

# Bound state effects in dark matter genesis: pushing the boundaries of higher excitations

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*based on 2112.01499 and 2308.01336  
in collaboration with T. Binder, M. Garny, S. Lederer, K. Urban*

**Jan Heisig (RWTH/UVA)**



Alexander von  
**HUMBOLDT**  
STIFTUNG

*IFJ Seminar, Kraków, Poland  
October 19, 2023*

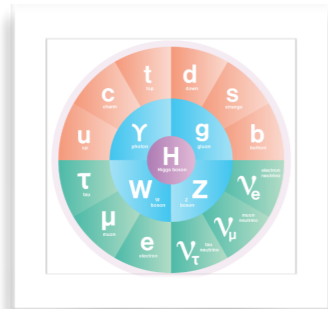
# Outline

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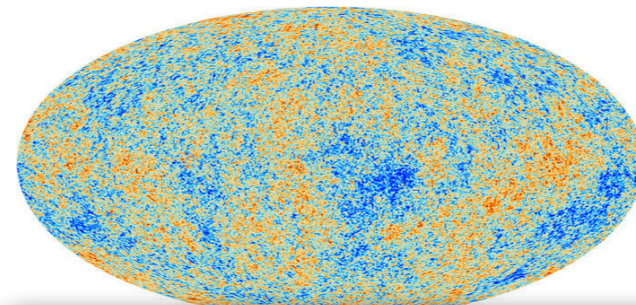
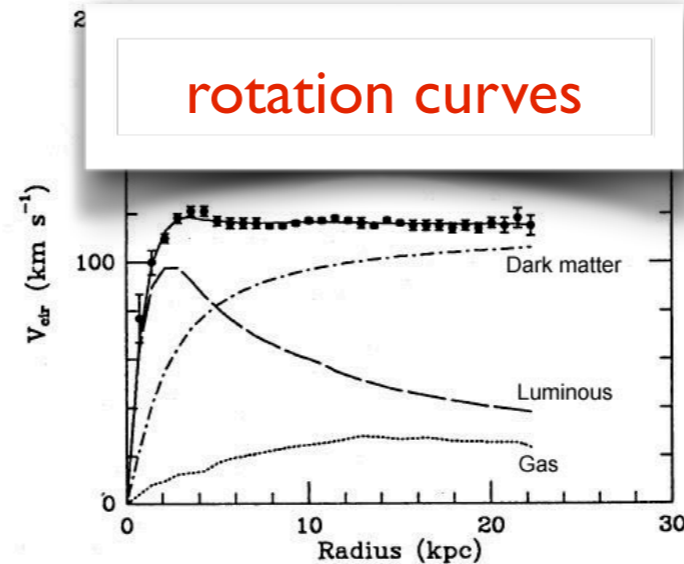
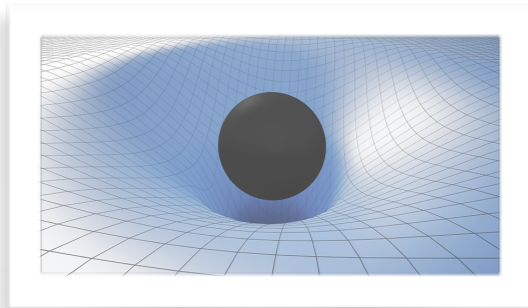
- Introduction to particle dark matter
  - Effects of (excited) bound state
- Implications for  $t$ -channel mediator models
  - Implications for LHC searches

# The phenomenon of Dark Matter

- More matter: gravitational effect on dynamics of visible matter
- Present on very different length scales



Standard model  
+  
general relativity



CMB anisotropies

bullet cluster



Distant Galaxy Lensed by Cluster Abell 2218 HST • WFPC2 • ACS

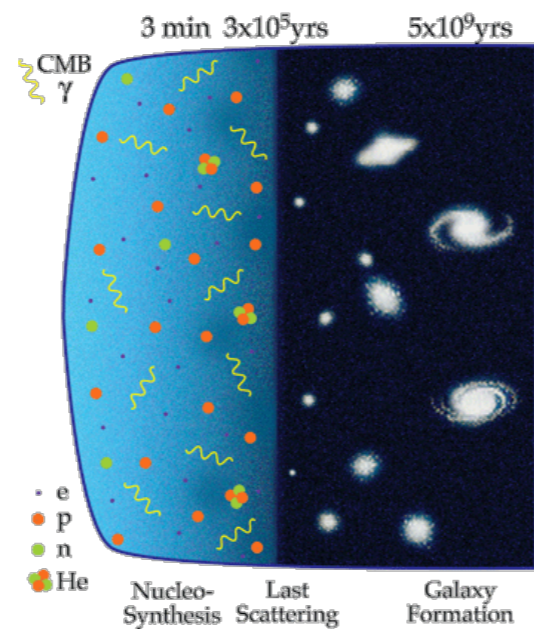


grav. lensing

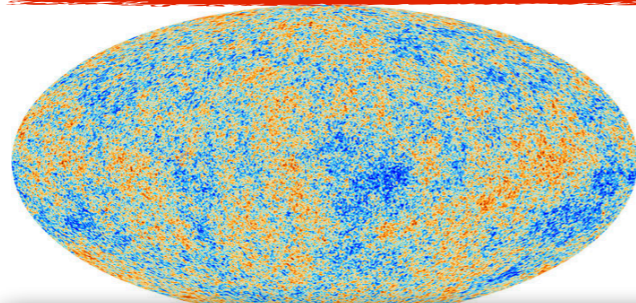
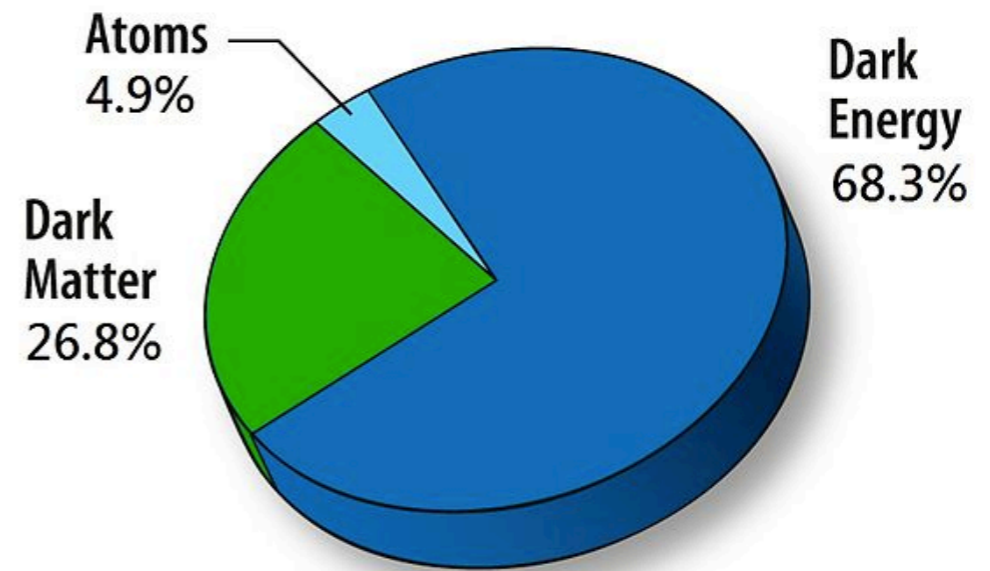


# The phenomenon of Dark Matter

- More matter: gravitational effect on dynamics of visible matter



Contributions  $\Omega_i$  to the energy density of the Universe today:



CMB anisotropies

TODAY

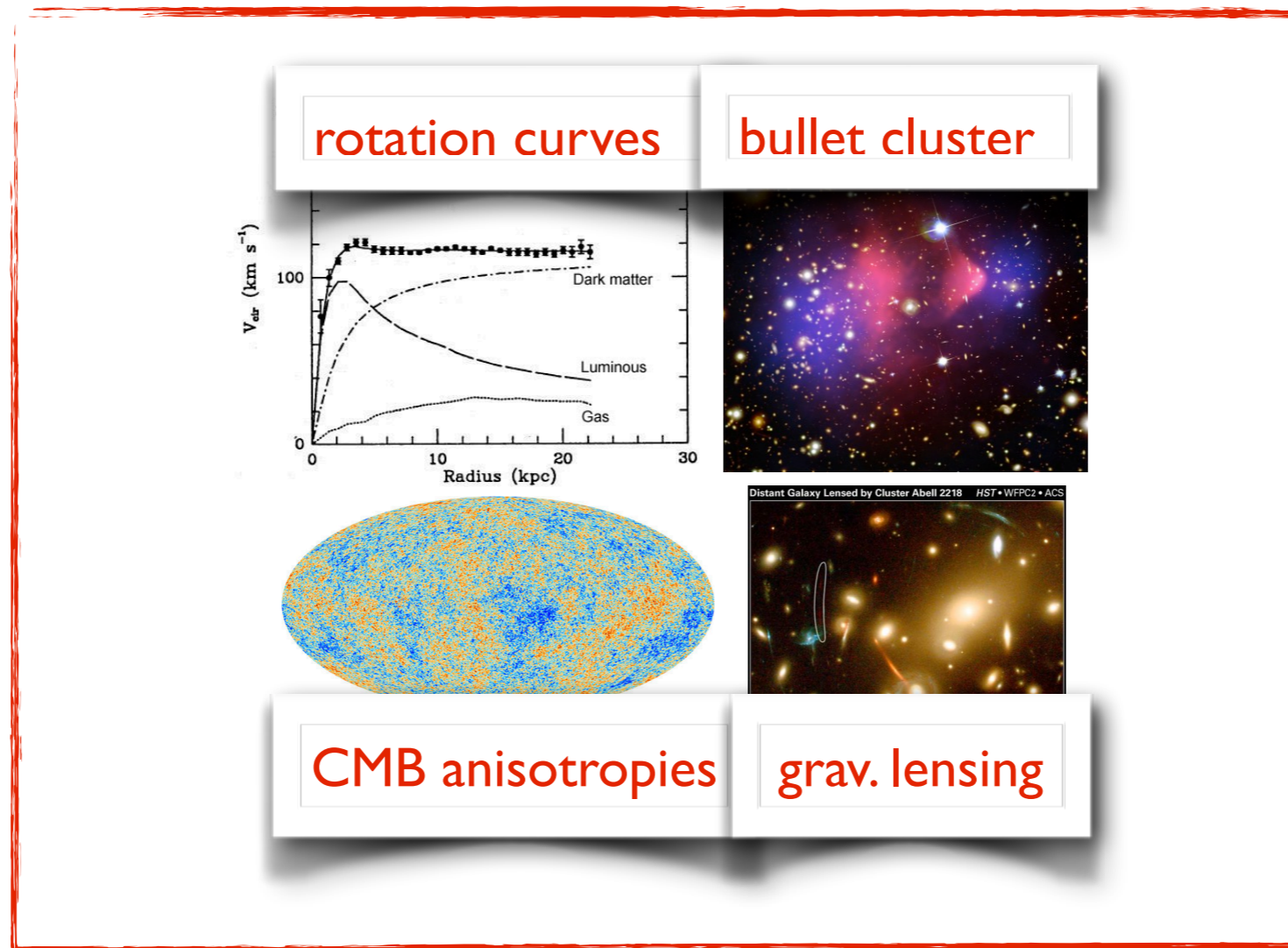
[Planck 2020]

$$\Omega_{\text{DM}} h^2 = 0.12 \pm 0.001$$



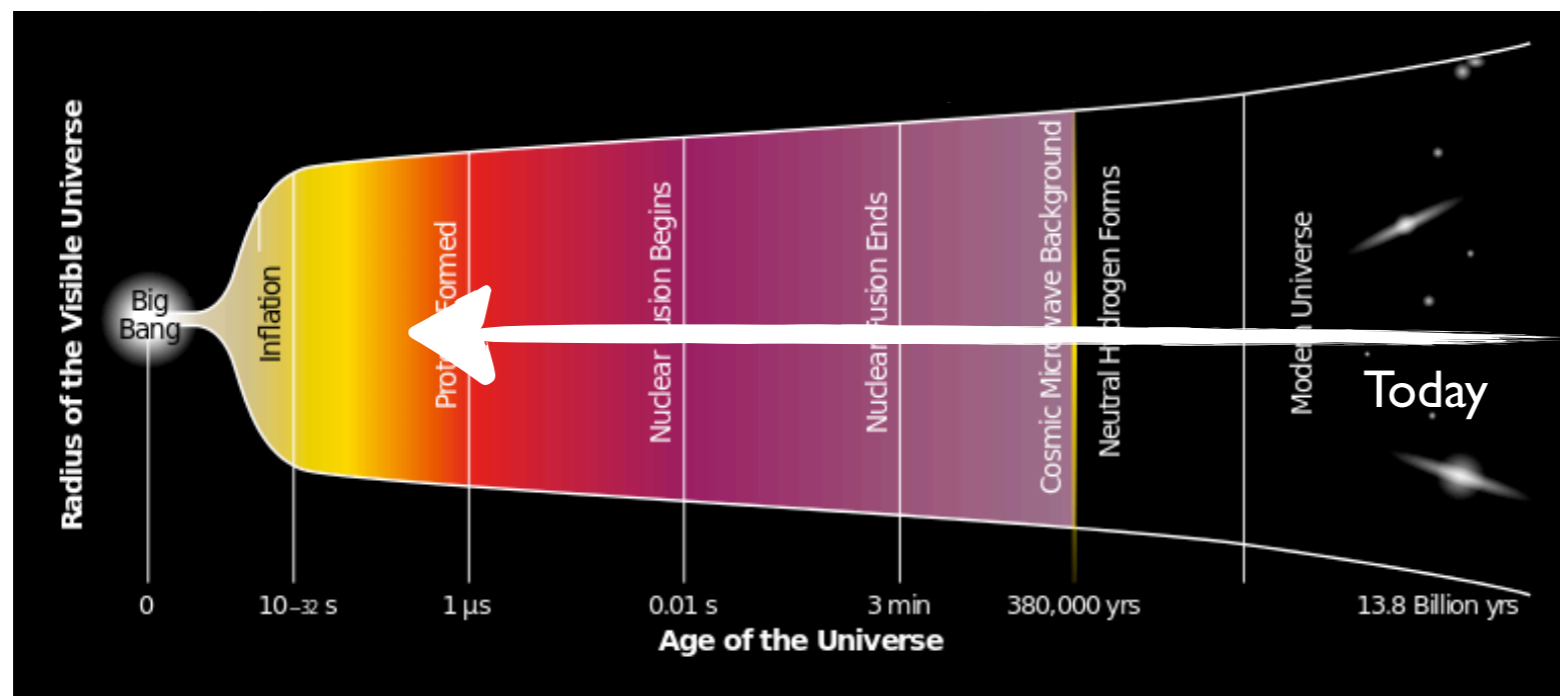
# Particle Dark Matter $\chi$

Explain appearance on all scales:  
Most plausible  $\rightarrow$  new particle



# Particle dark matter: a thermal relic

- Relic from thermal abundance
- Consider cosmological history of Universe:

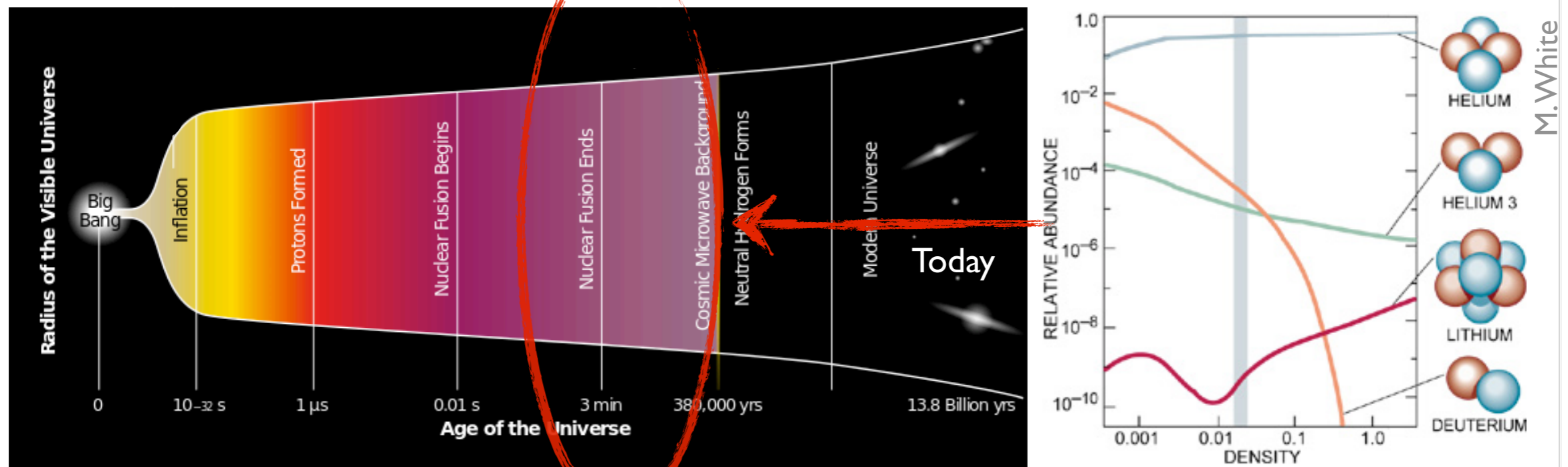


Particle physics  
+ cosmology:  
Extrapolate to early  
hot Universe  
⇒ Boltzmann Eqs.

Expansion with Hubble rate  $H$

# Particle dark matter: a thermal relic

- Successful example: Big Bang Nucleosynthesis

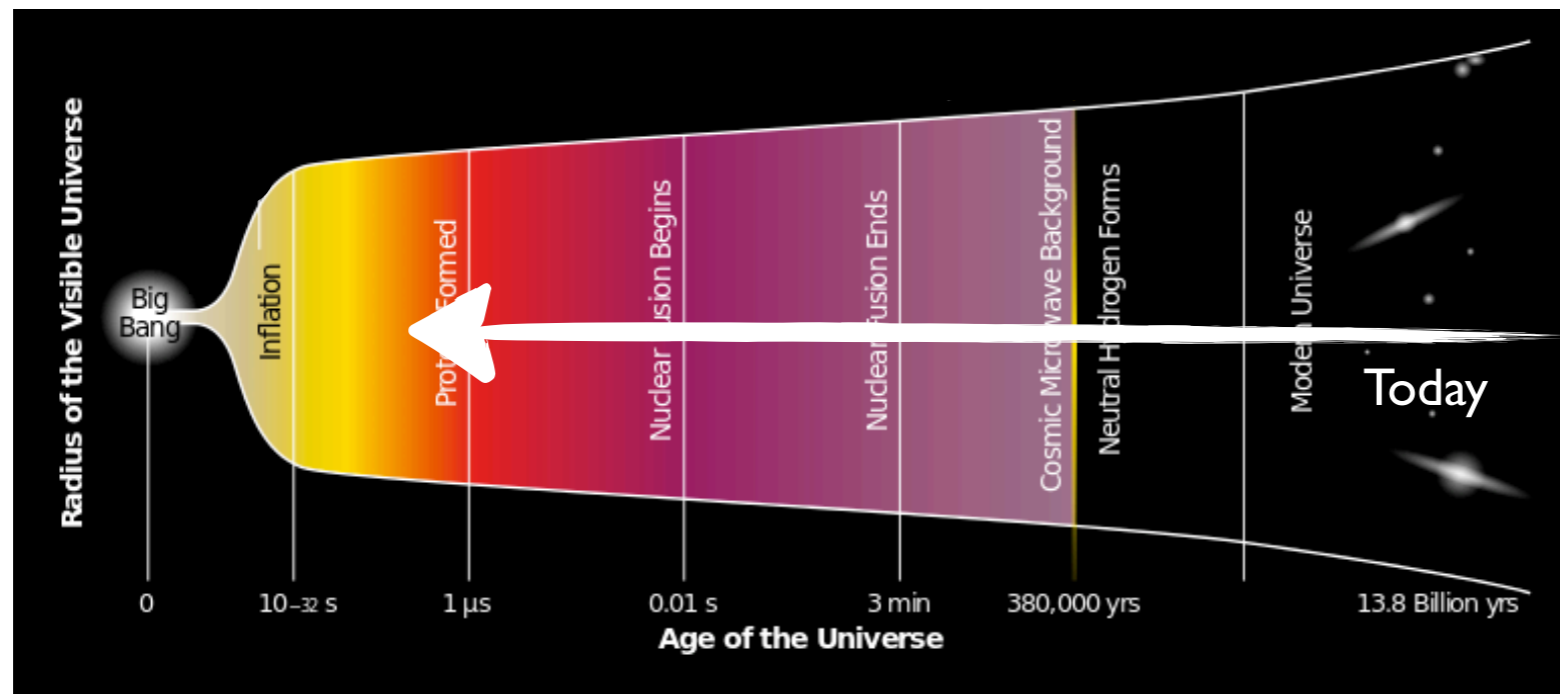


→ Explains primordial abundances of light elements



# Particle dark matter: a thermal relic

- Relic from thermal abundance
- Consider cosmological history of Universe:



Particle physics  
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⇒ Boltzmann Eqs.

Expansion with Hubble rate  $H$

# Boltzmann equations for particle densities

[Zel'dovich, Okun, Pikel'ner 1966; Lee, Weinberg 1977; Binetruy, Girardi, Salati 1984; Bernstein, Brown, Feinberg 1985; Srednicki, Watkins, Olive 1988; Kolb, Turner 1990; Griest, Seckel 1991; Gondolo, Gelmini 1991; Edsjo, Gondolo 1997]

$$\underline{E_\chi (\partial_t - H p \partial_p) f_\chi(p, t) = C [f_\chi]}$$

Relativistic Liouville operator for homogeneous, isotropic Universe

Collision operator



Cosmology

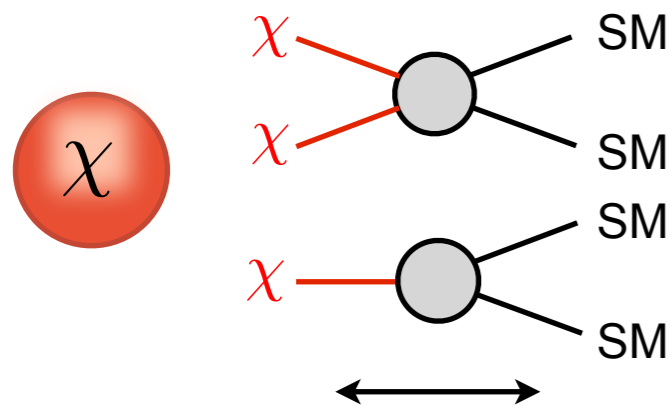


Particle Physics

$$H \quad \wedge \quad \Gamma$$

# Boltzmann equations for particle densities

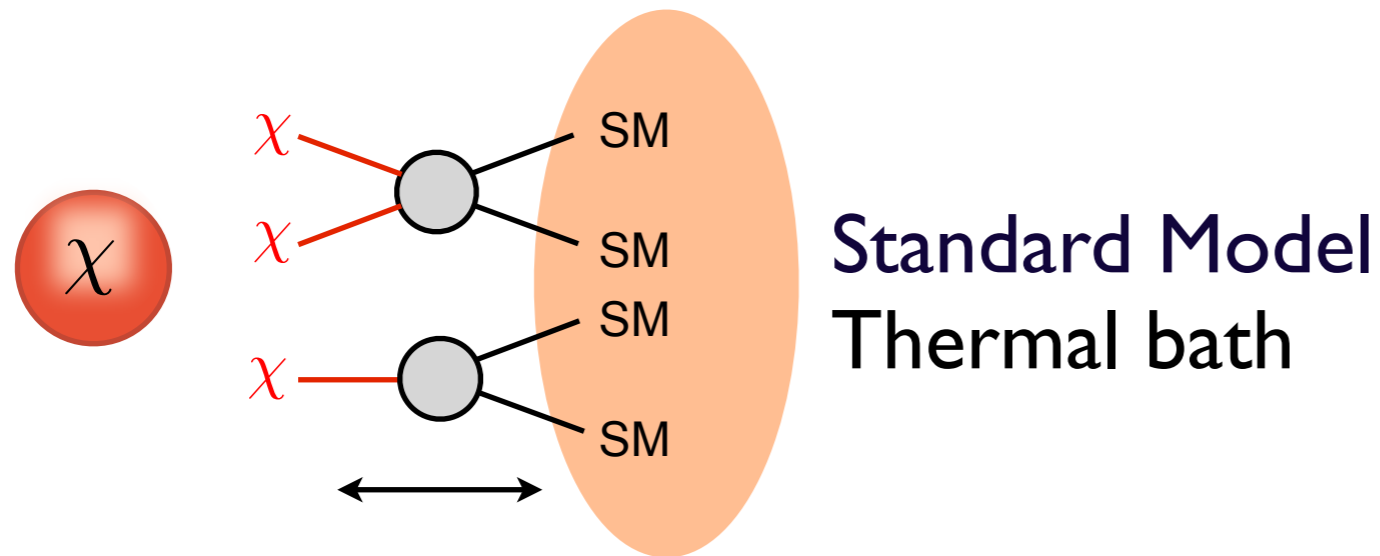
[Zel'dovich, Okun, Pikel'ner 1966; Lee, Weinberg 1977; Binetruy, Girardi, Salati 1984; Bernstein, Brown, Feinberg 1985; Srednicki, Watkins, Olive 1988; Kolb, Turner 1990; Griest, Seckel 1991; Gondolo, Gelmini 1991; Edsjo, Gondolo 1997]





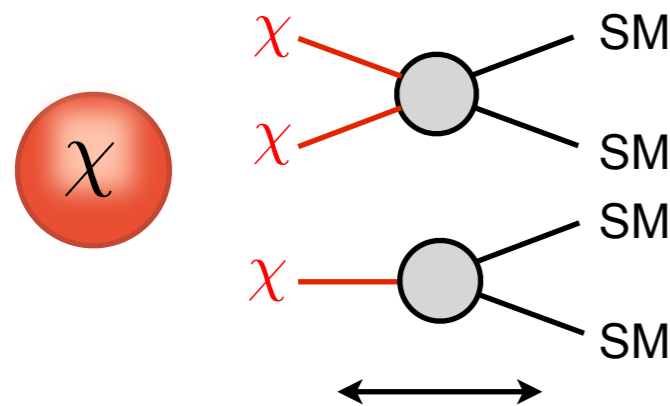
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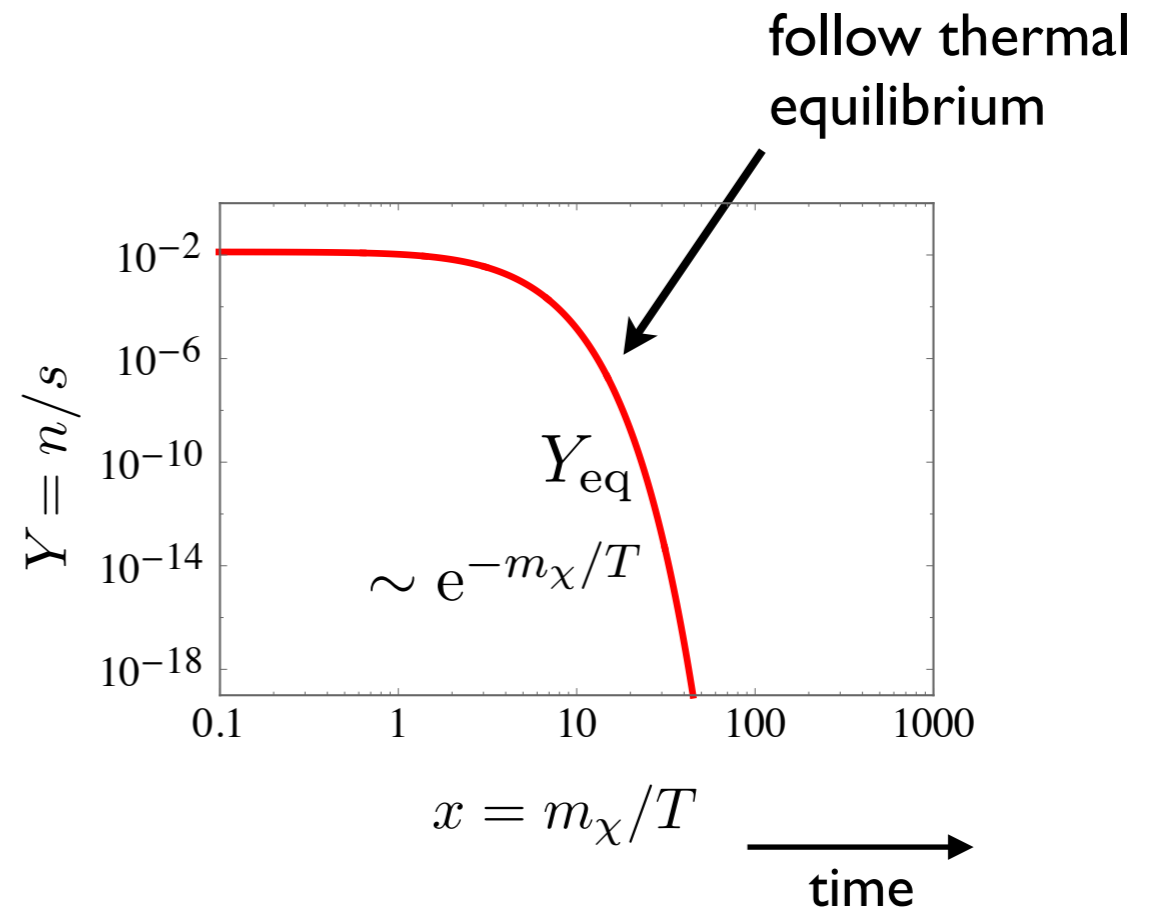
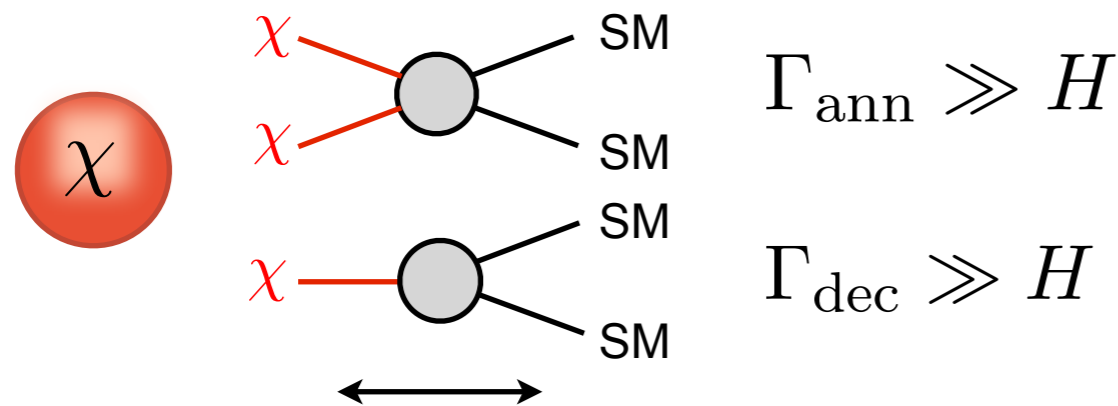
Thermal equilibrium:

$$\Gamma_{\text{ann}} = n_{\chi} \langle \sigma v \rangle_{\text{ann}} \gg H \quad \text{and/or}$$

$$\Gamma_{\text{dec}} \gg H$$

# Boltzmann equations for particle densities

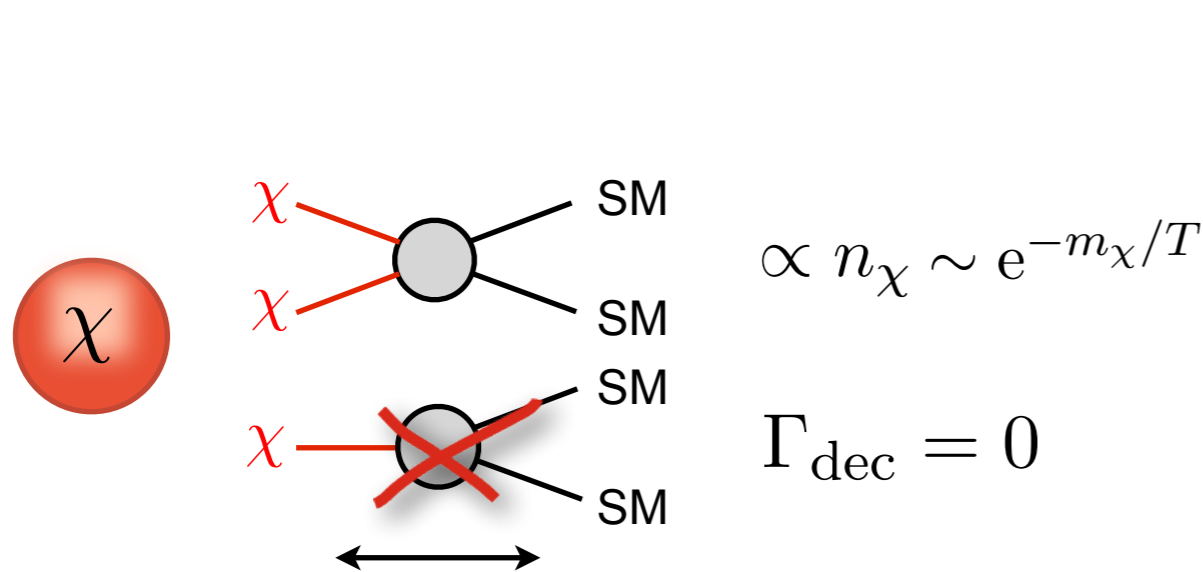
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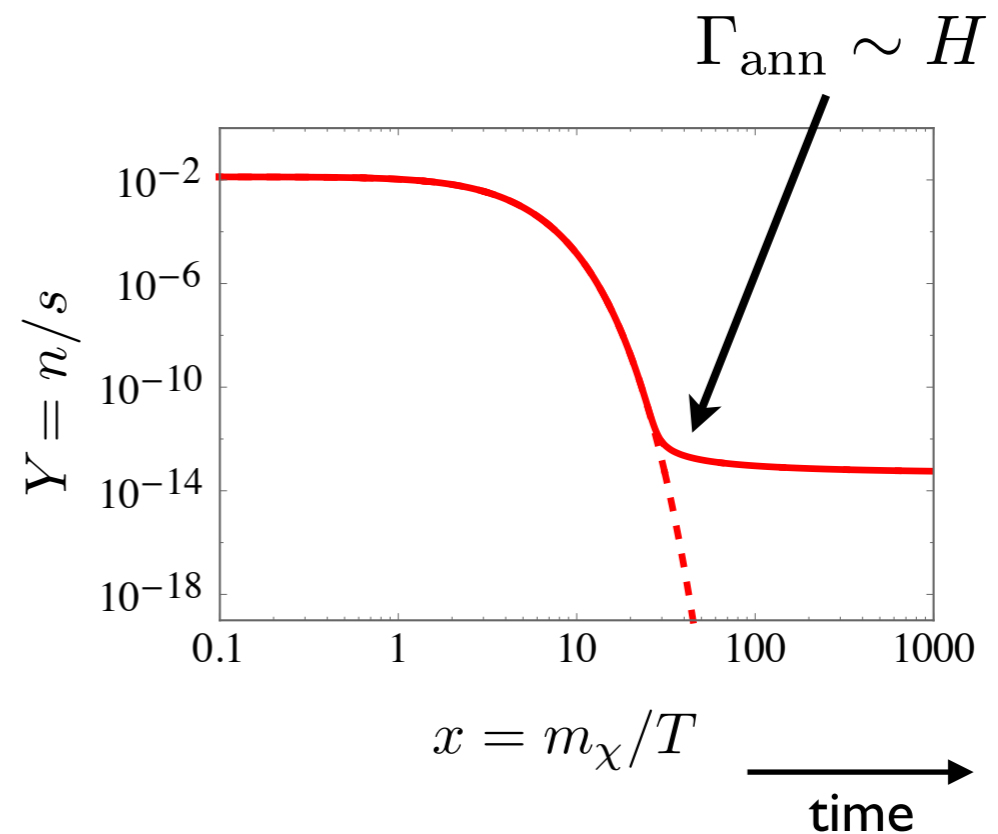
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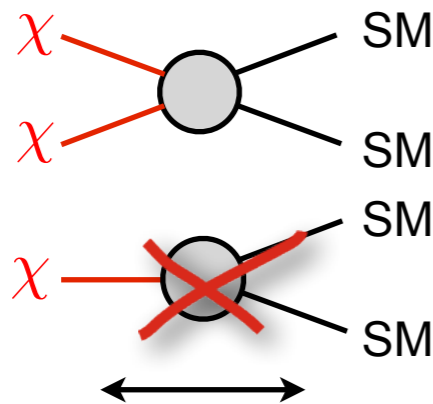
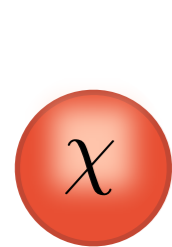
Forbid decay  
e.g. via  $Z_2$ -symmetry

$$\propto n_\chi \sim e^{-m_\chi/T}$$



# Boltzmann equations for particle densities

[Zel'dovich, Okun, Pikel'ner 1966; Lee, Weinberg 1977; Binetruy, Girardi, Salati 1984; Bernstein, Brown, Feinberg 1985; Srednicki, Watkins, Olive 1988; Kolb, Turner 1990; Griest, Seckel 1991; Gondolo, Gelmini 1991; Edsjo, Gondolo 1997]

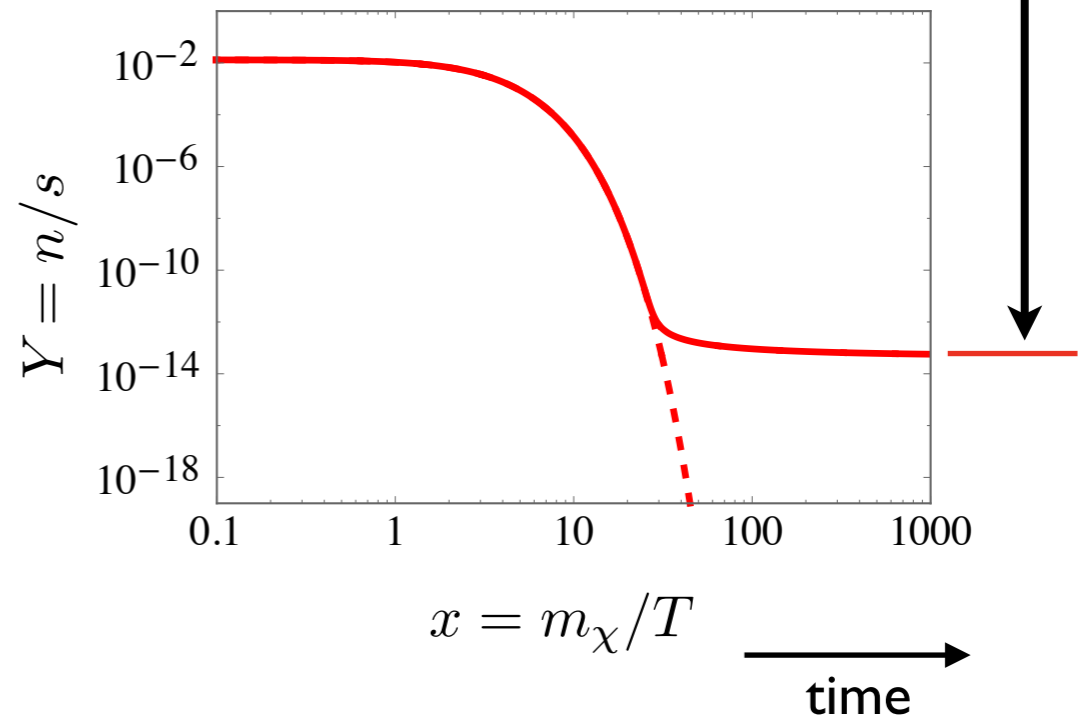


Forbid decay  
e.g. via  $Z_2$ -symmetry

$$\propto n_\chi \sim e^{-m_\chi/T}$$

$$\Gamma_{\text{dec}} = 0$$

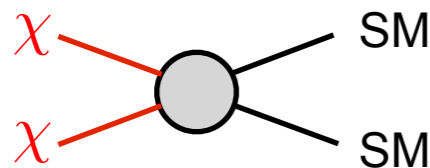
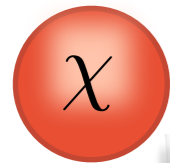
$$(\Omega h^2)_{\text{Planck}} \simeq 0.12$$



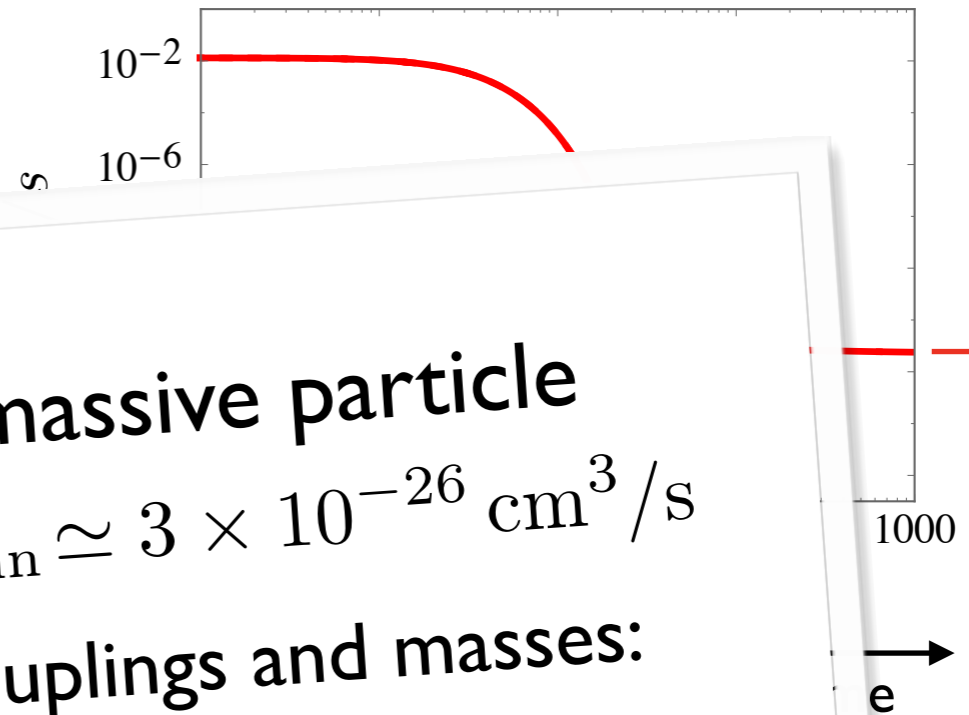
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$$(\Omega h^2)_{\text{Planck}} \simeq 0.12$$



$$\propto n_\chi \sim e^{-m_\chi/T}$$



**Weakly interacting massive particle**

$$\Omega_{\text{DM}} h^2 \simeq 0.12 \Rightarrow \langle \sigma v \rangle_{\text{ann}} \simeq 3 \times 10^{-26} \text{ cm}^3/\text{s}$$

$\Rightarrow$  roughly with weak couplings and masses:

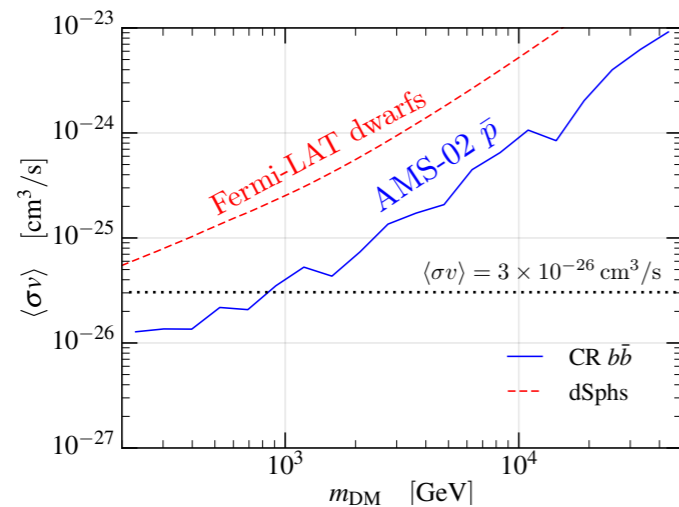
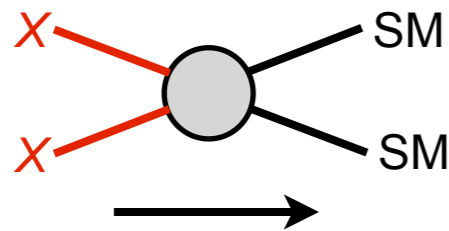
**"WIMP miracle"**

# WIMP paradigm attractive

- Works with simple/natural models ("WIMP miracle")
- Independent of largely unconstrained/unknown physics of the very early universe (inflation/reheating)
- Naturally provides perfectly cold dark matter (CDM)
- Testable at indirect detection, direct detection and collider experiments

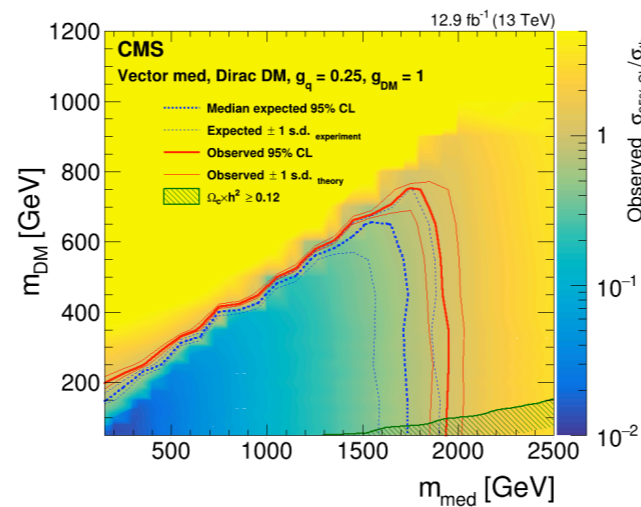
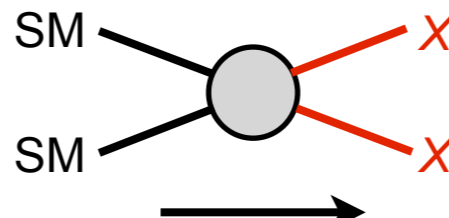
# WIMP Dark Matter: searches

## Indirect detection



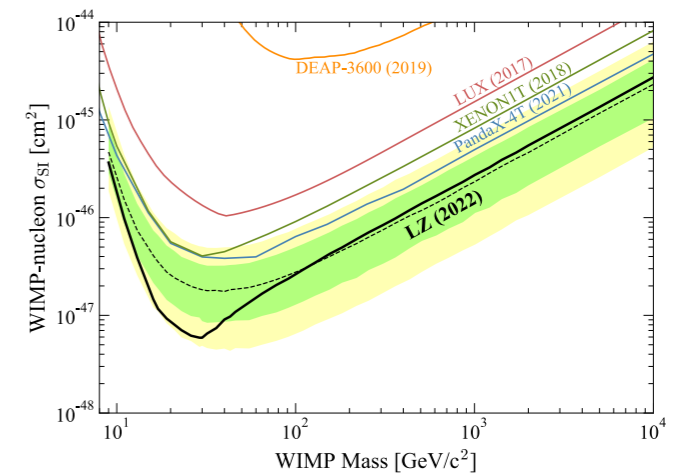
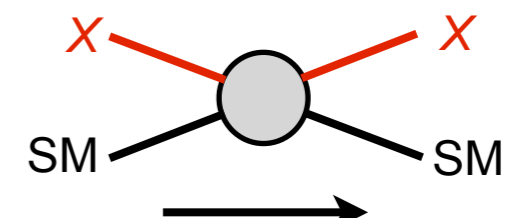
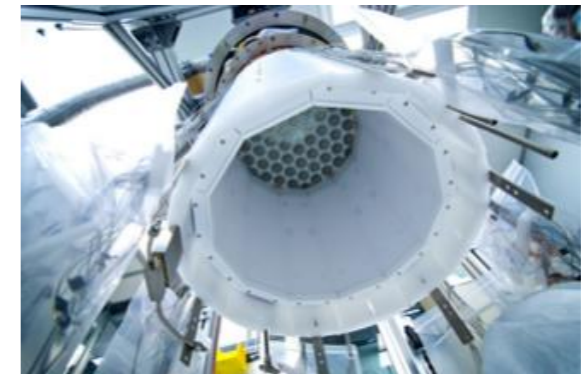
[Cuoco, JH, Korsmeier, Krämer 2017]

## Direct production



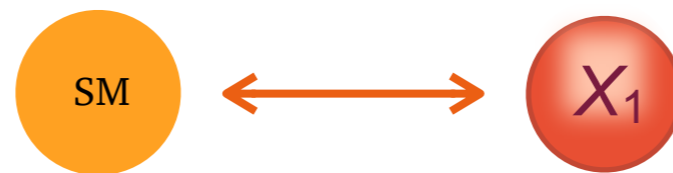
[CMS EXO-16-039]

## Direct detection



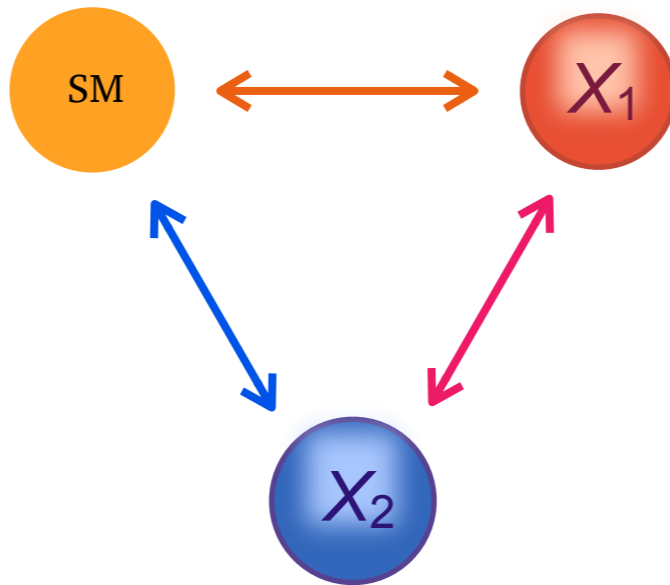
[LZ 2207.03764]

# Decouple constraints WIMP searches and relic density

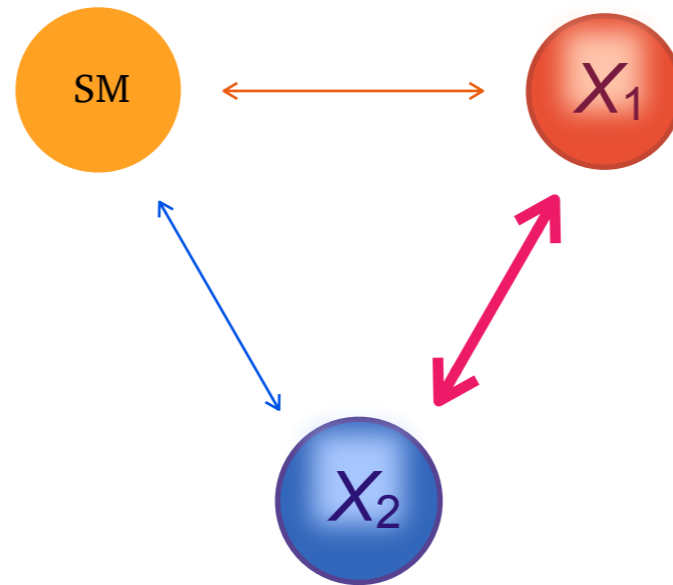




# Decouple constraints WIMP searches and relic density

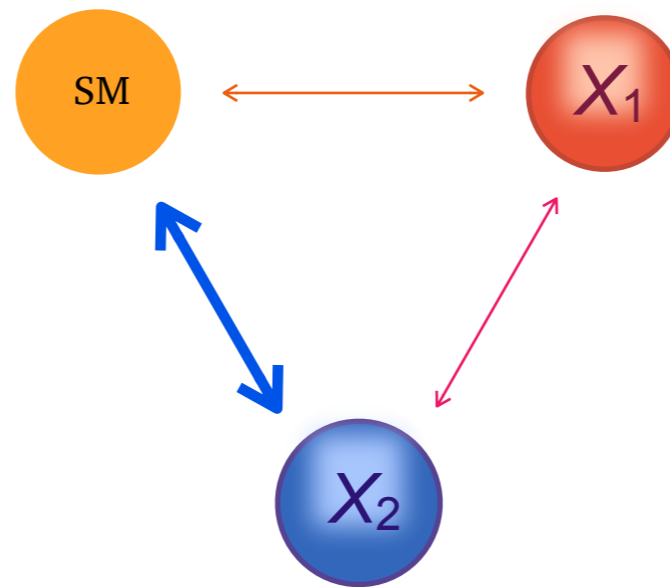


# Decouple constraints WIMP searches and relic density



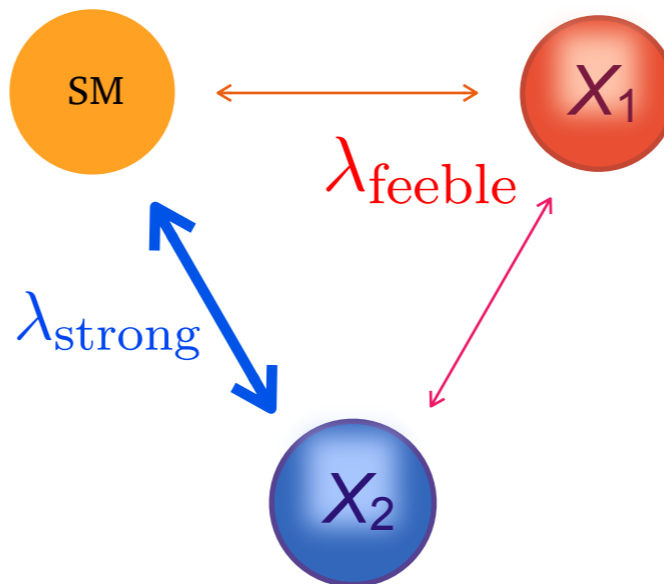
- Freeze-out in secluded sector

# Decouple constraints WIMP searches and relic density



- Variety of scenarios in  $t$ -channel mediator models ('charged parent particle model')

# Decouple constraints WIMP searches and relic density



- Large hierarchy of couplings:

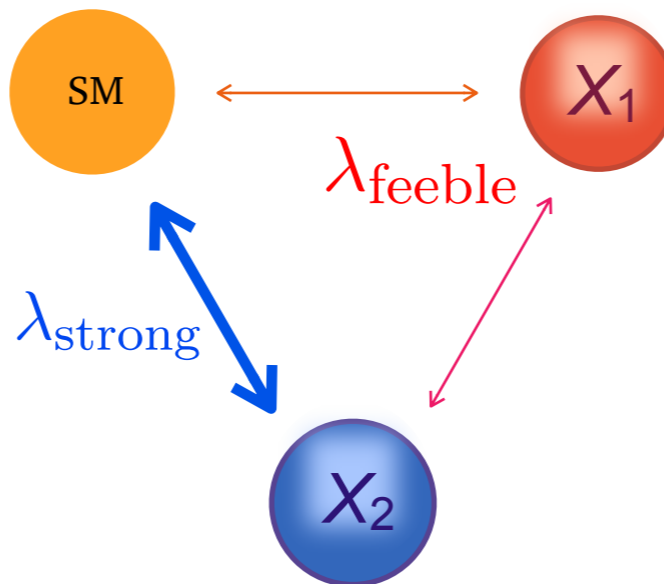
$$\lambda_{\text{strong}} \gg \lambda_{\text{feeble}}$$

- Interesting for the subject for two reasons:

(i) ‘strong’ coupling: significant bound state effects

(ii) very weak (‘feeble’) coupling: suppresses scattering and decay rates  
 $\Rightarrow$  typically prolonged freeze-out process

# Decouple constraints WIMP searches and relic density



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$$\lambda_{\text{strong}} \gg \lambda_{\text{feible}}$$

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(i) ‘strong’ coupling: significant bound state effects

(ii) very weak (‘feeble’) coupling: suppresses scattering and decay rates

⇒ typically prolonged freeze-out process

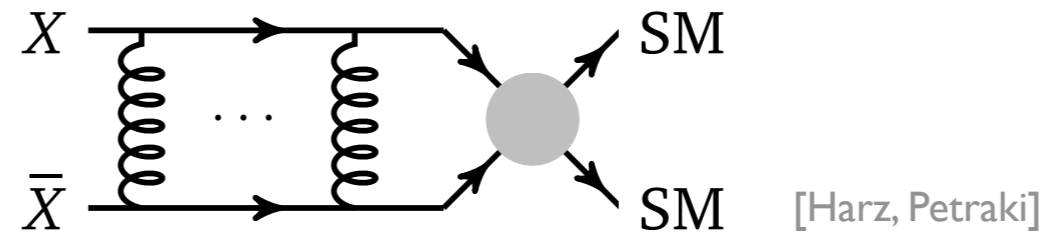
⇒ bound state at late times (low  $T$ ): excitations highly relevant

**Effects of (excited) bound state**



# Non-relativistic effects

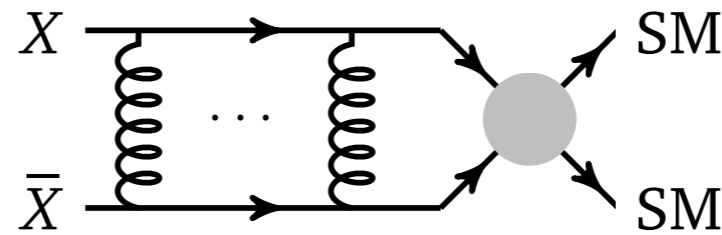
I. Sommerfeld effect:



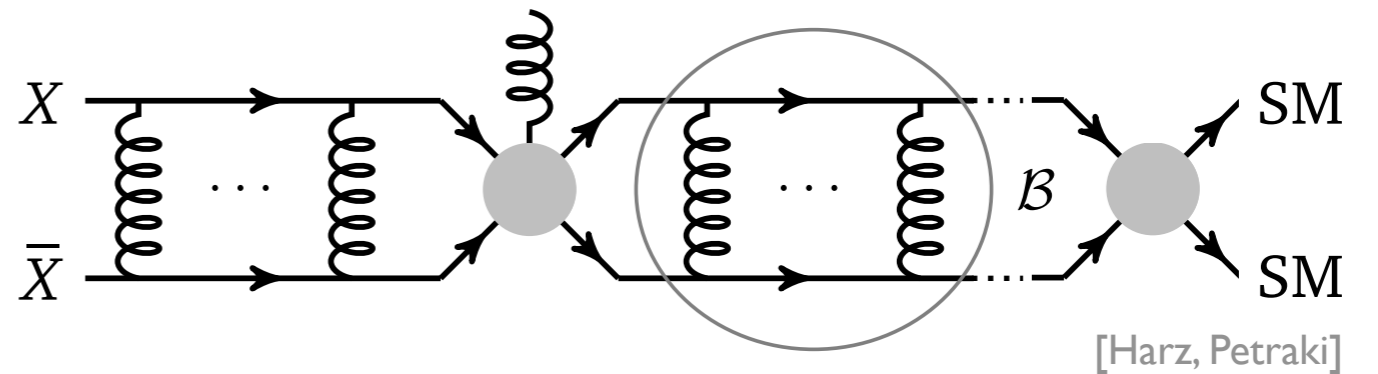
[see e.g. Hisano, Matsumoto, Nojiri hep-ph/0307216; Hisano, Matsumoto, Nojiri, Saito hep-ph/0412403; ...]

# Non-relativistic effects

1. Sommerfeld effect:



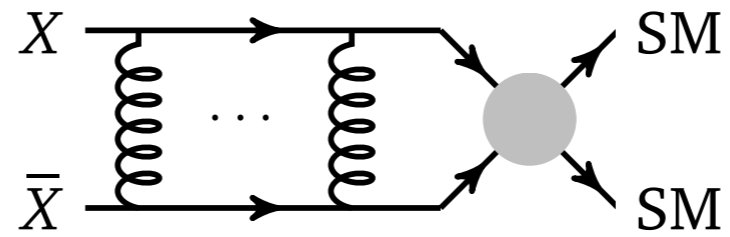
2. Bound state formation:



[see e.g. K. Petraki, M. Postma, M. Wiechers 1505.00109; S.P. Liew, F. Luo 1611.08133; J. Harz, K. Petraki 1805.01200; A. Mitridate, M. Redi, J. Smirnov, A. Strumia 1702.01141; T. Binder, B. Blobel, J. Harz, and K. Mukaida 2002.07145; ...]

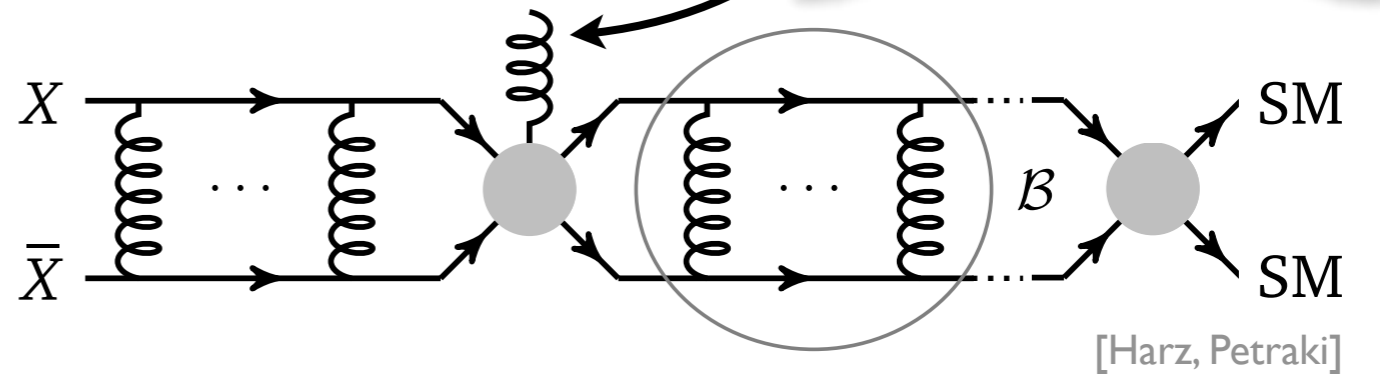
# Non-relativistic effects

1. Sommerfeld effect:



$$\Delta E = \frac{\mathbf{p}^2}{2\mu} + E_{\mathcal{B}}$$

2. Bound state formation:

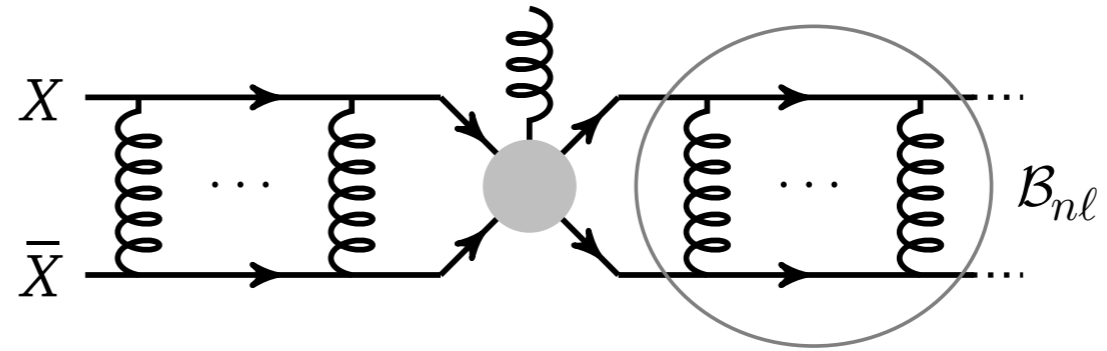


[Harz, Petraki]

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# Computation of bound state formation

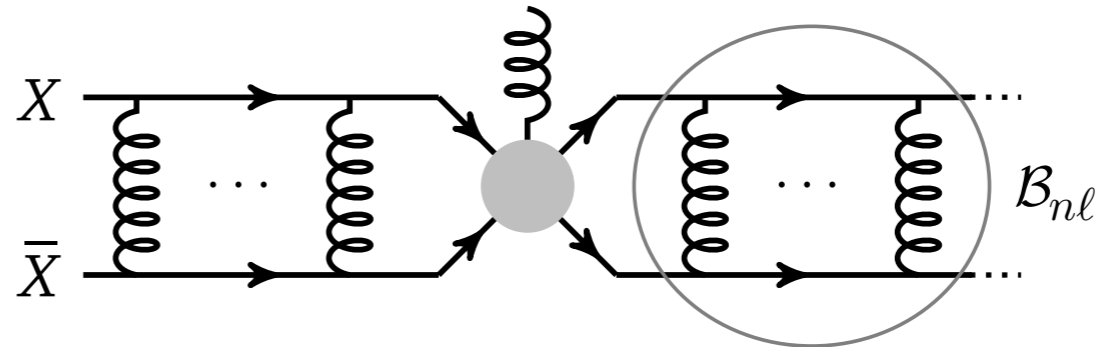
2. Bound state formation:



[Harz, Petraki]

# Computation of bound state formation

## 2. Bound state formation:



[Harz, Petraki]

- Force carrier = vector (gauge field):  
bound state formation  $X\bar{X} \sim$  dipole transition:

$$(\sigma v)_{nl} \propto |\langle \psi_{nl} | \mathbf{r} | \psi_{\mathbf{p}} \rangle|^2$$

$$g \mathbf{r} \cdot \mathbf{E}$$

[(Color-)electric dipole operator (in pNRQCD), see e.g. Yao+ 2019]

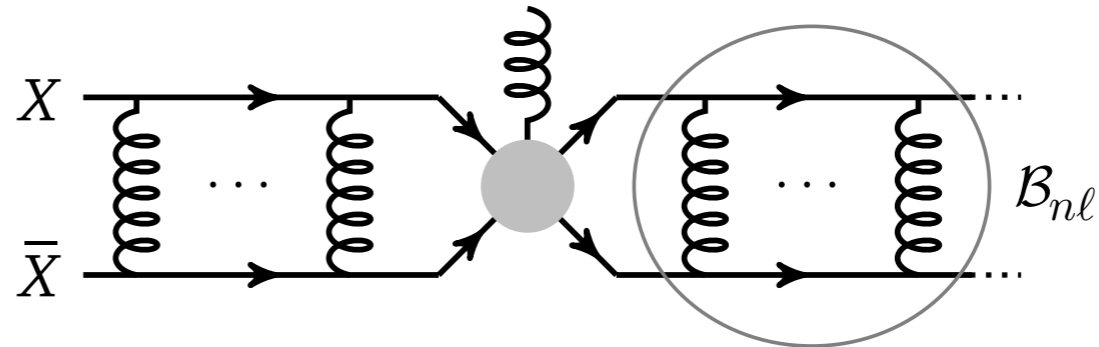
bound state

scattering state

Wave functions: solutions to two-body Schrödinger eqs.

# Computation of bound state formation

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[Harz, Petraki]

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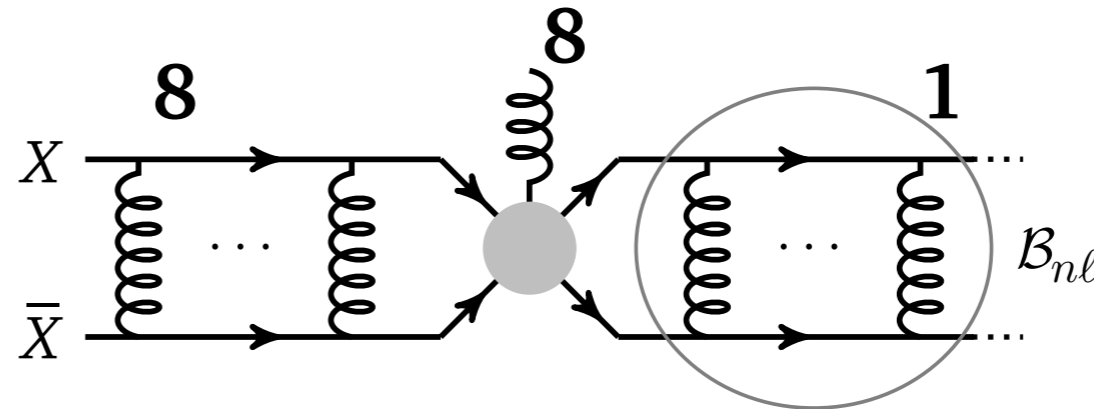
- Force carrier massless (unbroken gauge theory):

Coulomb limit:  $V_{[R]}(r) = -\frac{\alpha_{[R]}^{\text{eff}}}{r}$  in unconfined phase



# Computation of bound state formation

## 2. Bound state formation:



[Harz, Petraki]

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- Force carrier massless (unbroken gauge theory):

Coulomb limit:  $V_{[\mathbf{R}]}(r) = -\frac{\alpha_{[\mathbf{R}]}}{r}$ ,  $\alpha_{[\mathbf{R}]}^{\text{eff}} = \alpha_s \times \begin{cases} 4/3, & \mathbf{R} = 1, \\ -1/6, & \mathbf{R} = 8. \end{cases}$

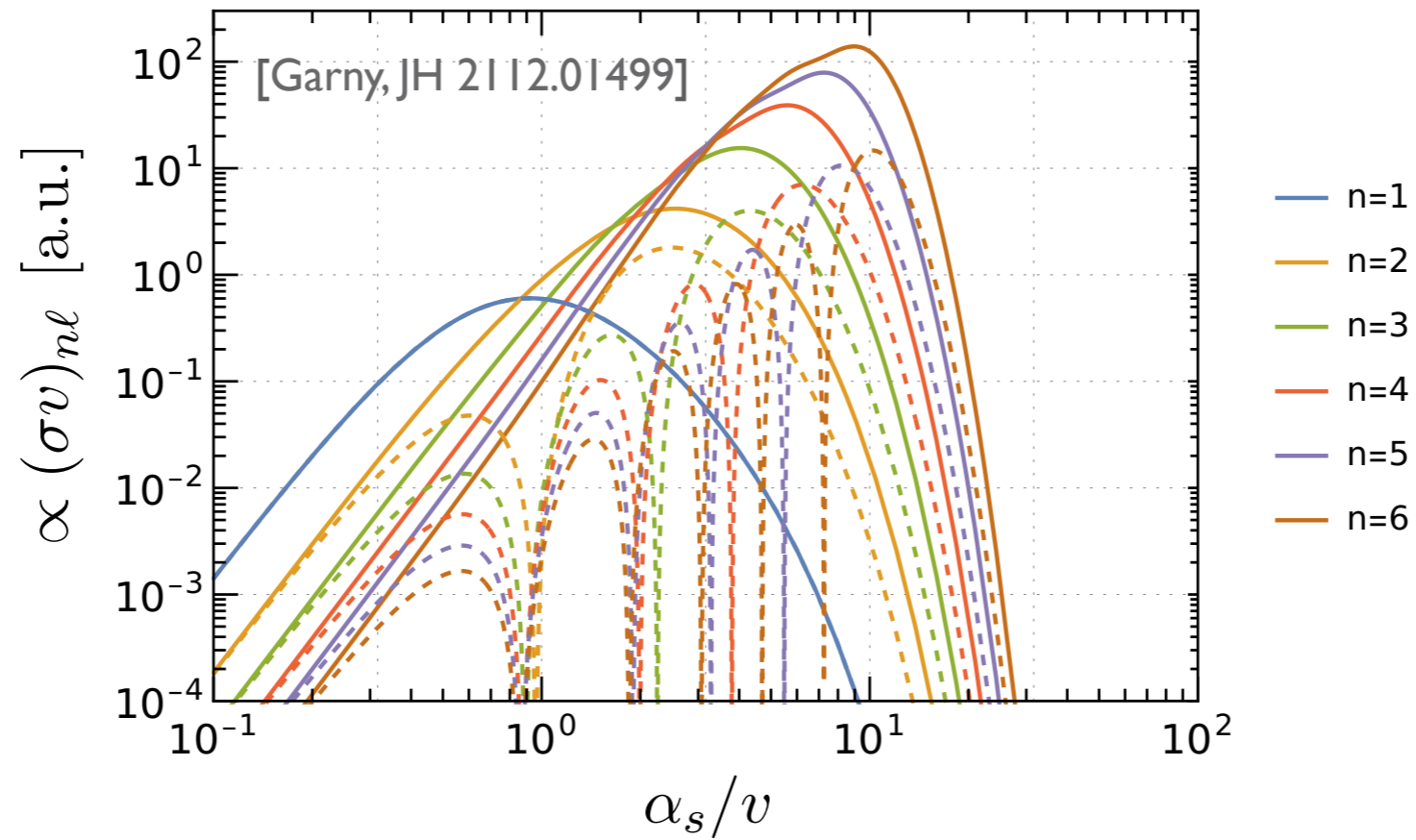
- Non-abelian gauge theory:

e.g.  $X$  in  $\mathbf{3}$  of  $SU(3)$ :  $\mathbf{3} \otimes \bar{\mathbf{3}} = \mathbf{1} \oplus \mathbf{8}$

involves repulsive potential

# Bound state formation cross section

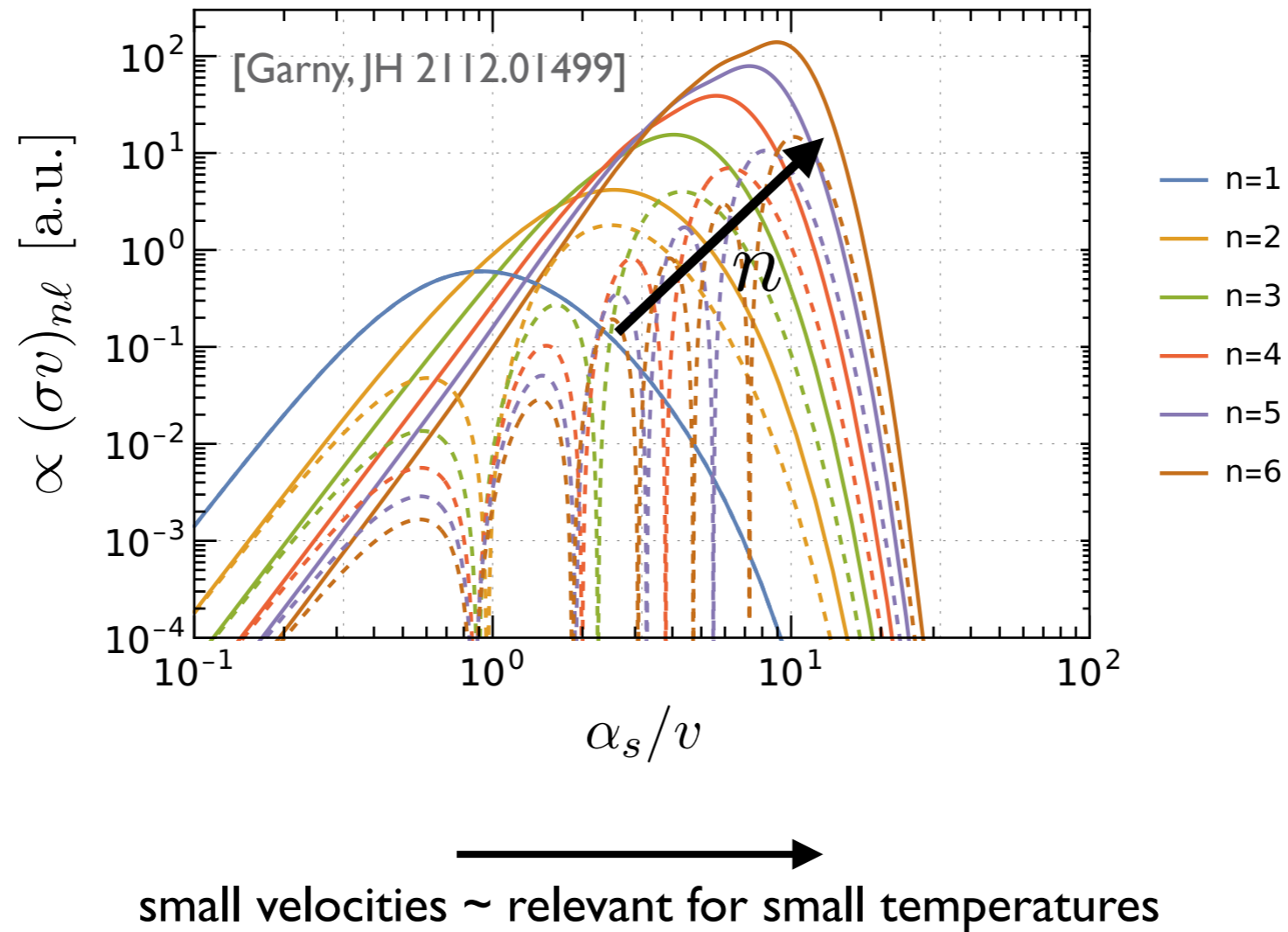
$$(\sigma v)_{nl} \propto |\langle \psi_{nl} | \mathbf{r} | \psi_{\mathbf{p}} \rangle|^2$$



→  
small velocities ~ relevant for small temperatures

# Bound state formation cross section

$$(\sigma v)_{nl} \propto |\langle \psi_{nl} | \mathbf{r} | \psi_{\mathbf{p}} \rangle|^2$$



# Partial-wave unitarity violating?

[Binder, Garny, JH, Lederer, Urban 2308.01336]

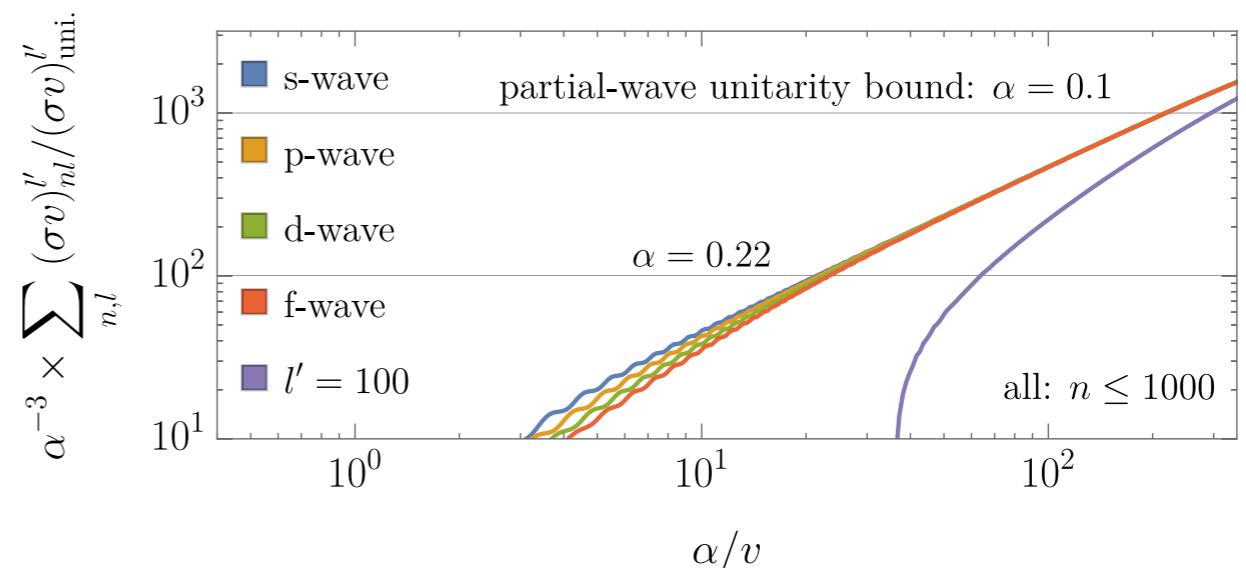
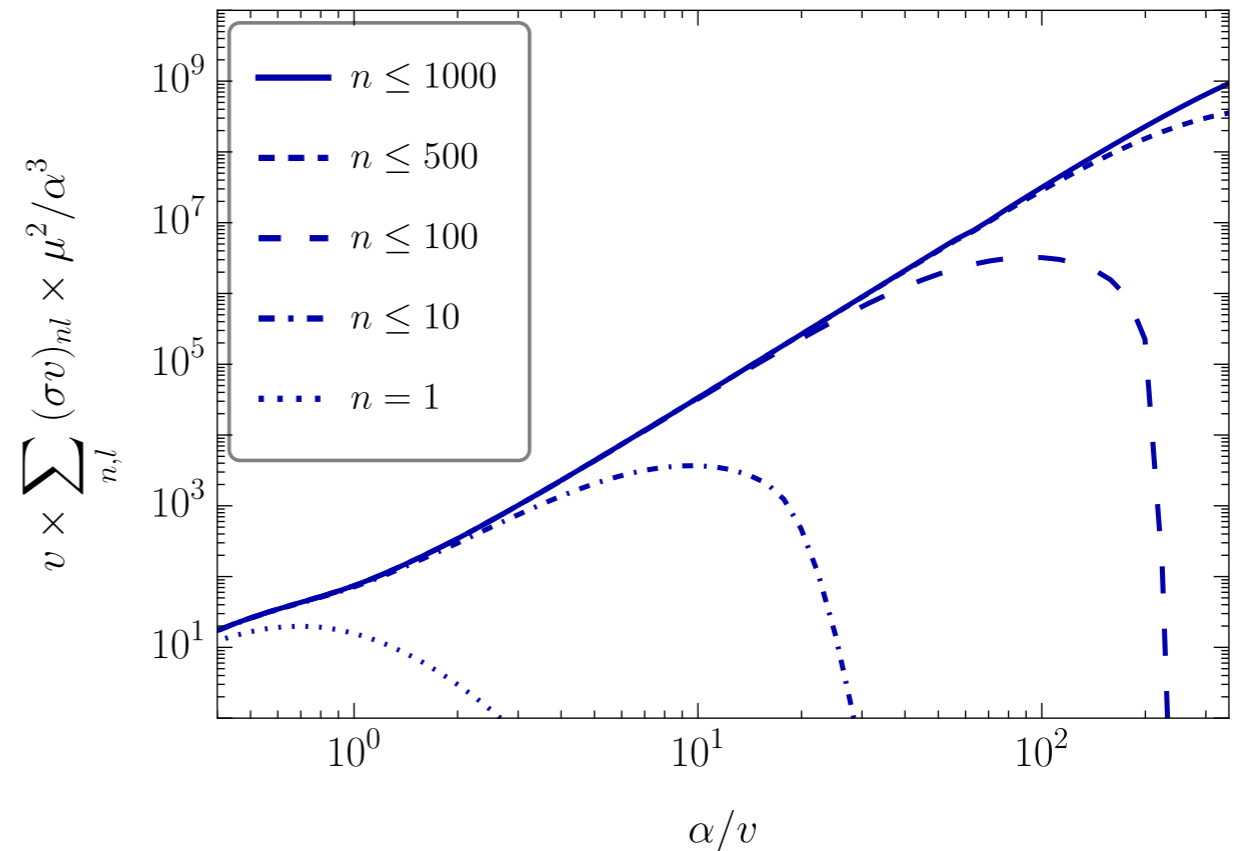
- Derived recursion formulas for highly efficient and stable computation of  $(\sigma v)_{nl}$  up to  $n = 1000$ ,  $\ell \leq n - 1$  (~million states)

- Impose partial-wave unitarity:  
[Griest, Kamionkowski 1990]

$$\begin{aligned}
 (\sigma v)^{\ell'} &= \sum_{n,\ell} (\sigma v)_{n\ell}^{\ell'} \\
 &\leq \frac{\pi(2\ell' + 1)}{\mu^2 v}
 \end{aligned}$$

[see also Harling, Petraki 2014;  
Baltes, Petraki 2017;  
Smirnov, Beacom 2019;  
Bottaro, Redigolo 2023]

Summed BSF cross section,  $SU(3)$



# Partial-wave unitarity violating?

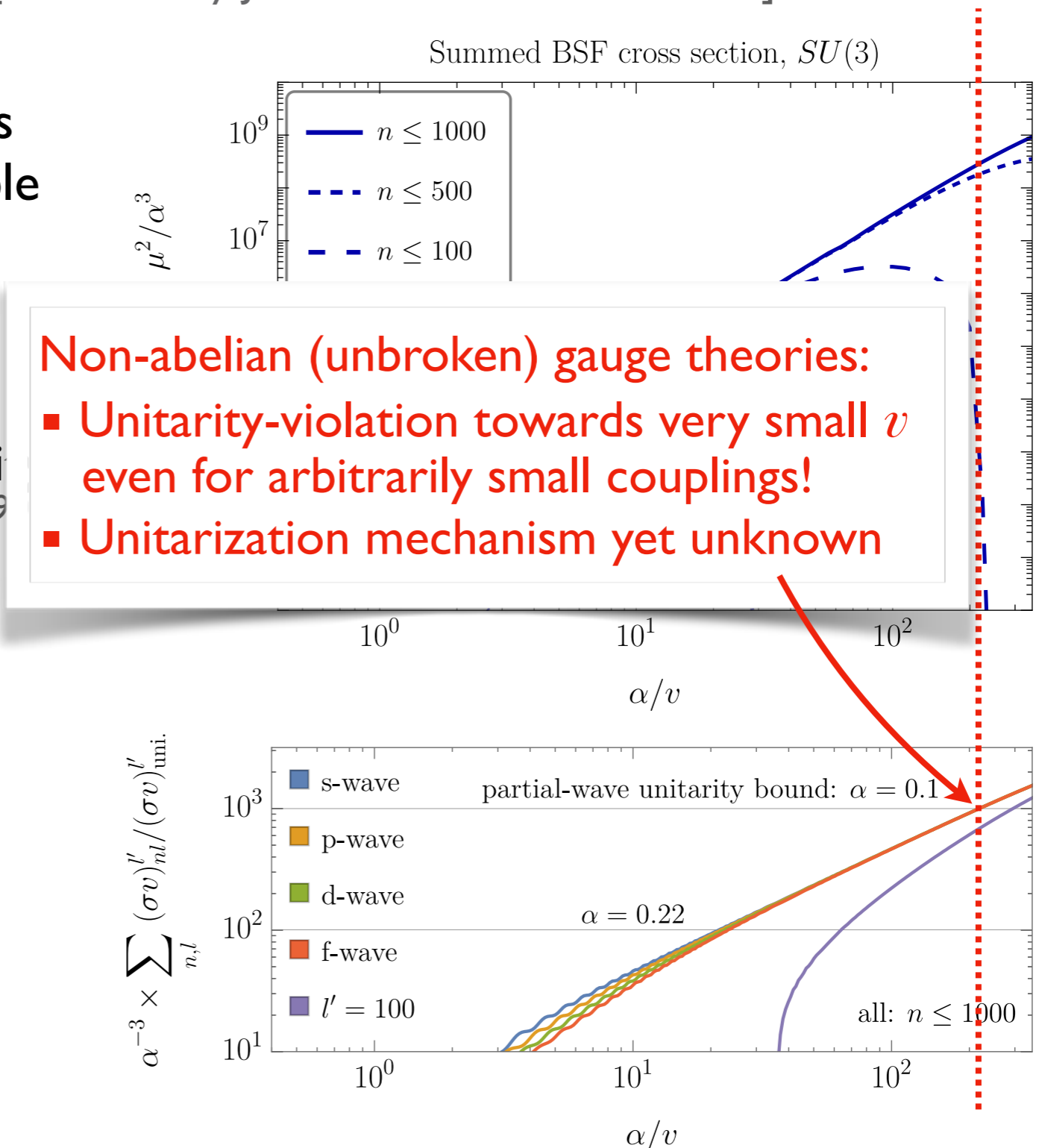
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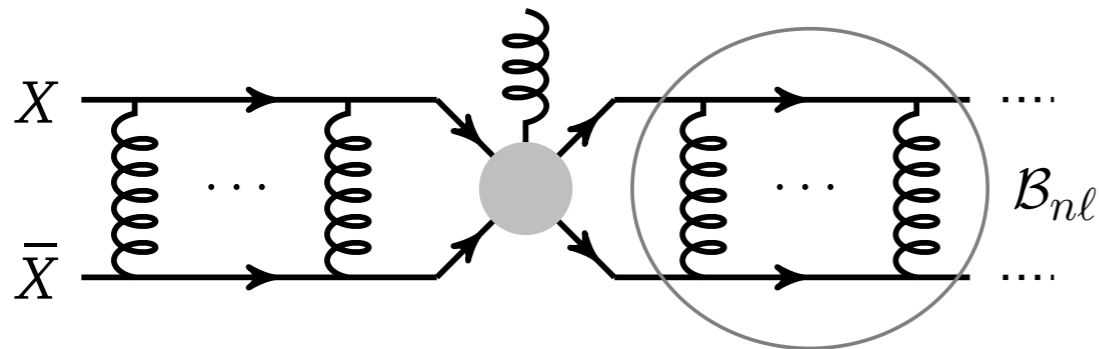
[Griest, Kamionkowski 1998]

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Smirnov, Beacom 2019;  
Bottaro, Redigolo 2023]



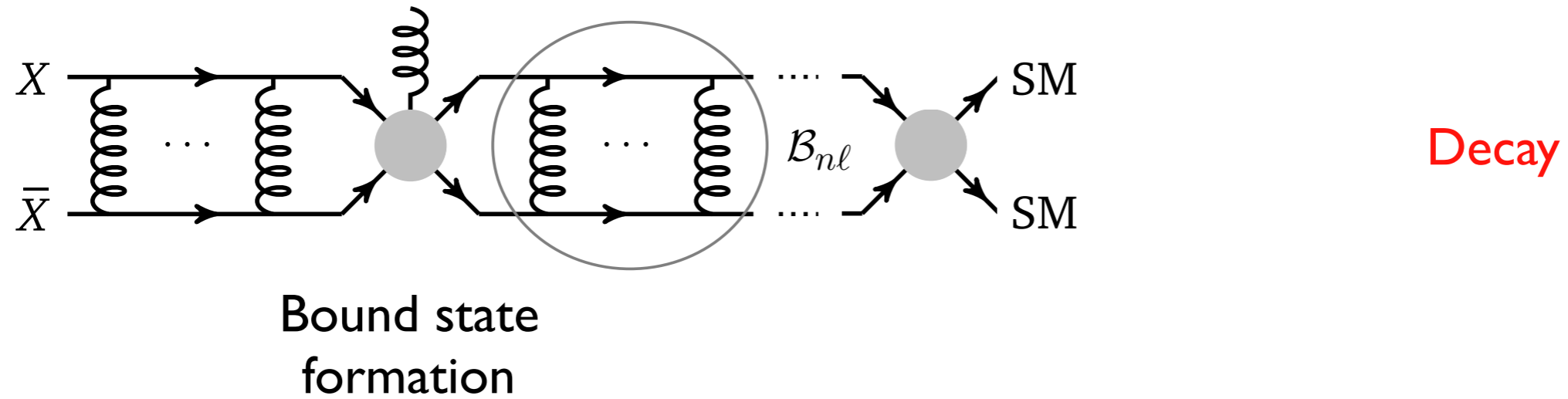
# Excited bound states in the thermal bath



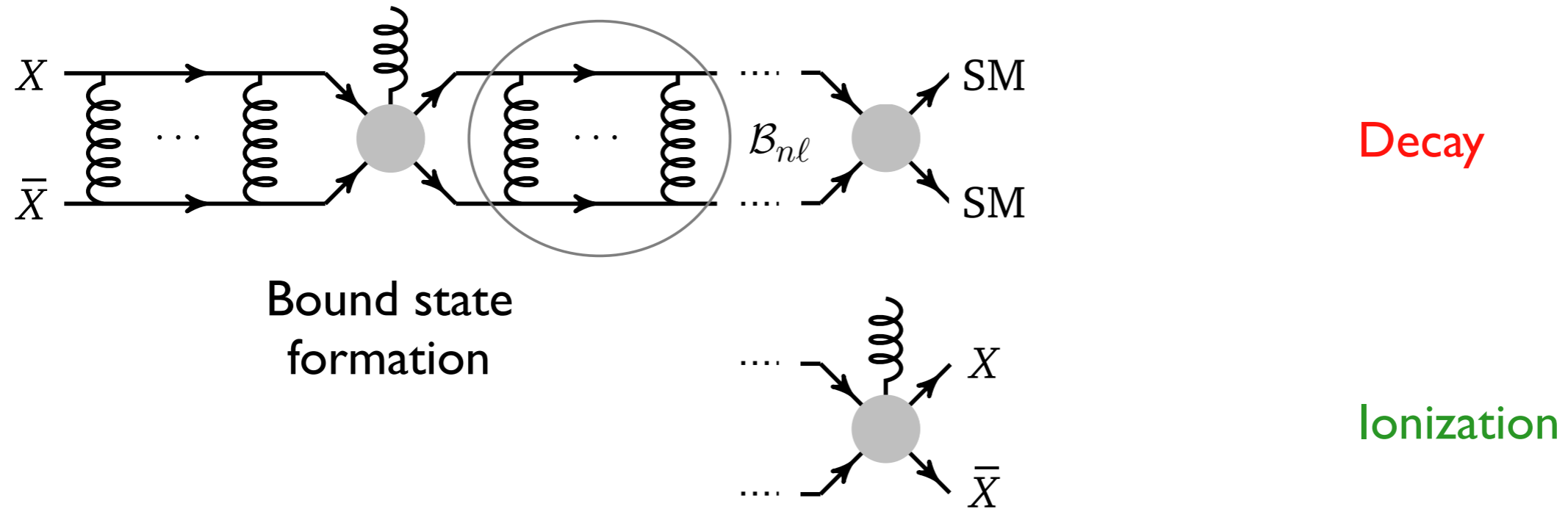
Bound state  
formation



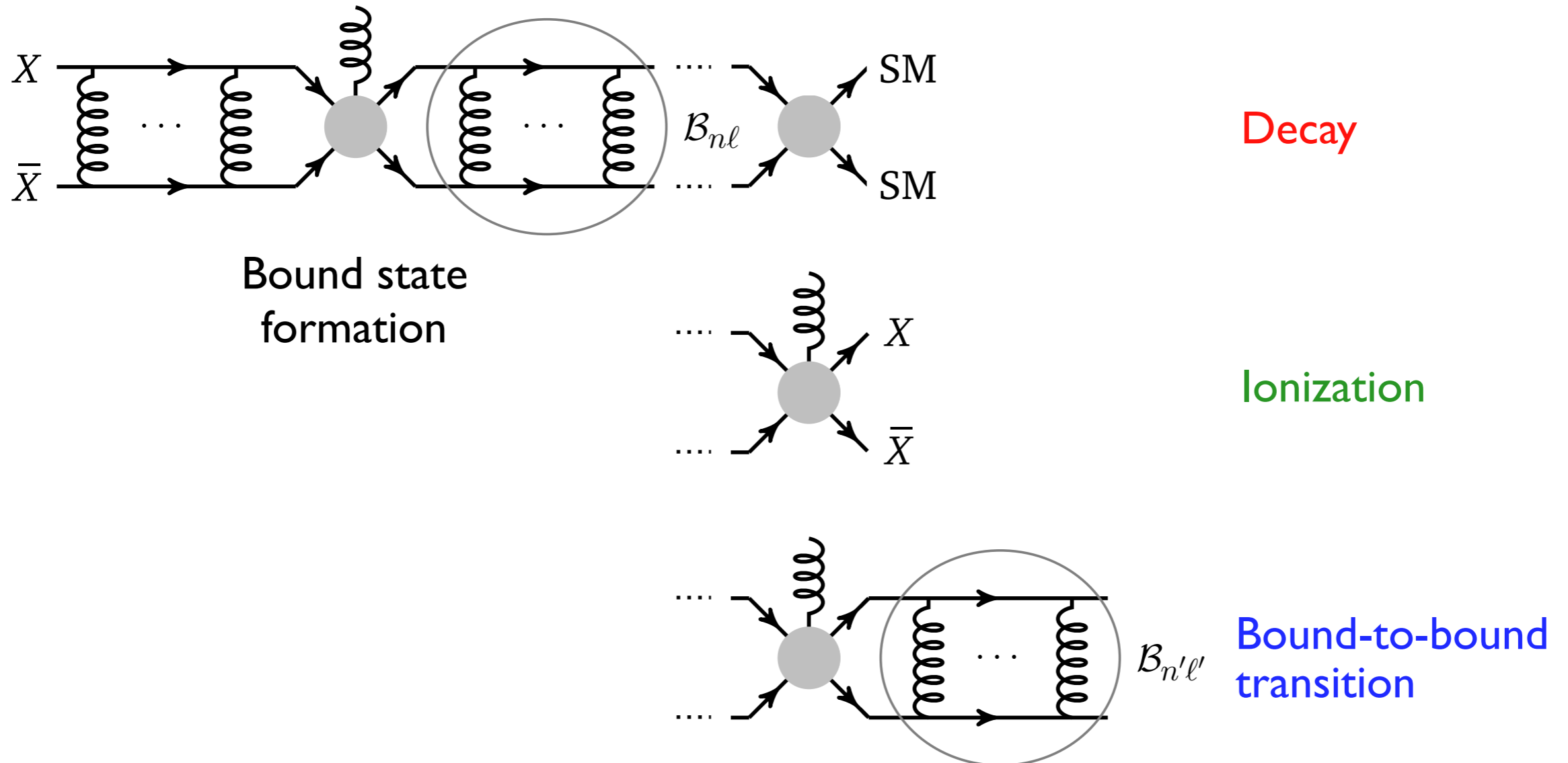
# Excited bound states in the thermal bath



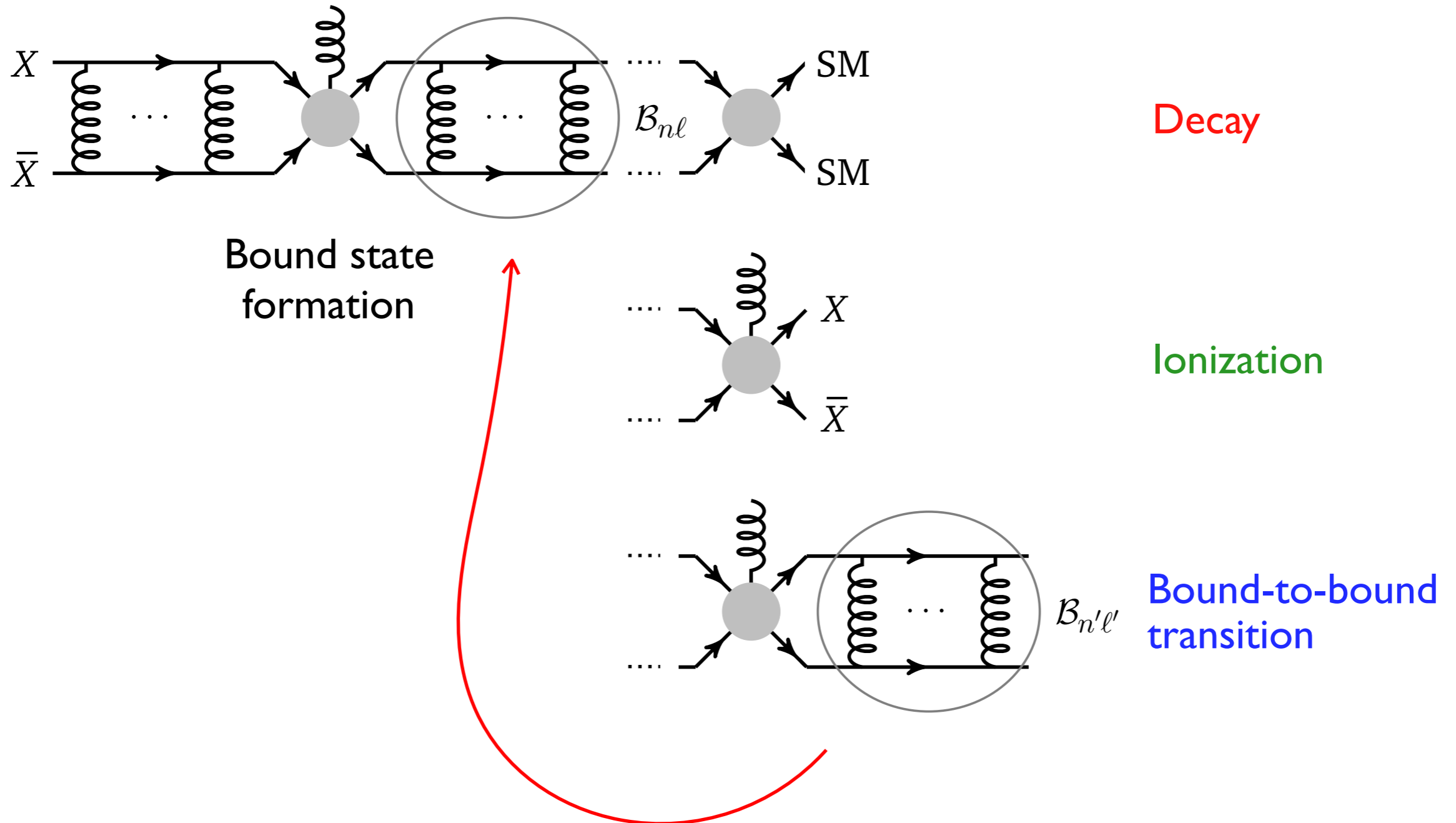
# Excited bound states in the thermal bath



# Excited bound states in the thermal bath



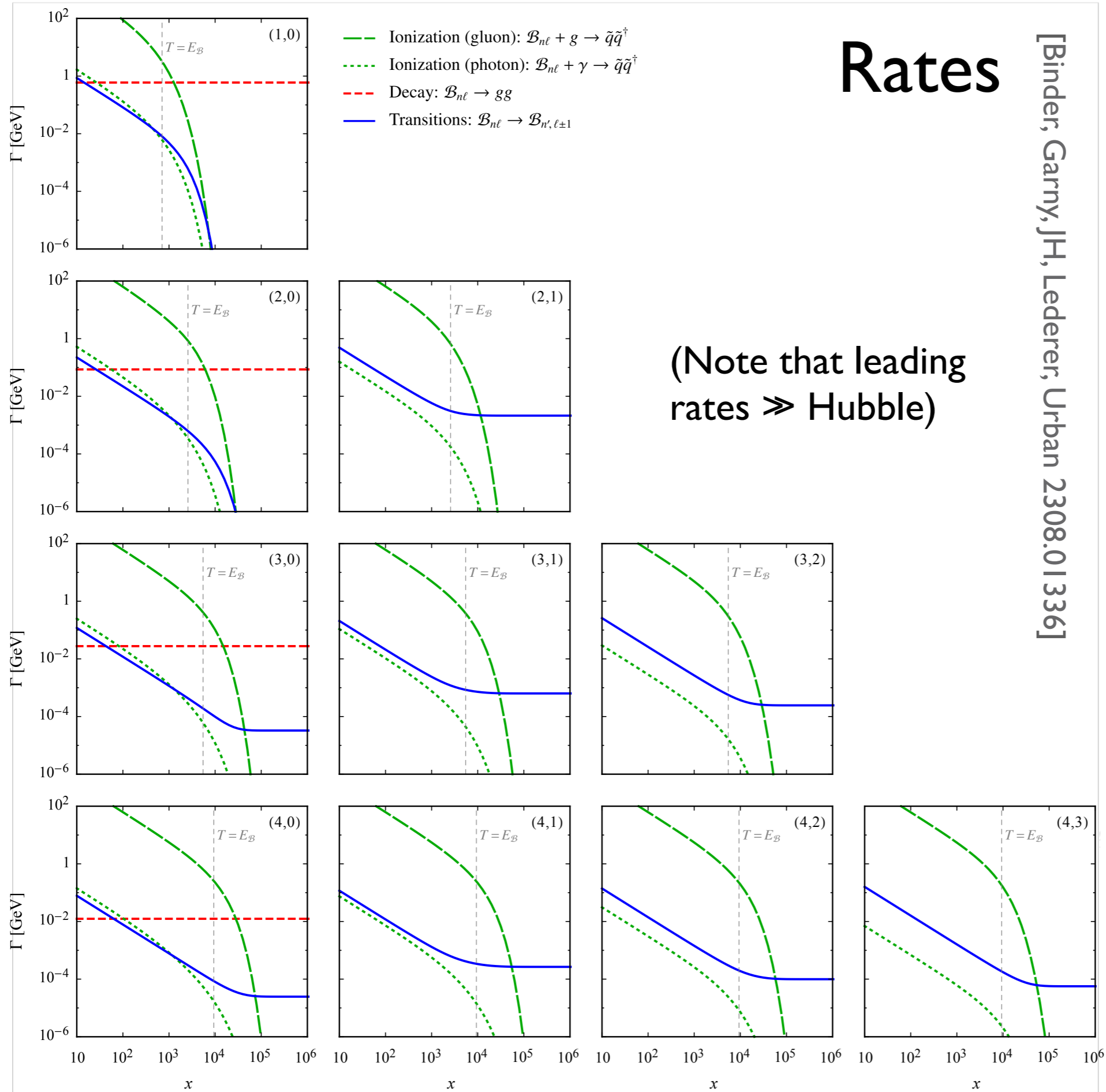
# Excited bound states in the thermal bath



# Rates

[Binder, Garny, JH, Lederer, Urban 2308.01336]

- Ionization (gluon):  $\mathcal{B}_{n\ell} + g \rightarrow \tilde{q}\tilde{q}^\dagger$
- - - Ionization (photon):  $\mathcal{B}_{n\ell} + \gamma \rightarrow \tilde{q}\tilde{q}^\dagger$
- - - Decay:  $\mathcal{B}_{n\ell} \rightarrow gg$
- Transitions:  $\mathcal{B}_{n\ell} \rightarrow \mathcal{B}_{n',\ell\pm 1}$



Decay

Ionization

Bound-to-bound transition

# Boltzmann equations including excitations

[Garny, JH 21 | 2.01499]

$X$  pair-annihilation  
(hard process+Sommerfeld)

Bound state  
formation

$$\textcircled{X} : \frac{dY_X}{dx} = \frac{1}{3Hs} \frac{ds}{dx} \left[ \frac{1}{2} \langle \sigma_{XX^+v} \rangle \left( Y_X^2 - Y_X^{\text{eq}2} \right) + \sum_i \frac{1}{2} \langle \sigma_{\text{BSF},iv} \rangle \left( Y_X^2 - Y_X^{\text{eq}2} \frac{Y_{\mathcal{B}_i}}{Y_{\mathcal{B}_i}^{\text{eq}}} \right) \right]$$

$$\textcircled{X} : \frac{dY_{\mathcal{B}_i}}{dx} = \frac{1}{3Hs} \frac{ds}{dx} \left[ \Gamma_{\text{ion}}^i \left( Y_{\mathcal{B}_i} - Y_{\mathcal{B}_i}^{\text{eq}} \frac{Y_X^2}{Y_X^{\text{eq}2}} \right) + \Gamma_{\text{dec}}^i \left( Y_{\mathcal{B}_i} - Y_{\mathcal{B}_i}^{\text{eq}} \right) - \sum_{j \neq i} \Gamma_{\text{trans}}^{j \rightarrow i} \left( Y_{\mathcal{B}_j} - Y_{\mathcal{B}_i} \frac{Y_{\mathcal{B}_j}^{\text{eq}}}{Y_{\mathcal{B}_i}^{\text{eq}}} \right) \right]$$

Ionization

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[see also Binder et al. 21 | 2.00042]

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0

Ionization

Decay

Bound-to-bound  
transition

→ Solve for  $Y_{\mathcal{B}_i}$  algebraically: set of linear equations!

[see also Binder et al. 21 | 2.00042]



# Boltzmann equations including excitations

[Garny, JH 21 | 2.0 | 499]

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0

Ionization

Decay

Bound-to-bound  
transition

Solve for  $Y_{\mathcal{B}_i}$  algebraically: set of linear equations!  
Reduce to single BME:

$$\textcircled{X} + \textcircled{X} : \frac{dY_X}{dx} = \frac{1}{3Hs} \frac{ds}{dx} \frac{1}{2} \langle \sigma_{XX^\dagger v} \rangle_{\text{eff}} (Y_X^2 - Y_X^{\text{eq}2})$$

where  $\langle \sigma_{XX^\dagger v} \rangle_{\text{eff}} = \langle \sigma_{XX^\dagger v} \rangle + \sum_i \langle \sigma_{\text{BSF},i v} \rangle R_i$ ,  $0 \leq R_i \leq 1$

[see also Binder et al. 21 | 2.00042]

# Boltzmann equations including excitations

[Garny, JH 21 | 2.0 | 499]

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0

Bound states effectively enhance  
the  $X$  pair-annihilation cross section

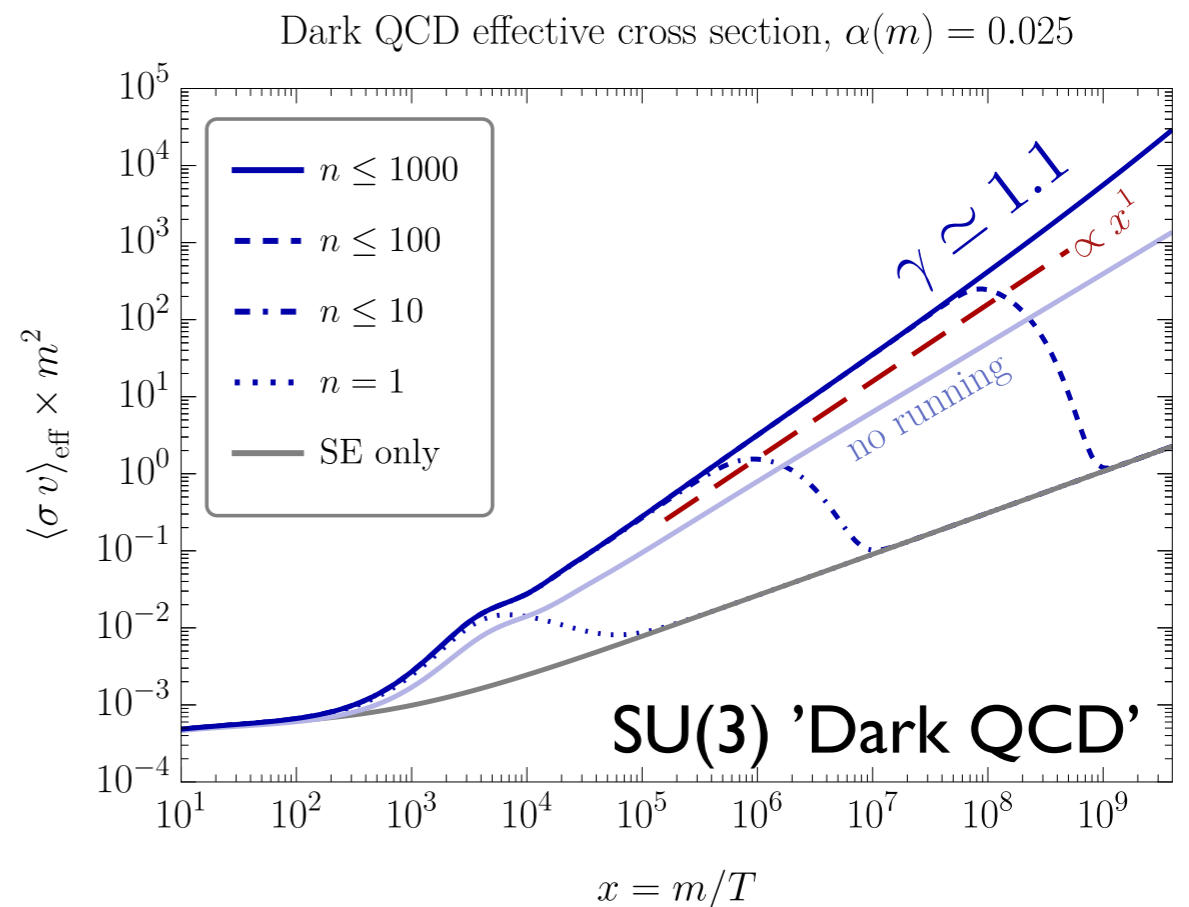
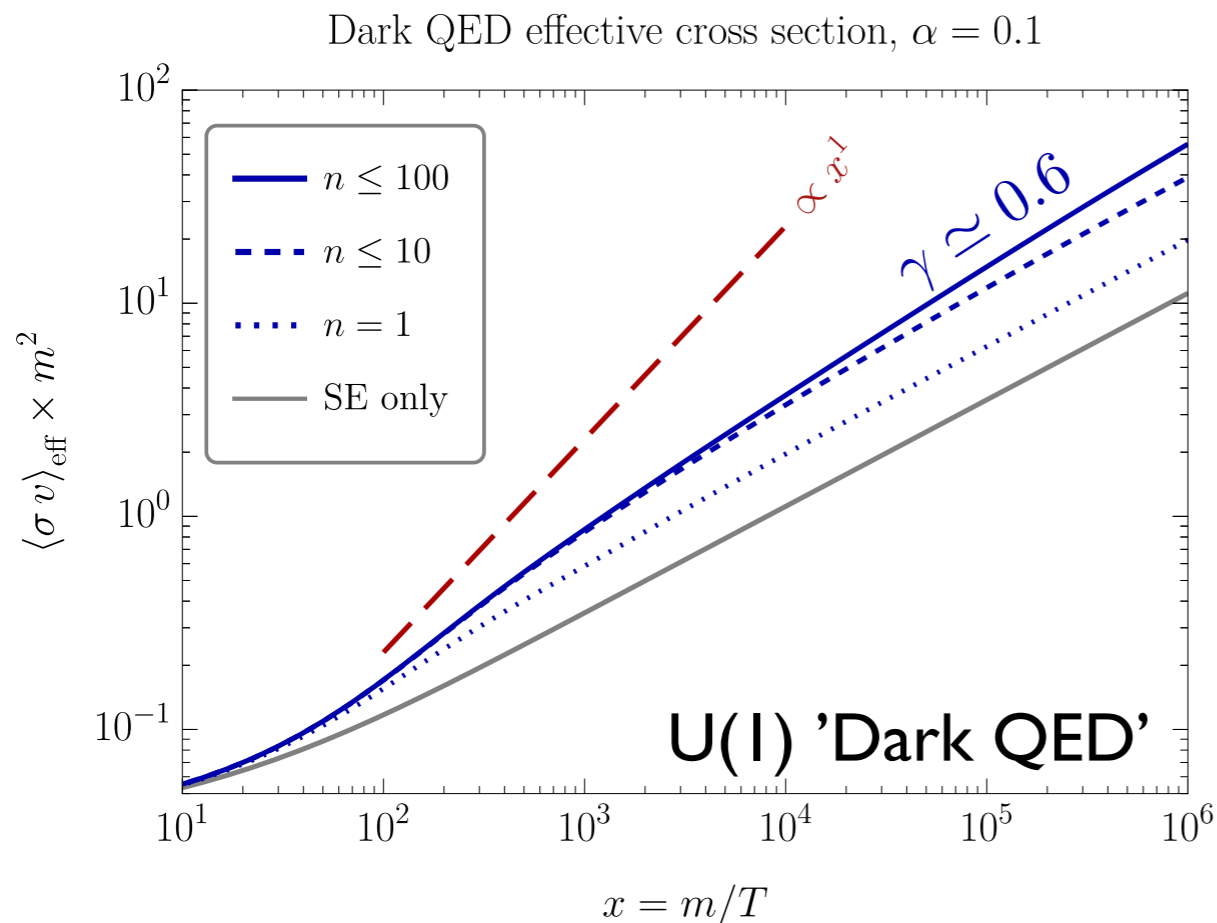
to-bound  
position

$$\textcircled{X} + \textcircled{X} : \frac{dY_X}{dx} = \frac{1}{3Hs} \frac{ds}{dx} \frac{1}{2} \langle \sigma_{XX^\dagger v} \rangle_{\text{eff}} (Y_X^2 - Y_X^{\text{eq}2})$$

where  $\langle \sigma_{XX^\dagger v} \rangle_{\text{eff}} = \langle \sigma_{XX^\dagger v} \rangle + \sum_i \langle \sigma_{\text{BSF},i v} \rangle R_i$ ,  $0 \leq R_i \leq 1$

# Effective annihilation cross section

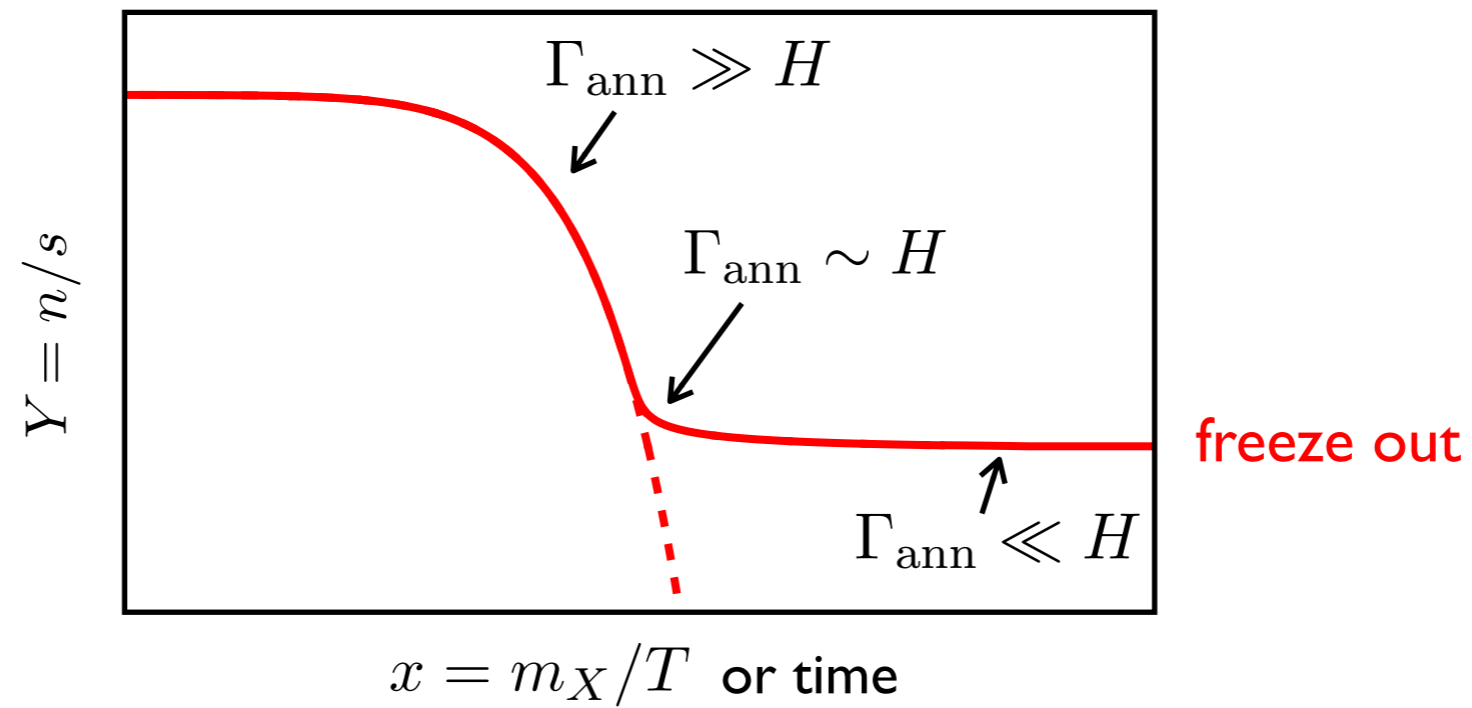
[Binder, Garny, JH, Lederer, Urban 2308.01336]



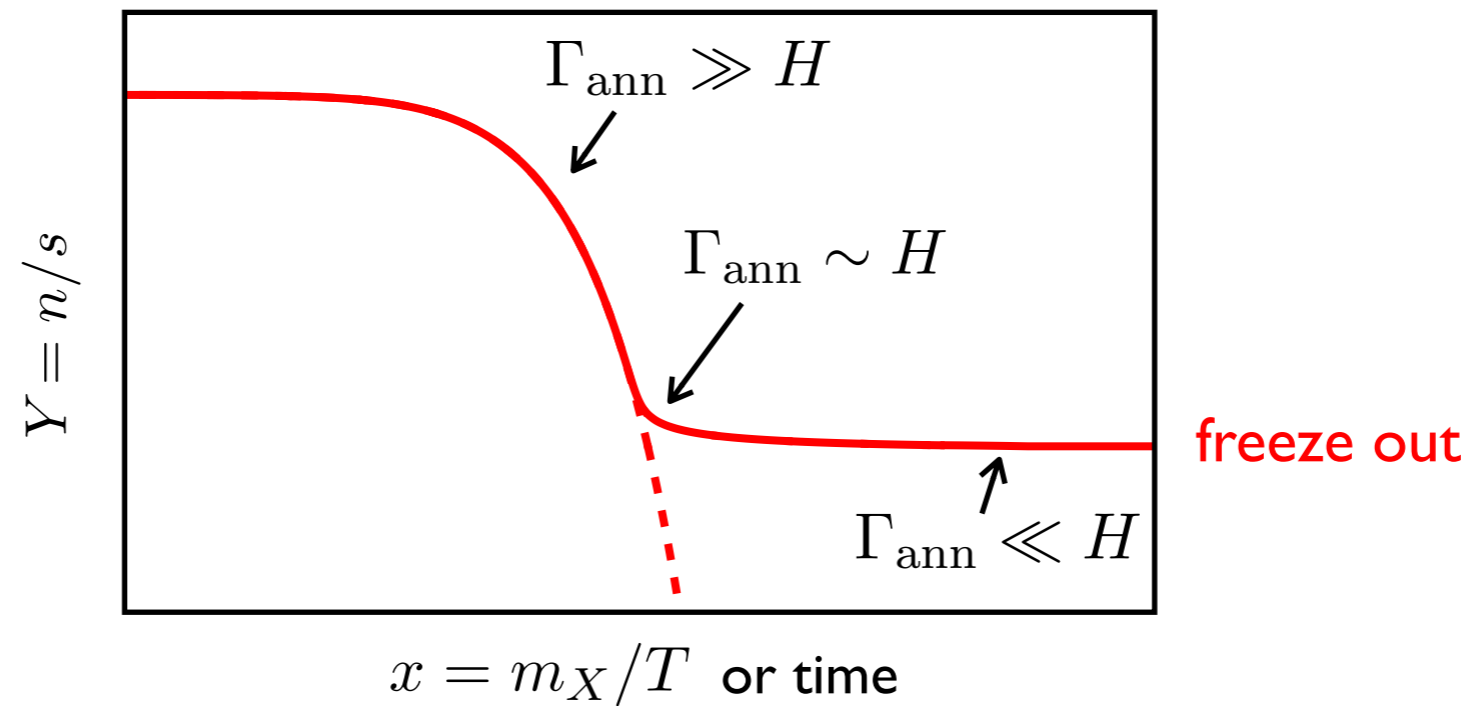
- Steepening of slope w.r.t. Sommerfeld only  $\gamma = 0.5$
- Note: no unitarity violation within unconfined phase
- No bound-to-bound transition in 'dark QCD' because no  $1 \rightarrow 1$  transitions mediated by a gluon (8)

$$\langle \sigma v \rangle_{\text{eff}} \propto x^\gamma$$

# Solutions to Boltzmann equations



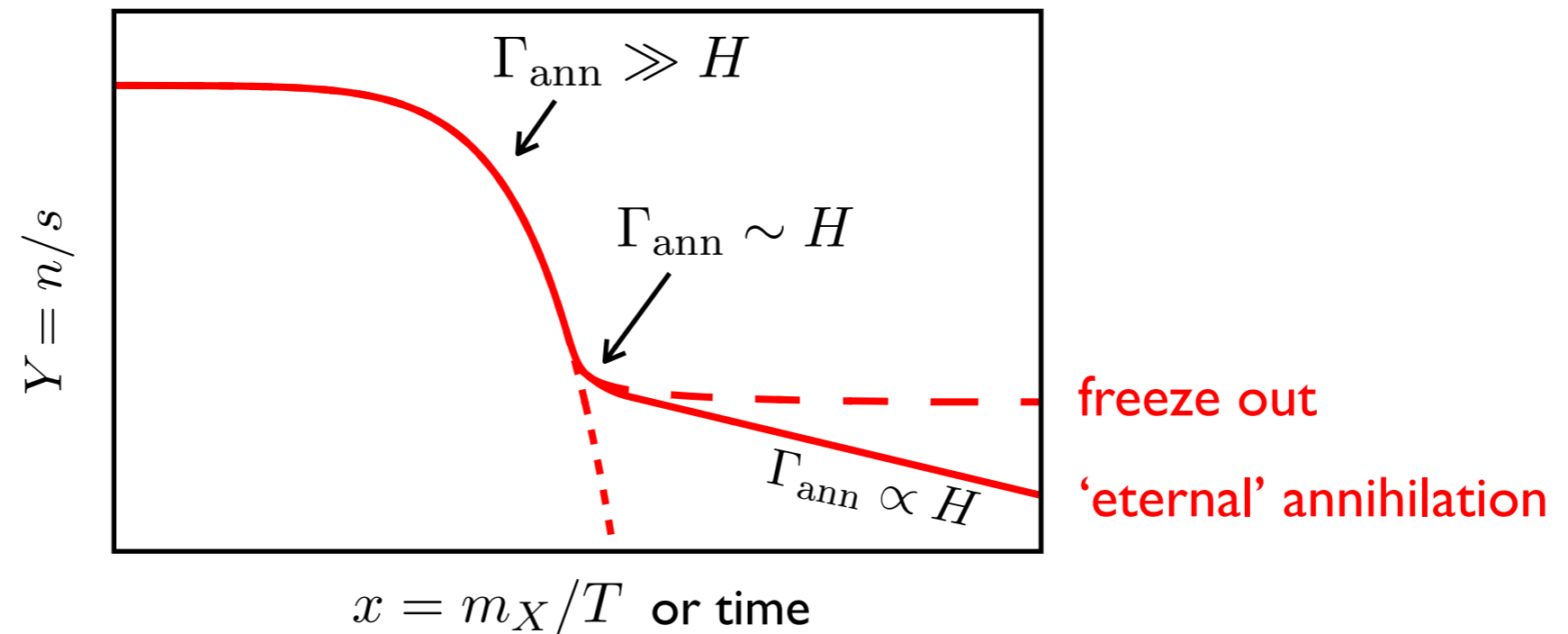
# Solutions to Boltzmann equations



$$\Gamma_{\text{ann}} = n_X \langle \sigma v \rangle_{\text{eff}} \quad \text{versus} \quad H \propto T^2$$

$$n \propto T^3 \quad \langle \sigma v \rangle_{\text{eff}} \propto T^{-\gamma} \quad \text{if} \quad \gamma < 1$$

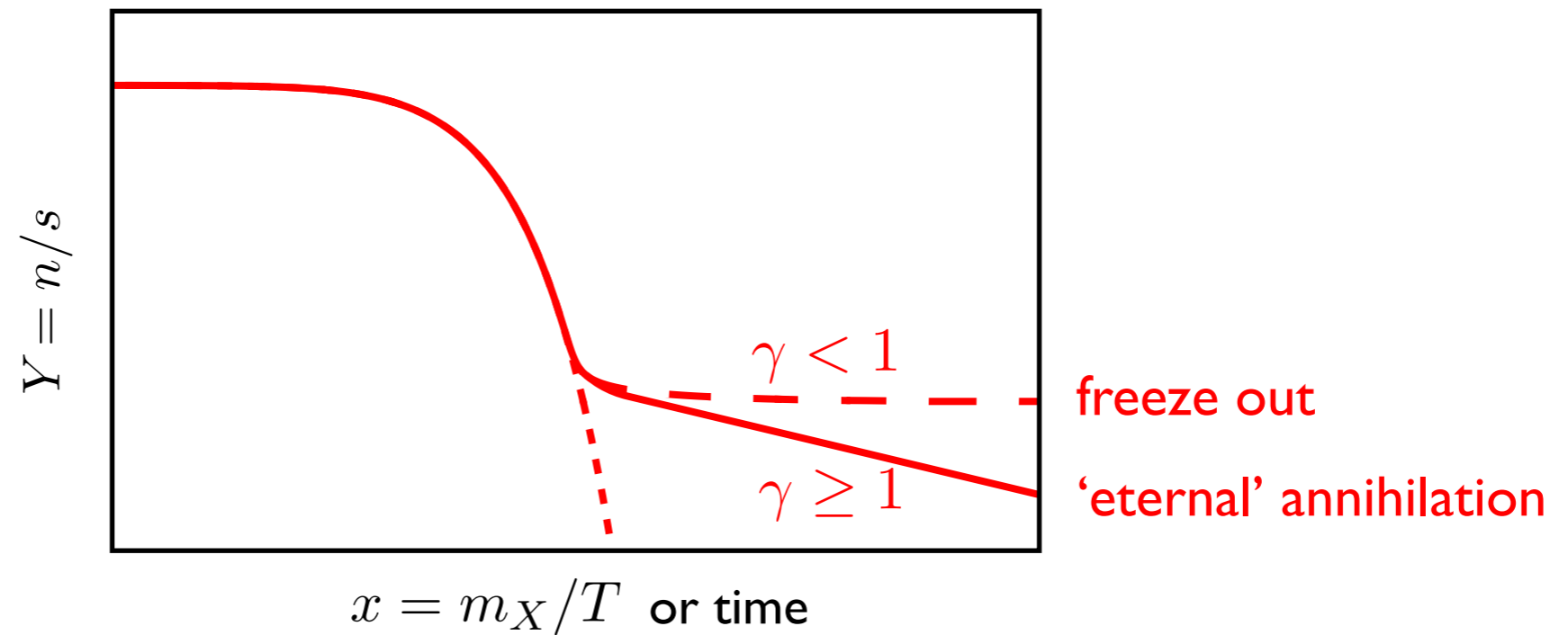
# Solutions to Boltzmann equations



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$$n \propto T^3 \quad \langle \sigma v \rangle_{\text{eff}} \propto T^{-\gamma} \quad \text{if } \gamma \geq 1$$

# Solutions to Boltzmann equations



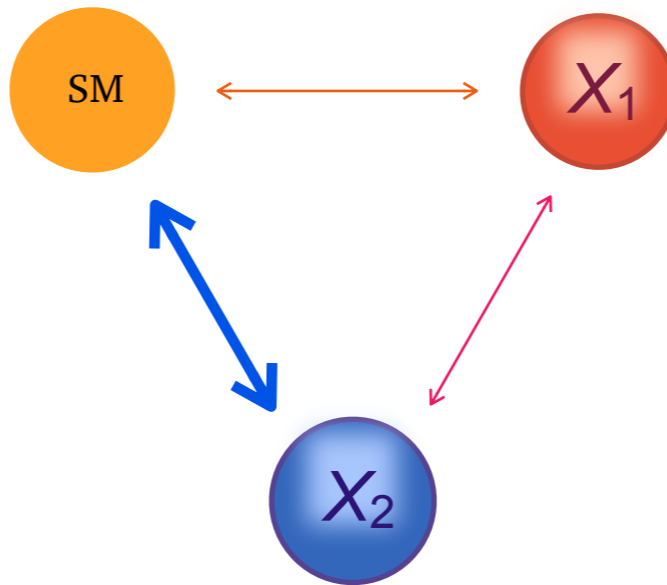
$$\Gamma_{\text{ann}} = n_X \langle \sigma v \rangle_{\text{eff}} \quad \text{versus} \quad H \propto T^2$$

$\swarrow$                        $\searrow$   
 $n \propto T^3$                        $\langle \sigma v \rangle_{\text{eff}} \propto T^{-\gamma}$

# Implications for *t*-channel mediator models

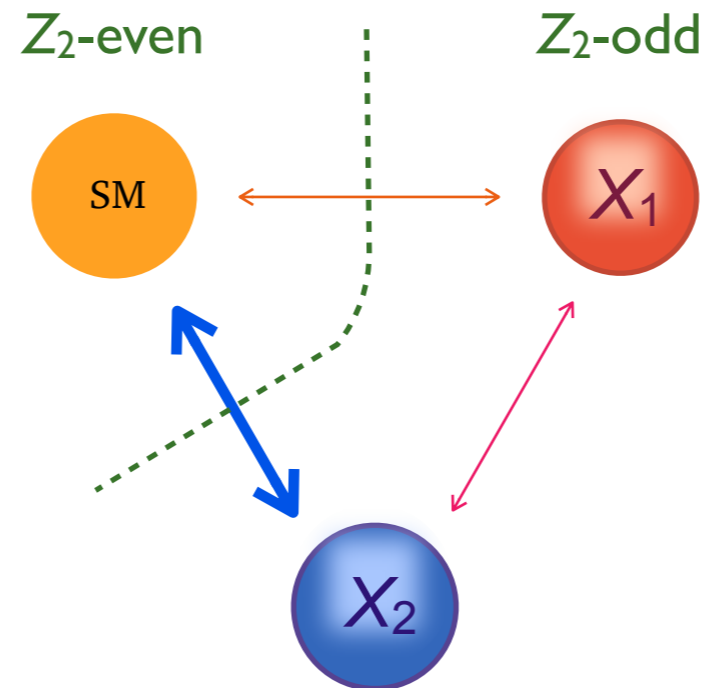


# $t$ -channel mediator dark matter model (aka 'charged parent particle model')



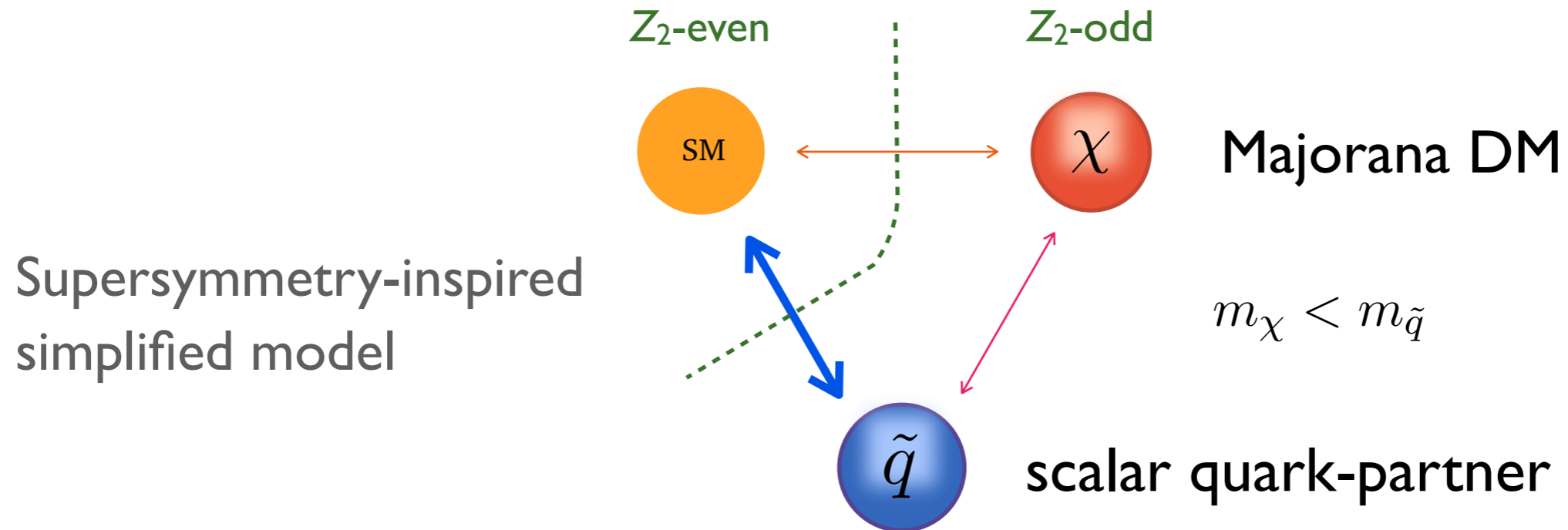
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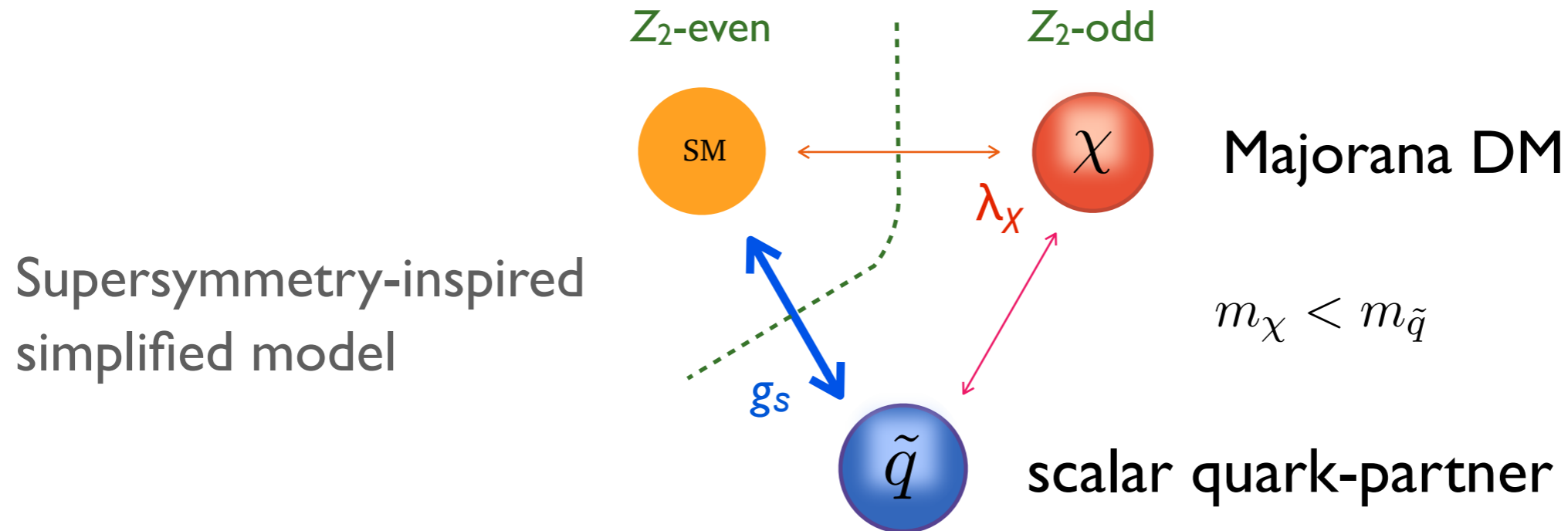
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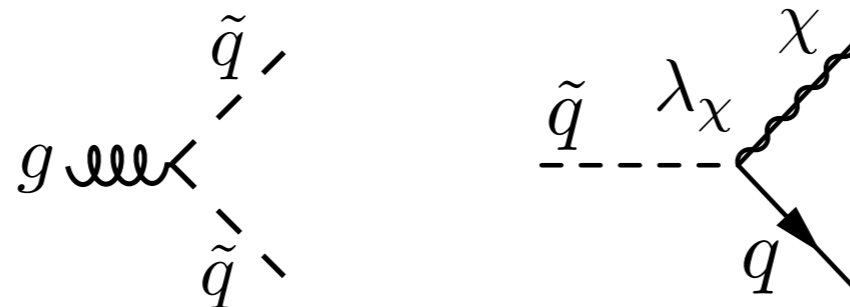


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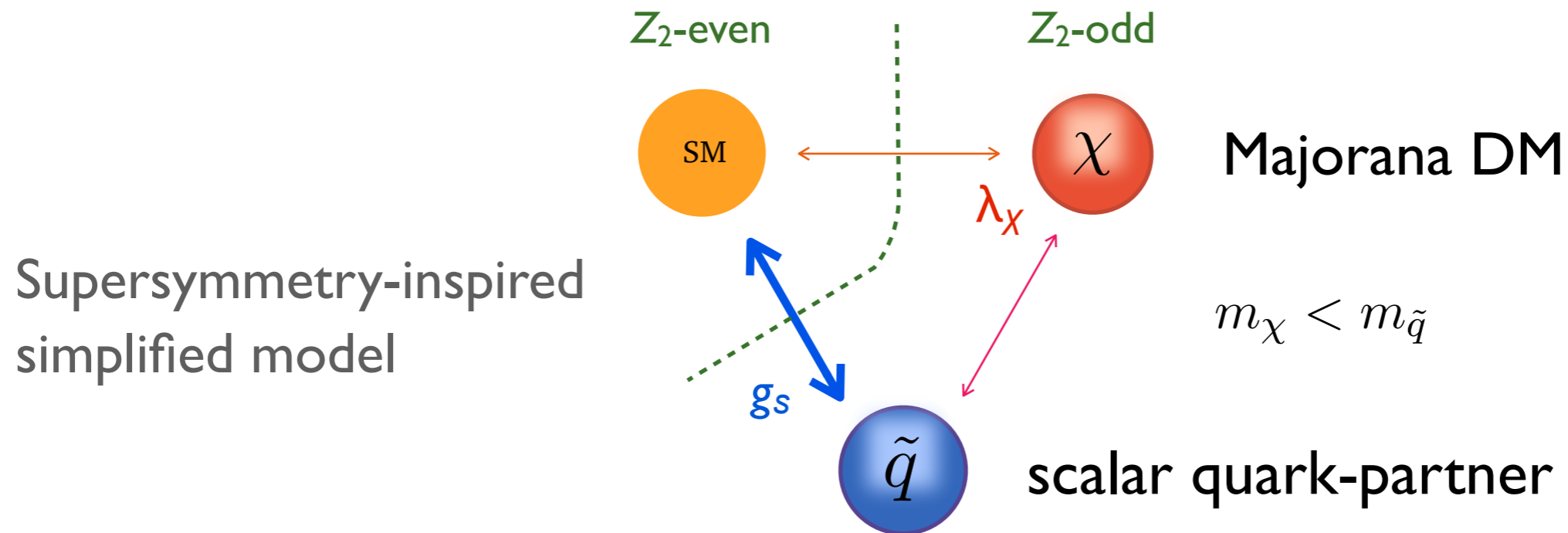


$$\mathcal{L}_{\text{int}} = |D_\mu \tilde{q}|^2 - \lambda_\chi \tilde{q} \tilde{q} \frac{1 - \gamma_5}{2} \chi + \text{h.c.}$$

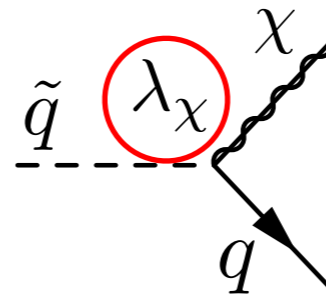
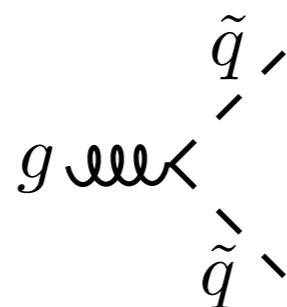


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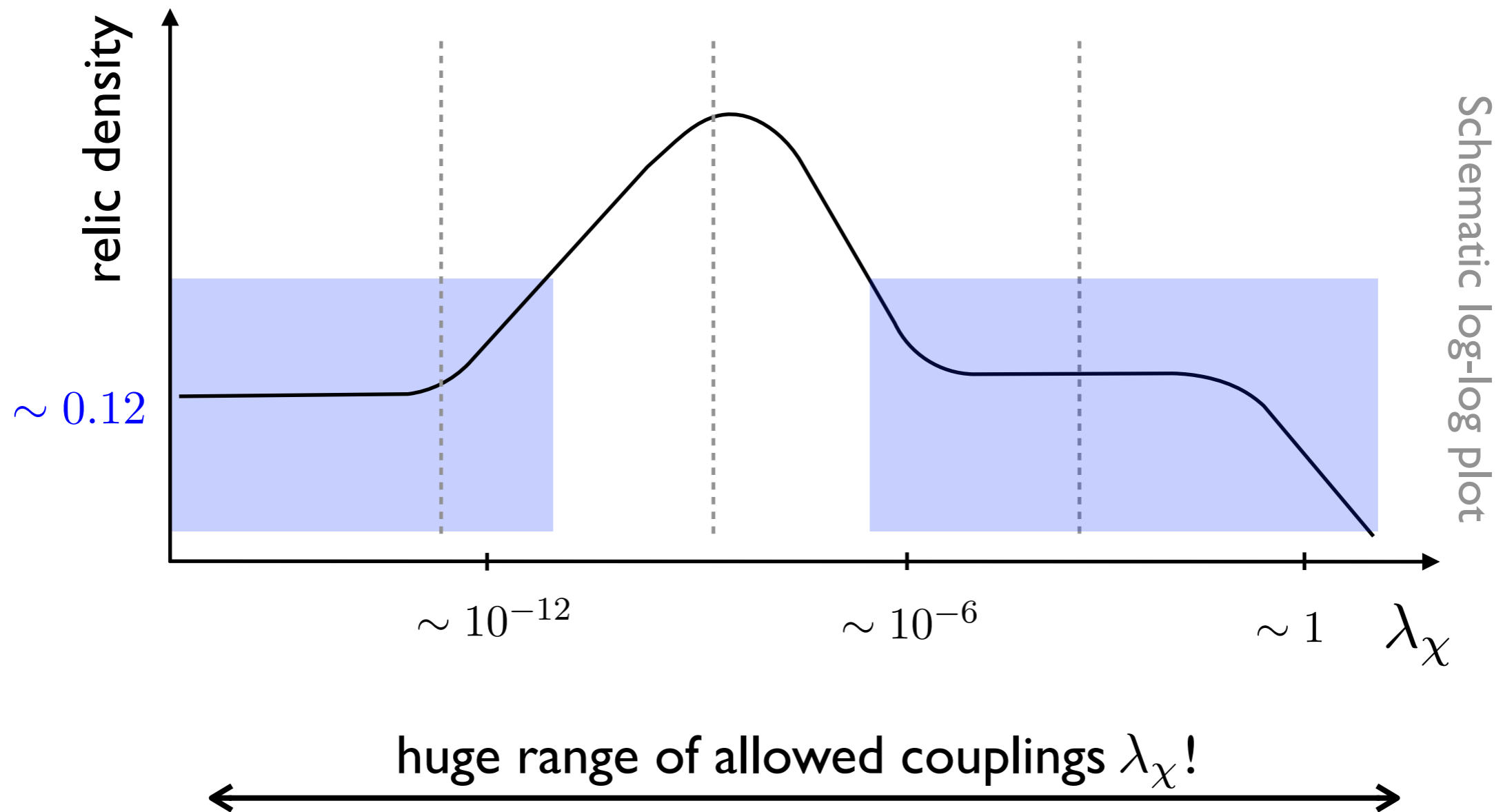


$$\mathcal{L}_{\text{int}} = |D_\mu \tilde{q}|^2 - \lambda_\chi \tilde{q} \tilde{q} \frac{1 - \gamma_5}{2} \chi + \text{h.c.}$$

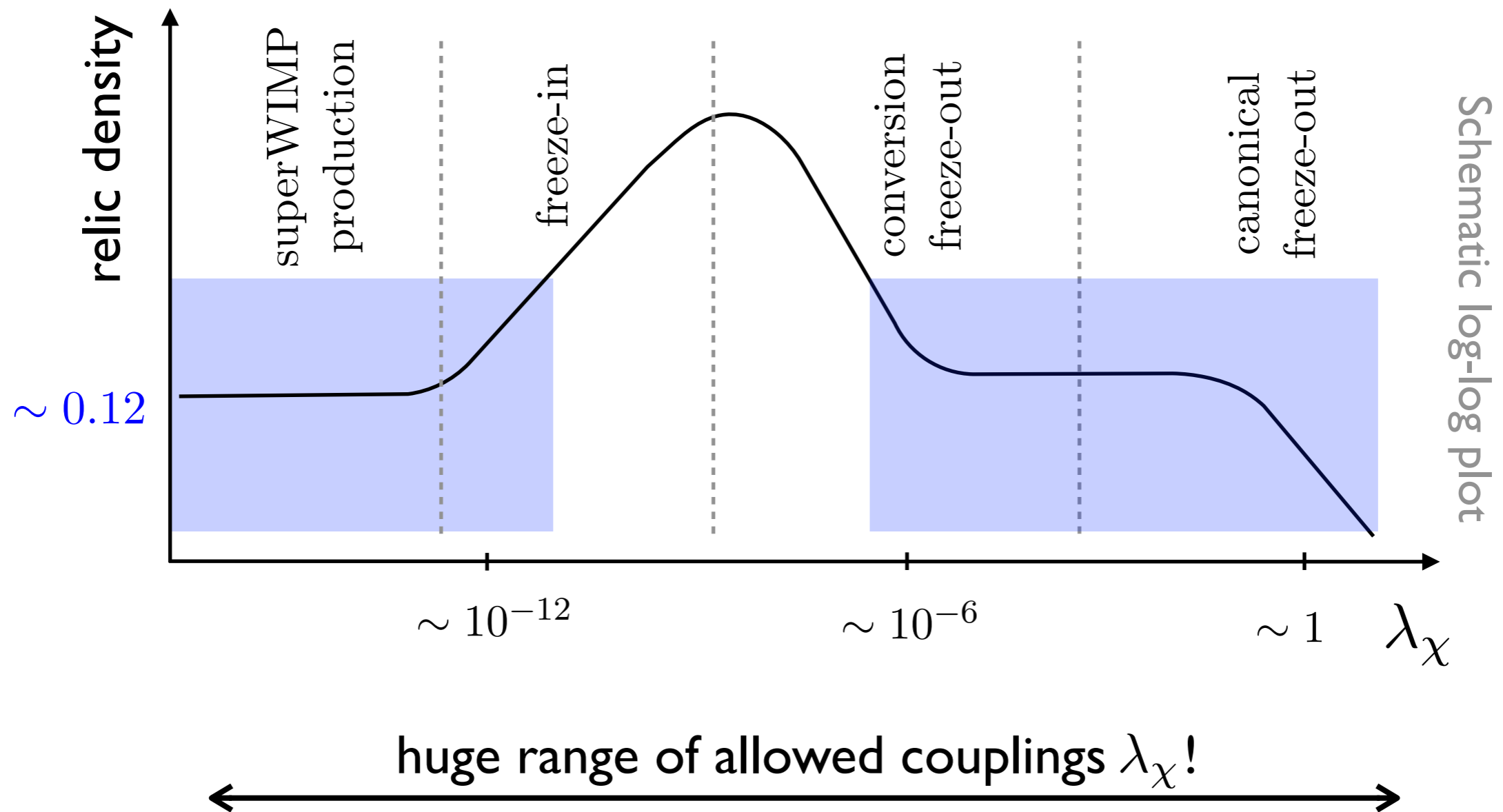


Unlike MSSM,  $\lambda_\chi$  free parameter

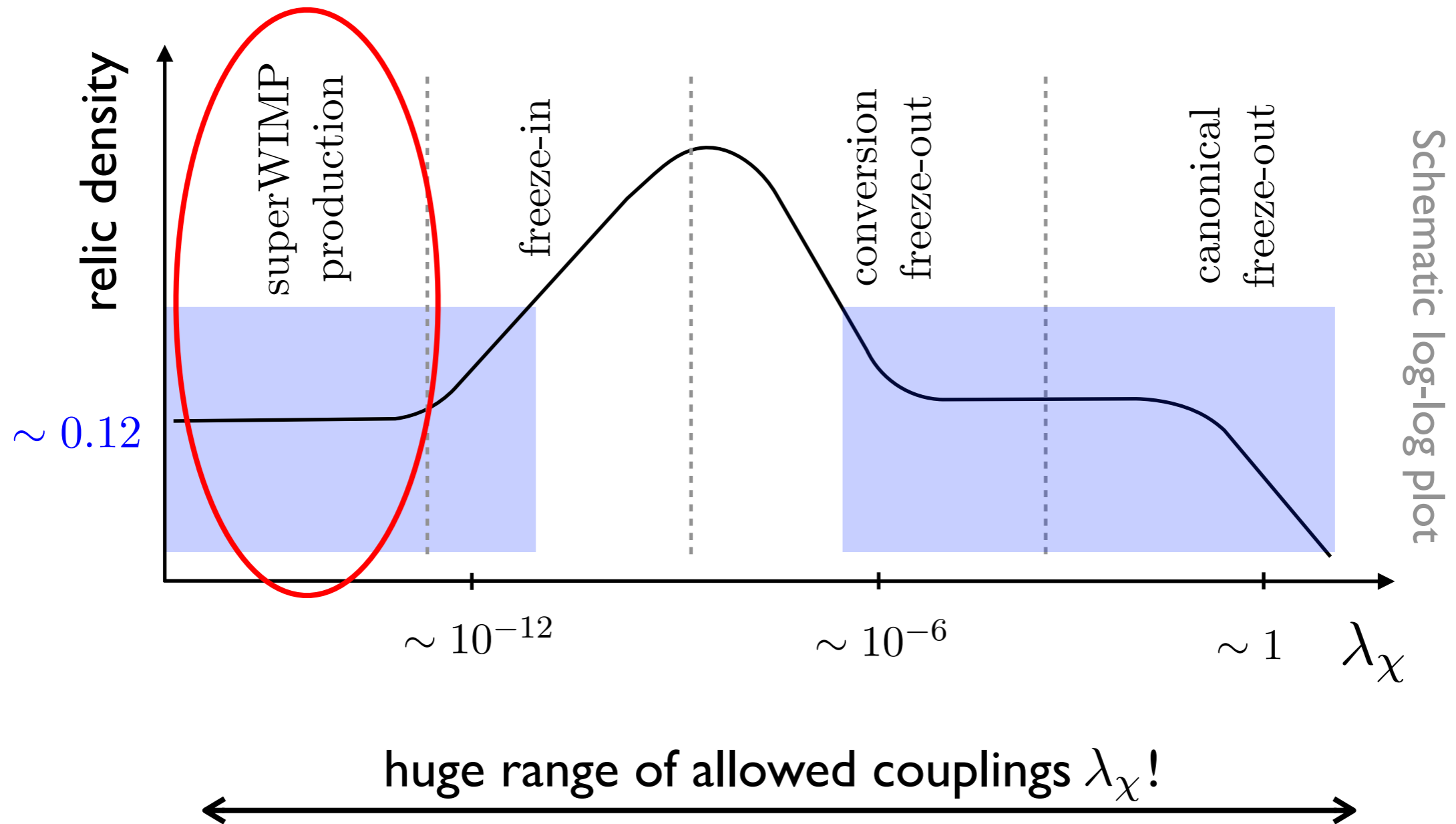
# *t*-channel models: Dark matter production scenarios



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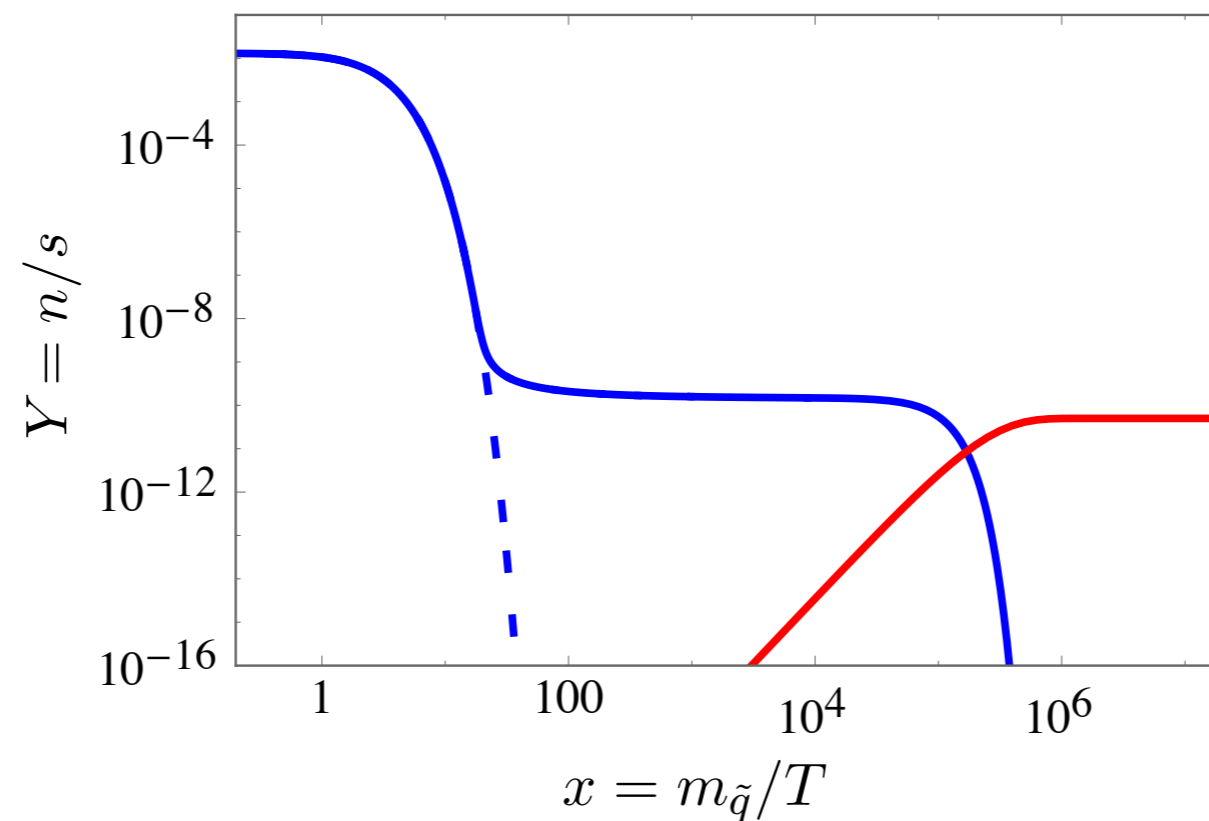




# superWIMP scenario

[Covi et al. 1999; Feng et al. 2003]

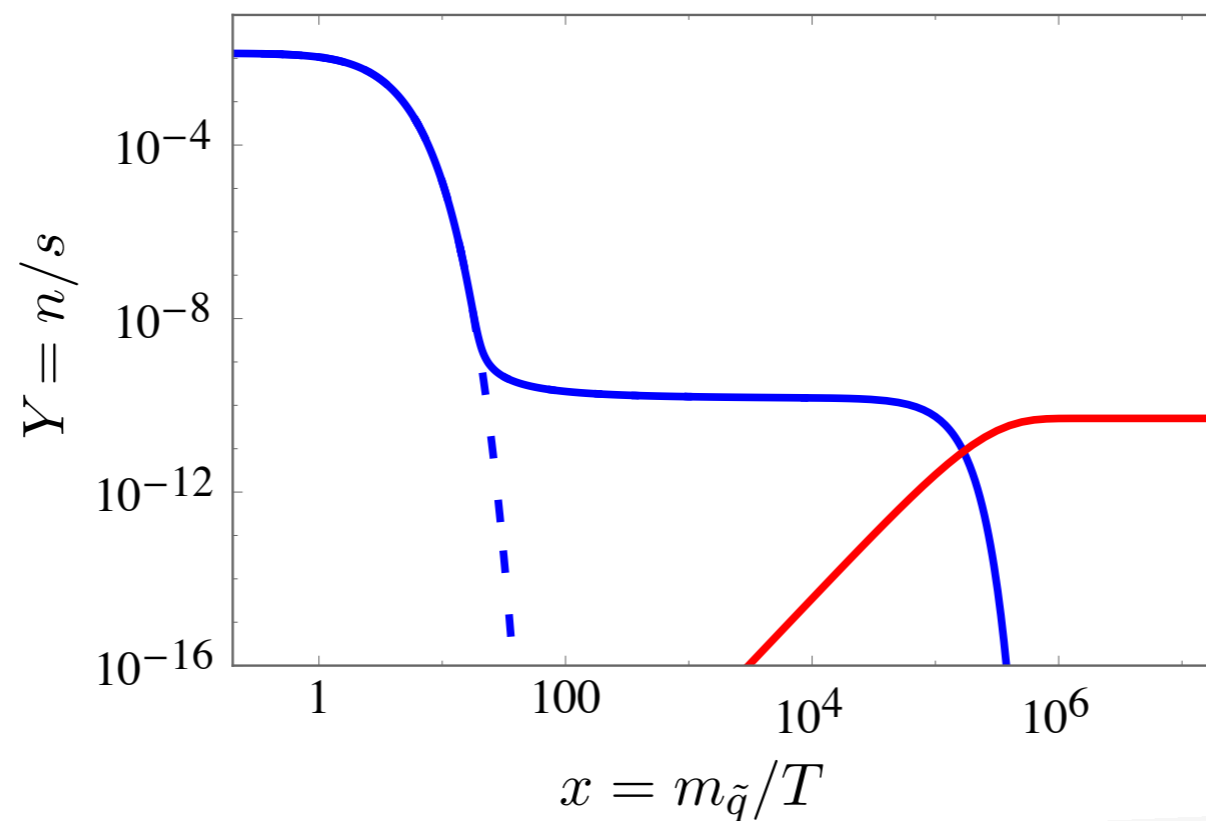
- Coupling  $\lambda_\chi \sim 10^{-16} - 10^{-10}$ , dark matter thermally decoupled
- Produced in late decays of frozen out  $\tilde{q}$   
⇒ Relic density independent of  $\lambda_\chi$



# superWIMP scenario

[Covi et al. 1999; Feng et al. 2003]

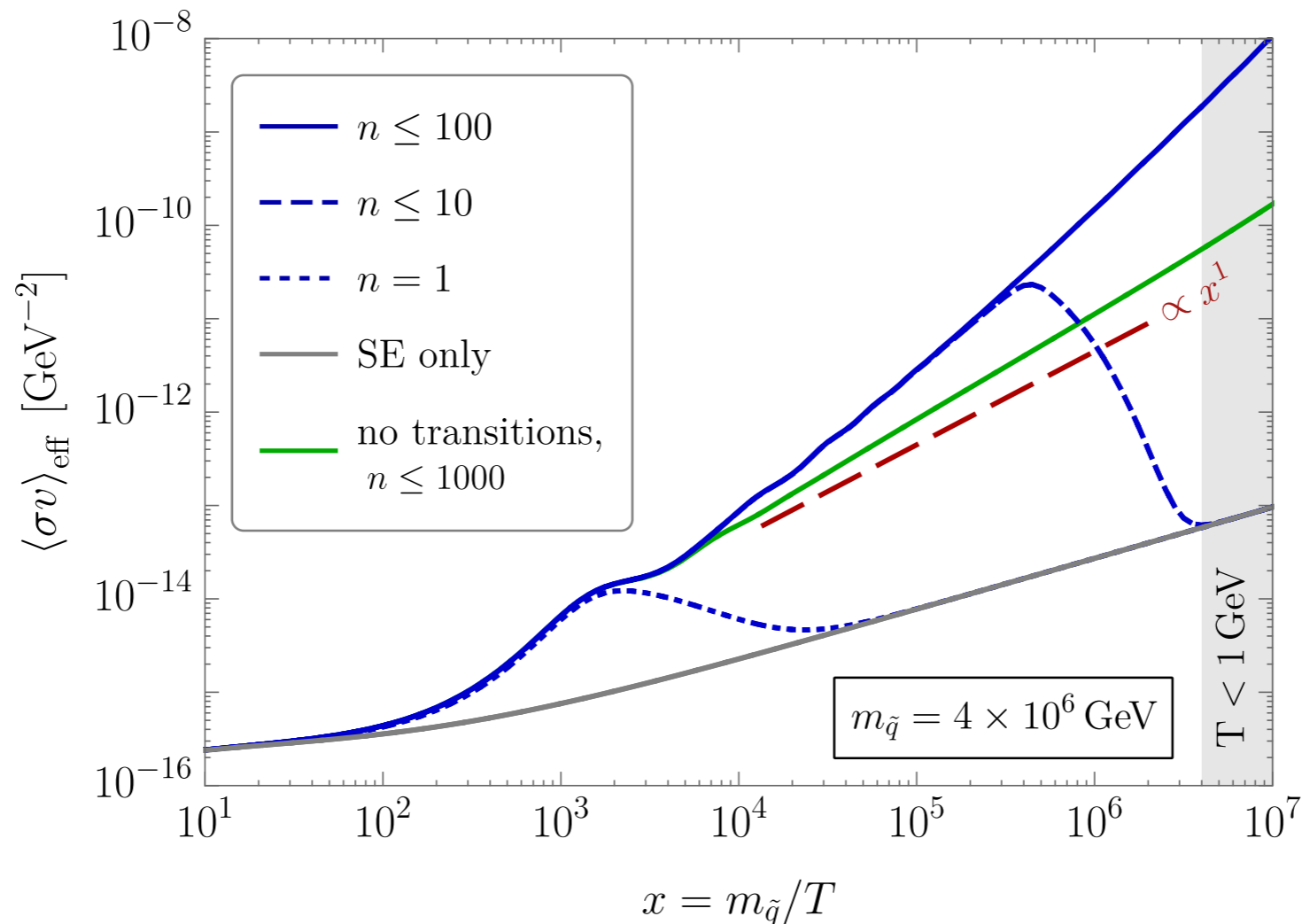
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Does the picture change with bound states?

# Effective annihilation cross section

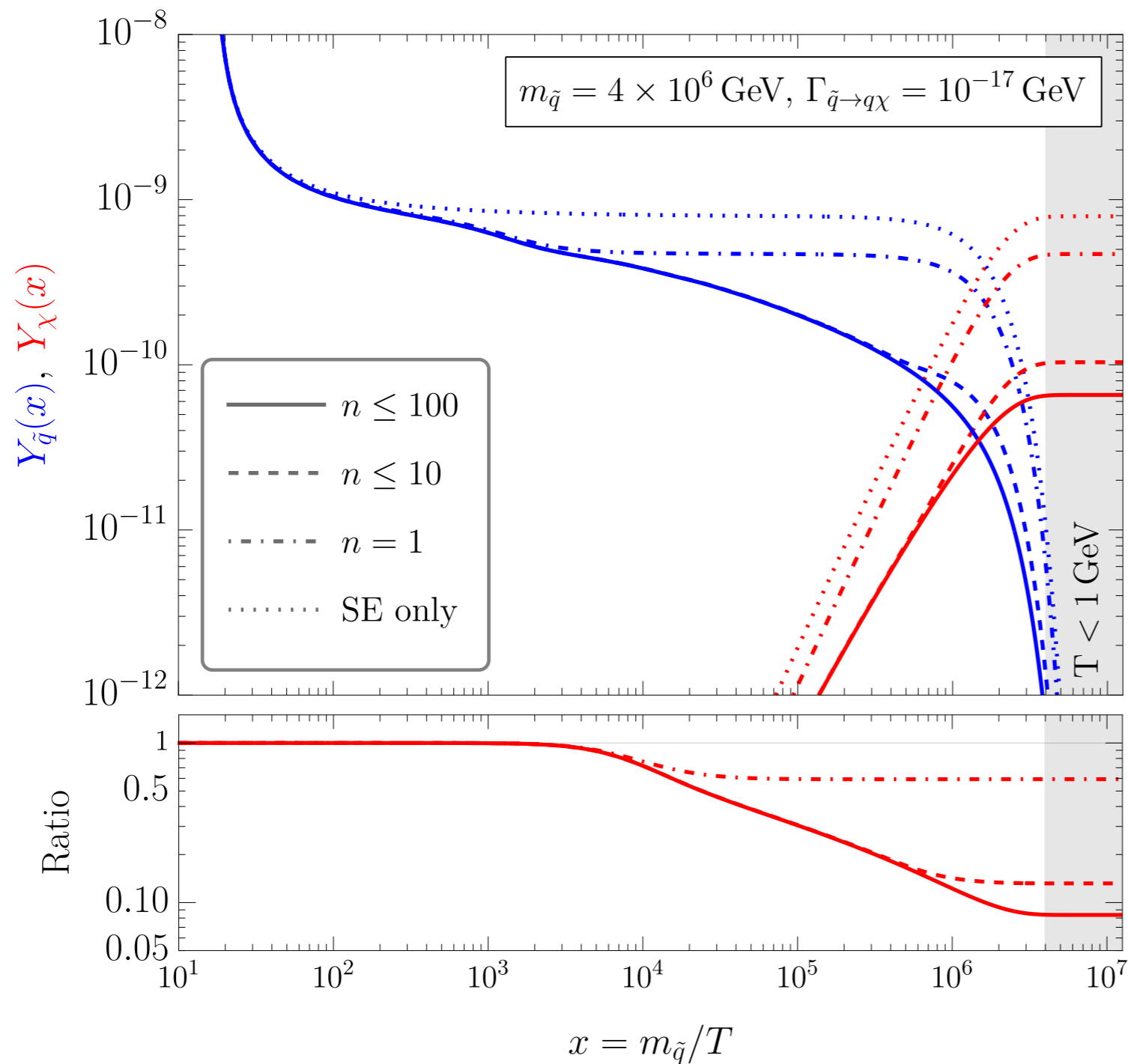
[Binder, Garny, JH, Lederer, Urban 2308.01336]



- $\tilde{q}$  also electrically charged: dipole transitions  $\mathcal{B}_{nl} \leftrightarrow \mathcal{B}_{n'l'}$  via photons  
 $\Rightarrow$  High impact on slope, super-critical (even w/o running)
- No unitarity-violating XS involved in unconfined phase, i.e.  $T > \Lambda_{\text{QCD}}$

# Impact on the relic abundance

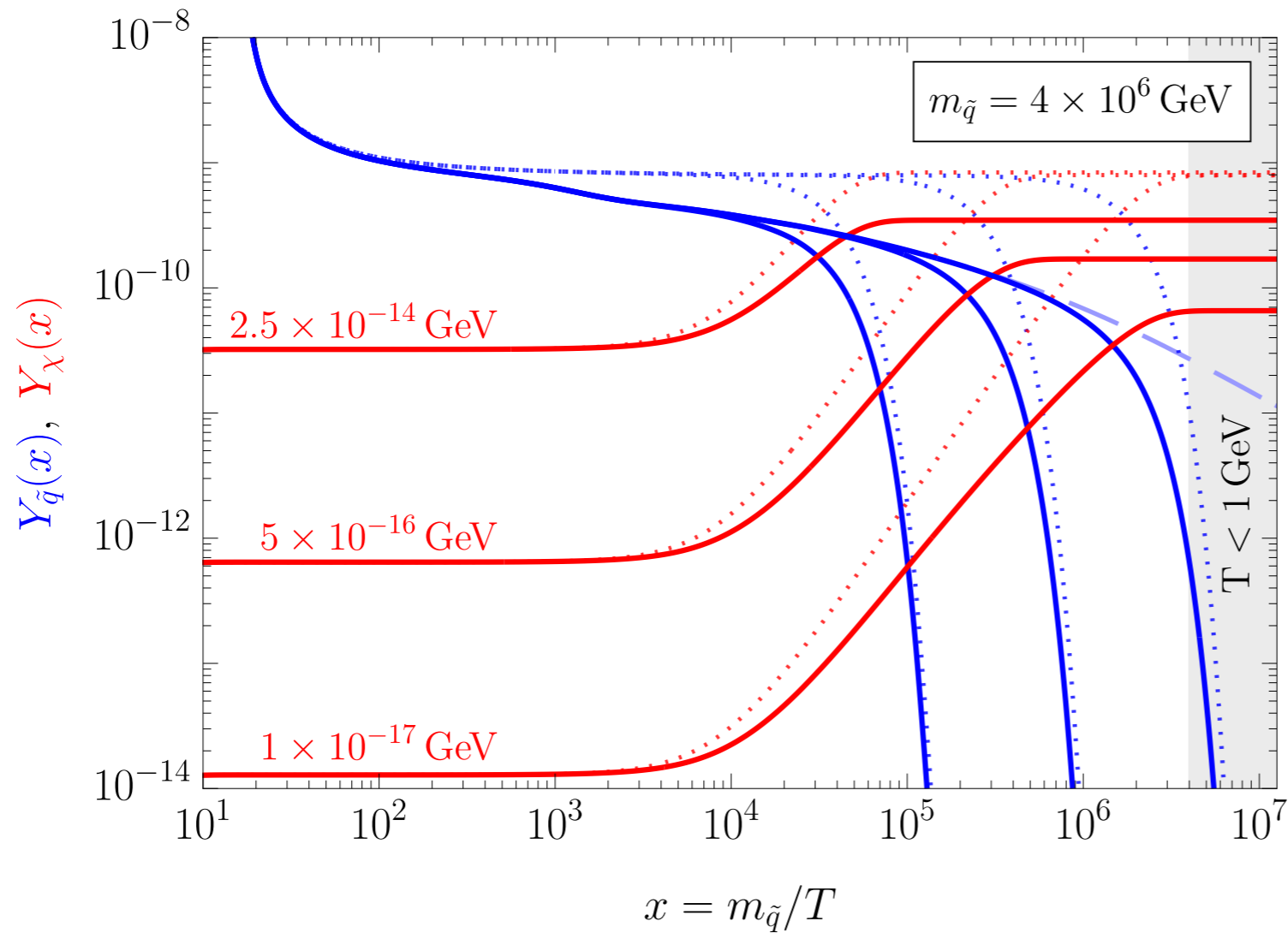
[Binder, Garny, JH, Lederer, Urban 2308.01336]



- Excited bound states highly relevant
- No freeze-out before decay ( $T_{\text{decay}} > 1 \text{ GeV}$ )
- Order-of-magnitude effect on relic density

# Impact on the relic abundance

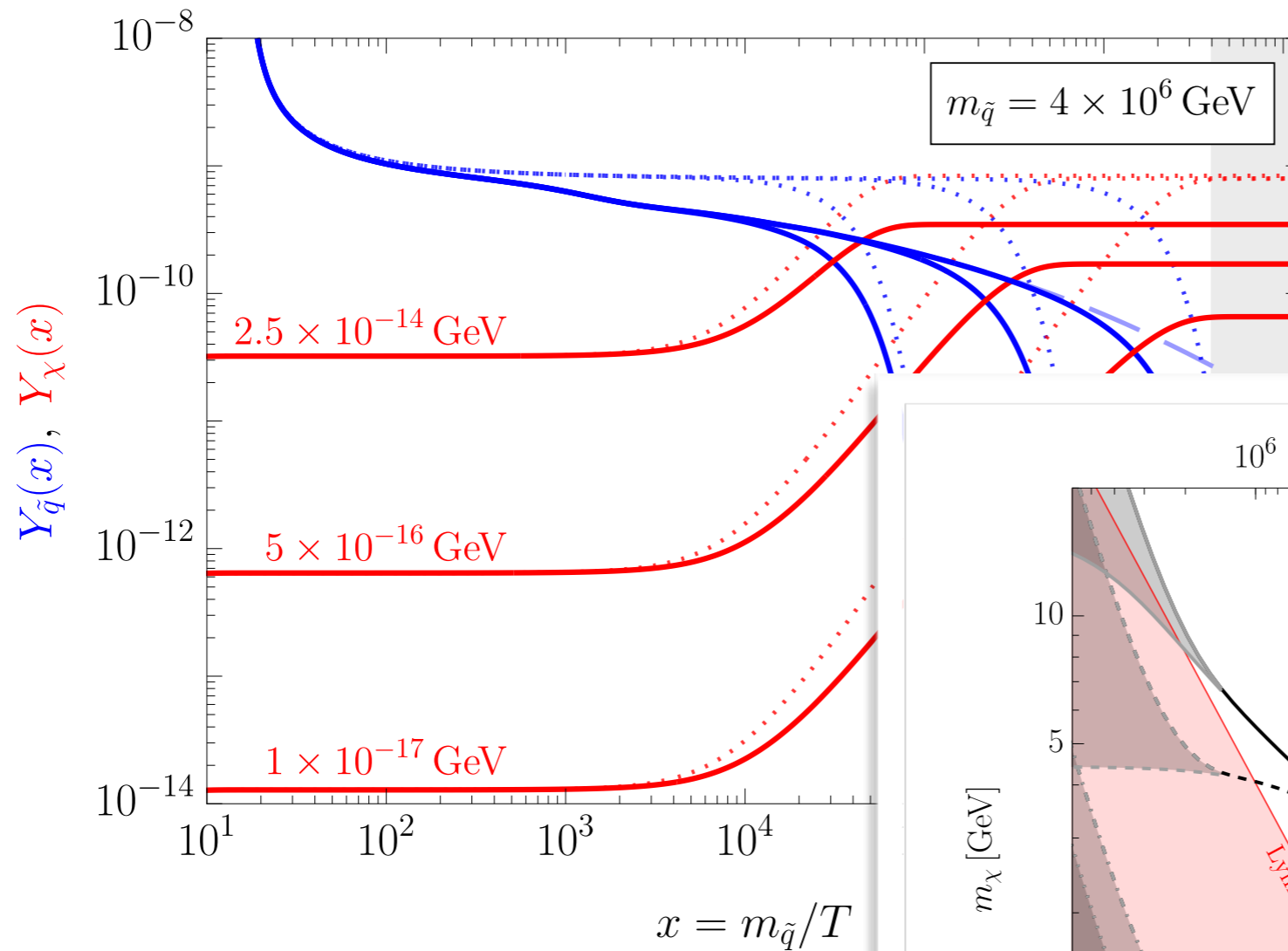
[Binder, Garny, JH, Lederer, Urban 2308.01336]



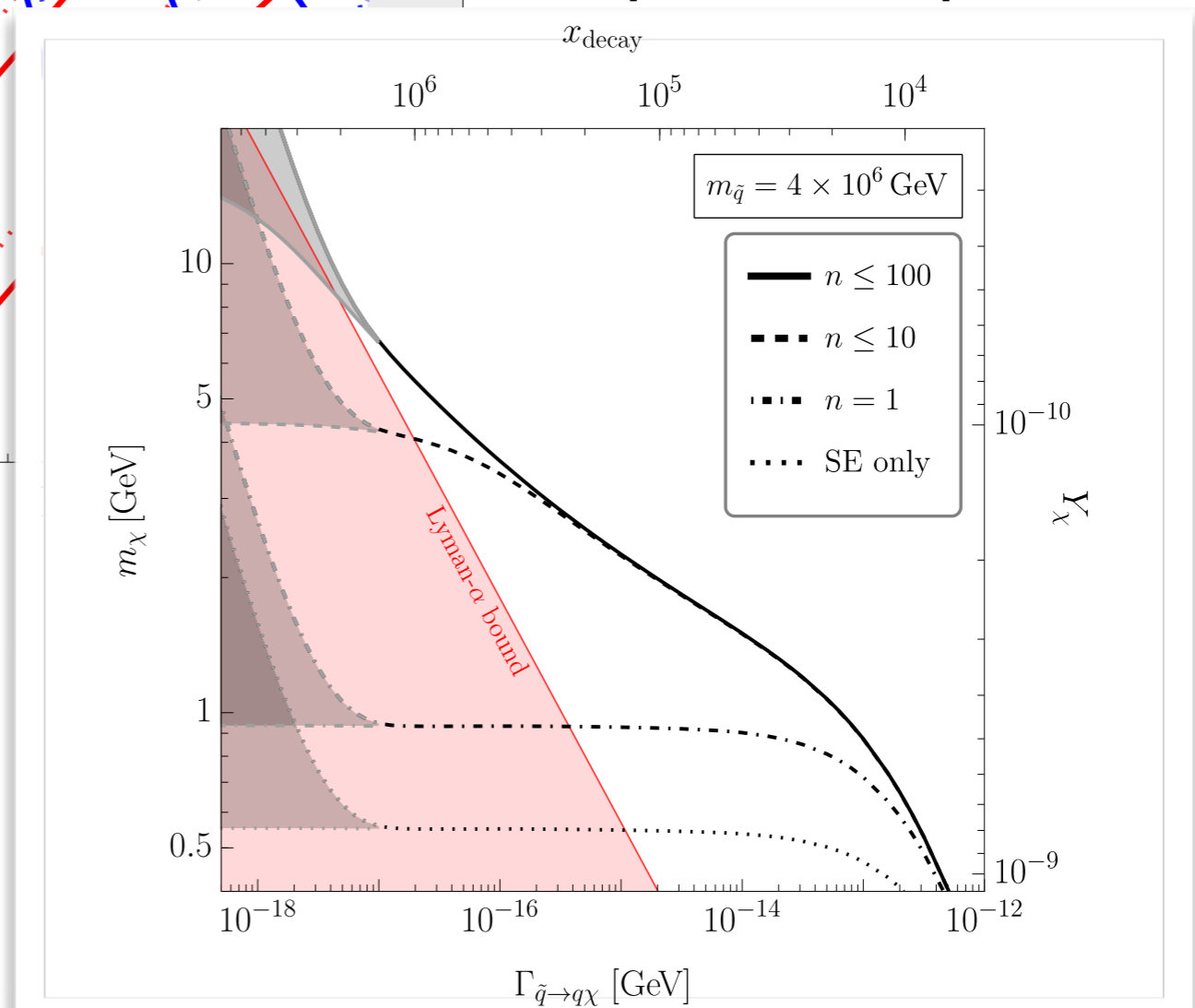
} Bound states introduce dependence on  $\lambda_{\chi}$  in superWIMP production!

# Impact on the relic abundance

[Binder, Garny, JH, Lederer, Urban 2308.01336]

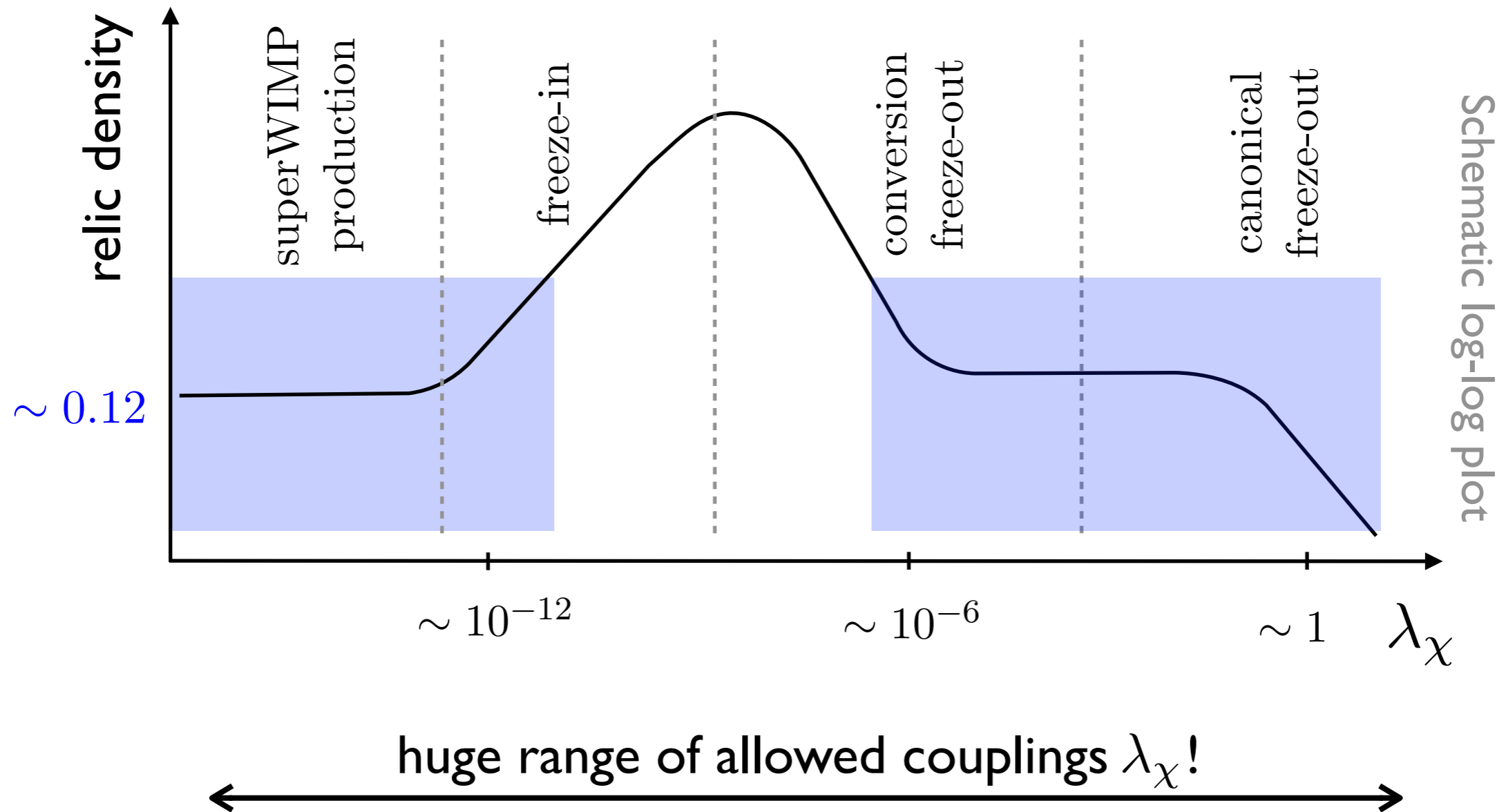


Bound states introduce dependence on  $\lambda_{\chi}$  in superWIMP production!

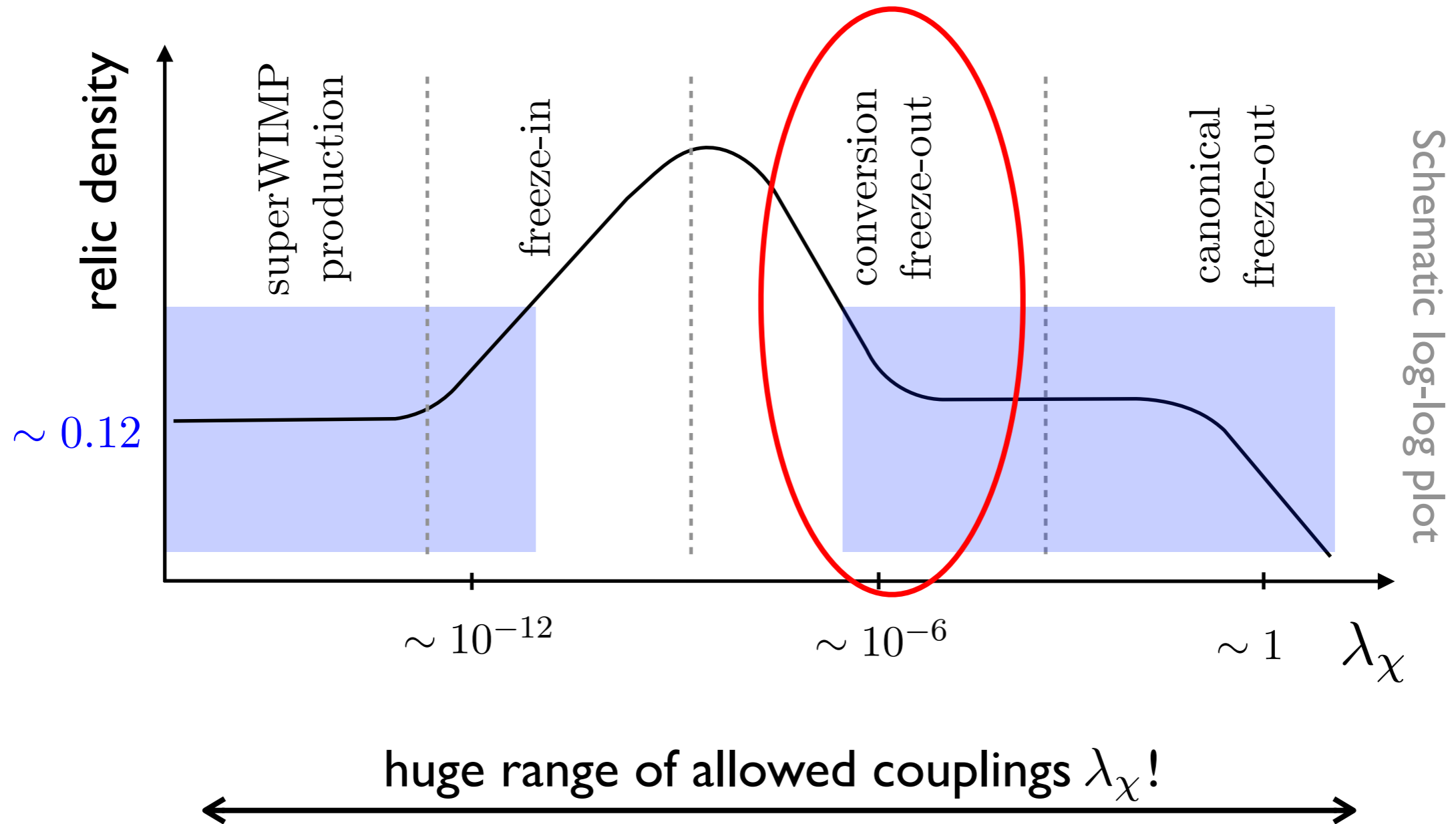


Relevant for constraints from cosmological structure formation (Lyman-alpha forest observations)

# *t*-channel models: Dark matter production scenarios



# *t*-channel models: Dark matter production scenarios



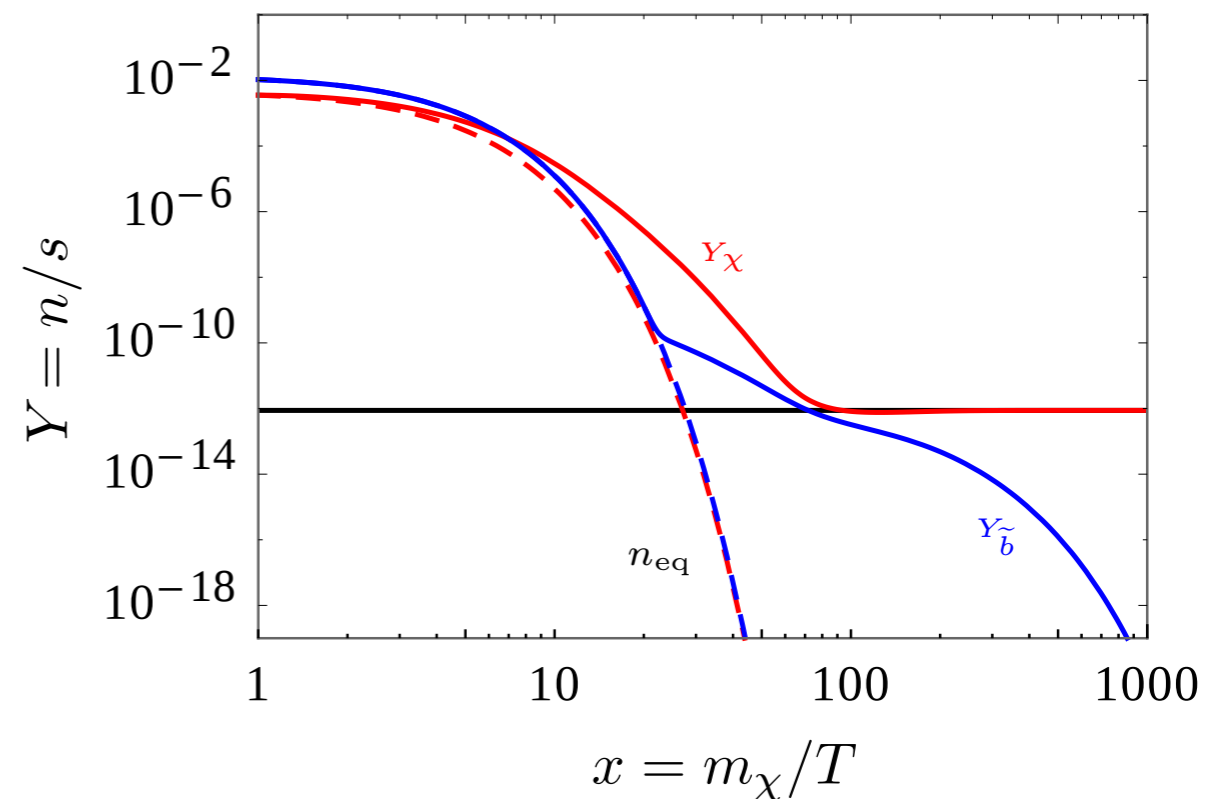
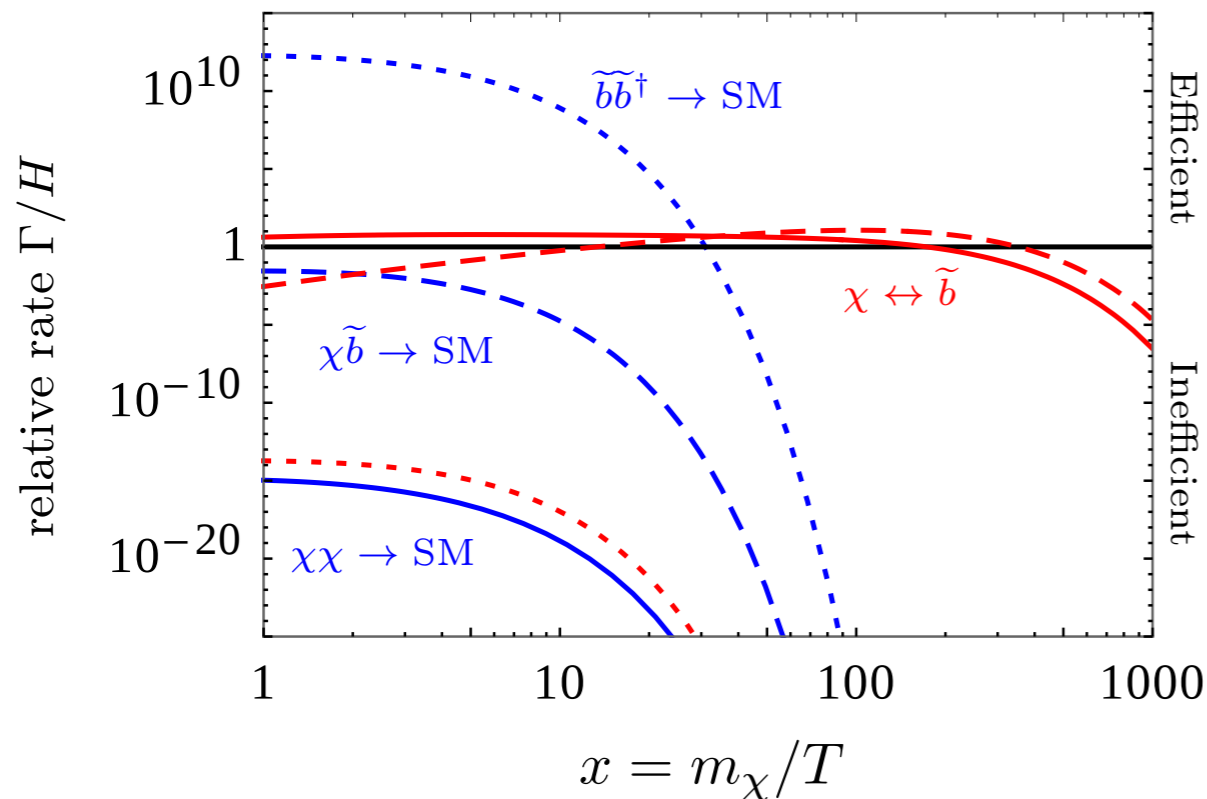
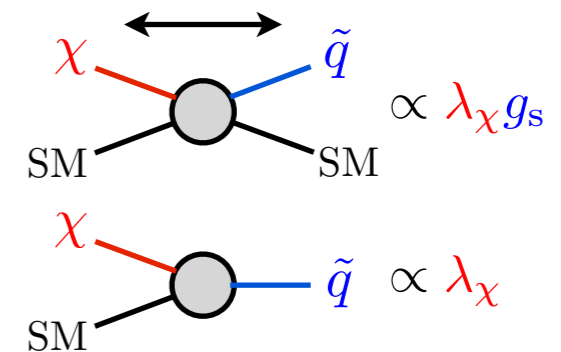


# Conversion-driven freeze-out

[Garny, JH, Lülfi, Vogl 2017; D'Agnolo, Pappadopulo, Ruderman 2017]

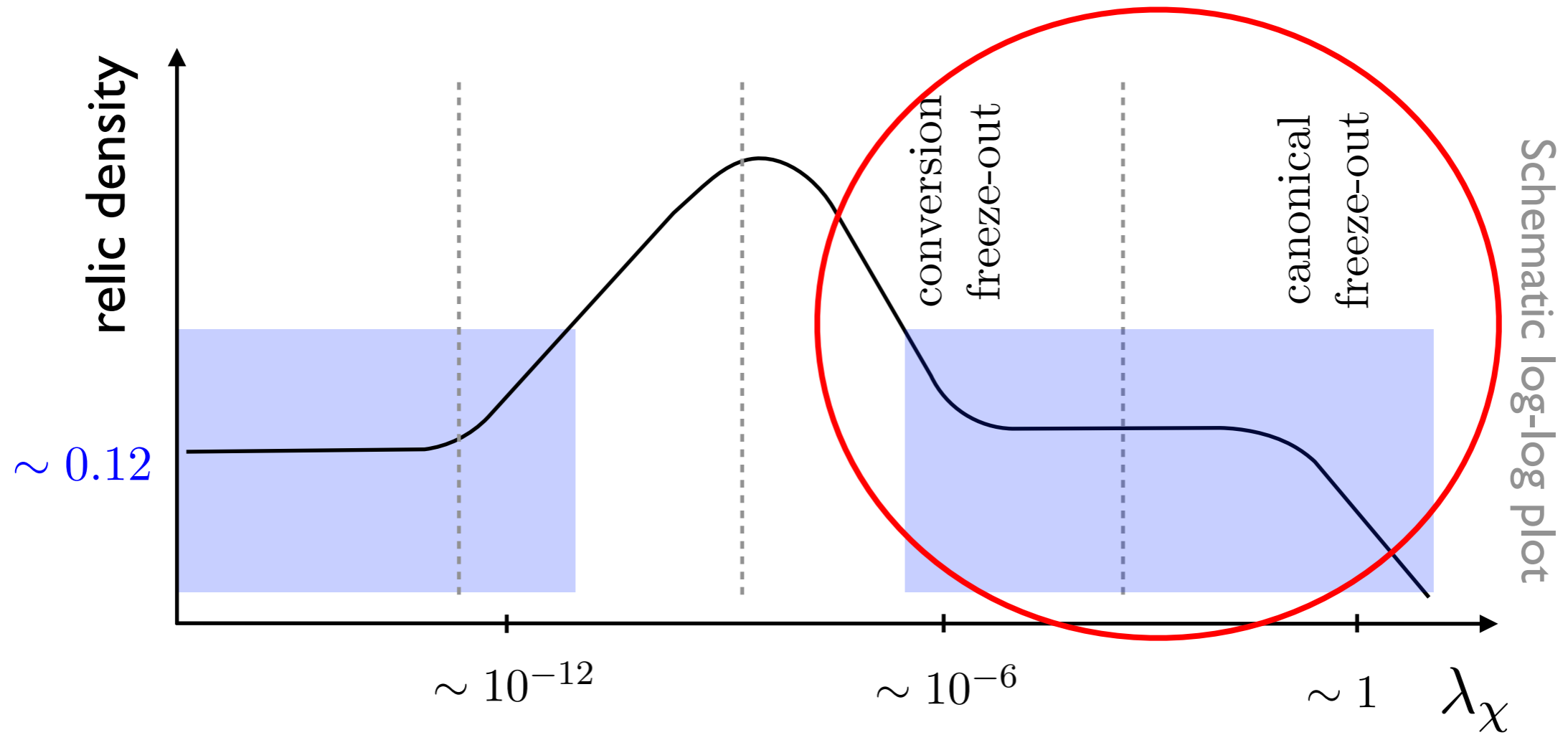
- Coupling  $\lambda_\chi \sim 10^{-6}$  just large enough to thermalize dark matter
- Conversions on the edge of being efficient  
 $\Gamma \sim H$ , initiate chemical decoupling  
 $\Rightarrow$  prolonged freeze-out process

Conversion processes:



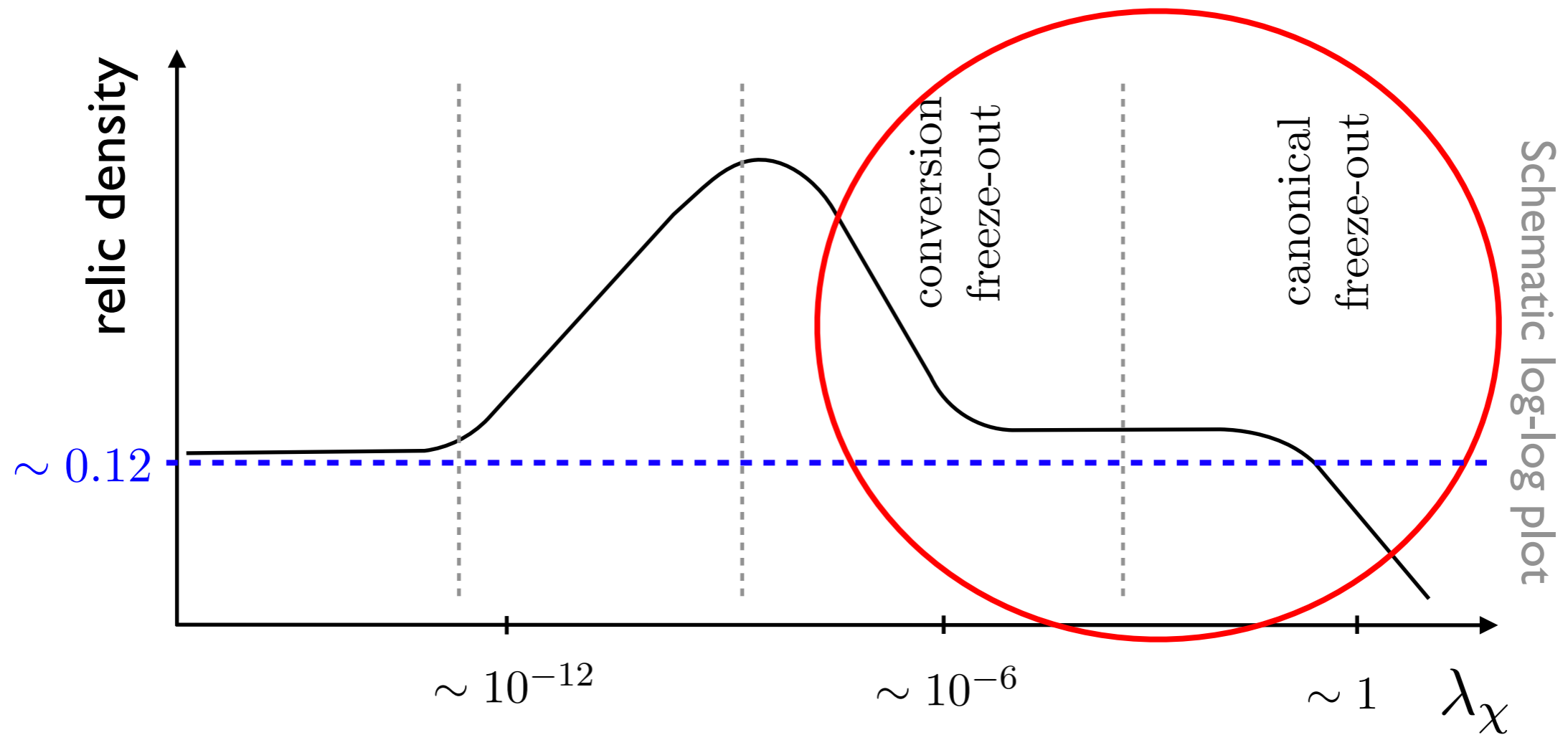
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[Garny, JH, Lülf, Vogl 2017; D'Agnolo, Pappadopulo, Ruderman 2017]



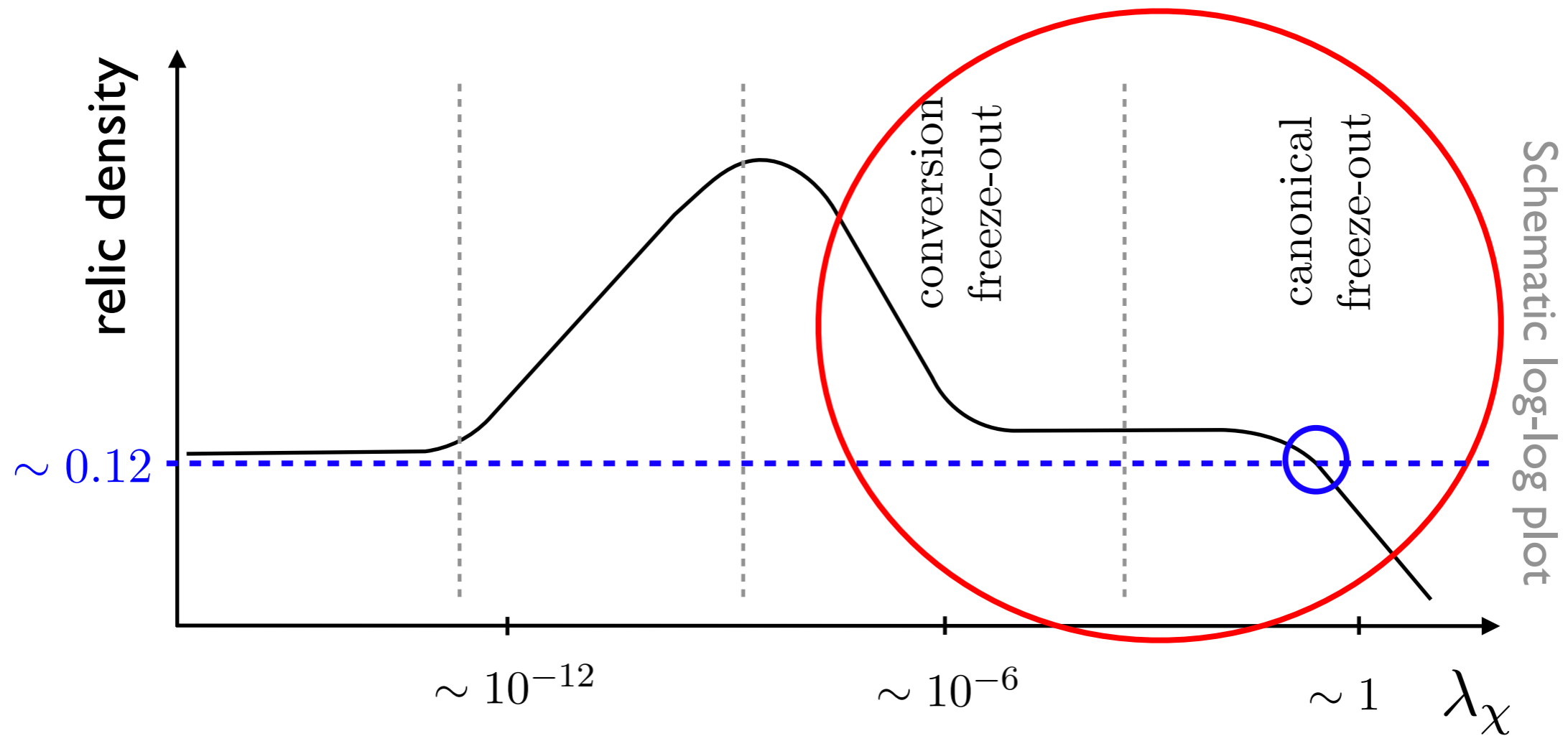
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[Garny, JH, Lulf, Vogl 2017; D'Agnolo, Pappadopulo, Ruderman 2017]



# Conversion-driven freeze-out

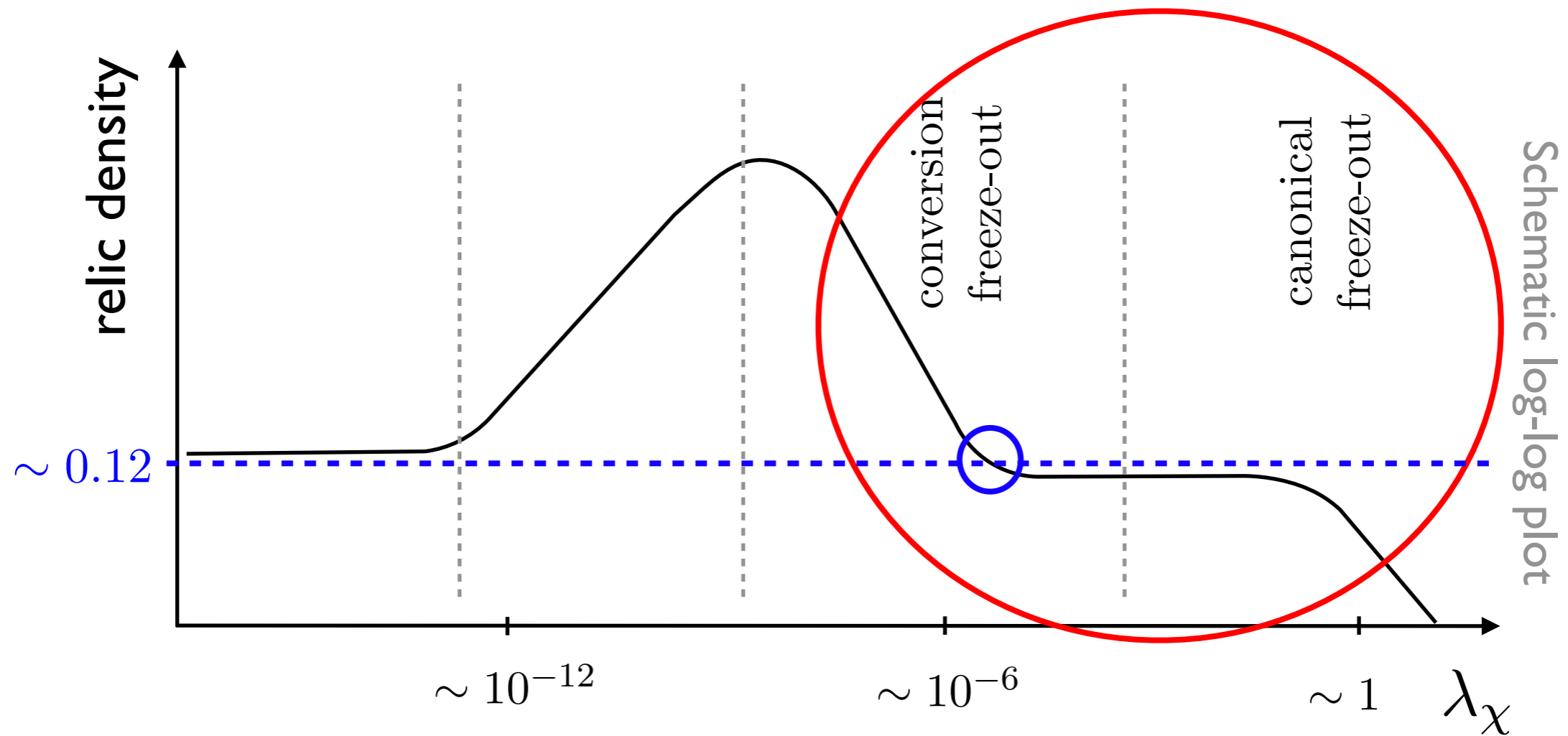
[Garny, JH, Lülf, Vogl 2017; D'Agnolo, Pappadopulo, Ruderman 2017]



Large mass splittings  $\Delta m = m_{\tilde{q}} - m_\chi$  require large  $\lambda_\chi$

# Conversion-driven freeze-out

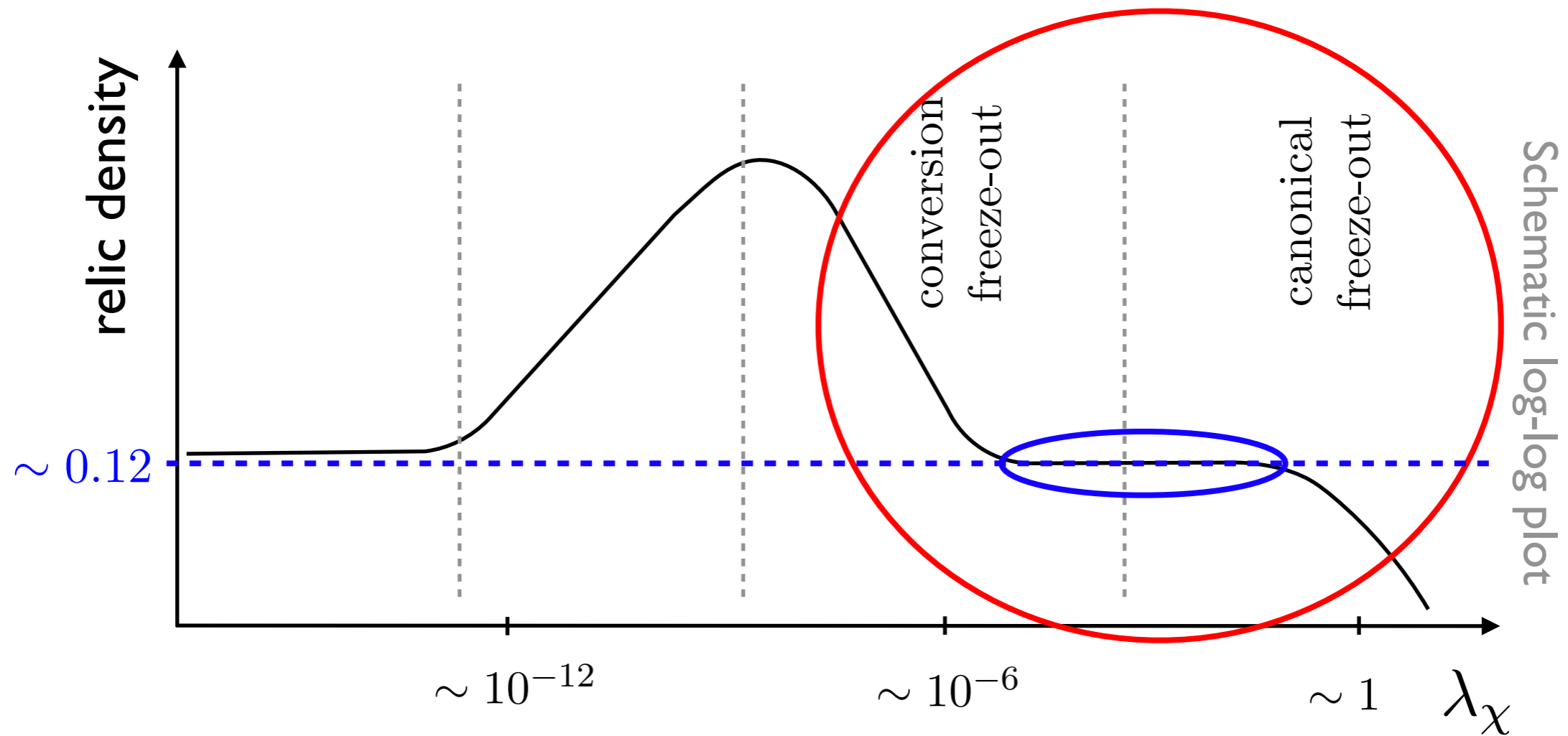
[Garny, JH, Lülf, Vogl 2017; D'Agnolo, Pappadopulo, Ruderman 2017]



Smaller mass splittings  $\Delta m = m_{\tilde{q}} - m_\chi$  require much smaller  $\lambda_\chi$ : sudden drop in the coupling!

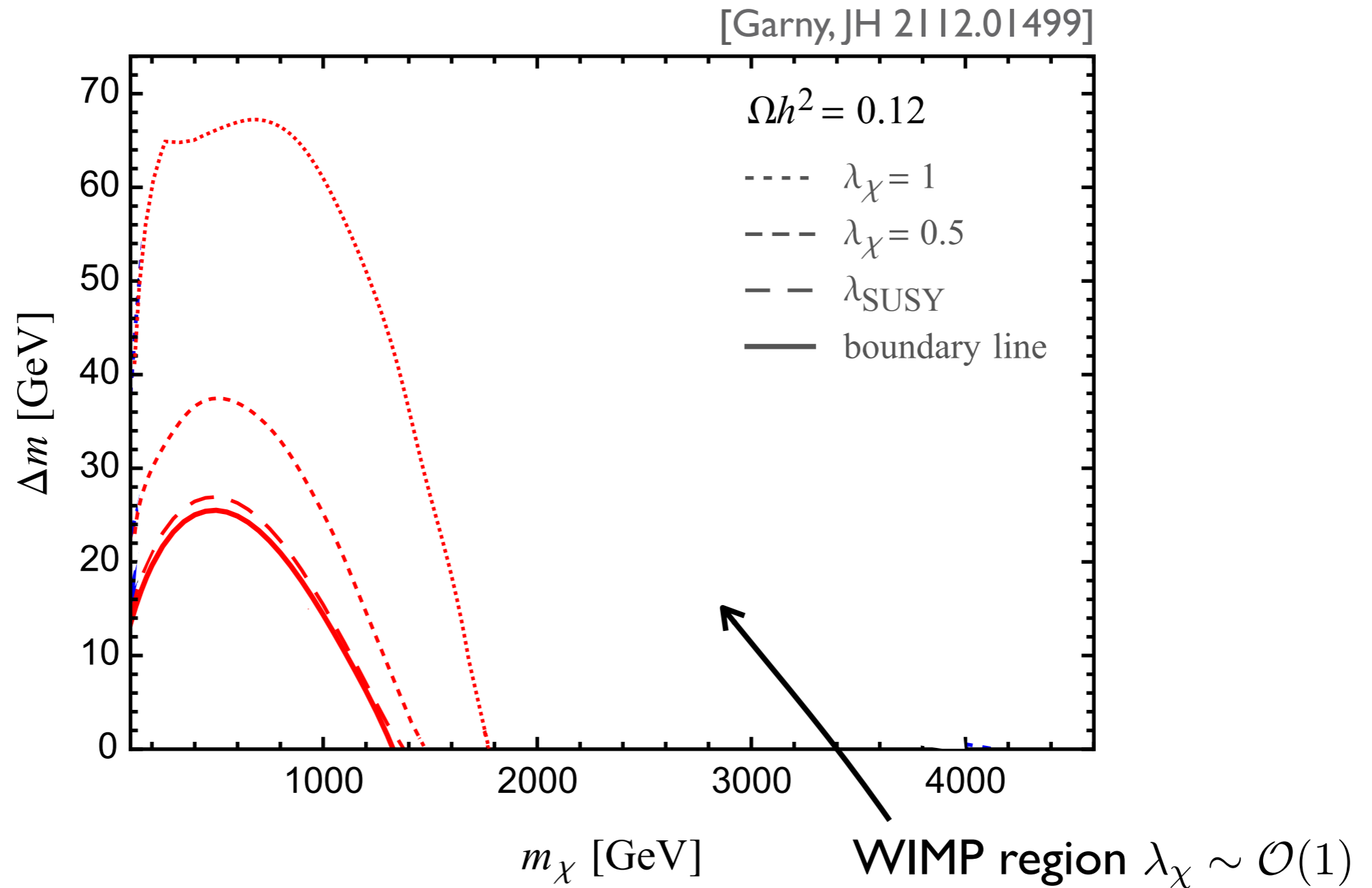
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[Garny, JH, Lülf, Vogl 2017; D'Agnolo, Pappadopulo, Ruderman 2017]

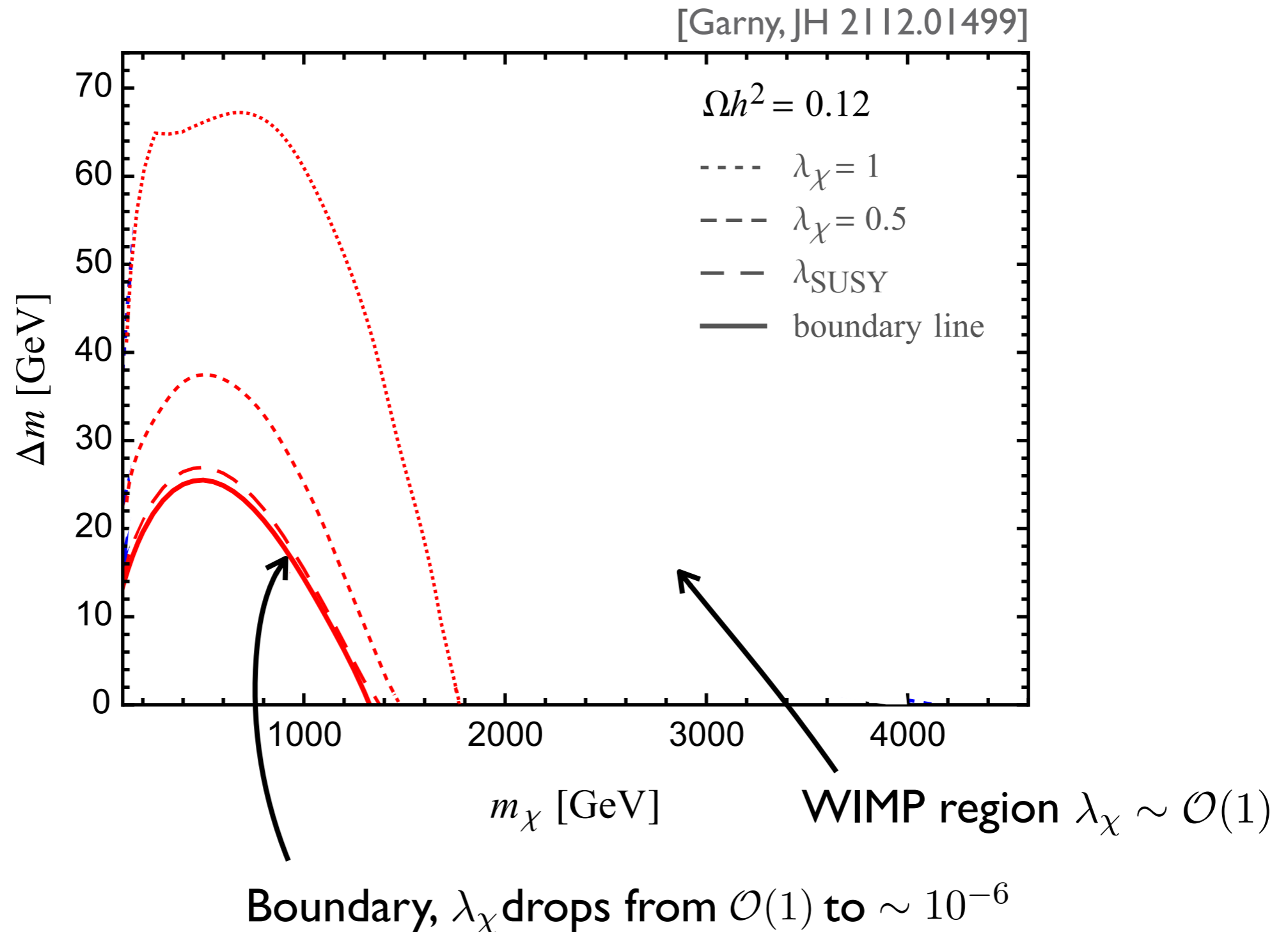


Boundary between WIMP and conversion-driven region  
at characteristic  $\Delta m = m_{\tilde{q}} - m_\chi$

# Bound state effects on the parameter space



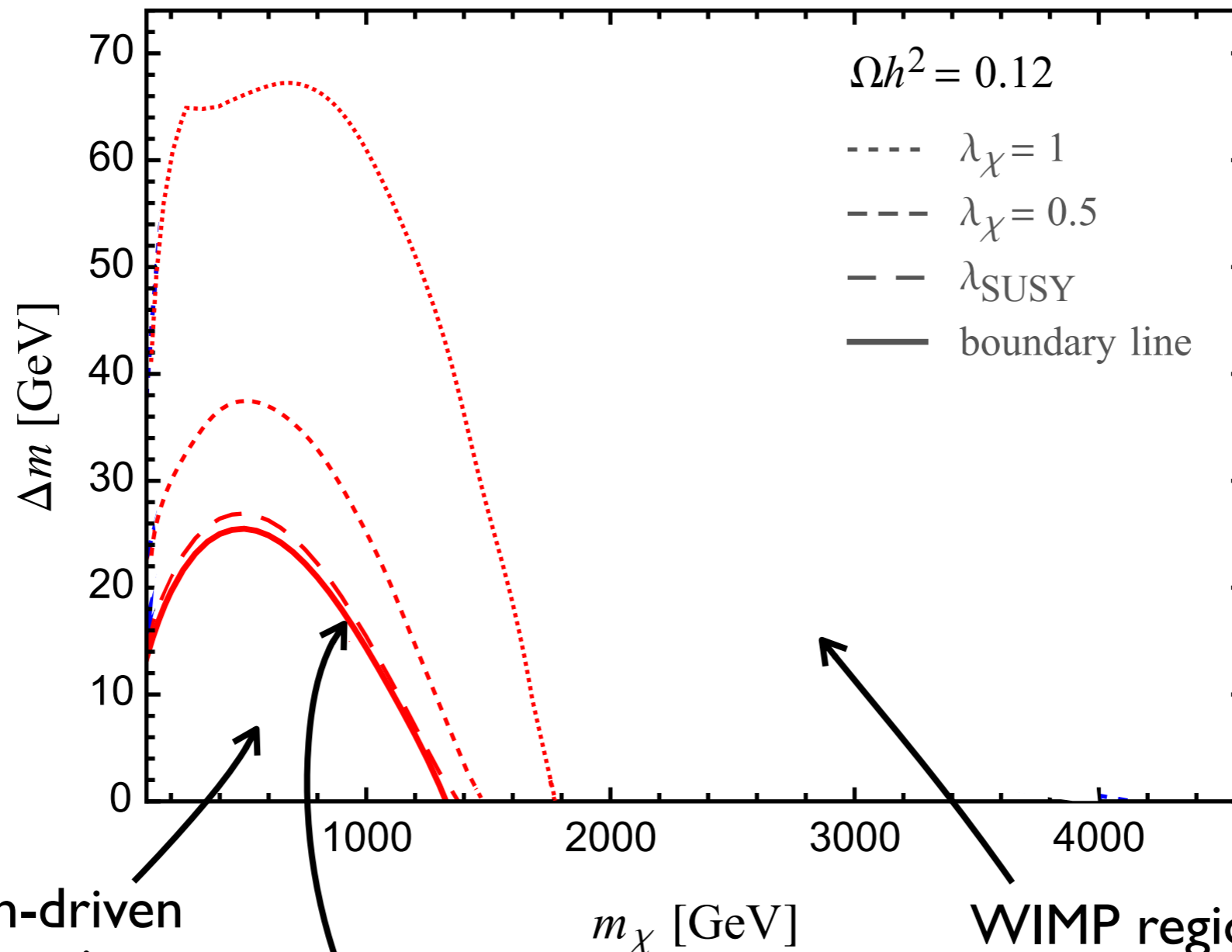
# Bound state effects on the parameter space





# Bound state effects on the parameter space

[Garny, JH 2112.01499]



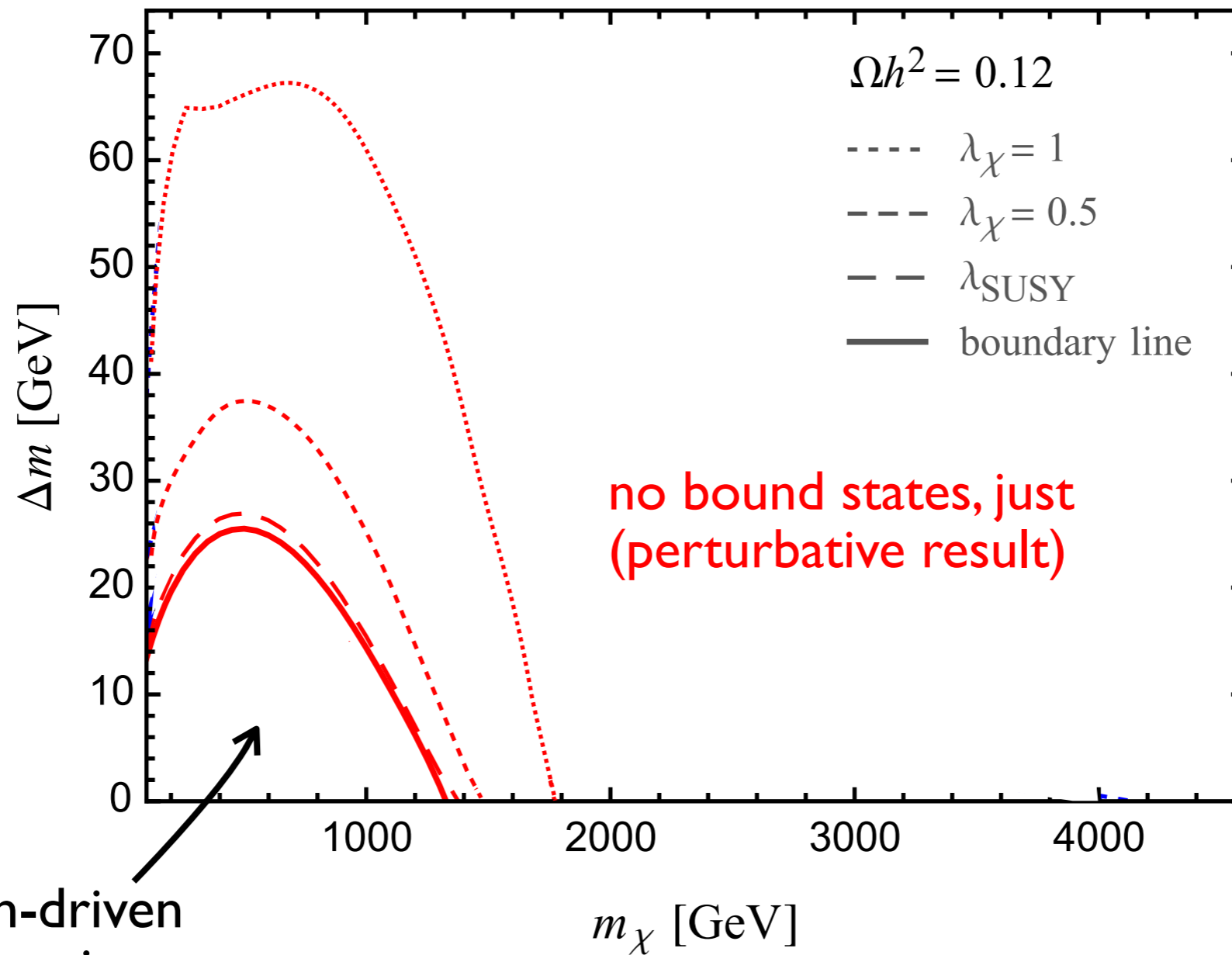
Conversion-driven  
freeze-out region,  
 $\lambda_\chi \sim 10^{-6}$

Boundary,  $\lambda_\chi$  drops from  $\mathcal{O}(1)$  to  $\sim 10^{-6}$

WIMP region  $\lambda_\chi \sim \mathcal{O}(1)$

# Bound state effects on the parameter space

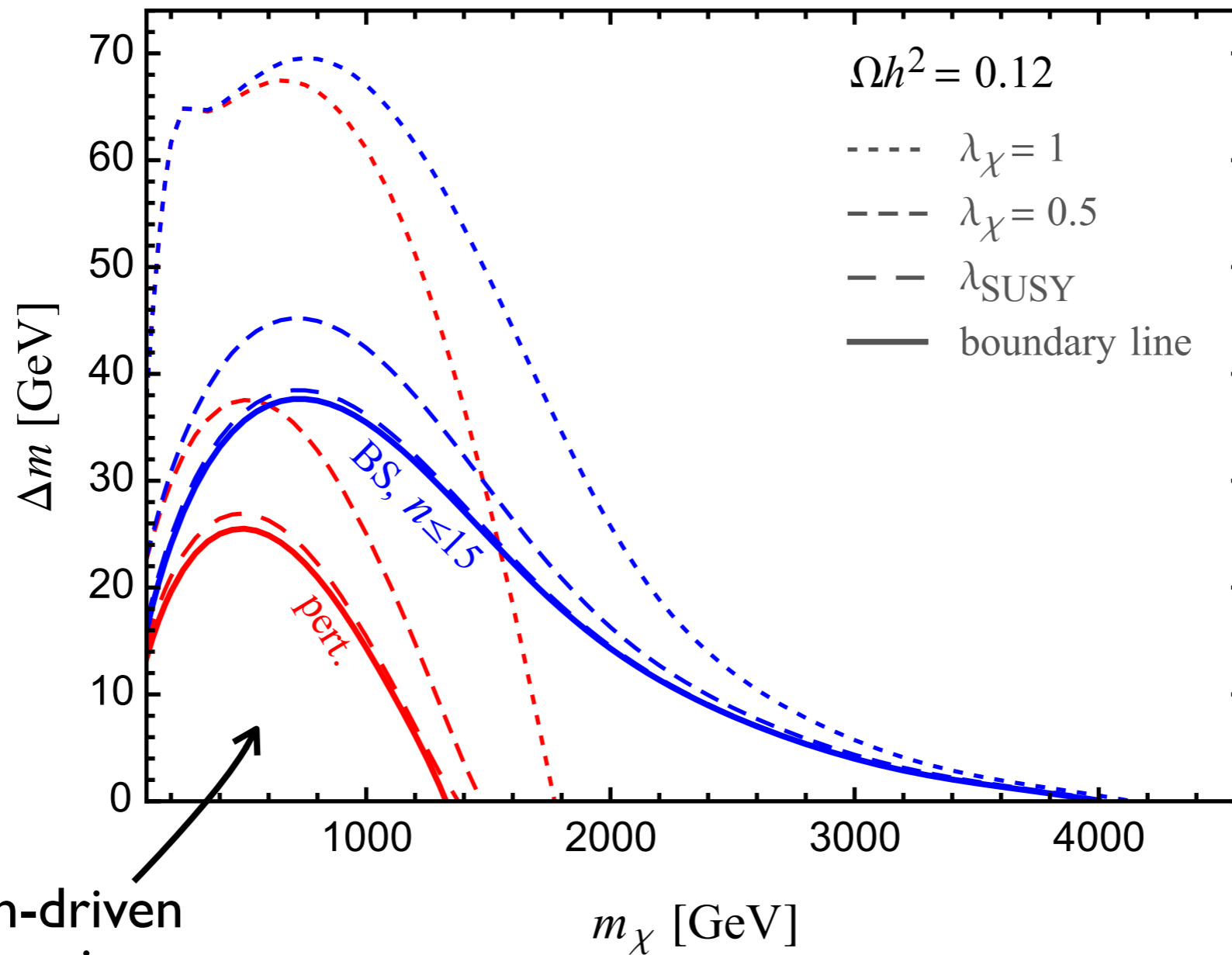
[Garny, JH 2112.01499]



Conversion-driven  
freeze-out region,  
 $\lambda_\chi \sim 10^{-6}$

# Bound state effects on the parameter space

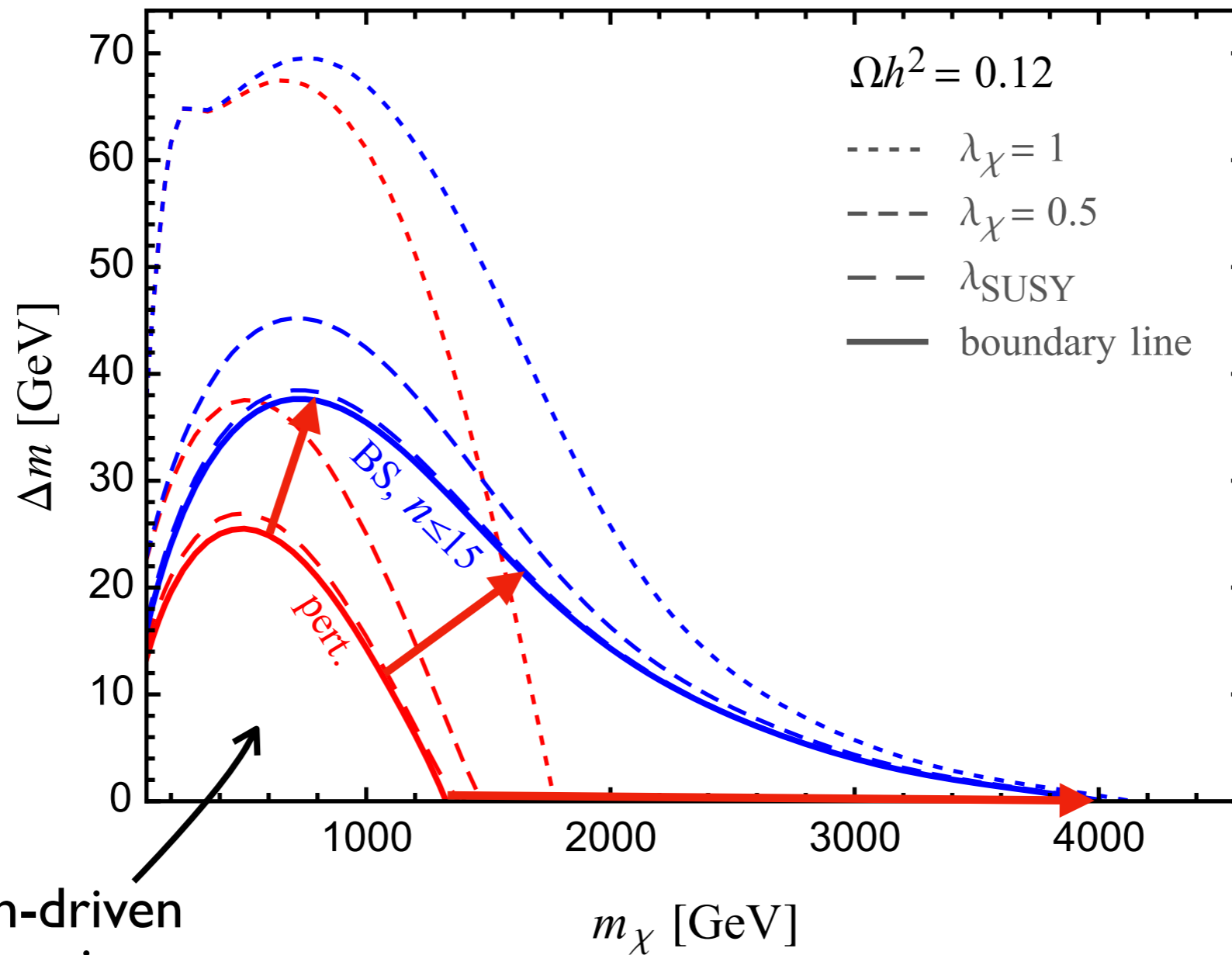
[Garny, JH 2112.01499]



Conversion-driven  
freeze-out region,  
 $\lambda_\chi \sim 10^{-6}$

# Bound state effects on the parameter space

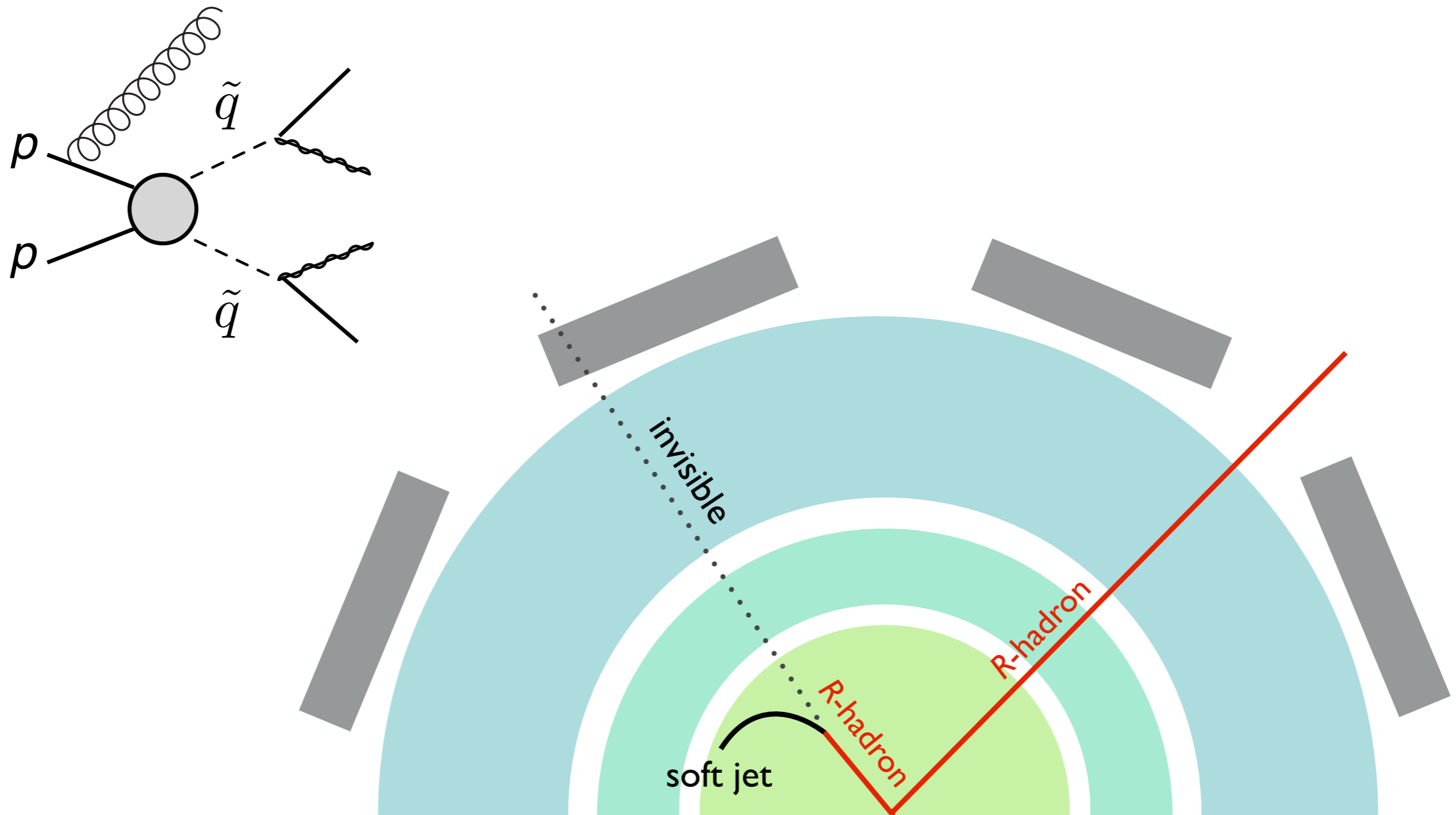
[Garny, JH 2112.01499]



Conversion-driven  
freeze-out region,  
 $\lambda_\chi \sim 10^{-6}$

# Implications for LHC searches

# Feeble couplings: Long-lived particles at LHC



# Lifetime in conversion scenario

Conversion rate on the edge of being efficient:

$$\Gamma_{\text{conv}} \sim H$$

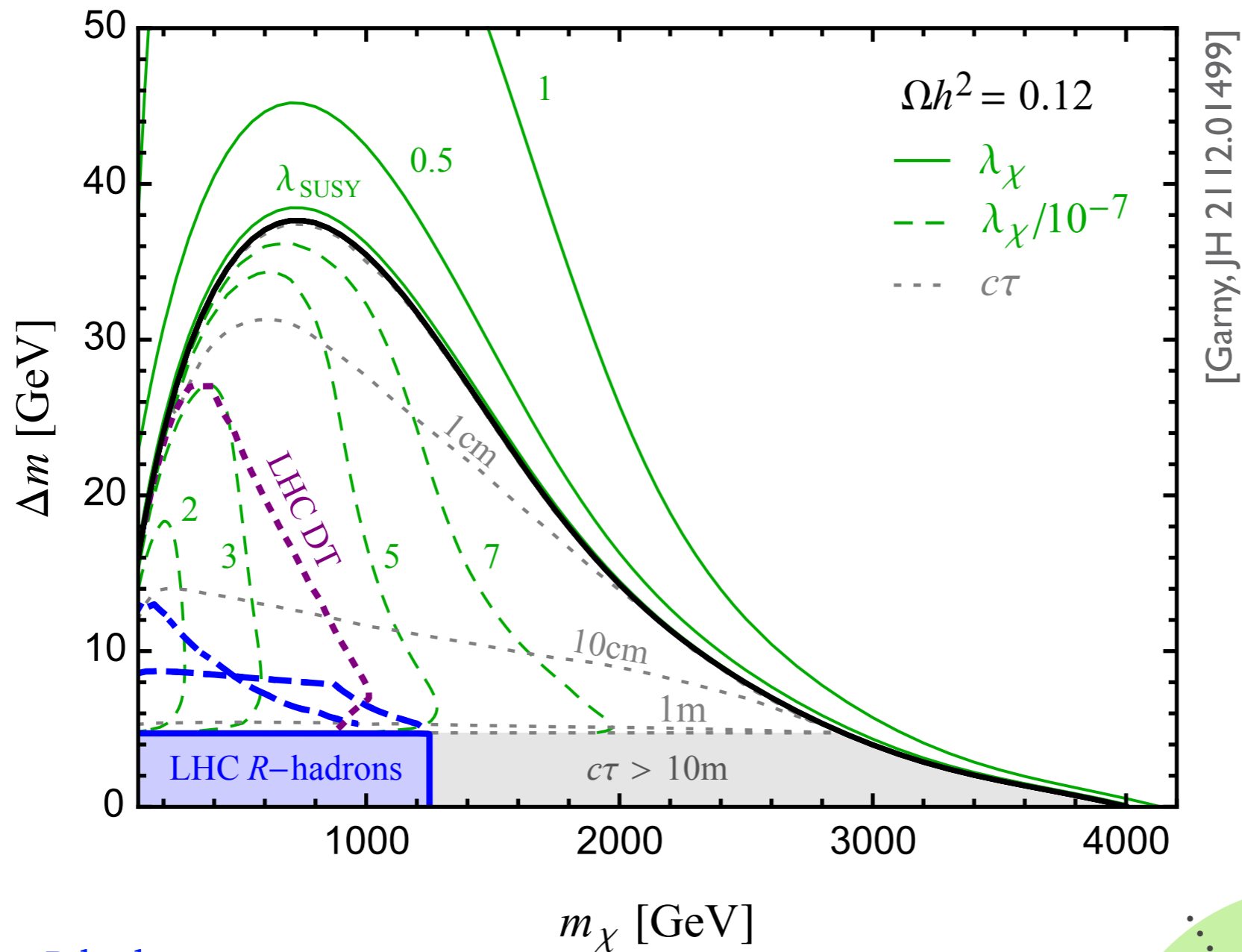
$$\Rightarrow \Gamma_{\text{dec}} \lesssim H$$

$$c\tau \gtrsim H^{-1} \simeq 1.5 \text{ cm} \left( \frac{(100 \text{ GeV})^2}{T^2} \right)$$

$$T \lesssim (10-100) \text{ GeV}$$

**$\Rightarrow$  Long-lived particles (LLPs) at LHC!**

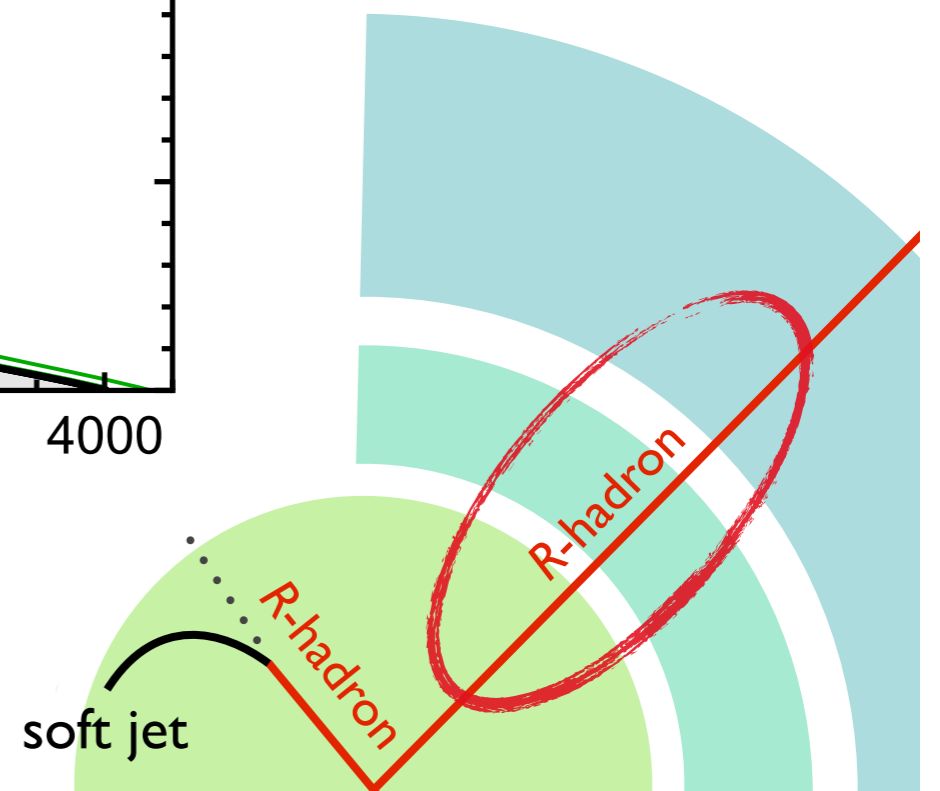
# Collider constraints



LHC – R-hadrons:

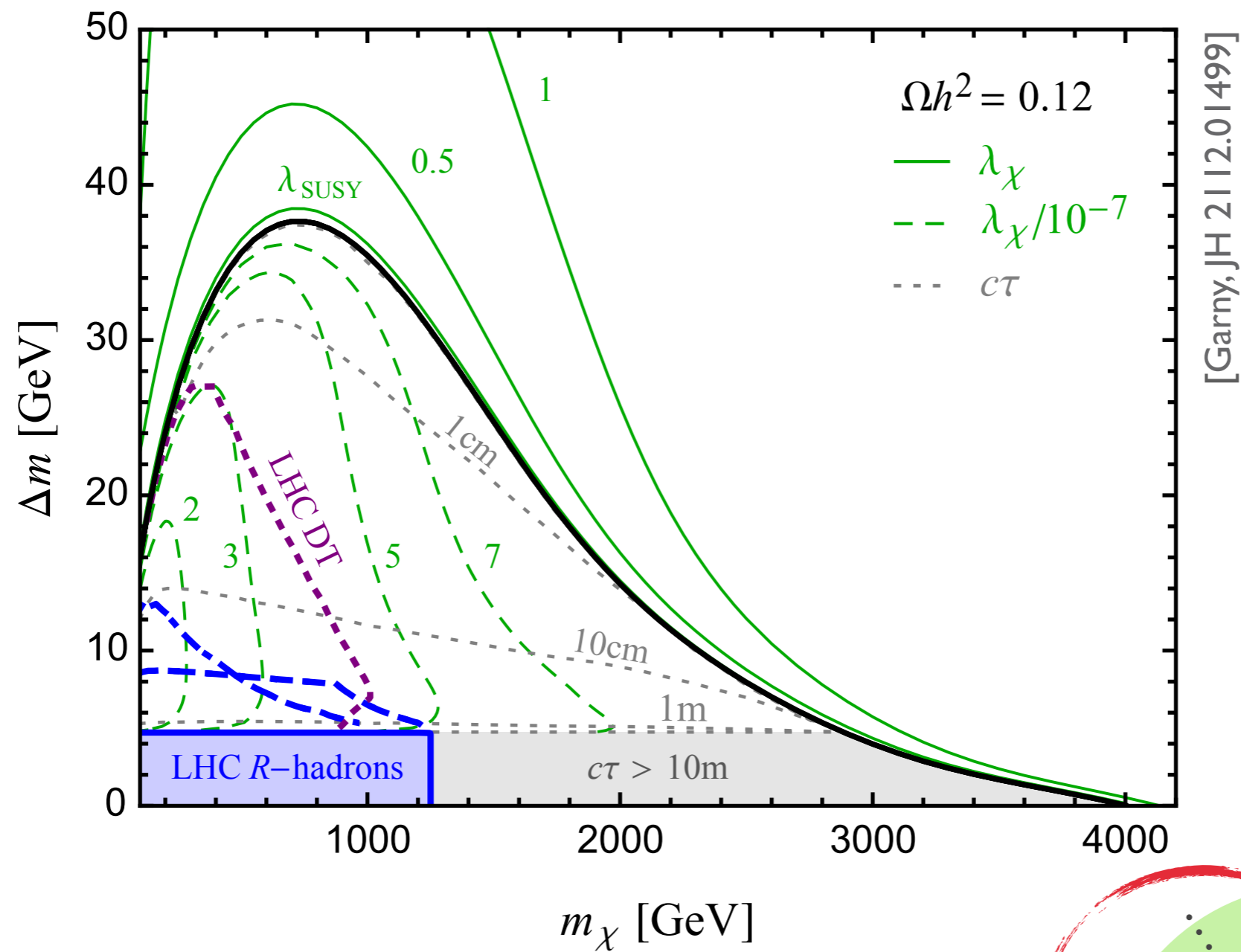
ATLAS [1902.01636, 1808.04095 approximate reinterpretation]

CMS [CMS-PAS-EXO-16-036, recasting from 1705.09292]

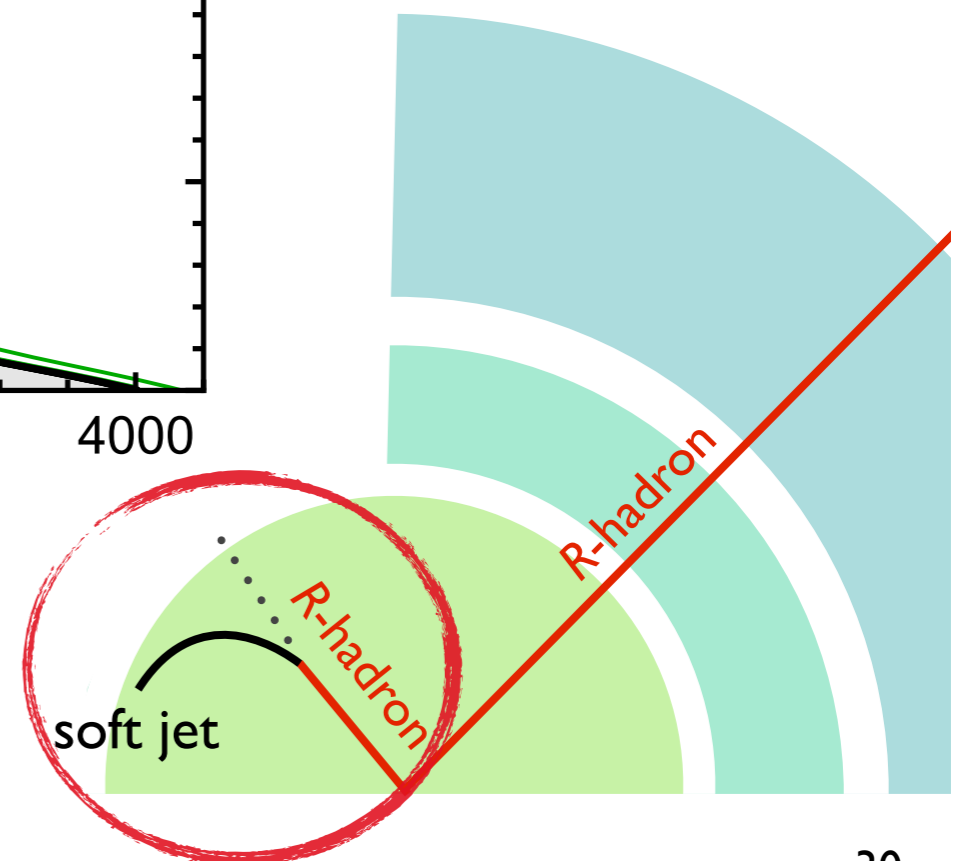




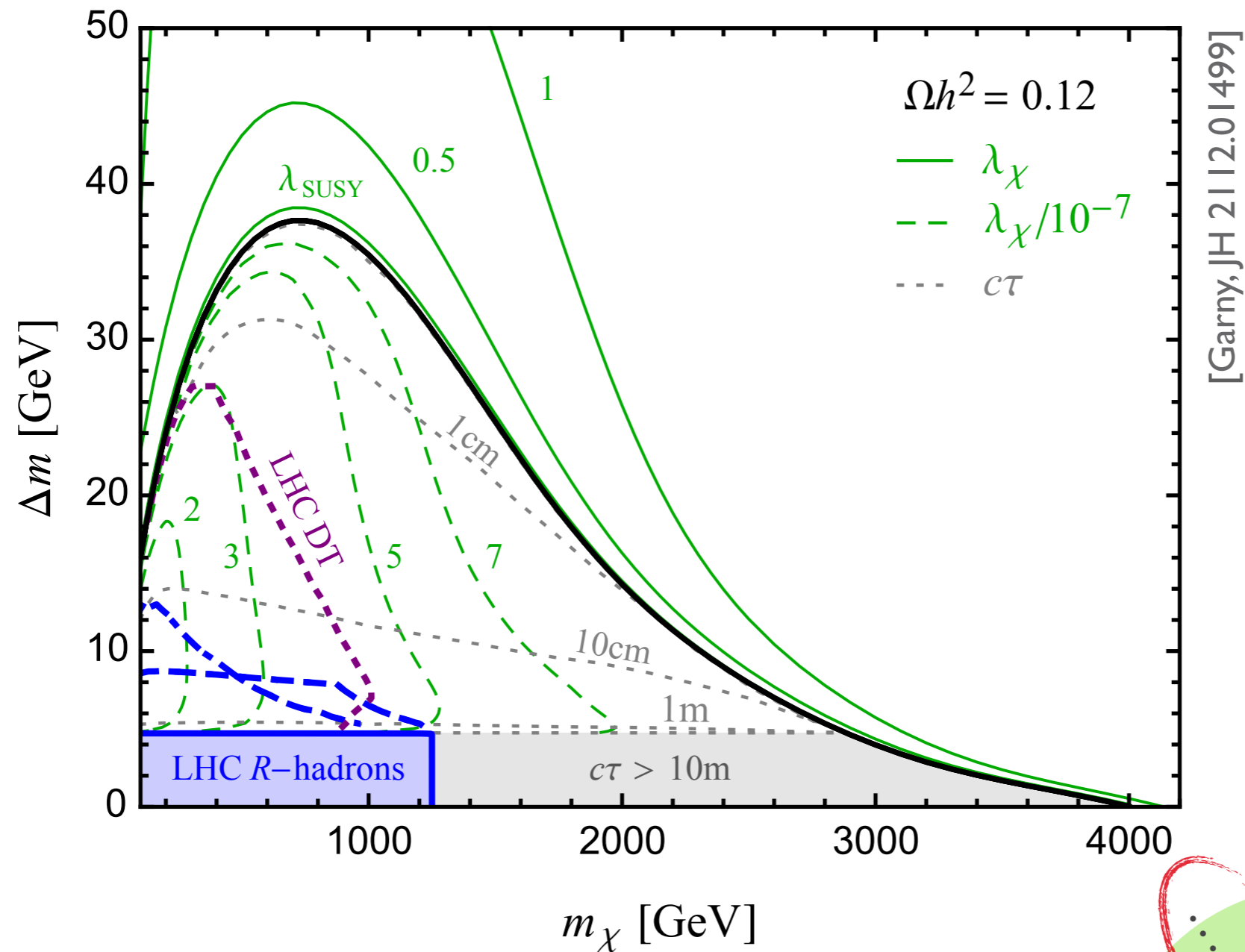
# Collider constraints



[Garny, JH 2112.01499]

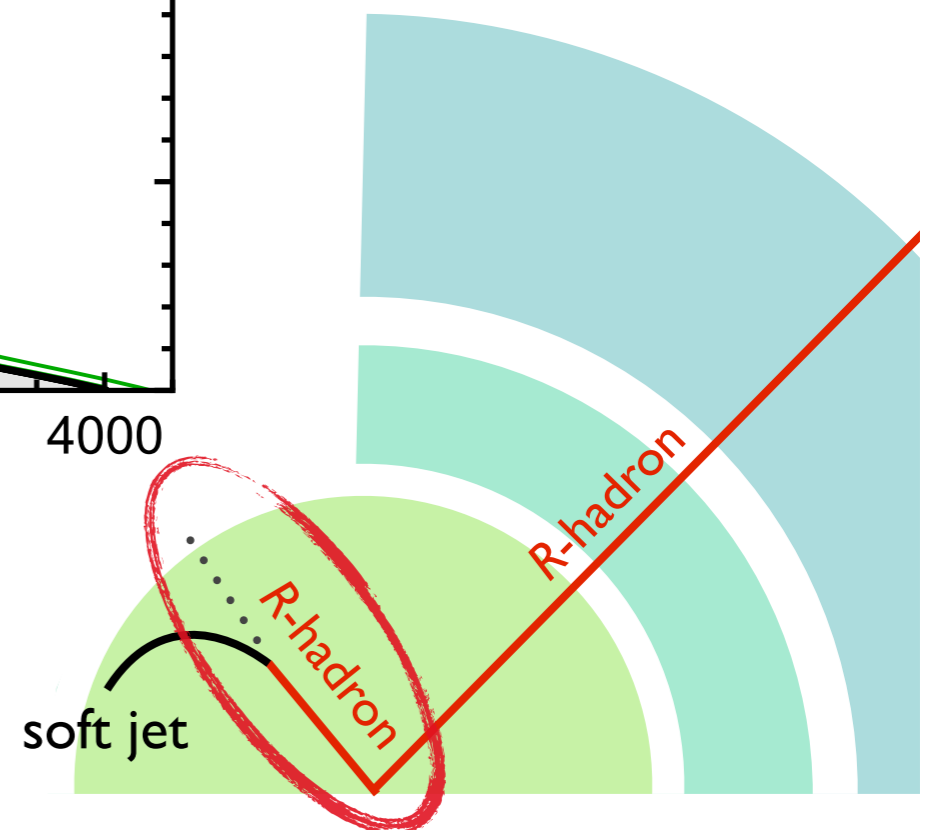


# Collider constraints

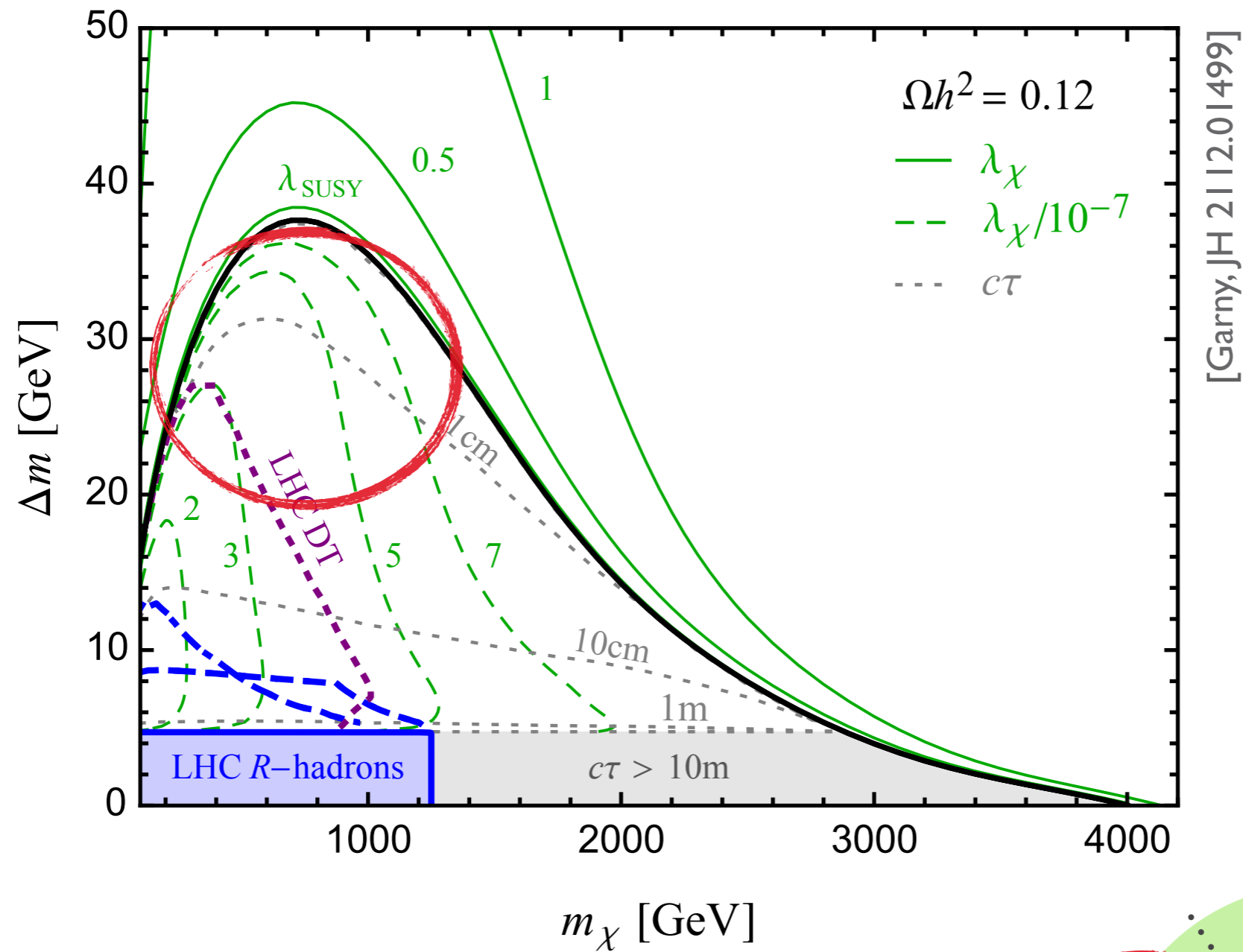


[Garny, JH 2112.01499]

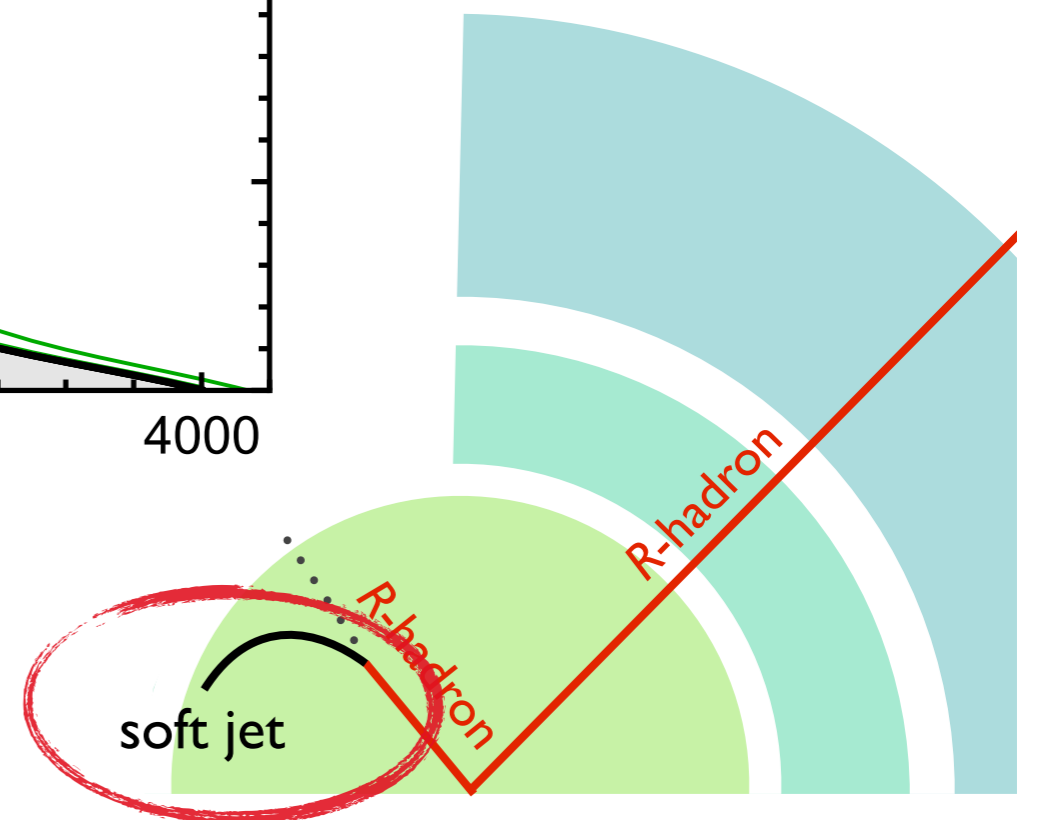
LHC – DT: ATLAS Disappearing-track search  
[1712.02118, recasting from 2002.12220, 7]



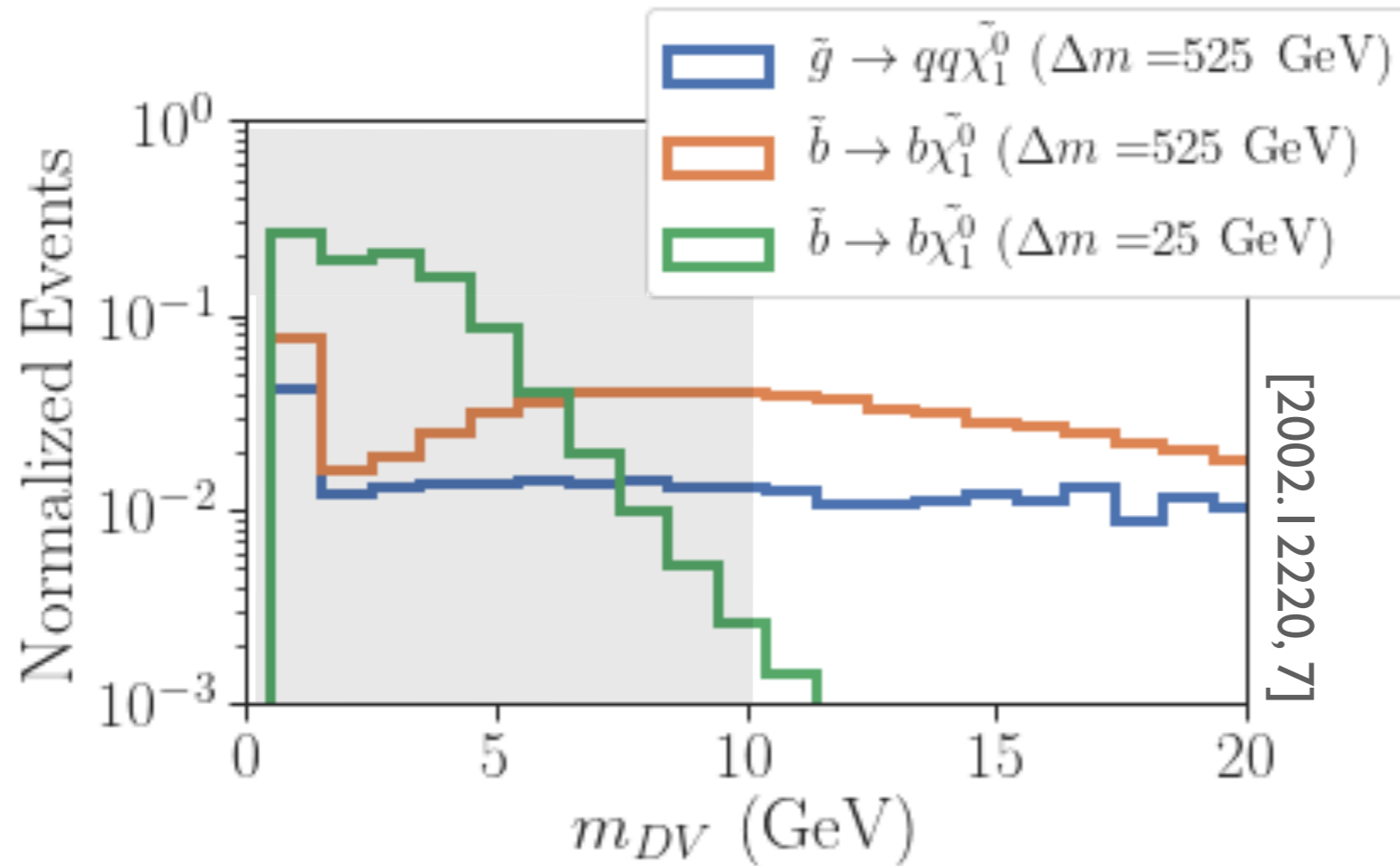
# Collider constraints



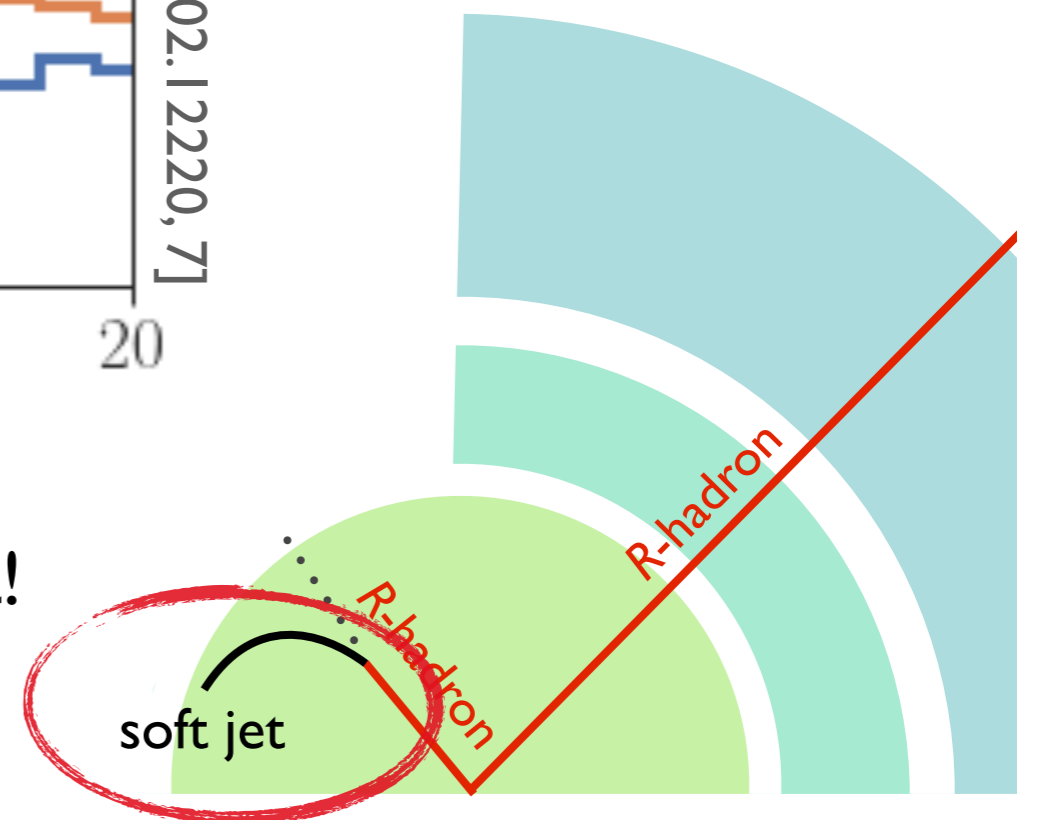
[Garny, JH 2112.01499]



# Collider constraints

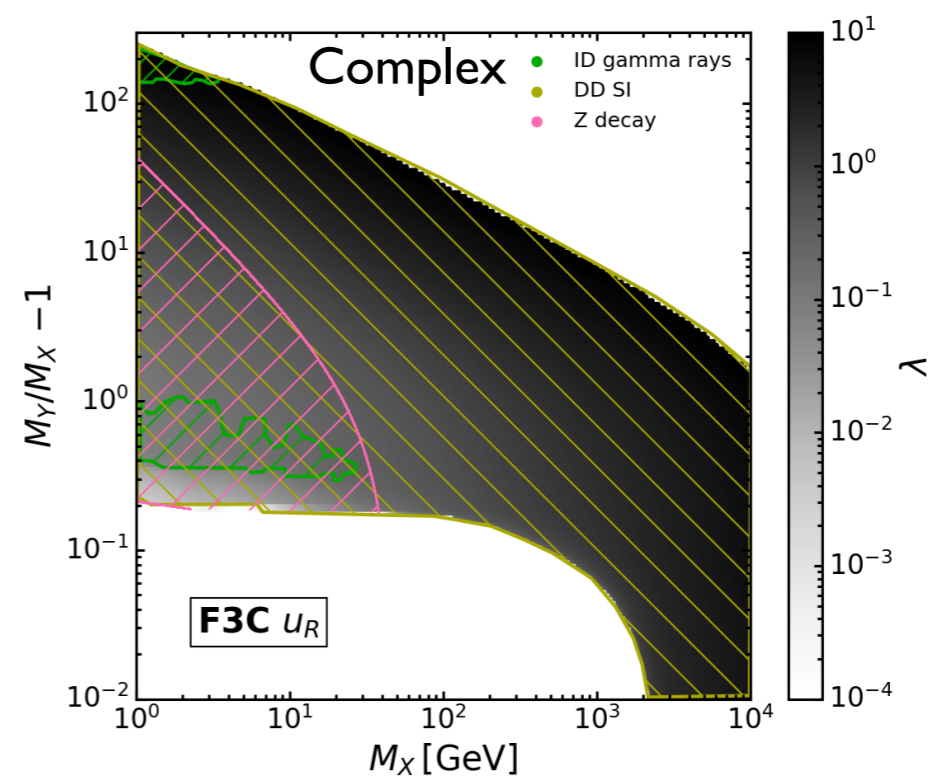
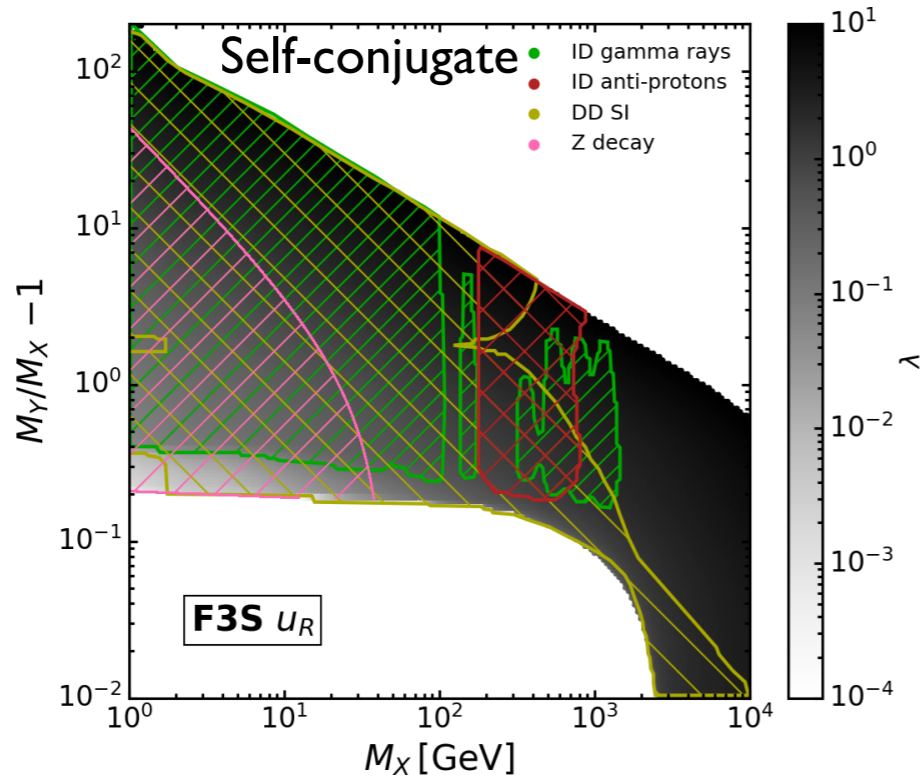


- Displaced jets+MET suffers from  $m_{inv}$ -cut!  
[ATLAS, 1710.04901]

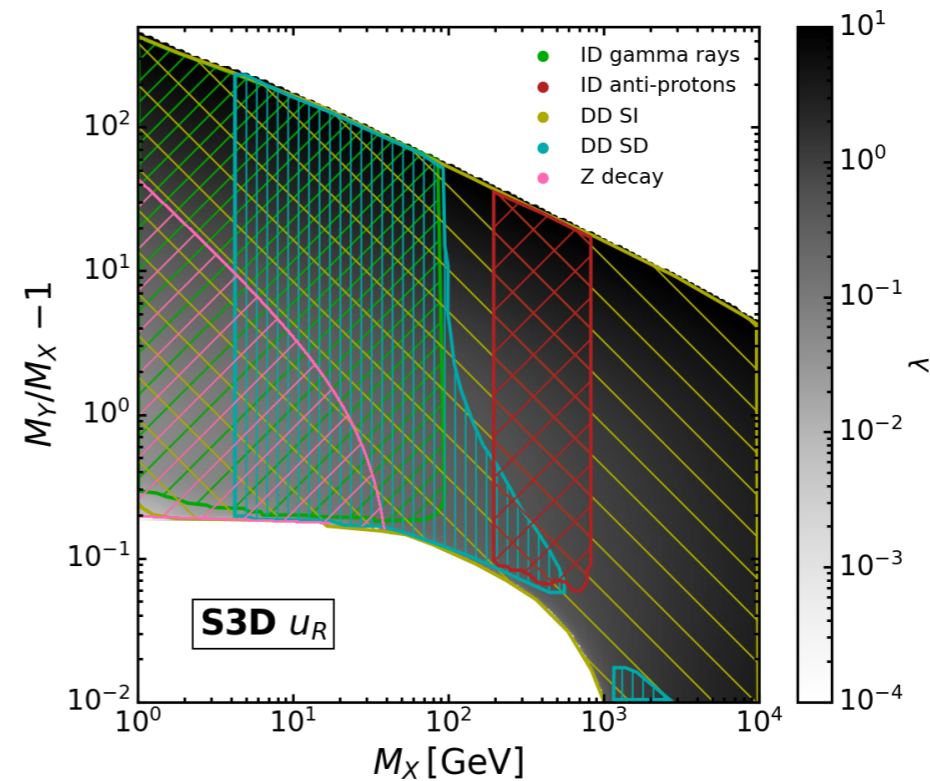
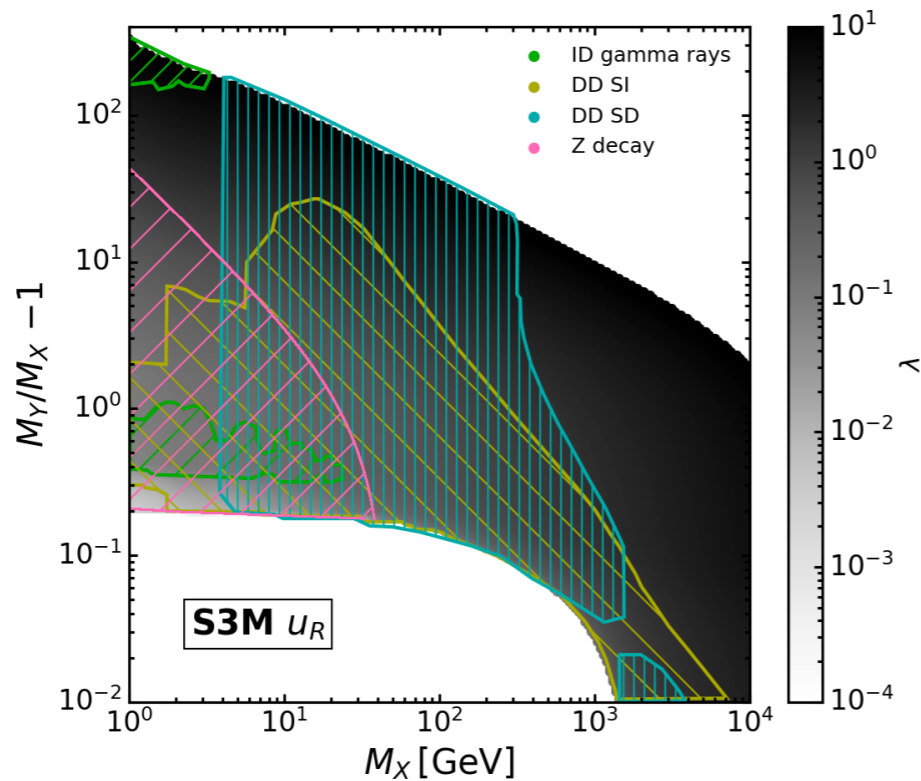


# Relevance for current searches

Scalar DM



Fermionic DM

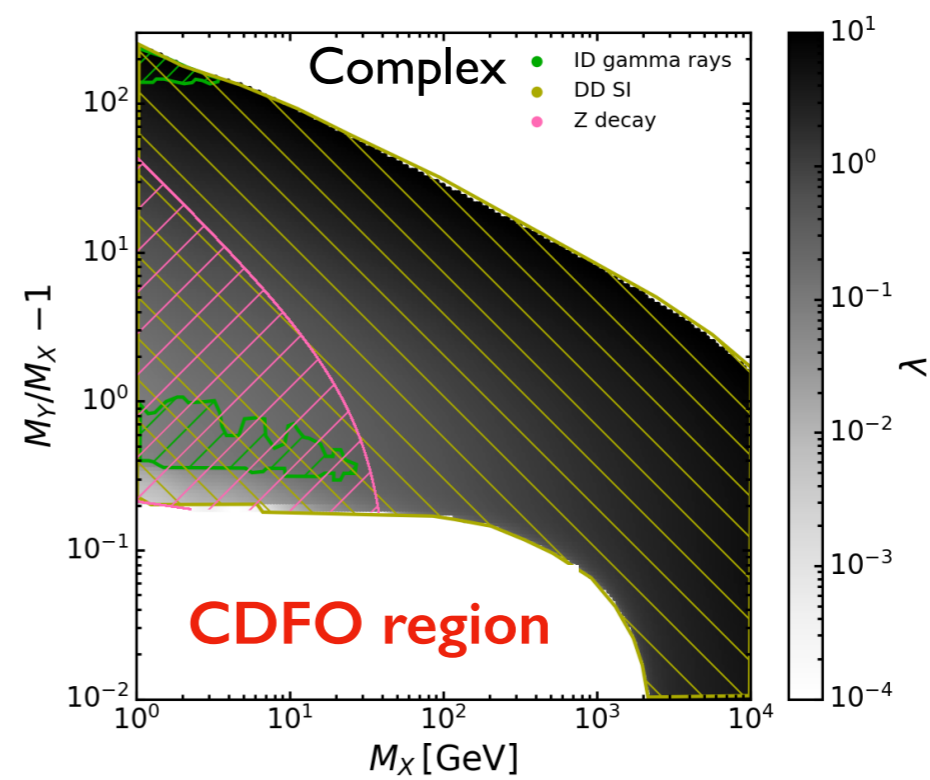
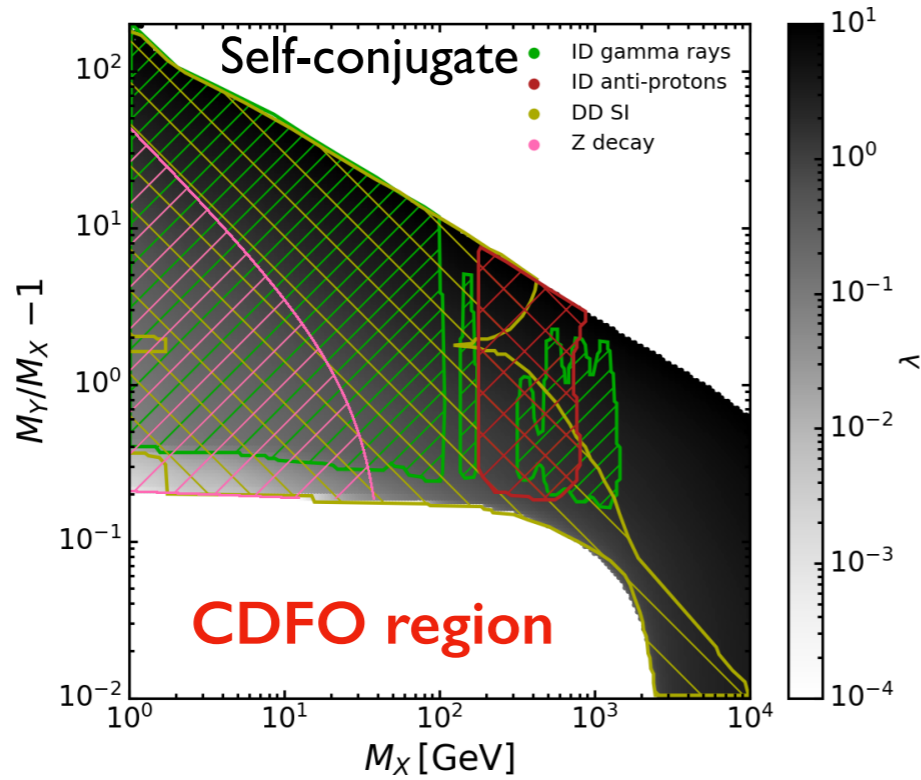


[Arina, Fuks, JH, Krämer, Mantani, Panizzi; 2307.10367]

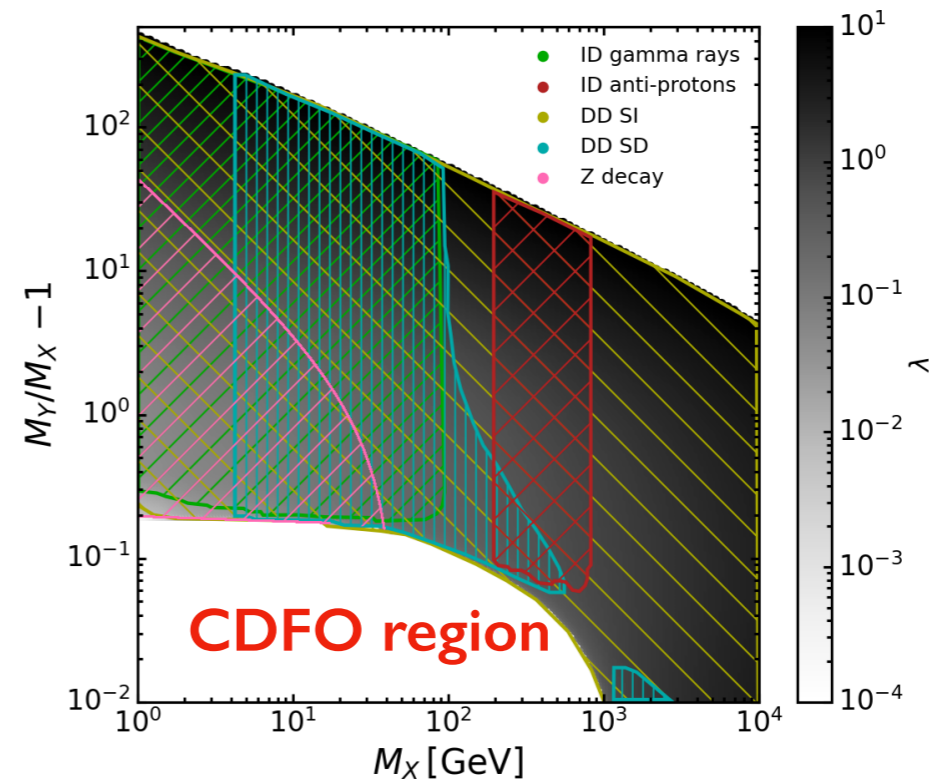
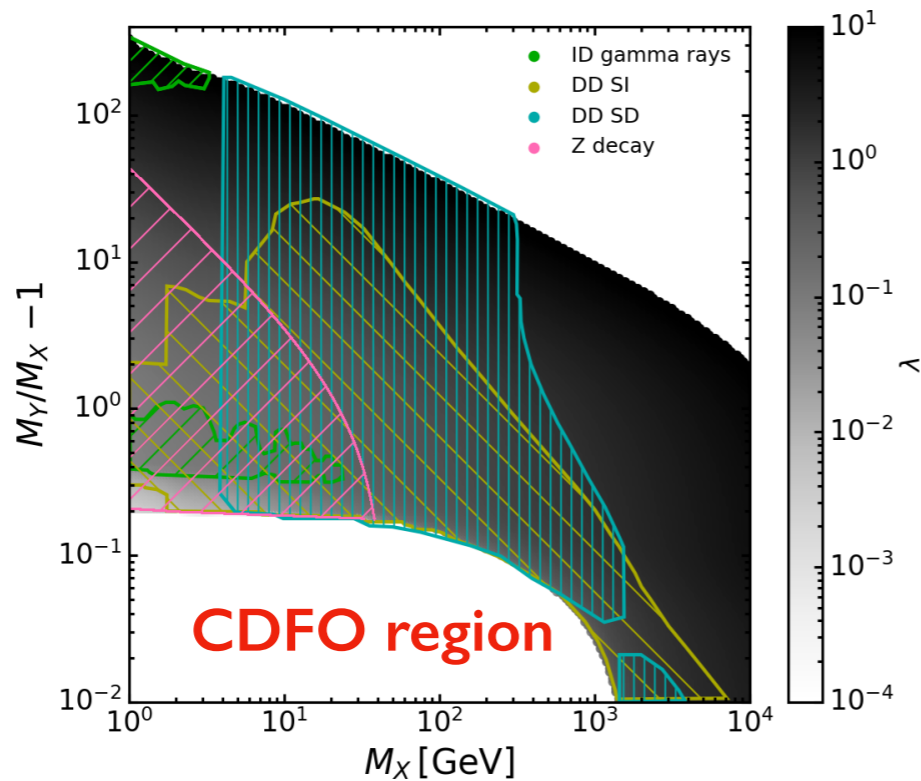


# Relevance for current searches

Scalar DM

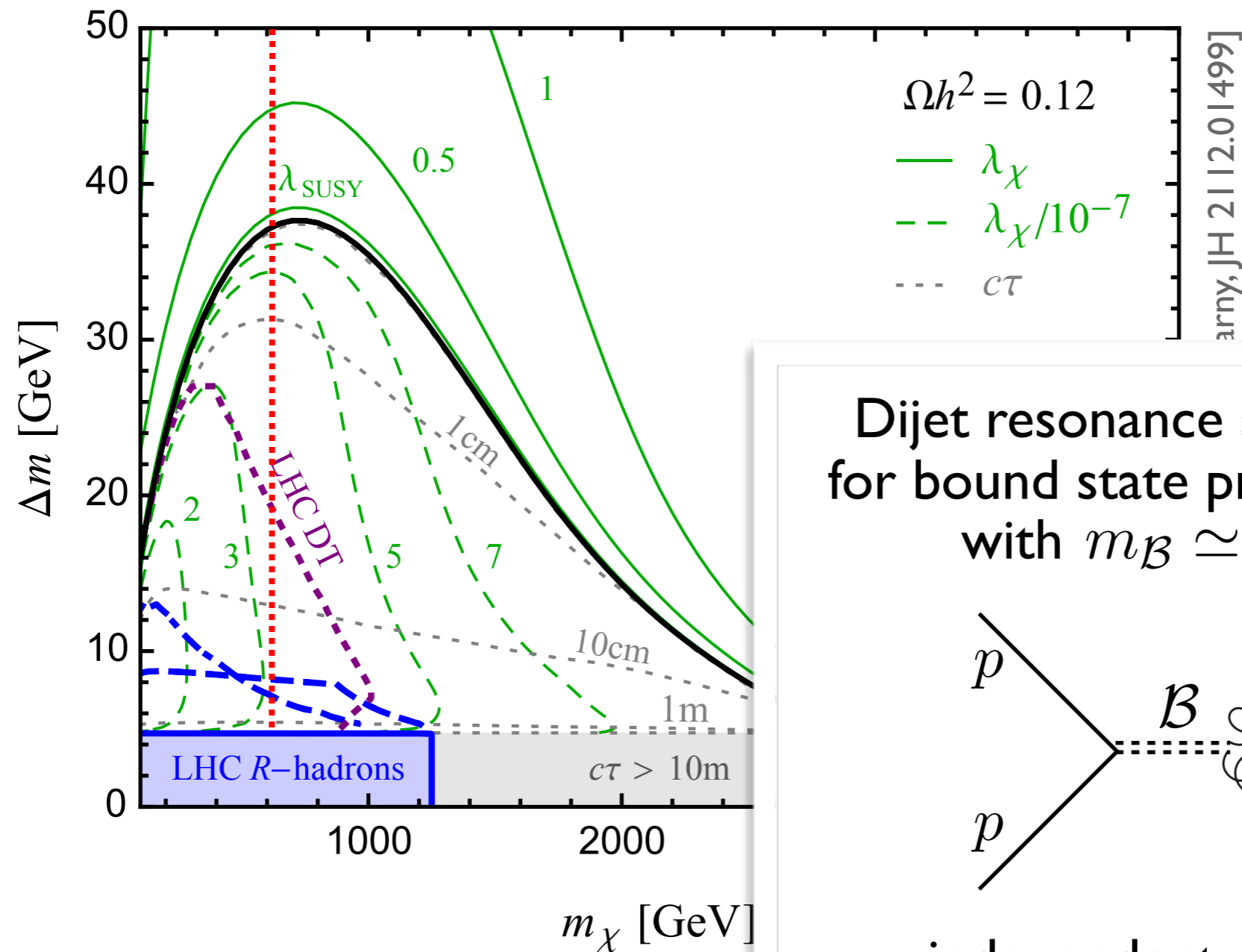


Fermionic DM

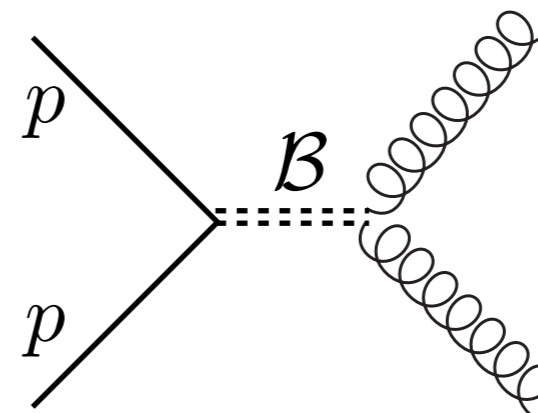


[Arina, Fuks, JH, Krämer, Mantani, Panizzi; 2307.10367]

# Collider constraints



Dijet resonance searches  
for bound state production  
with  $m_B \simeq 2m_{\tilde{q}}$



independent of lifetime  
and  $\Delta m$



# Summary

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- Dark matter elusive: systematically explore mechanisms of DM genesis
  - Consider minimal extensions to SM with large hierarchy in couplings  $\lambda_{\text{strong}} \gg \lambda_{\text{feeble}}$
  - Prolonged freeze-out dynamics:  
⇒ effects of excited bound states highly relevant
  - Non-abelian theory: ‘eternal’ annihilation
  - superWIMP scenario: Relic density does depend on decay rate
  - Conversion-driven freeze-out: parameter space largely enhanced
  - Interesting prospects for long-lived particle searches
  - Open problem: Unitarization of BSF cross section
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