

# Recent Results from WIMP-search analysis of CDMS-2 data

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for the CDMS collaboration

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# CDMS Collaboration



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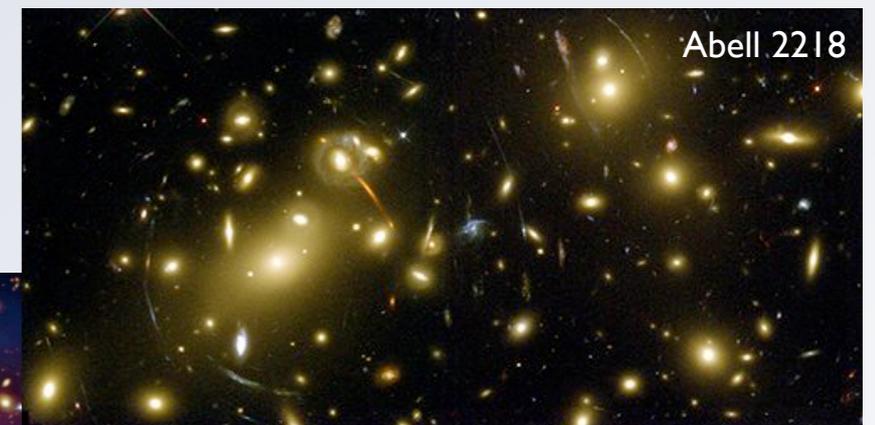
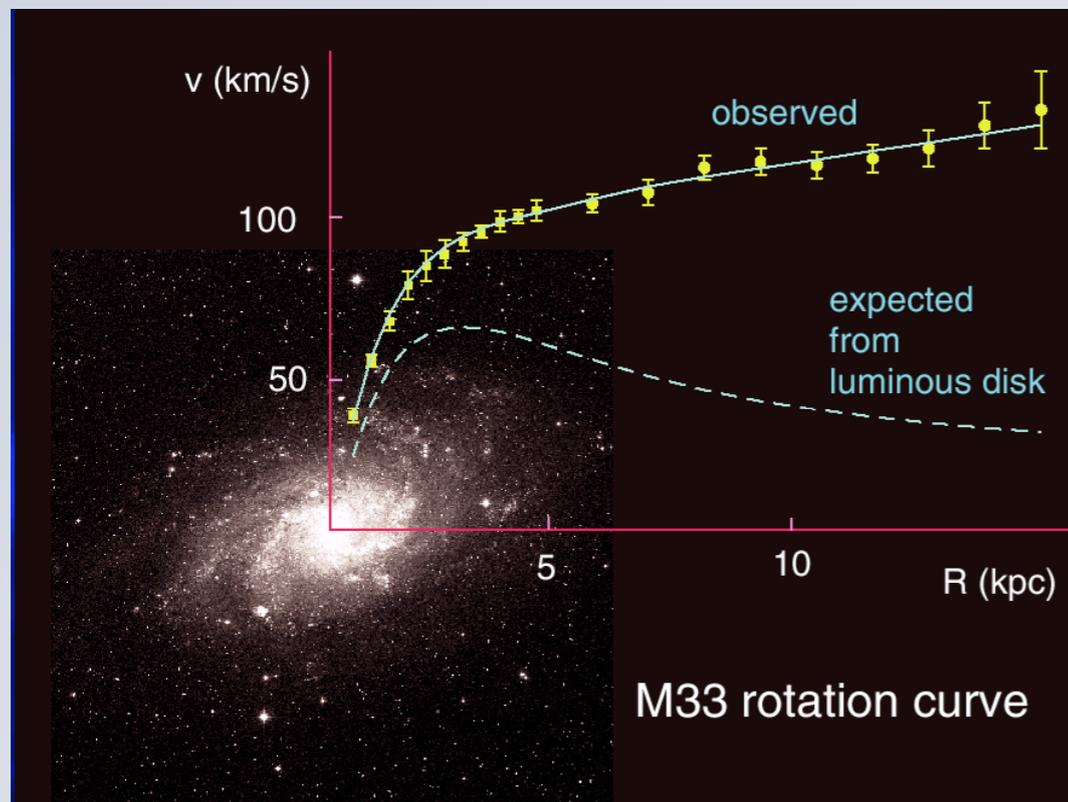
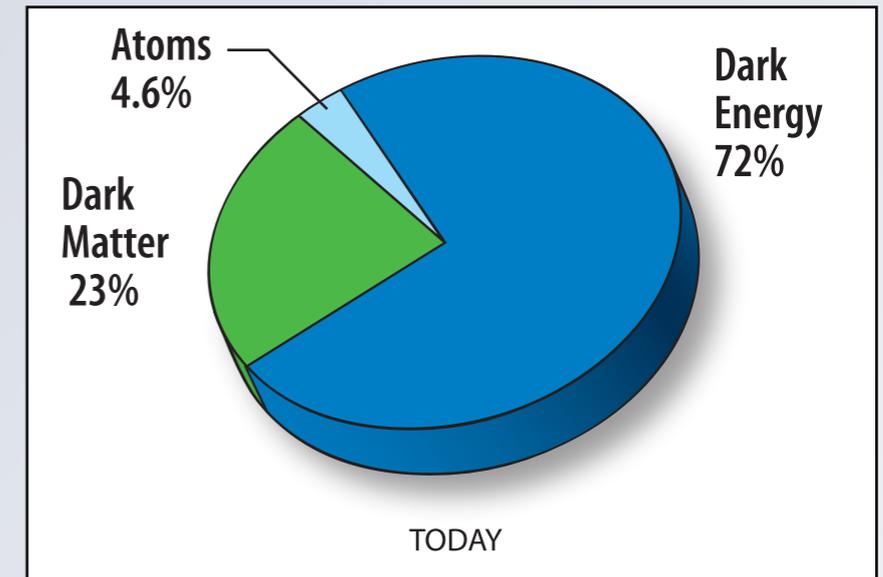
[P. Cushman](#), [L. Dong](#), [M. Fritts](#), [V. Mandic](#), [X. Qiu](#), [O. Kamaev](#),  
[A. Reissetter](#)

## University of Zurich

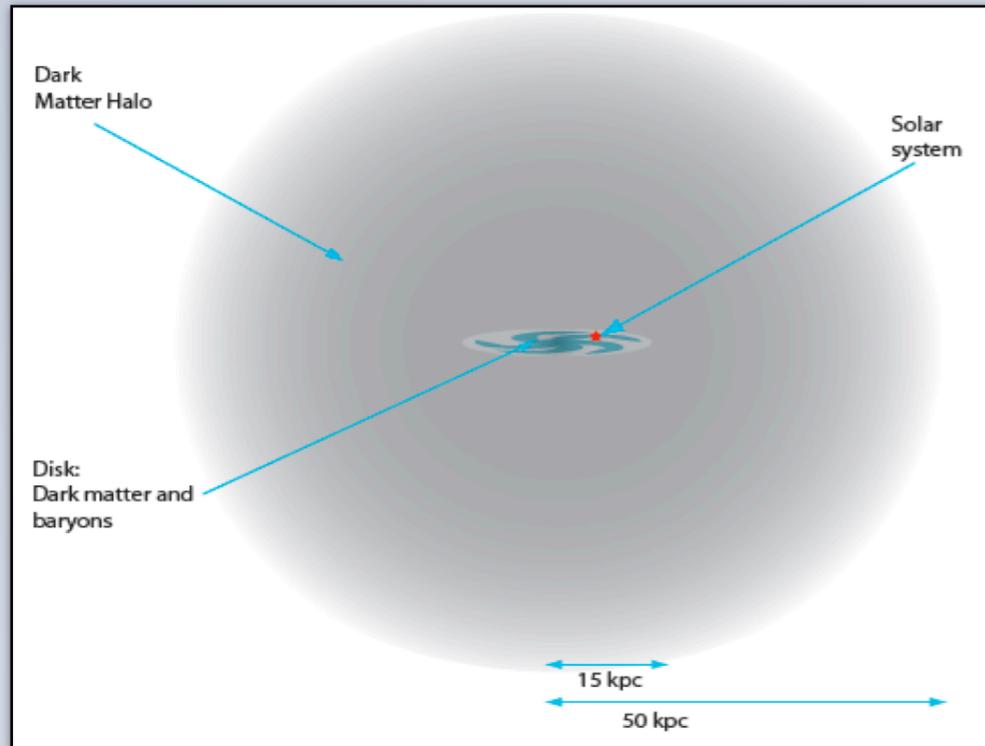
[S. Arrenberg](#), [T. Bruch](#), [L. Baudis](#), [M. Tarka](#)

# Motivating Dark Matter & WIMPs

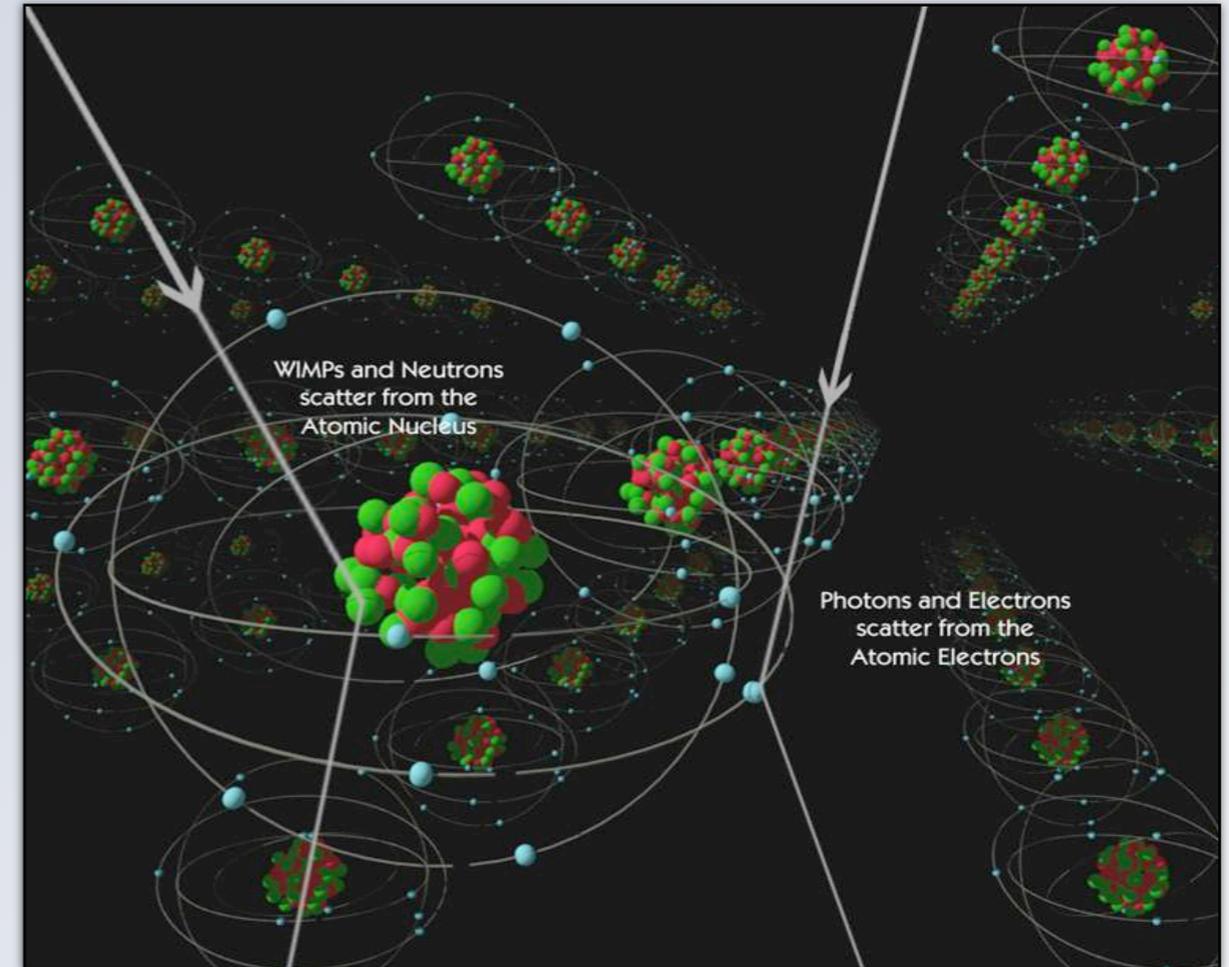
- Evidence for dark matter on various scales
  - Galaxies
  - Galaxy clusters
  - Large scale
- WIMPs (among other models) are particularly well motivated



# Direct Detection of WIMPs



**Spherical Isothermal Halo**  
**Max. Boltz.  $\nu$  distribution,**  
 **$\langle v \rangle \sim 230 \text{ km/s}$**



## Build a good mousetrap!

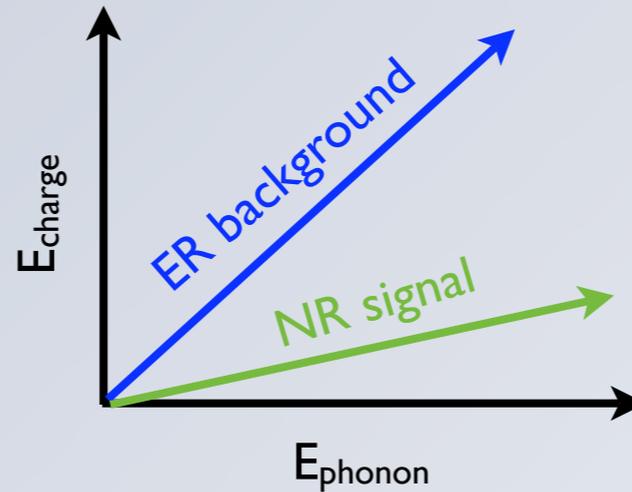
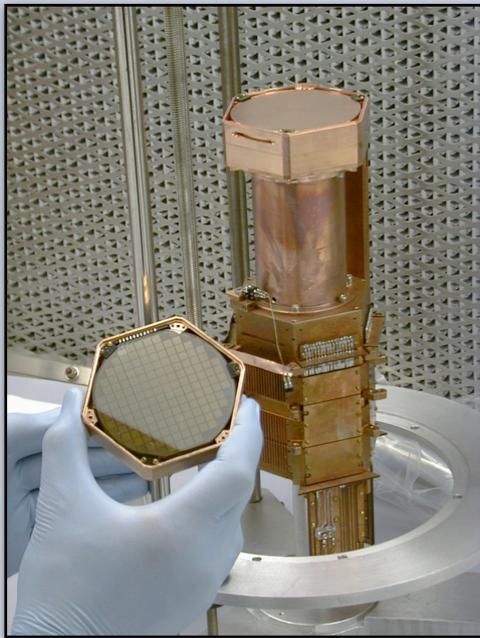
- Choose target material to “see” recoils
- Discriminate NRs from ERs
- Reduce background

RARE EVENT SEARCH

**“See” Nuclear Recoils**  
 **$\langle E_r \rangle \sim 30 \text{ keV}$ ,  $\sim 1 \text{ event/kg/100 days}$**

# CDMS-2 in a nutshell

**Ge & Si** target masses



**Allow <1 background event to maximize discovery potential**

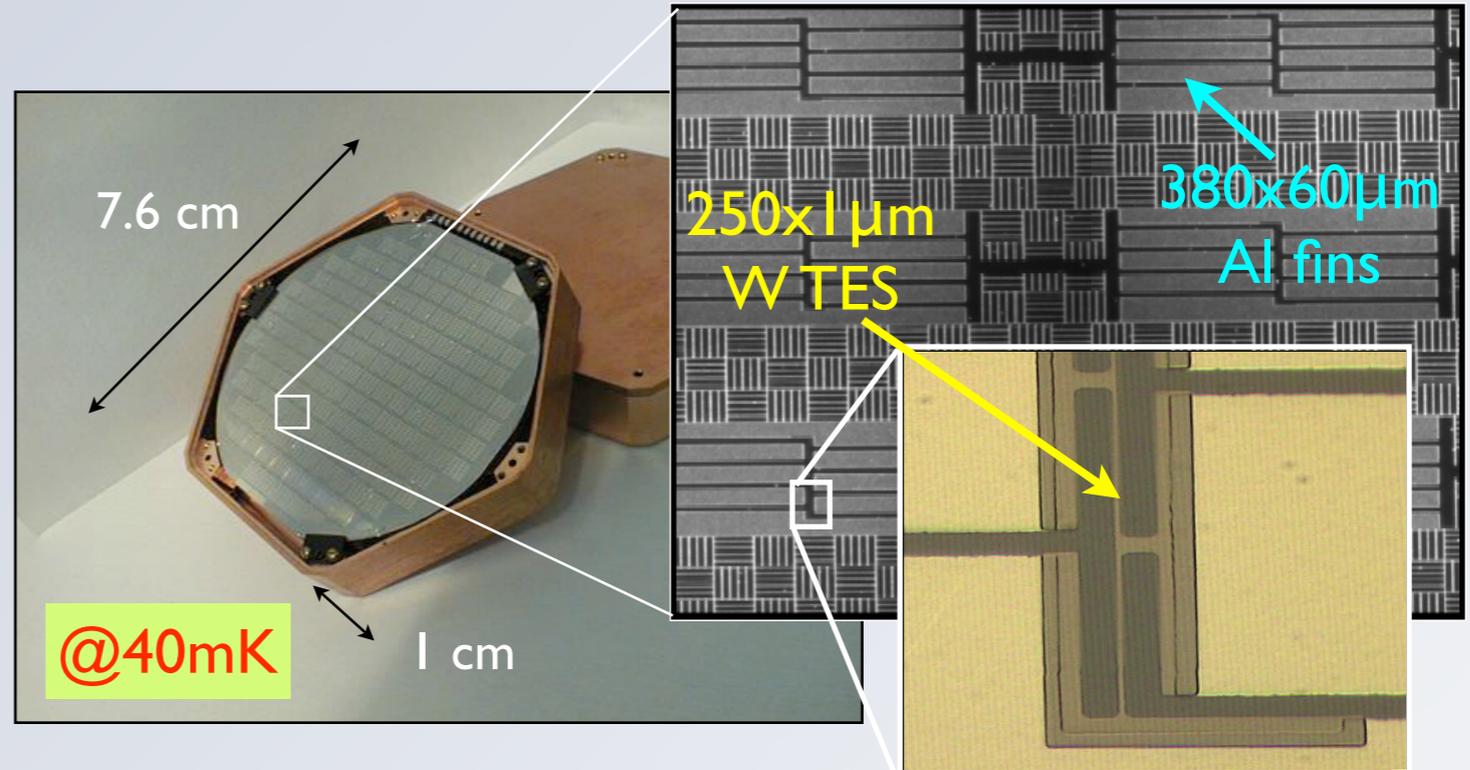
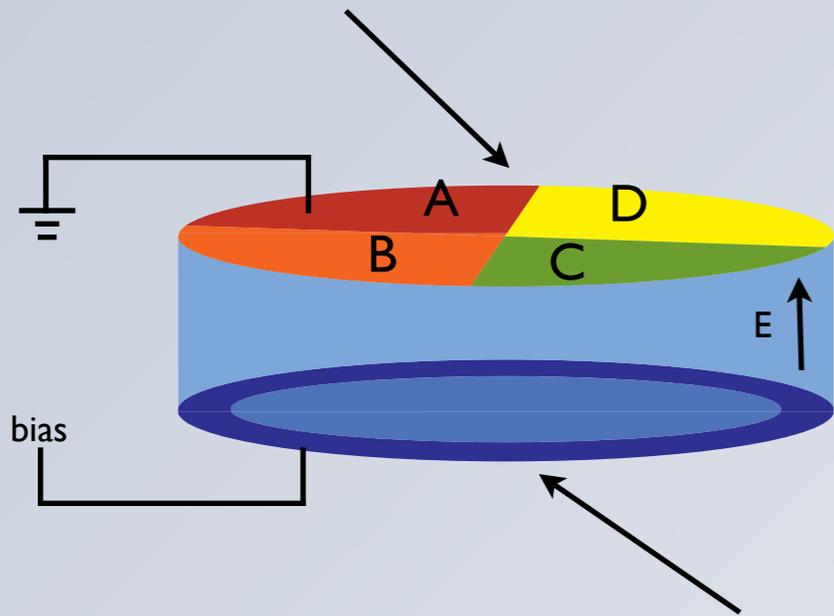
Event by event **discrimination** of nuclear and electron recoils using **ionization** and **phonon** signals



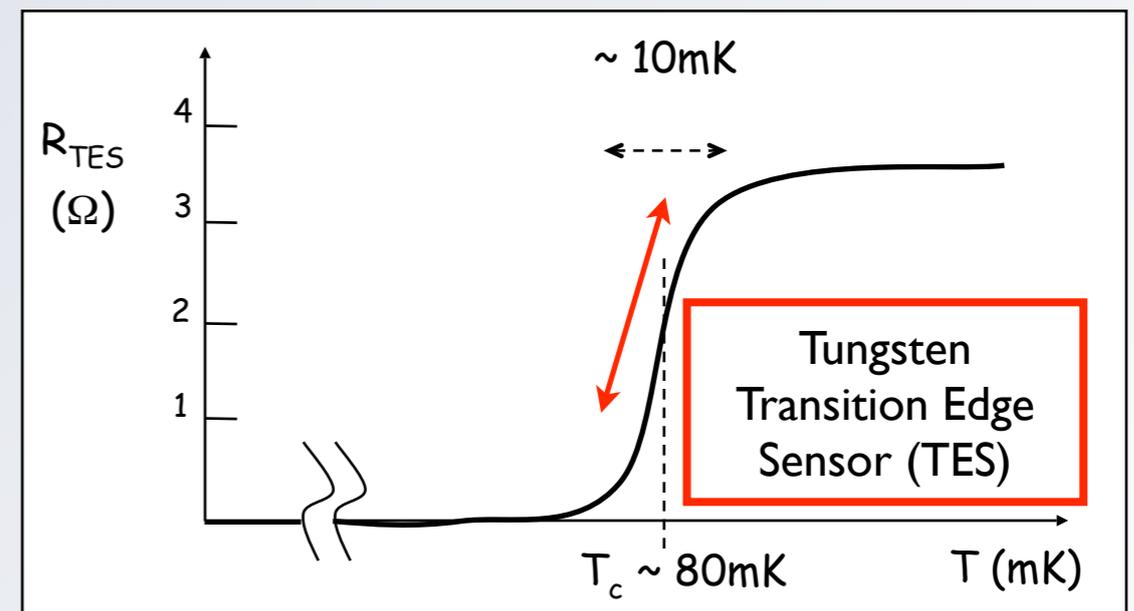
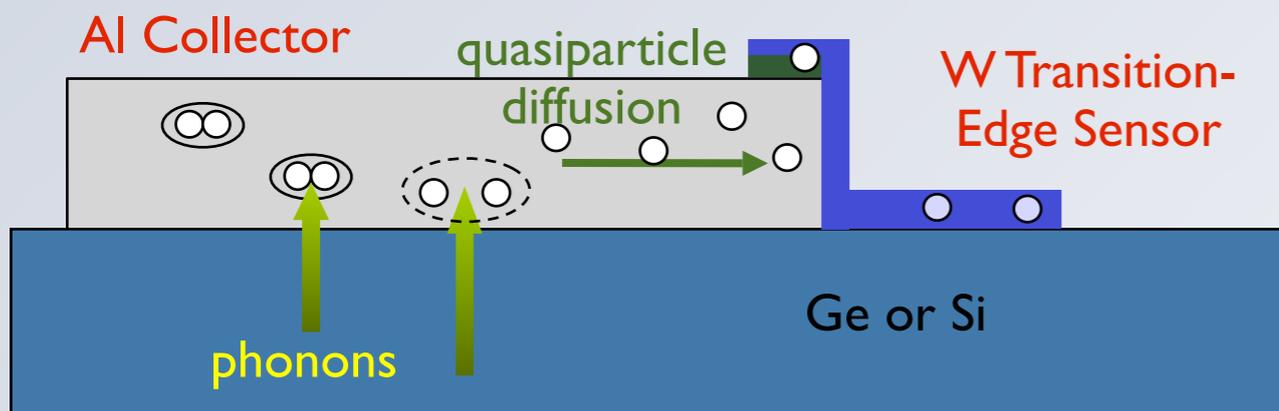
**Control Backgrounds** by going underground, using clean materials and shielding

# ZIP: Z-sensitive Ionization & Phonon Detectors

**Phonon readout:** 4 quadrants of superconducting Al phonon absorbers and W Transition-Edge Sensors

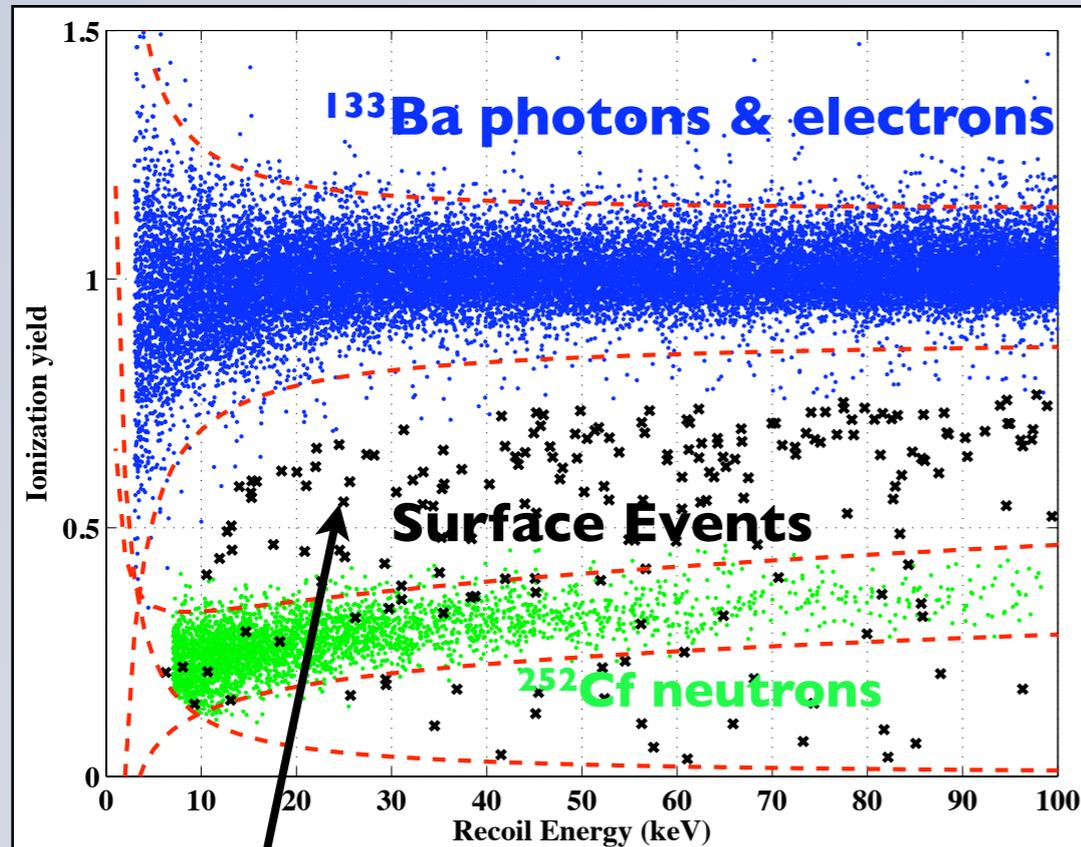


**Charge readout:** 2 concentric electrodes for measurement and fiducial volume

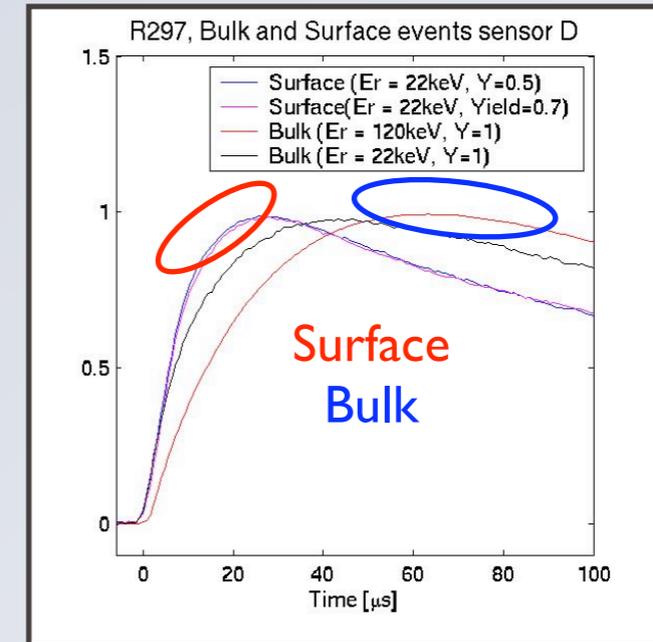


# Basic Discrimination principles

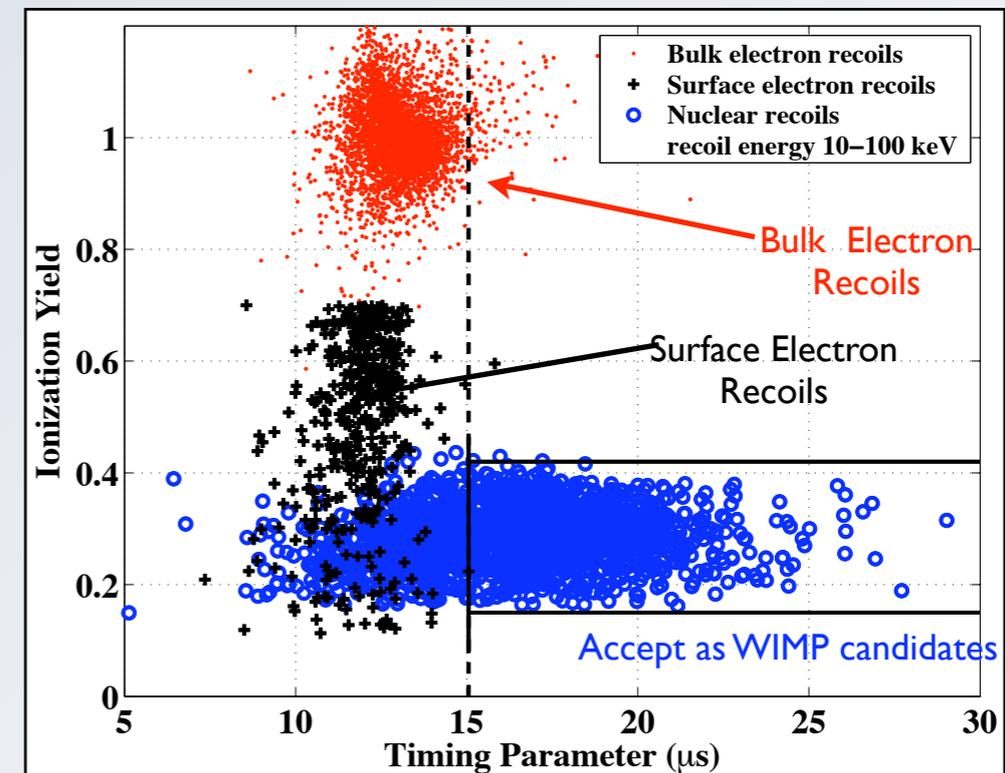
Reject bulk electron recoils using  $\gamma = Q/E_{\text{recoil}}$



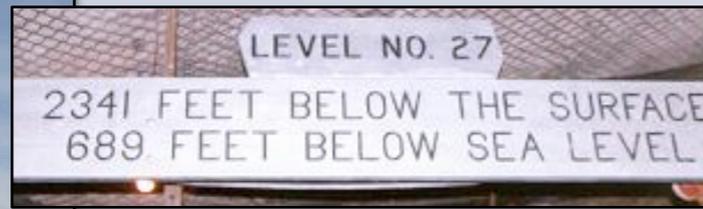
- Events within 10 micron “dead layer” have poor charge collection
- Electrons, x-rays low energy gammas from contamination of surfaces surrounding detectors



Reject surface events using promptness of phonon pulse



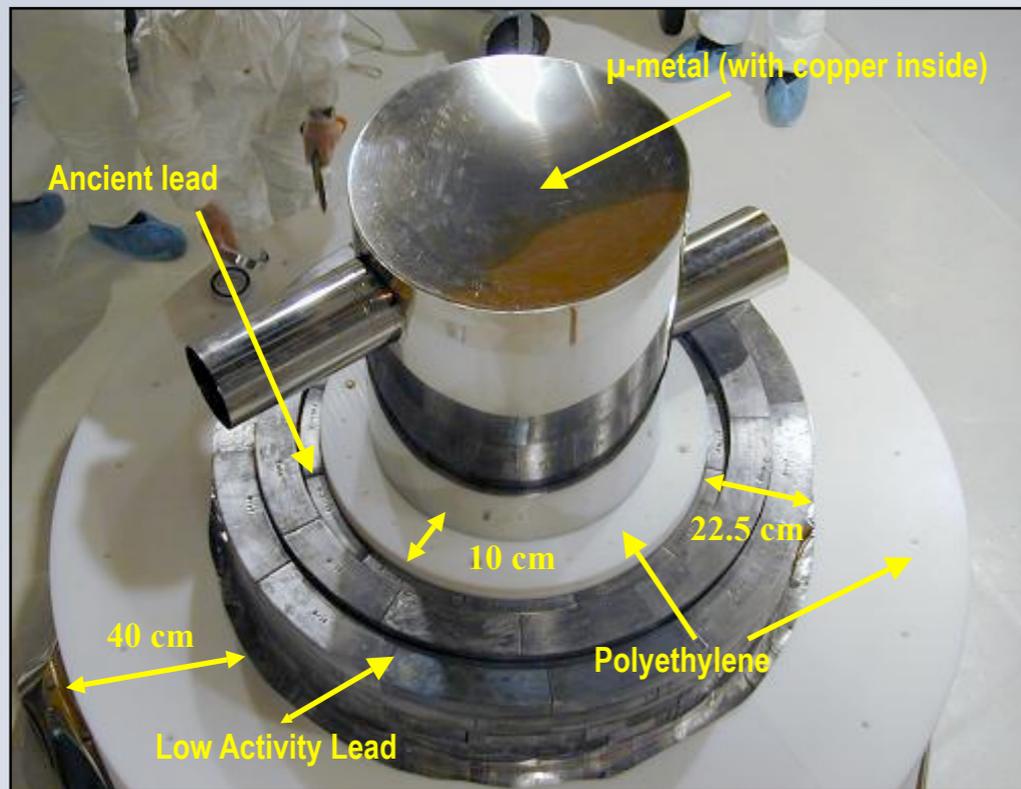
# Background Control in CDMS-2



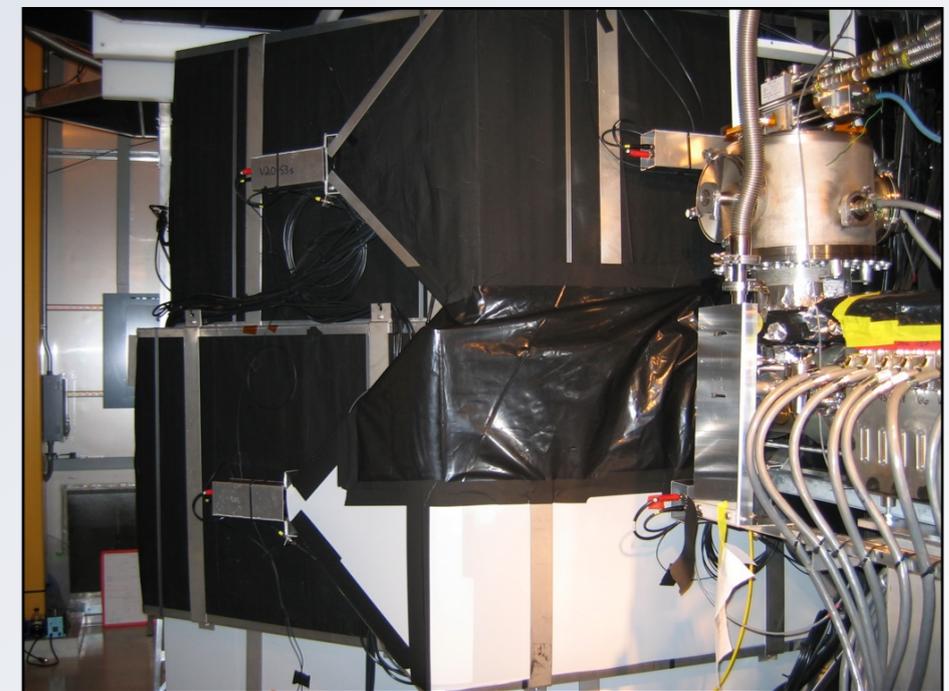
Soudan Underground Lab



RF shielded class  
10,000 clean room

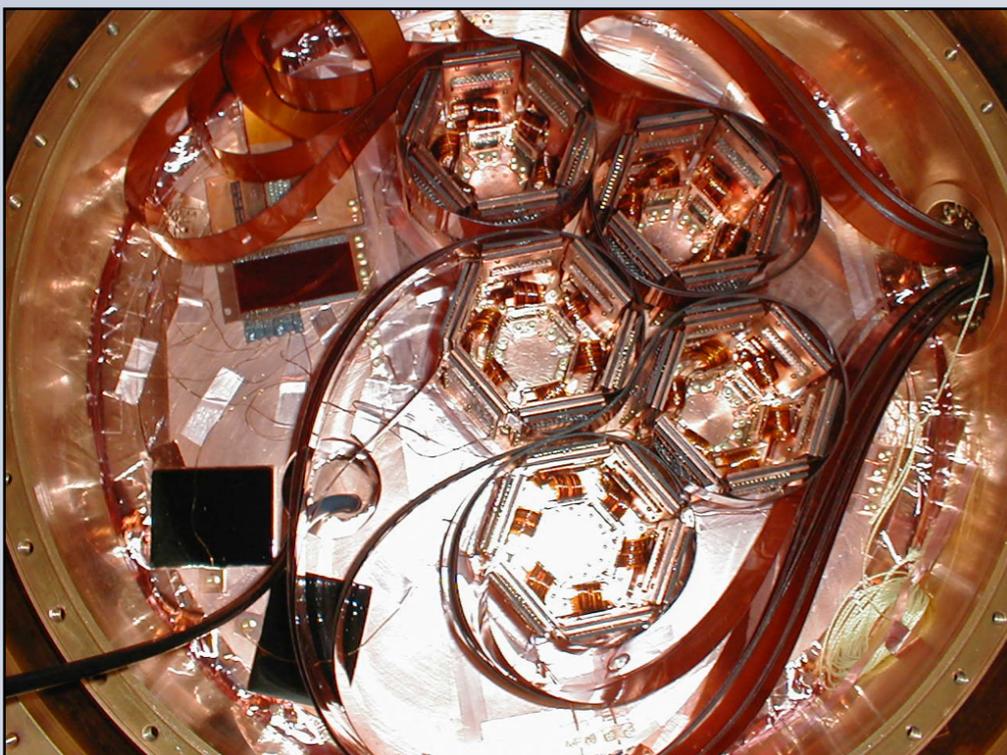


Passive shielding



Plastic scintillator muon veto

# CDMS-2 @ Soudan



	T1	T2	T3	T4	T5
Z1	G6	S14	S17	S12	G7
Z2	G11	S28	G25	G37	G36
Z3	G8	G13	S30	S10	S29
Z4	S3	S25	G33	G35	G26
Z5	G9	G31	G32	G34	G39
Z6	S1	S26	G29	G38	G24

Side View

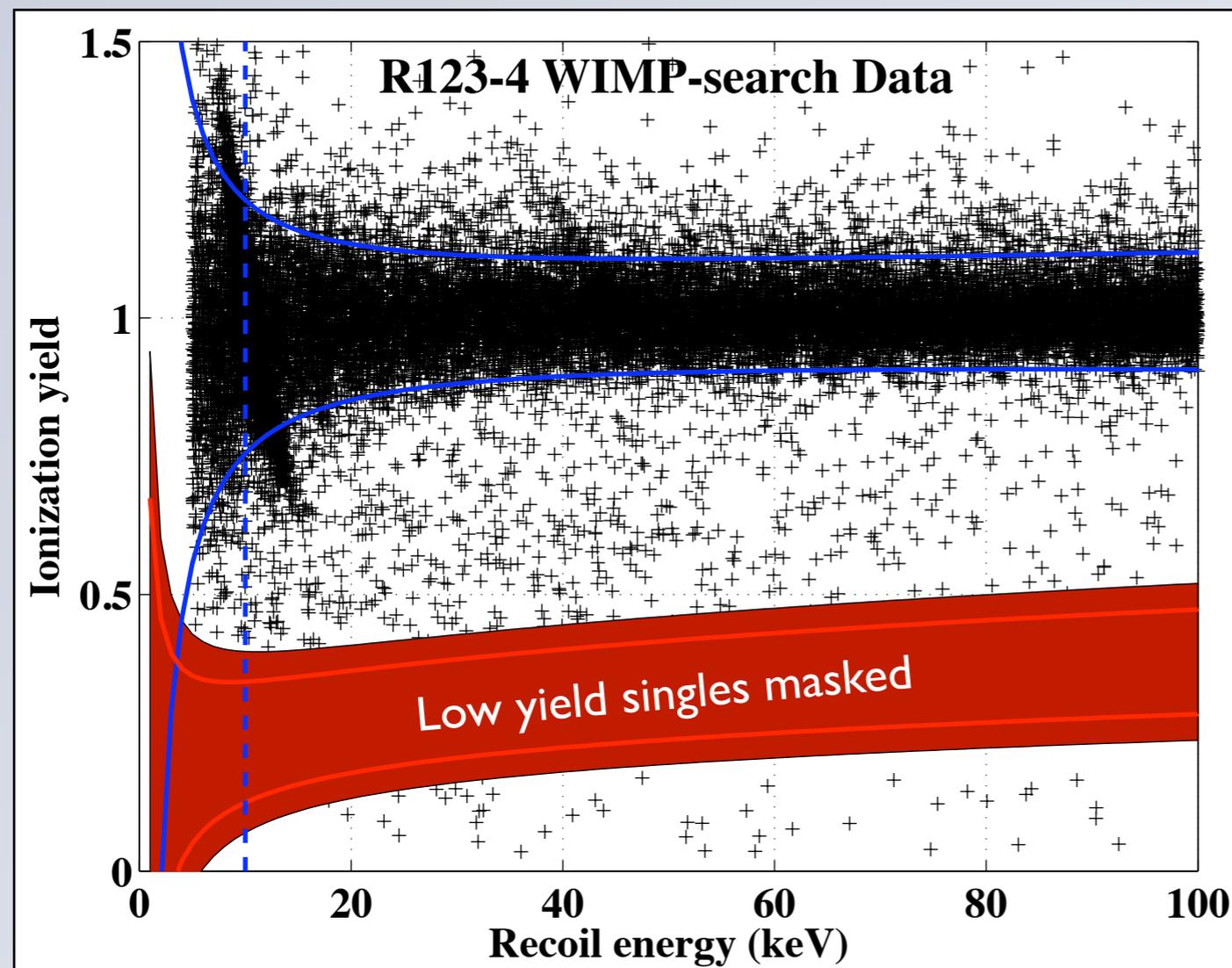
- 30 detectors installed and operating in Soudan since June 2006.
  - 4.75 kg of Ge, 1.1 kg of Si
- Seven Total Data Runs:
  - R123 - R124:
    - taken: (10/06 - 3/07) (4/07 - 7/07)
    - exposure: ~400 kg-d (Ge “raw”)
    - PRL 102, 011301 (2009)
  - R125 - R128
    - taken: (7/07 - 1/08) (1/08 - 4/08)  
(5/08 - 8/08) (8/08 - 9/08)
    - exposure: ~ 750 kg-d (Ge “raw”)
    - Under Analysis
  - R129:
    - taken: (11/08 - 3/09)

# CDMS-2 : First Five Tower Result

PRL 102, 011301 (2009)

## Blind Analysis:

Event selection and efficiencies were calculated without looking at the signal region of the WIMP-search data.



## Event Selection:

- Energy threshold ( $10-100$  keV)
- Veto-anticoincident
- Single-scatter
- Inside fiducial volume
- 2-sigma Nuclear Recoil
- Phonon timing

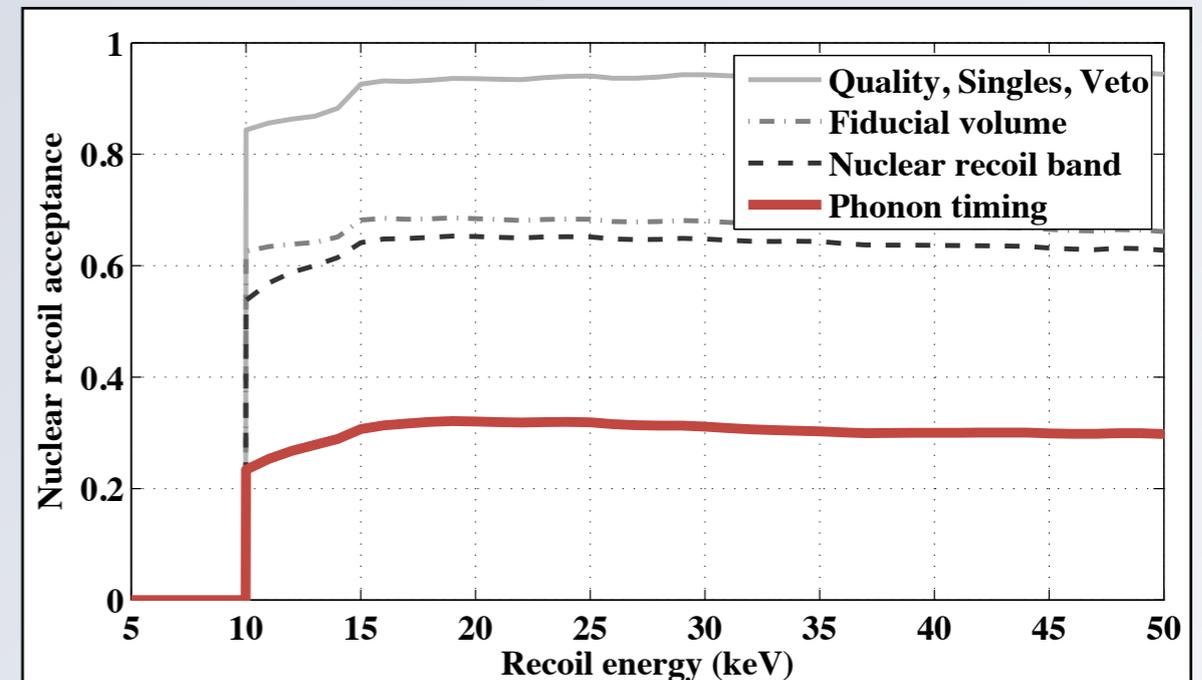
# CDMS-2 : First Five Tower Result

PRL 102, 011301 (2009)

## Surface Background

*Estimated number of background events to pass surface cut in Ge*

$$0.6^{+0.5}_{-0.3} (\text{stat.})^{+0.3}_{-0.2} (\text{syst.})$$



## Neutron Background

*Poly Cu ( $\alpha, n$ ):  $< 0.03$*

*Pb (fission):  $< 0.1$*

*Cosmogenic:  $< 0.1$  (MC 0.03-0.05)*

398 raw kg-d  
121 kg-d WIMP equiv. @ 60  
GeV/c<sup>2</sup> (10 - 100 keV  
analysis energy range)

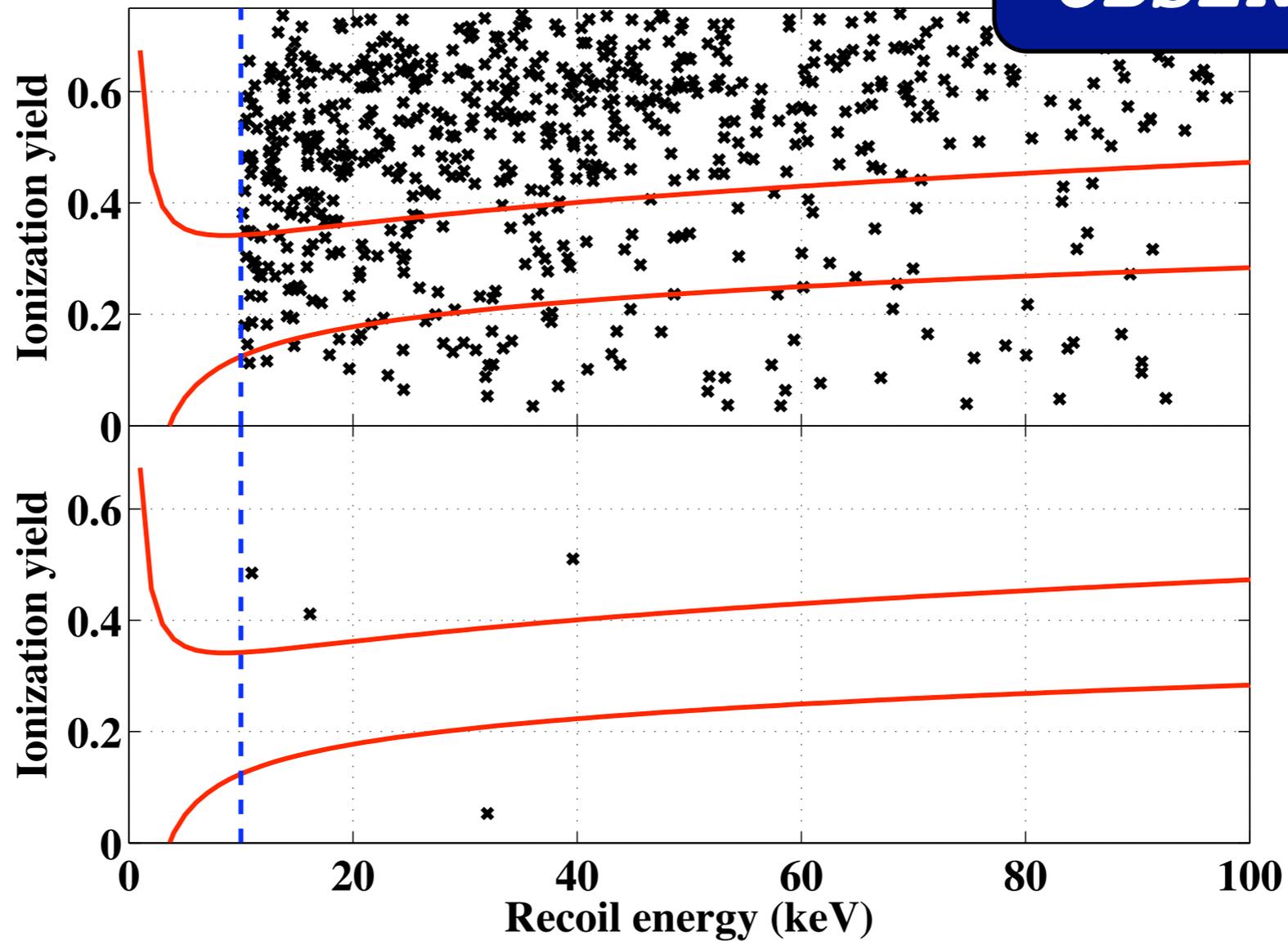
# CDMS-2 : First Five Tower Result

PRL 102, 011301 (2009)

**NO EVENTS  
OBSERVED!**

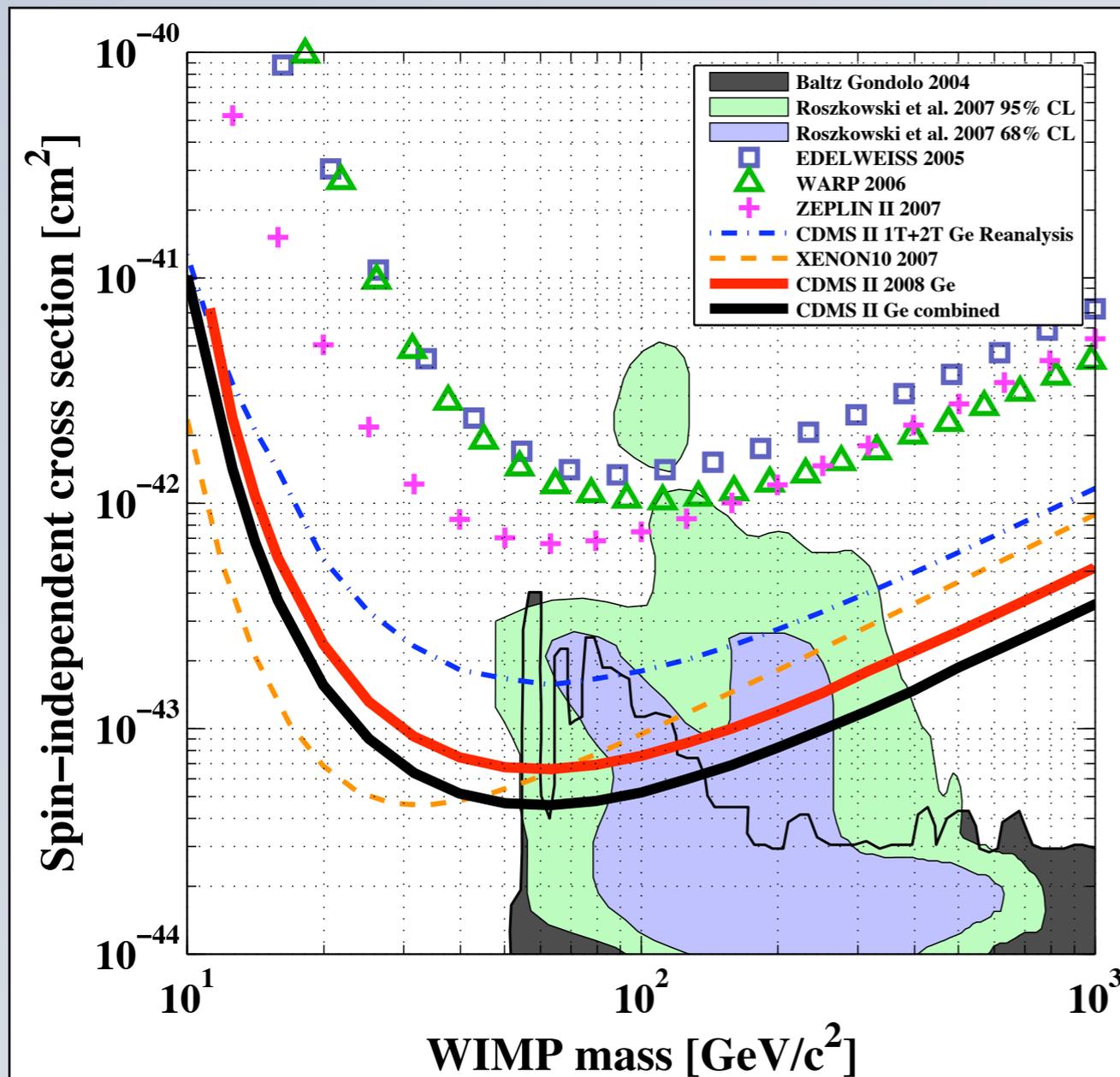
FAIL  
timing cut

PASS  
timing cut



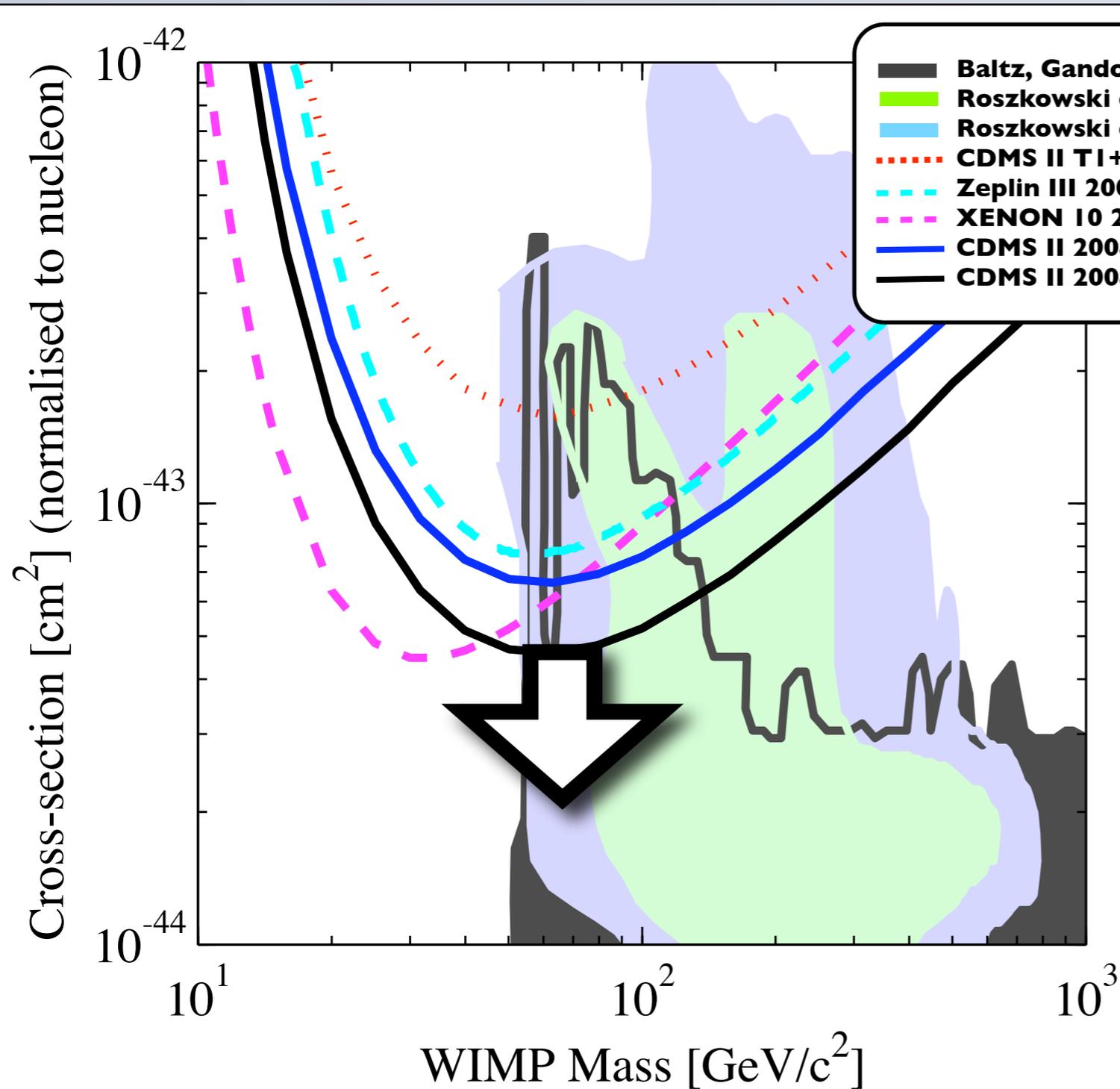
# CDMS-2 : First Five Tower Result

PRL 102, 011301 (2009)



Upper limit at the 90% C.L. on the WIMP-nucleon cross-section is  $4.6 \times 10^{-44} \text{ cm}^2$  for a WIMP of mass **60 GeV/c<sup>2</sup>**

# CDMS-2 : Projected Sensitivity (2009)



## Raw Exposure

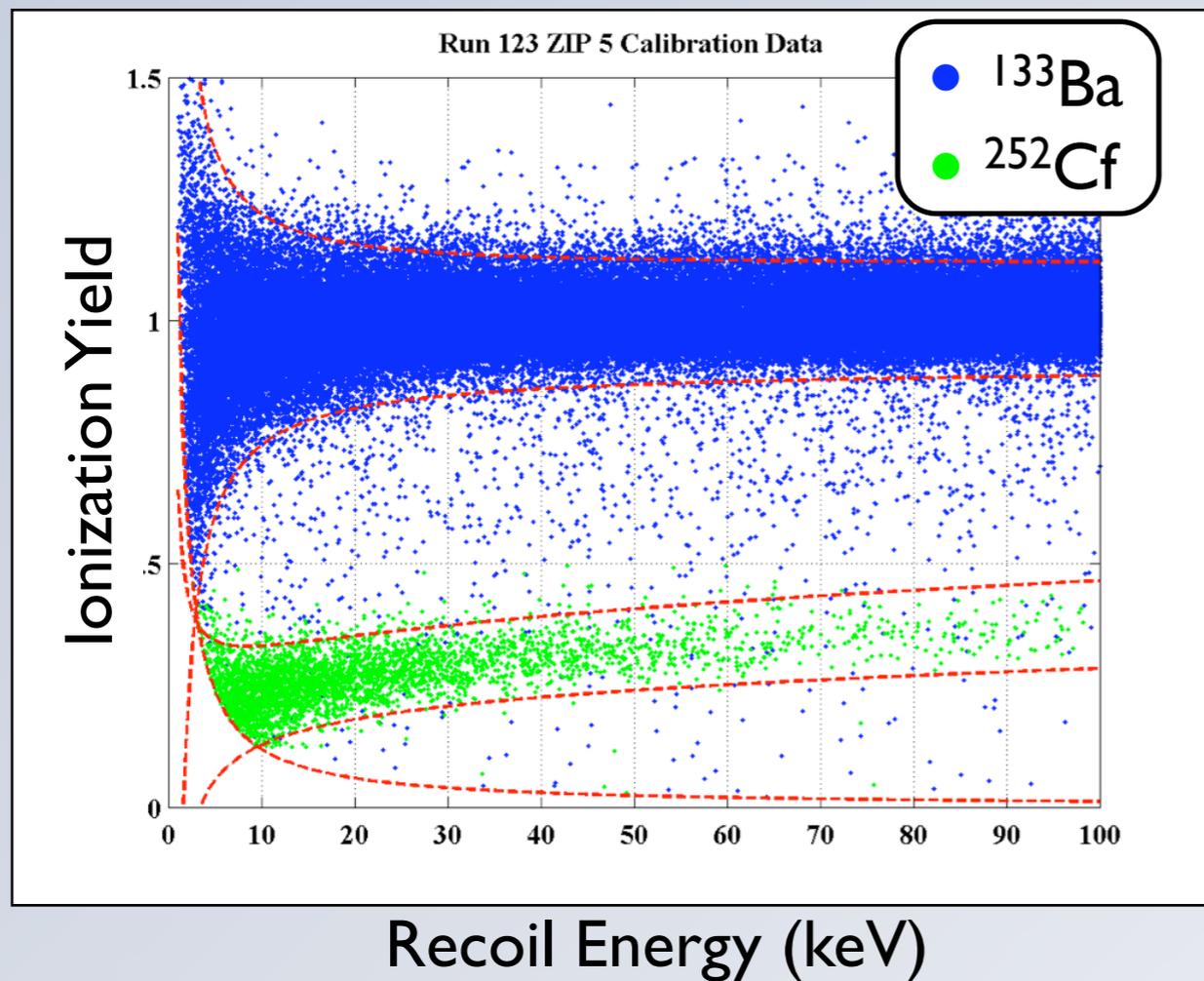
- R118-R119 =  $\sim 120$  kg-d
- Run 123-124 =  $\sim 400$  kg-d
- Run 125-128 =  $\sim 750$  kg-d

**$\sim 2.5$  times more  
total exposure**

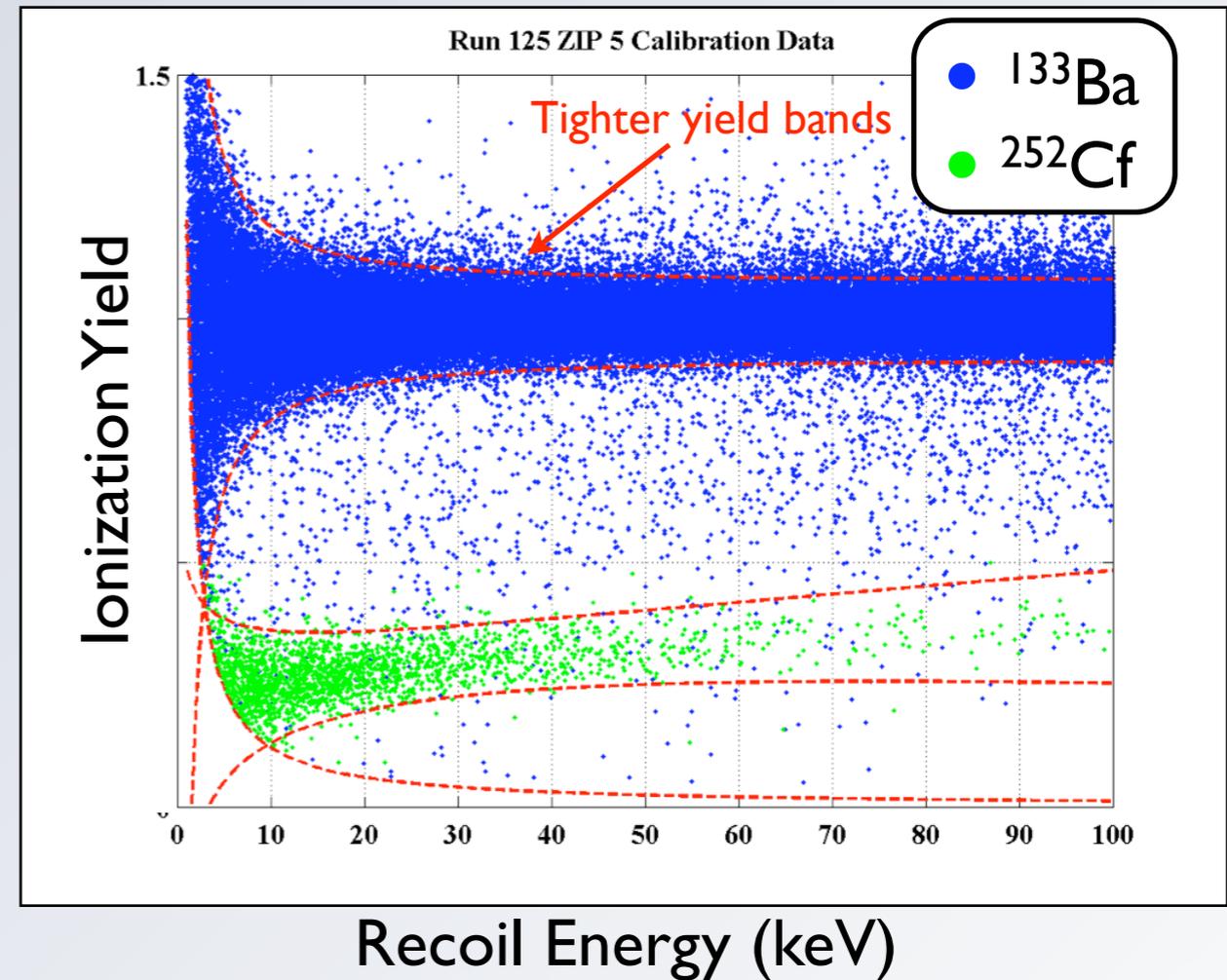
Results expected  
late Summer 2009

# CDMS-2: Calibration Data Preview (2009)

## Previous Analysis PRL 102, 011301 (2009)



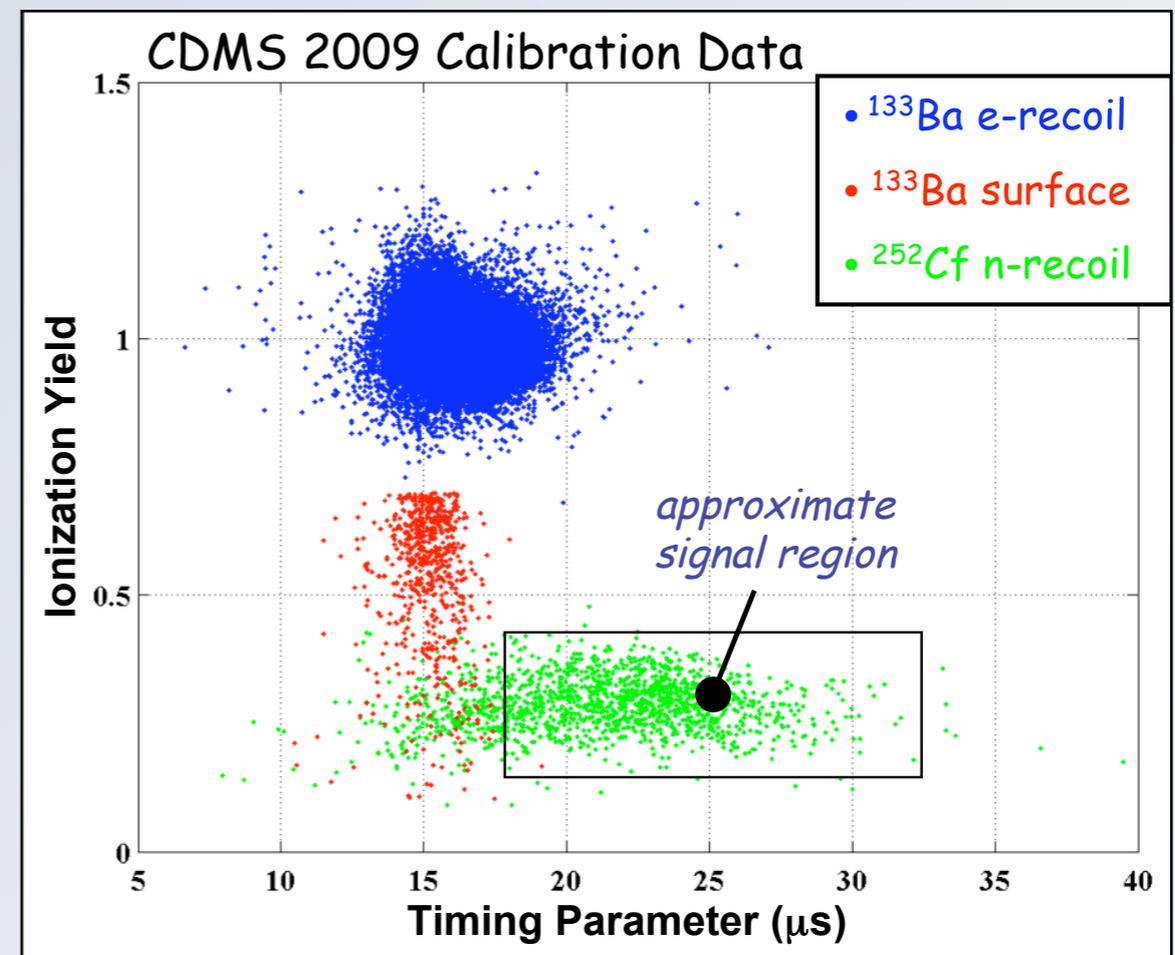
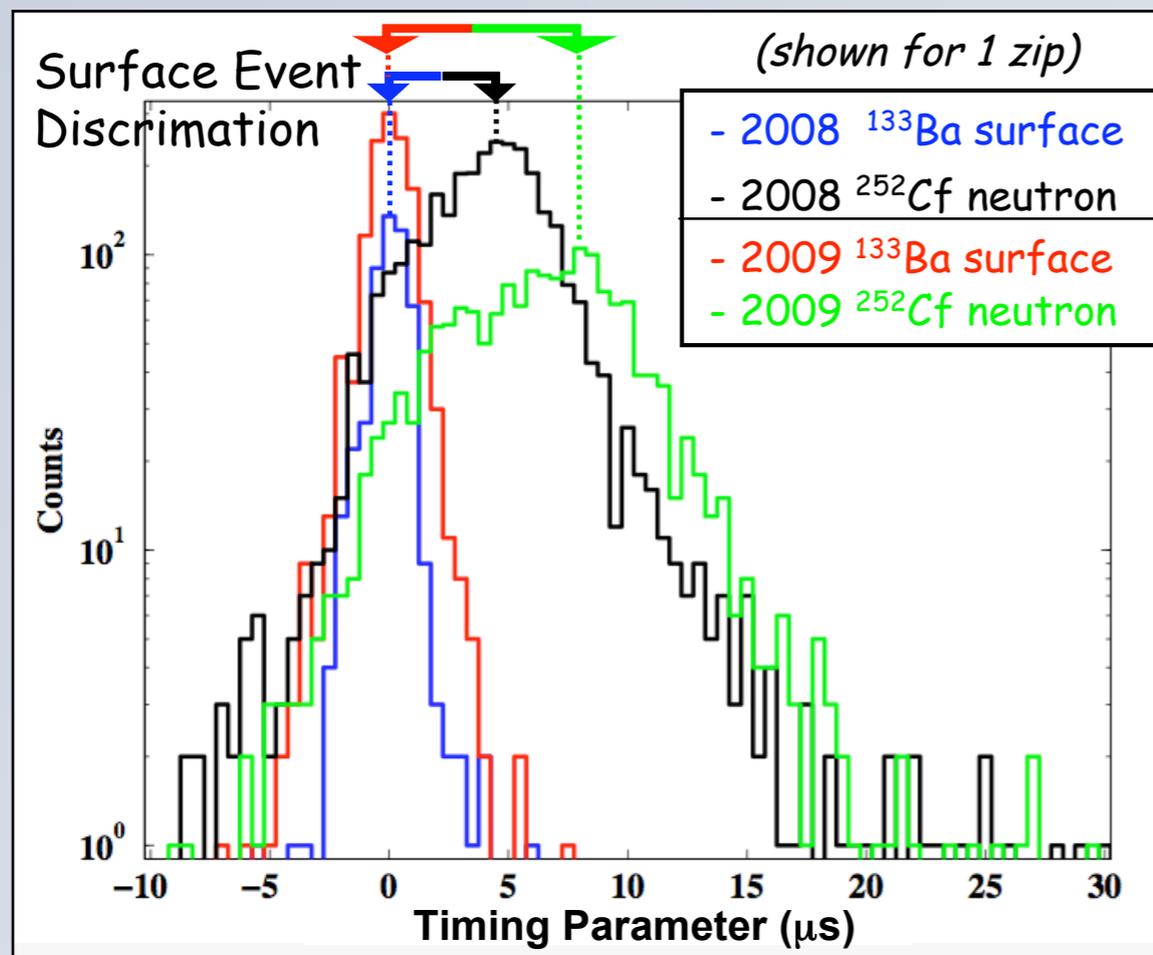
## Current Analysis



Data Quality and preliminary cuts defined!

# CDMS-2: Timing Discrimination Preview (2009)

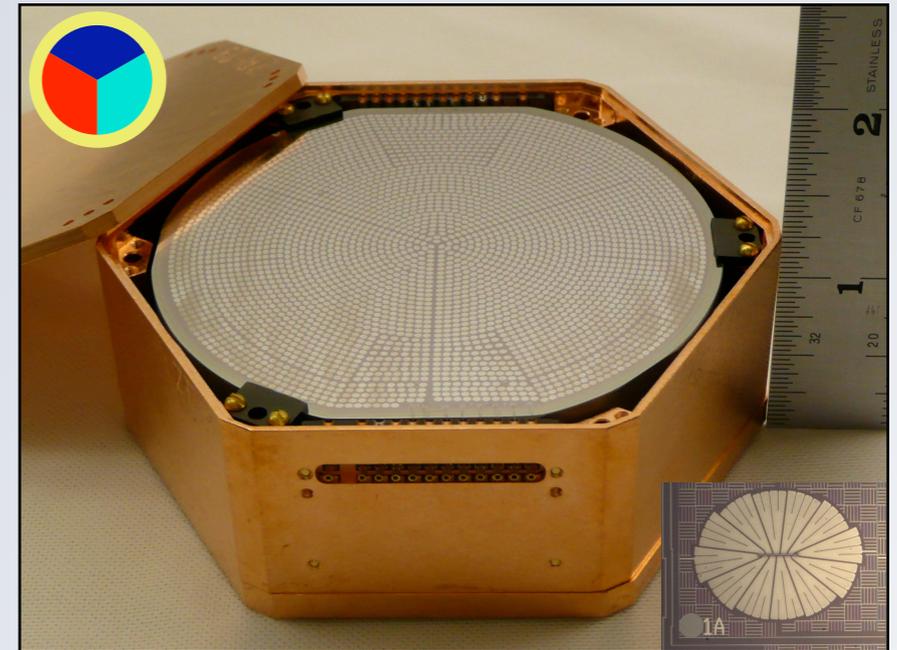
Aiming for x2-3 better surface event rejection  
to keep expected background < 1 event



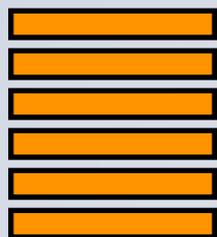
Stay Tuned...

# SuperCDMS Soudan

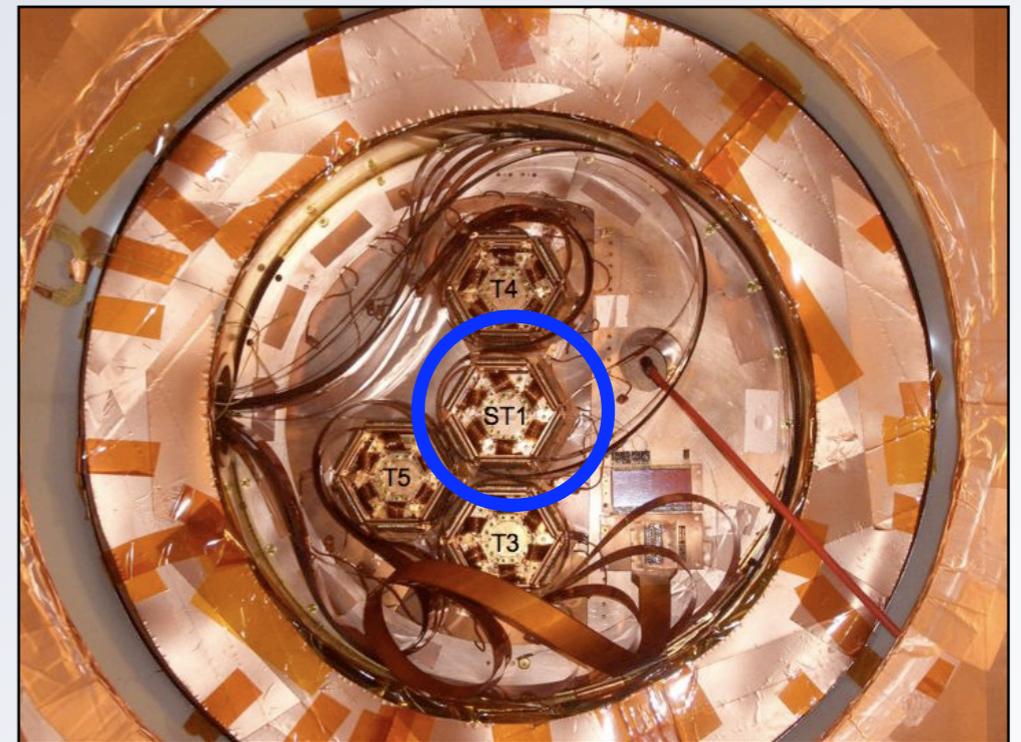
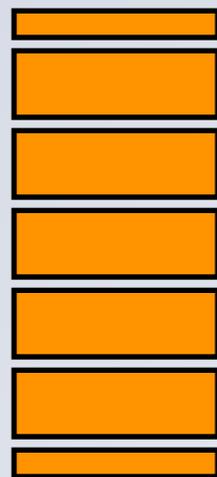
- CDMS-2 target exposure acquired
- 15kg Ge experiment to succeed CDMS-2 with target reach of  $5 \times 10^{-45} \text{ cm}^2$  at  $60 \text{ GeV}/c^2$
- Detectors x2.5 more massive
- Phonon and Ionization sensors modified for better surface event rejection
- First SuperTower already installed



CDMS-2 Tower



SuperTower



# Summary

- Data taken between Oct. 2006 and July 2007 has been analyzed and a cross section limit of  $< 4.6 \times 10^{-44} \text{cm}^2$  (90% CL) was placed for a WIMP of mass 60 GeV/c<sup>2</sup>.
- CDMS II finished taking data on March 18, 2009. We are currently analyzing the last data sets.
- SuperCDMS Soudan will replace CDMS-2 and will extend reach to  $5 \times 10^{-45} \text{cm}^2$  at 60 GeV/c<sup>2</sup>.
- The first SuperTower has been built and has been commissioned at Soudan.