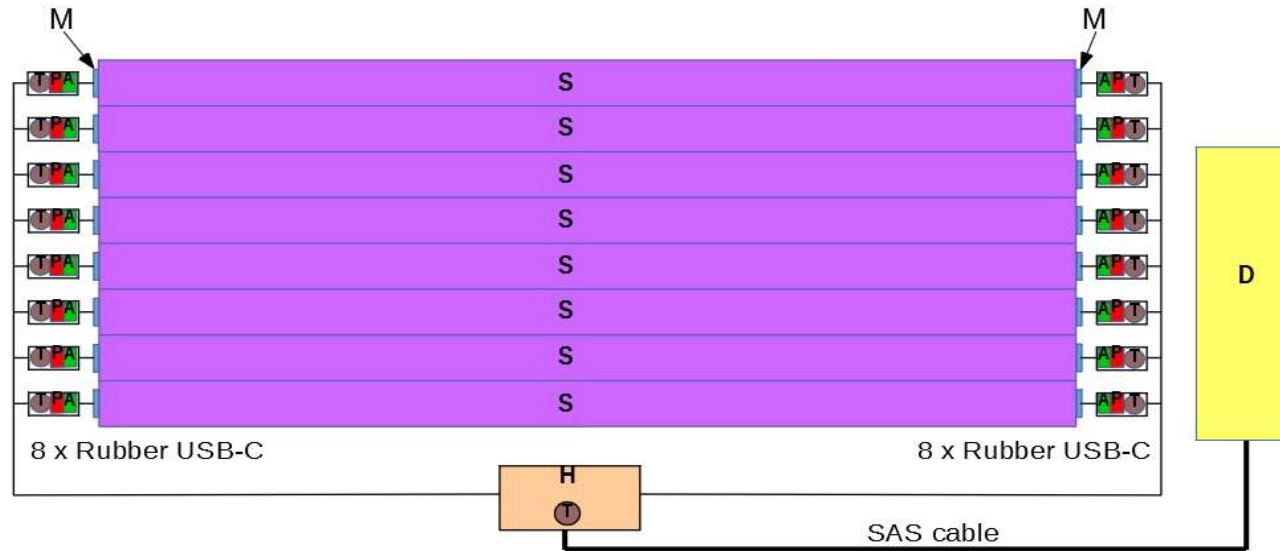
**2 sections (2 x 8 scintillators)  
+ AFE + DSP + DAQ,**



# 2. Detector



## MCORD Section



Position resolution  
In X axis – up to 5 cm  
In Y axis – 5-10 cm

Time Resolution –  
about 1 ns

Legend: **S** (violet) – plastic scintillator, **M** (blue) – SiPM, **P** (red) – power supply with temperature compensation circuit, **T** (brown) – temperature sensor, **A** (green) – amplifier, **H** (orange) – Passive Signal Hub & Power Splitter, **D** (yellow) – MicroTCA system with ADC boards.

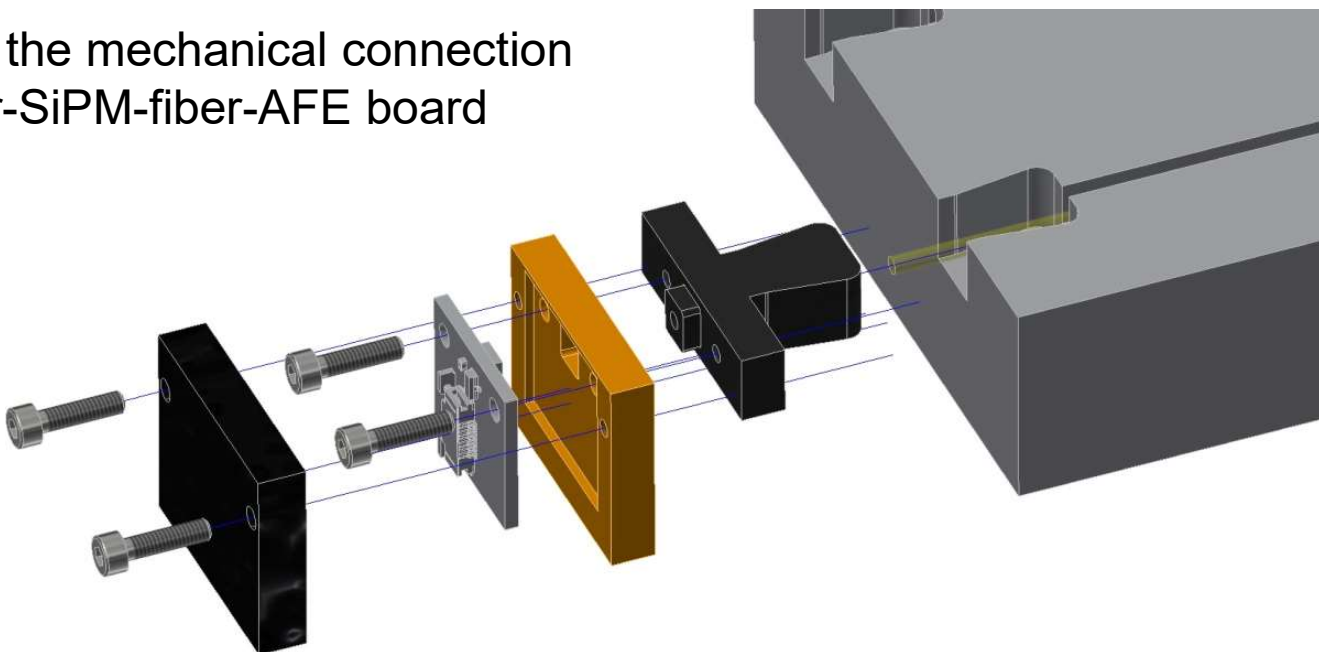
**Single MCORD section**  
**1744 x 735(675) x 50**  
**[mm]**



## 2. Detector



Project of the mechanical connection  
scintillator-SiPM-fiber-AFE board



**Plastic scintillator:**

polystyrene (Nuvia)

162 x 7.2 x 2.2 cm

**WLS fiber:**

2 mm dia. (Kuraray)

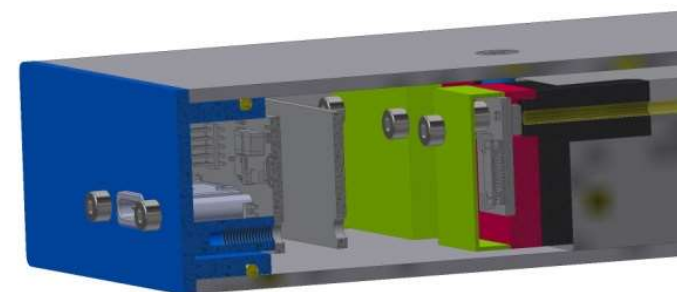
**SiPM (MPPC):**

3x3 mm<sup>2</sup> (Hamamatsu)

**Housing:**

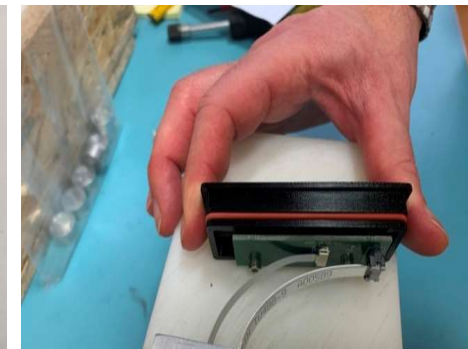
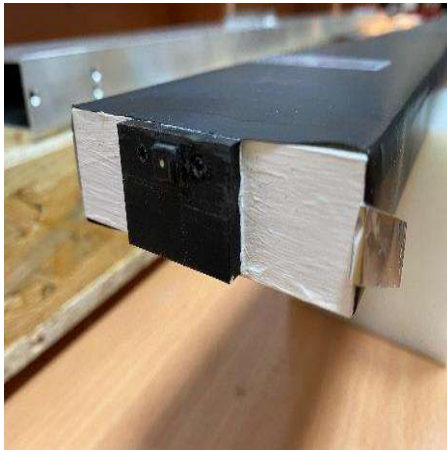
aluminum profile

174 x 8 x 3 cm





## 2. Detector Slab manufacturing



## MCORD single detector assembly

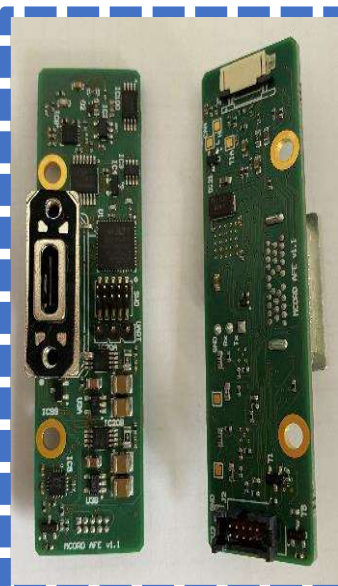
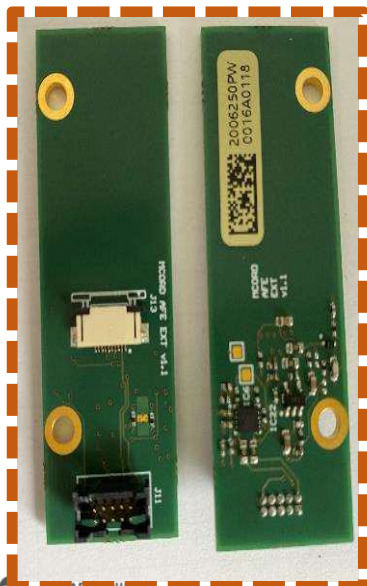
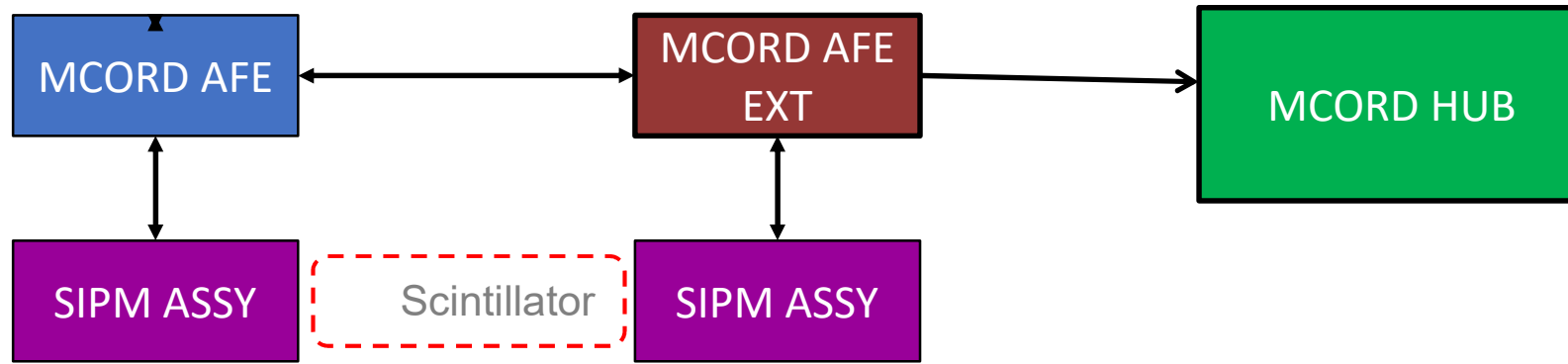


M.Bielewicz, 15.1.2024,  
The 1st Visegrad CREDO Workshop

# 3. Analog Front End and HUB



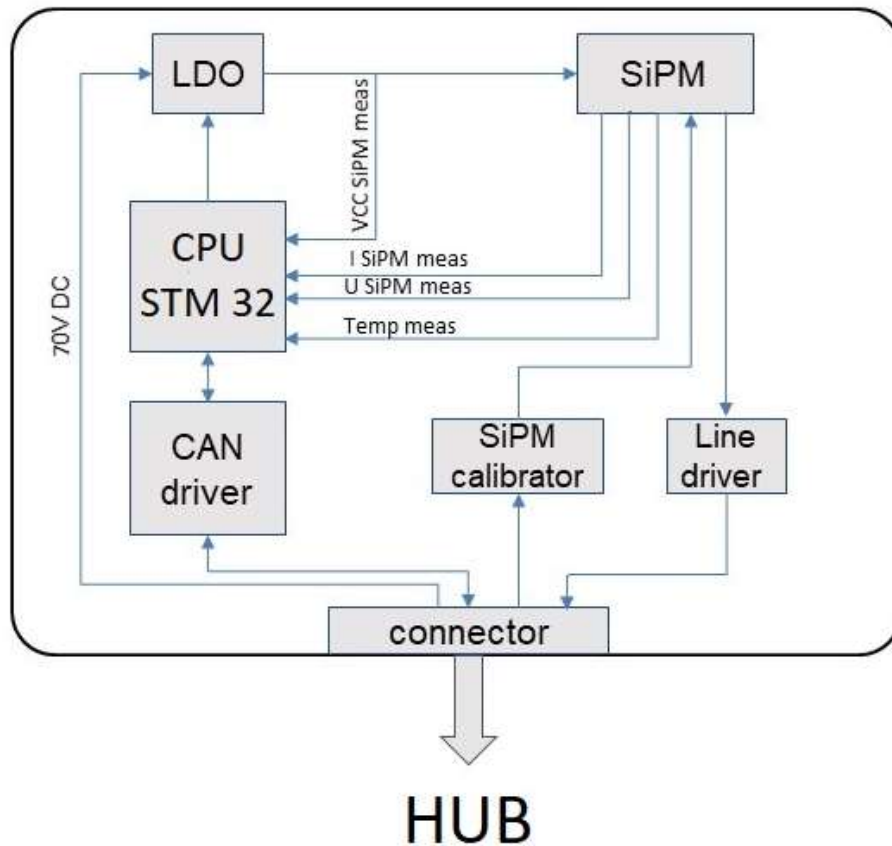
The main boards ver.3 :



# 3. Analog Front End - functionality



- Voltage controller for SiPMs and Amplifier physical signal
- Access to all settings and data from HUB via CAN-bus interface
- Protection for AFE



## ➤ Main blocks

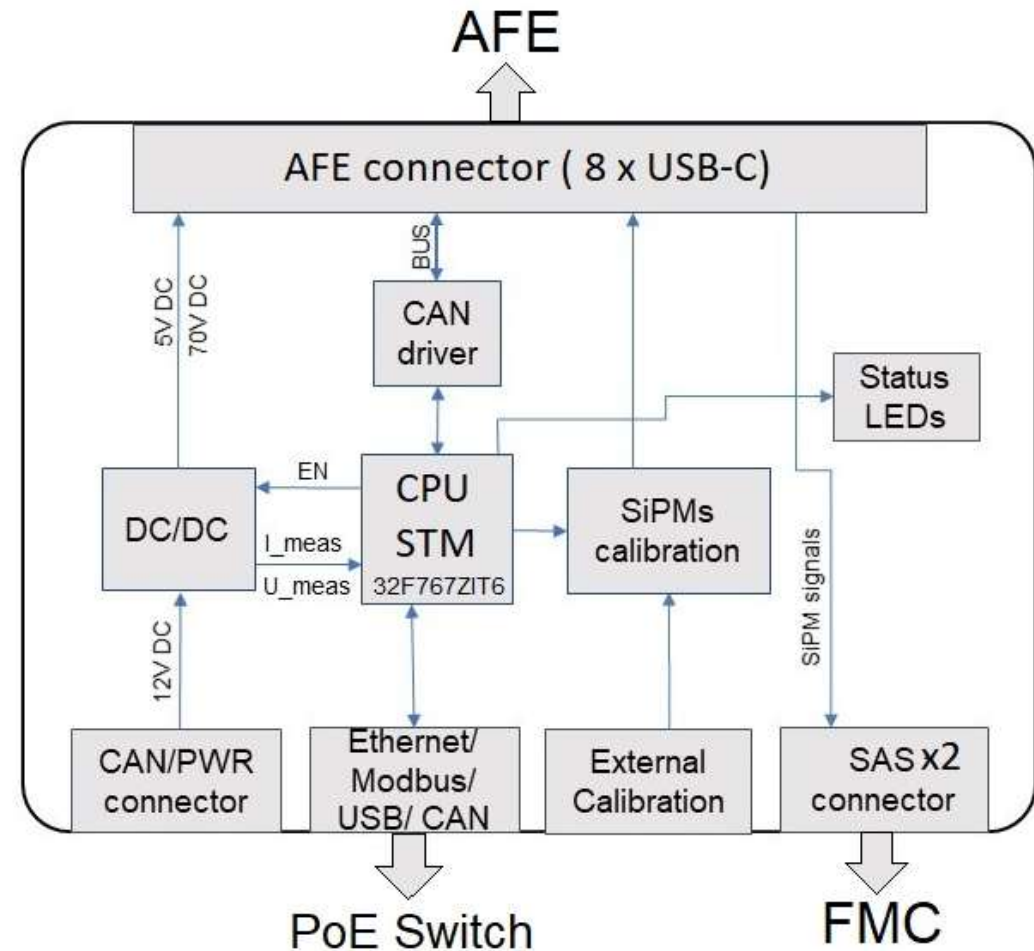
- Embedded CPU (STM32F072CBU6)
- **Temperature sensor (LM45)**
- SiPM voltage controller + LDO (Low Dropout Regulator)
- **SiPM calibrator**
- SiPM signal transmitter to HUB (differential signal)
- **CAN network driver**
- **Measurements (12 bit ADC)**
  - 2 x SiPM voltage
  - 2x SiPM current
  - 2 x SiPM VCC voltage
  - 2 x SiPM temperature
- **Control (8 bit DAC)**
  - 2 x SiPM voltage



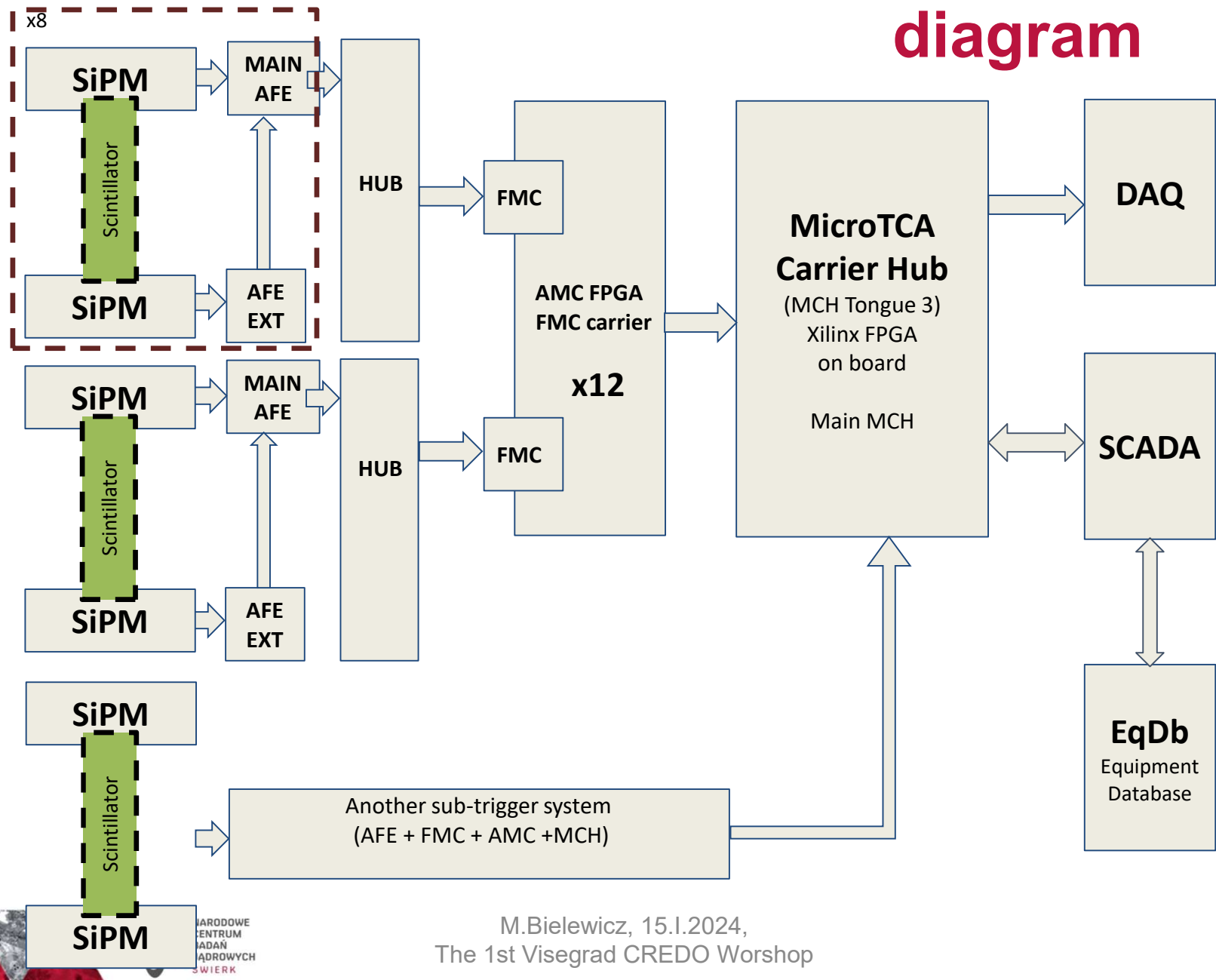
### 3. HUB - functionality



- Mikro PYTHON programing
- PoE supply
- **Generation of 5V and 70V**
- ETH <-> CAN
- Distribution of signals from AFE to SAS cables
- Status LEDs on AFE ASSY and HUB for quick fault identification
- **Generation of calibration signals to AFE**
- STM32 CPU with microPython



# 3. MCORD readout system schematic diagram



## 3. Triger



# Data processing

- Latency estimation for **L1 trigger** (event without parameters)
  - ✓ AFE cabling 8ns/m, with 10m cabling latency is 80ns
  - ✓ ADC + SERDES latency: 400nsEstimated total latency: **about 1us**

Latency estimation for **L2 trigger** (event with parameters)

- ✓ MGT latency: 500ns
  - ✓ Algorithm latency : 2-5us
  - ✓ Formatter and transmitter latency: 1us
- Estimated total latency:
- 3.5 – 7.5us**

Latency estimation for **L3 trigger** (between MTCA systems)

- ✓ MGT latency: 500ns
  - ✓ Fiber latency: 500ns + 8ns/m
  - ✓ Algorithm latency : 2-5us
  - ✓ Formatter and transmitter latency: 1us
- Estimated total latency:
- 10 – 15us**



# 3. Laboratory tests



## Measuring system

AFE Board

AFE Hub

SAS to BCN converter

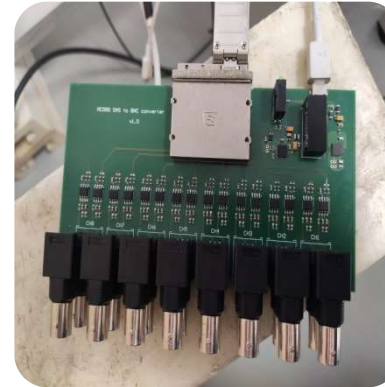
Digitizer



Plastic scintillator in an aluminum housing with an AFE amplification system and a Hamamatsu MPPC photodetector



Managed control system for AFE power supplies mounted in boards. Up to 8 boards can be connected once



Converter of signals received by SAS cable to appropriate single BNC channels for each MPPC



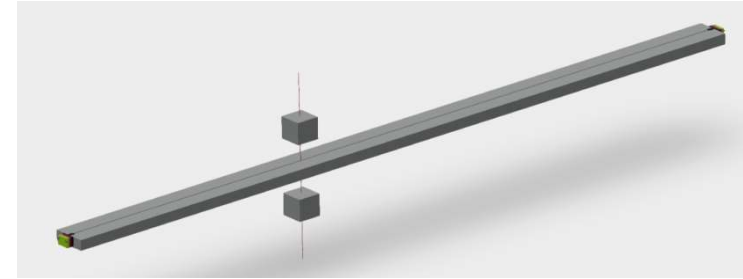
Digital multi-channel amplitude acquirer by CAEN for analysis of received signals



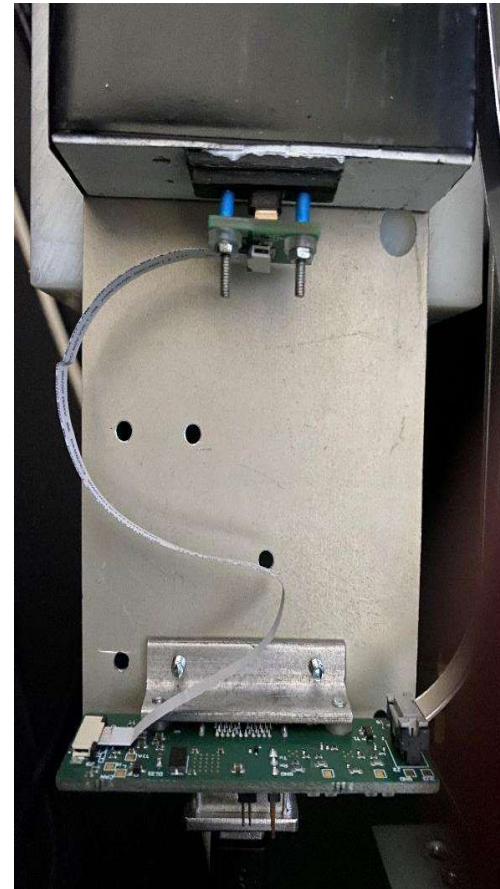
# 3. Laboratory tests – 1st step



One Plastic MCORD detector  
+ 2 plastic hodoscopes (muon trigger)  
+ DAQ: CAEN DT5730

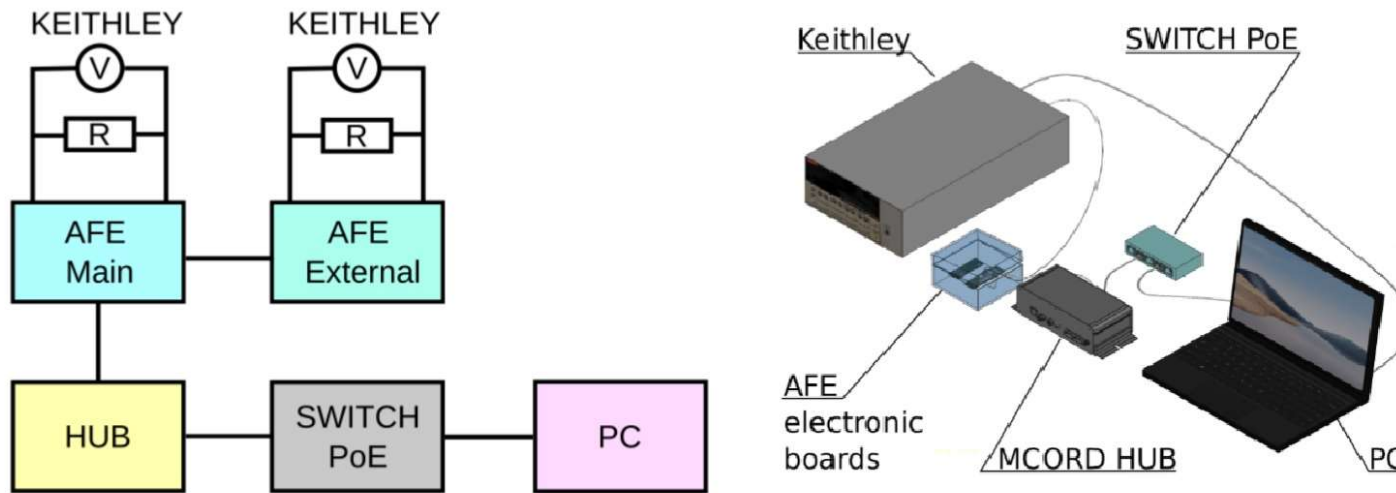


**BLACK BOX test setup**

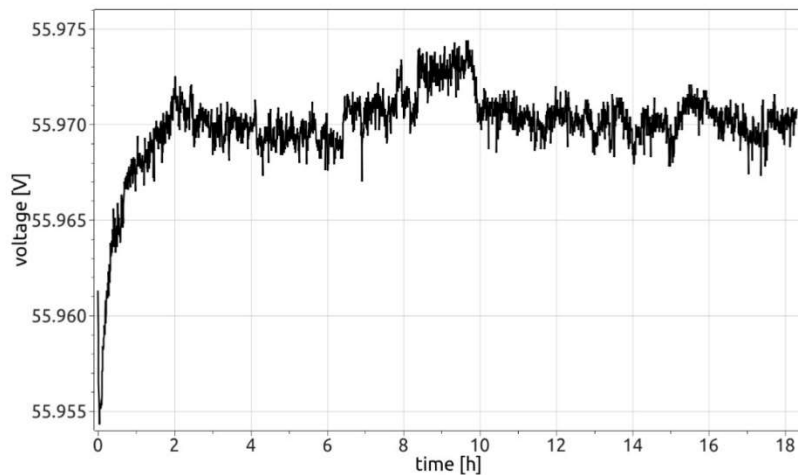




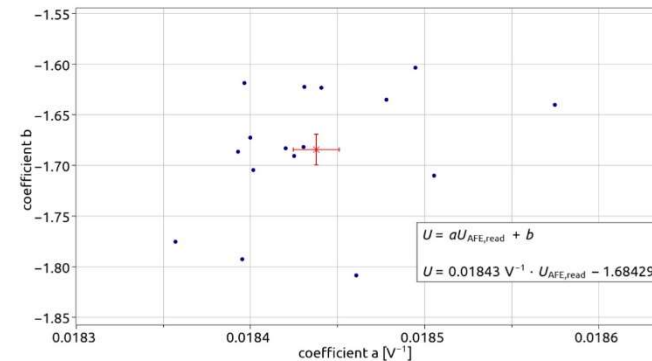
# 3. Laboratory tests – 2nd step



Schematic diagram of the measuring station for calibration measurements



Voltage  $U$  during heat-up of main AFE measured using the Keithley voltmeter.



Internal voltmeter calibration coefficient distribution. The scale is limited to a narrow range to show small differences in presented values. The average value is marked with a red marker.

## Calibration of A/D and D/A converters.



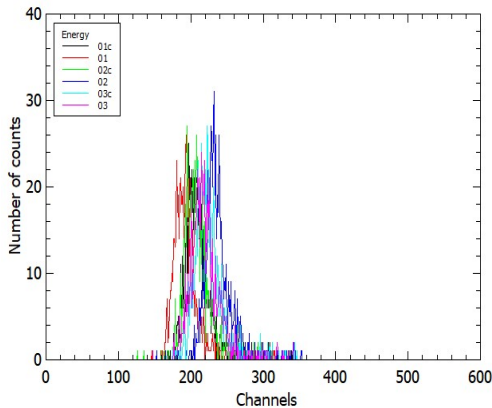
# 3. Laboratory tests – 3th step



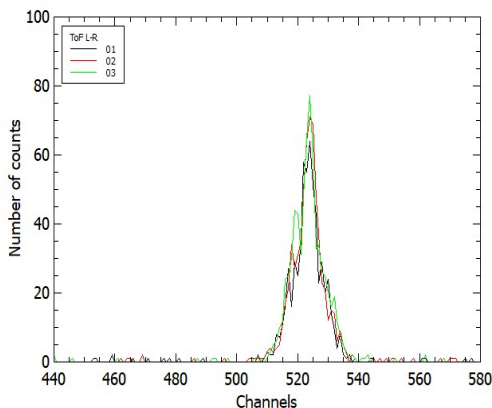
**3x plastic MCORD detectors**  
**+ 2x plastic hodoscopes (muon triggers)**  
**+ DAQ: CAEN DT5730**



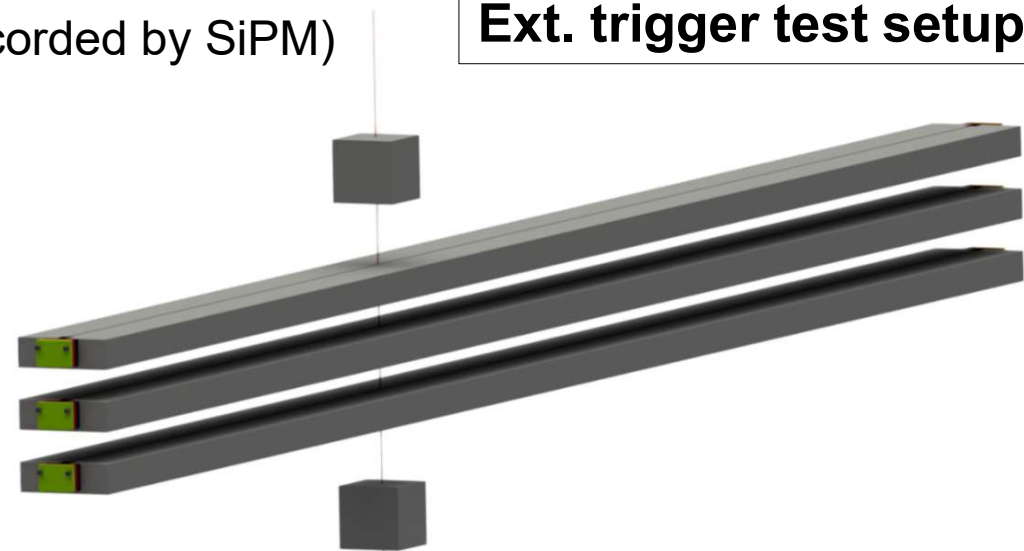
**Ext. trigger test setup**



Energy  
(amplitude recorded by SiPM)



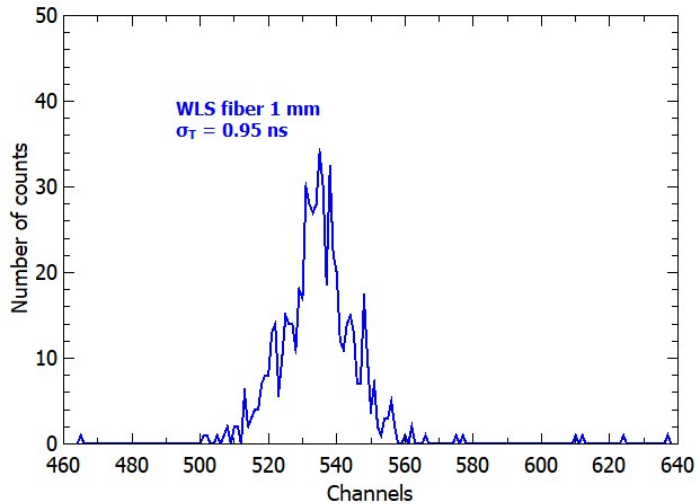
ToF  
(between both ends of a scintillator) -> Time resolution



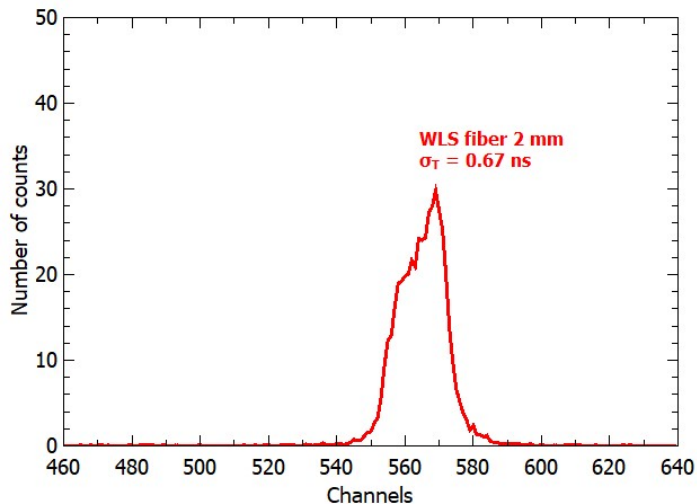
**Plastic (162 x 7.2 x 2.2 cm) + WLS fiber (1 mm) + 2x MPPC 3 x 3 mm (pixel size 75um)**  
**Hodoscopes: plastic (5 x 5 x 5 cm) + PMT (2" dia) → 99,5% efficiency**



### 3. Laboratory tests – 3th step



WLS fiber (1 mm)  
**CRT ( $\sigma$ ) = 0.95 ns  $\implies \sigma_x = 7.1$  cm**

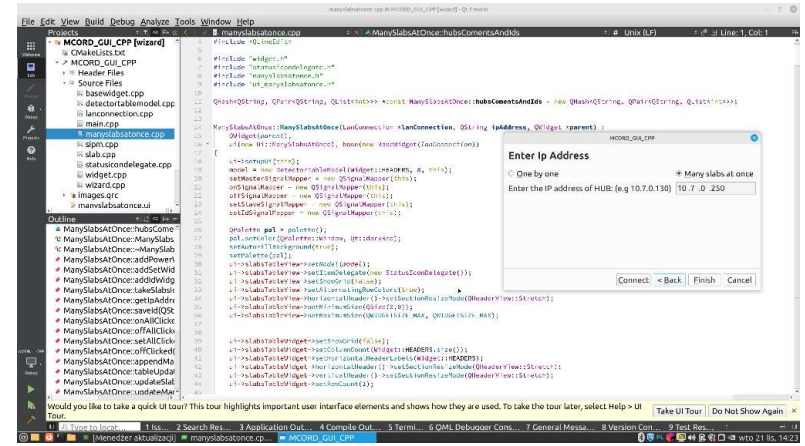
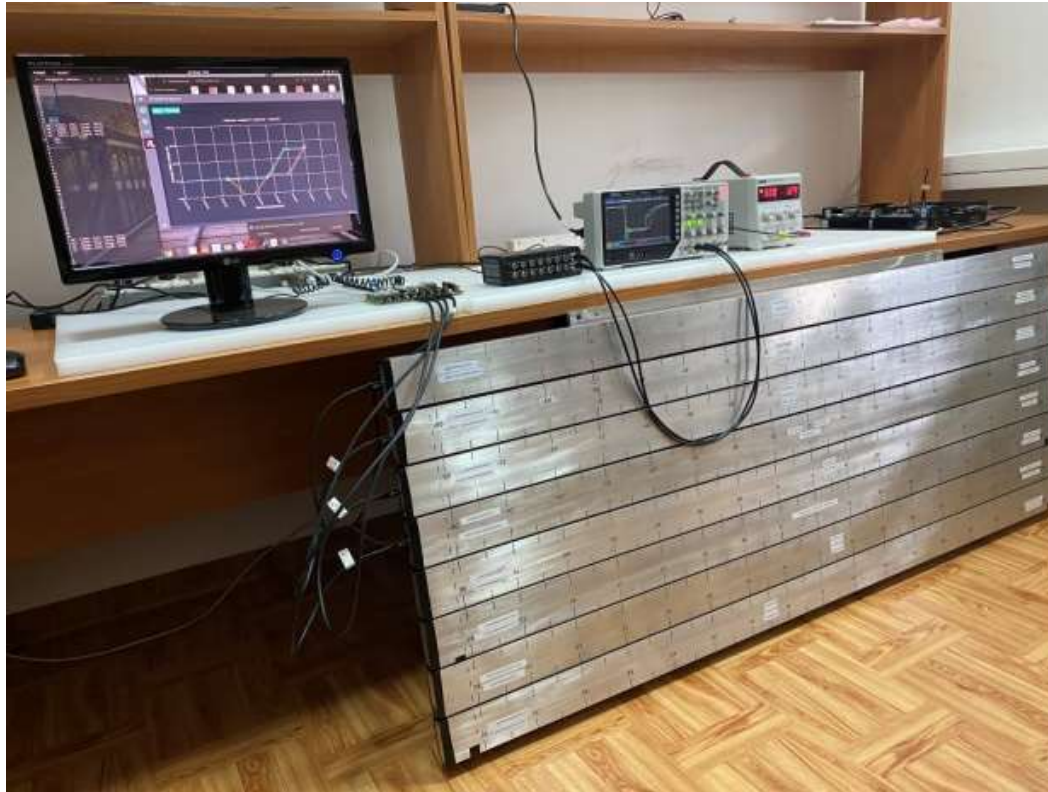


WLS fiber (2 mm)  
**CRT ( $\sigma$ ) = 0.67 ns  $\implies \sigma_x = 5.1$  cm**

**(!) improved timing resolution for 2 mm WLS fiber (!)**



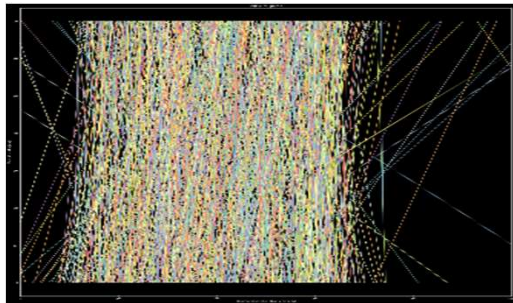
# 3. Laboratory tests – 4th step



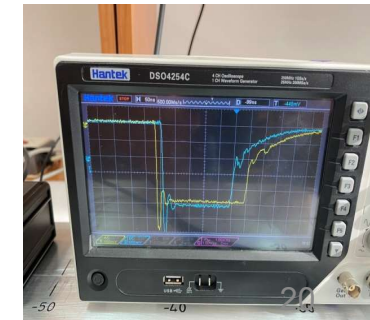
Detection Slabs

Section No. 1 Connected by LAN to IP: 10.7.0.250

Slab No.	Status	Power	Type	Set SIPM Volt.		UVI		I(nA)		T(C)
				Set All	Chart U(V)	Chart I(nA)	Chart T(C)			
16	Set	On	Master	53	52.49	50.28	56.1	68.0	29.37	
17	Set	On	Slave	50	50.03	57.01	1162.1	56.1	33.88	
18	Set	On	Master	50	50.34	50.34	68.0	68.0	29.28	
31	Set	On	Slave	50	50.30	50.30	56.1	68.0	31.38	
	Set	On	Master		OFF	OFF	56.1	68.0	29.12	
	Set	On	Slave		OFF	OFF	56.1	68.0	32.59	
	Set	On	Master		OFF	OFF	0.0	0.0	0.00	
	Set	On	Slave		OFF	OFF	0.0	0.0	0.00	
	Set	On	Master		OFF	OFF	0.0	0.0	0.00	
	Set	On	Slave		OFF	OFF	0.0	0.0	0.00	
	Set	On	Master		OFF	OFF	0.0	0.0	0.00	
	Set	On	Slave		OFF	OFF	0.0	0.0	0.00	



## MCORD demonstrator and GUI program



# 4. Earthquake - Cosmic ray correlation



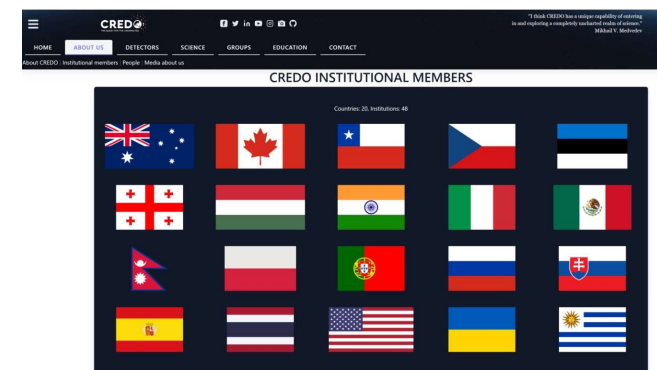
## Earthquake Precursors

Large scale correlations between cosmic rays and earthquakes presumably related to earthquakes precursors has been observed. The found periodicity is rather similar to the sun spots solar cycle.

Cosmic ray data correspond to the measurements at the Pierre Auger observatory in Malargüe, Argentina, whereas seismic data is taken from Moscow and Oulu stations located in Russia and Finland, respectively.

A  $6\sigma$  correlation effect has been observed in a period of about 4.5 years. Details can be found in a the publication

**Observation of large scale precursor correlations between cosmic rays and earthquakes with a periodicity similar to the solar cycle, P. Homola, et al., *J ATMOS SOL-TERR PHY*[70] Vol. 247 (2023) 106068**



# 4. CREDO-MEXICO-CHILE program



**COSMO-SEISMIC TASK**

The global scale cosmic ray research has recently been focused on the search for so-called cosmo-seismic correlations that are correlations between cosmic ray rates and seismic activity.

**OUR CONCEPT**   **WHAT INSPIRED US**   **RESEARCH PERFORMED**

Mass movements inside the Earth that lead to earthquakes simultaneously cause temporary changes in gravitational and geomagnetic fields. These changes are propagating at the speed of light and can potentially be observed on the surface of the planet earlier than earthquakes. It can be possible, for example, by registering changes in the frequency of detection of secondary cosmic radiation, which is very sensitive to the geomagnetic conditions.

The scope of research and the list of institutions are discussed. Currently, we are preparing the text of the **Memorandum of Understanding**, which will become the basis for obtaining the grant.

**Recent Developments within The Cosmic-Ray Extremely Distributed Observatory(CREDO),**  
 O. Suchanov, P. Homola, D.E. AlvarezCastillo, M. Bielewicz et al.,  
*Revista MEX FIS[20] Vol. 4 No 2 (2023) 1*

The use of the Mexican and Chile cosmic ray observatory + several MCORD coincidence detectors installed in universities + several smaller detectors for schools etc. - Creation of a measurement network collecting data simultaneously with seismographic stations



# 5. CDR publictation



## Conceptual design report of the MPD Cosmic Ray Detector (MCORD)

Published 25 November 2021 • © 2021  
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[Journal of Instrumentation](#)  
[Volume 16 November 2021](#)

<https://doi.org/10.1088/1748-0221/16/11/P11035>

**Jinst** PUBLISHED BY IOP PUBLISHING FOR SISSA MEDIALAB  
RECEIVED: July 7, 2021  
ACCEPTED: November 11, 2021  
PUBLISHED: November 25, 2021

**Conceptual design report of the MPD Cosmic Ray Detector (MCORD)**

M. Bielewicz,<sup>a,c,\*</sup> A. Bancer,<sup>a</sup> M. Barabanov,<sup>a</sup> A. Chlopik,<sup>a,c</sup> M. Czarnynoga,<sup>b,e</sup> D. Dabrowski,<sup>b,e</sup> A. Dudzinski,<sup>a</sup> A. Dziedzic,<sup>a</sup> M. Grodzicka-Kobyłka,<sup>a</sup> J. Grzyb,<sup>a</sup> K. Grodzicki,<sup>a</sup> E. Jaworska,<sup>a</sup> P. Kankiewicz,<sup>c</sup> G. Kasprówicz,<sup>b</sup> A. Kisiel,<sup>b,e</sup> P. Kolasinski,<sup>b</sup> S. Kowalski,<sup>f</sup> M. Krakowiak,<sup>a</sup> M. Kulich,<sup>d</sup> M. Kutyla,<sup>b,c</sup> J. Lukasik,<sup>g</sup> B. Maksiak,<sup>a</sup> J. Marzec,<sup>b</sup> T. Matulewicz,<sup>d</sup> S. Mianowski,<sup>a</sup> M. Milewicz-Zalewska,<sup>b,e</sup> M. J. Peryt,<sup>b,c</sup> M. Piatek,<sup>c</sup> A. Pollo,<sup>a</sup> K. Pozniak,<sup>b</sup> F. Protoklitow,<sup>b,e</sup> D. Pszczel,<sup>a</sup> S. Pulawski,<sup>f</sup> R. Romaniuk,<sup>b</sup> K. Roslon,<sup>b,e</sup> M. Rybczynski,<sup>c</sup> J. Rządiewicz,<sup>a</sup> S. Satybaldyeva,<sup>a</sup> D. Shchegolev,<sup>c</sup> I. Shmyrev,<sup>b,e</sup> M. Sitek,<sup>a</sup> M. Sowinski,<sup>b</sup> G. Stefanek,<sup>c</sup> J. Stepaniak,<sup>a</sup> E. Strugalska-Gola,<sup>a</sup> L. Swiderski,<sup>a</sup> A. Syntfeld-Kazuch,<sup>a</sup> T. Szczesniak,<sup>a</sup> M. Szuta,<sup>a</sup> A. Vodopyanov,<sup>c</sup> D. Wielanek,<sup>b</sup> Z. Włodarczyk,<sup>c</sup> K. Wojcik<sup>f</sup> and W. Zabolotny<sup>b</sup>

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<sup>g</sup>Institute of Nuclear Physics PAN, Radzikowskiego 152 str., 31-342, Krakow, Poland  
E-mail: [Narcin.Bielewicz@ncbj.gov.pl](mailto:Narcin.Bielewicz@ncbj.gov.pl)

ABSTRACT: This report presents a concept of constructing a detector dedicated for detection of muons observed during measurements carried out at the MPD (Multi-Purpose Detector) detector that is currently under construction at the NICA facility, Russia, Dubna. It has been proposed to design and build an additional detector that will complement the current MPD set and increase its measurement capabilities. The main goal of this project is to provide information from cosmic muons that pass

\*Corresponding author.

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2021 JINST 16 P11035



# 5. Electronic analysis publications



## Practical Implementation of an Analogue and Digital Electronics System for a Modular Cosmic Ray Detector—MCORD

Published 22 March 2023

*Electronics* **2023**, 12(6), 1492;

<https://doi.org/10.3390/electronics12061492>

APPLICATION OF ARTIQ CONTROL SYSTEM IN MODULAR COSMIC RAY DETECTOR – MCORD  
DOI 10.15199/13.2023.8.19, ELEKTRONIKA 8/2023, p.93

electronics

MDPI

Article  
**Practical Implementation of an Analogue and Digital Electronics System for a Modular Cosmic Ray Detector—MCORD**

Marcin Bielewicz <sup>1,\*</sup>, Aleksandr Bancer <sup>1</sup>, Andrzej Dziejdz <sup>1</sup>, Jarosław Grzyb <sup>1</sup>, Elżbieta Jaworska <sup>1</sup>, Grzegorz Kasprzowicz <sup>2</sup>, Michał Kieca <sup>1</sup>, Piotr Kolasinski <sup>2</sup>, Michał Kuc <sup>1</sup>, Michał Kuklewski <sup>2</sup>, Marcin Pietrak <sup>1</sup>, Krzysztof Pozniak <sup>2</sup>, Maciej Sitek <sup>1</sup>, Mikołaj Sowinski <sup>2</sup>, Łukasz Świdorski <sup>1</sup>, Agnieszka Syntfeld-Kazuch <sup>1</sup>, Jarosław Szewinski <sup>1</sup> and Wojciech Marek Zabolotny <sup>2</sup>

<sup>1</sup> National Centre for Nuclear Research, 05-400 Otwock, Poland  
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**Abstract:** A Modular Cosmic Ray Detector (MCORD) was prepared for use in various physics experiments. MCORD detectors can be used in laboratory measurements or can become a part of large measurement sets. MCORD can be used as a muon detector, a veto system, or a tool supporting the testing and calibration of other detectors. MCORD can also work as a stand-alone device for scientific and commercial purposes. The basic element of MCORD is one section consisting of eight oblong scintillators with a double-sided light reading performed by silicon photomultipliers (SiPMs). This work presents a practical description of testing, calibration, and programming of analogue and digital electronics modules. The characterisation and calibration methods of the analogue front-end electronic modules, the obtained results, and their implementation into an operating system are presented. In addition, we describe the development environment and the procedures used to prepare our kit for practical use. The architecture of the FPGAs is also presented with a description of their programming as a data-collecting system in a simple coincidence circuit. We also present the possibilities of extending the data analysis system for large experiments.

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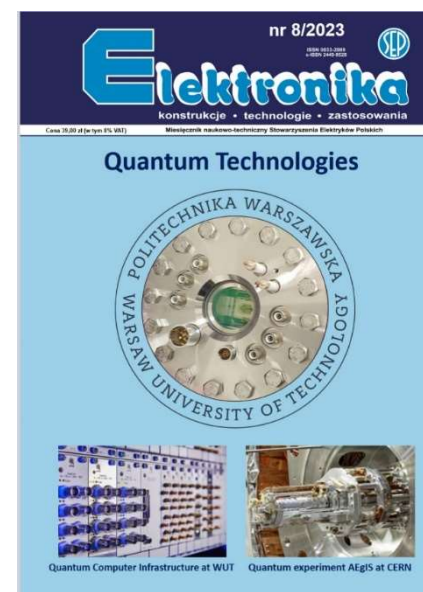
**Citation:** Bielewicz, M.; Bancer, A.; Dziejdz, A.; Grzyb, J.; Jaworska, E.; Kasprzowicz, G.; Kieca, M.; Kolasinski, P.; Kuc, M.; Kuklewski, M.; et al. Practical Implementation of an Analogue and Digital Electronics System for a Modular Cosmic Ray Detector—MCORD. *Electronics* 2023, 12, 1492. <https://doi.org/10.3390/electronics12061492>

Academic Editor: Spyridon

**Keywords:** cosmic ray; muon detector; modular detector; silicon photomultiplier; detector control system; AFE; FPGA; trigger

### 1. Introduction

The idea of building a new cosmic ray detector was born in 2018 in connection with the construction of the MPD (Multi-Purpose Detector) [1,2] for the NICA collider (Dubna, Russia). The currently designed system is being prepared for use in other large experiments such as NA61/SHINE for educational and training purposes, and for commercial use in the



M.Bielewicz, 15.I.2024,  
The 1st Visegrad CREDO Workshop





# Thank You for Attention!

